

SUSY and Cosmo WG: summary

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LAPTH

Status of CMSSM

ILC measurements

DM and ILC

Conclusion

SUSY

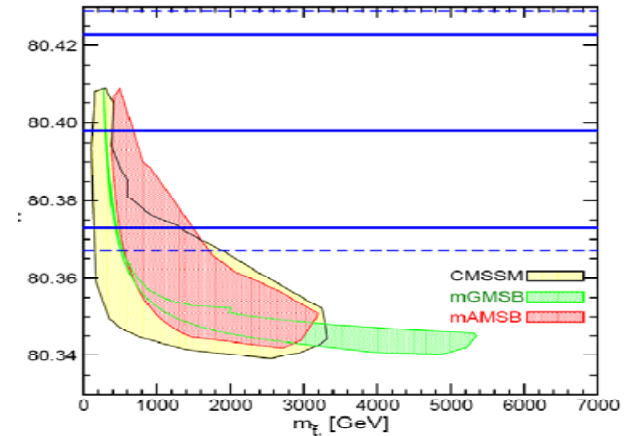
- One attractive extension of SM
- Symmetry fermion -boson
- Solution to gauge hierarchy problem
- Radiative EW symmetry breaking –light Higgs

- Gauge coupling unification
- Could be link to string theories

- Cold dark matter candidate if R-parity conservation
- MSSM, CMSSM, extensions...

Constraints on SUSY – status

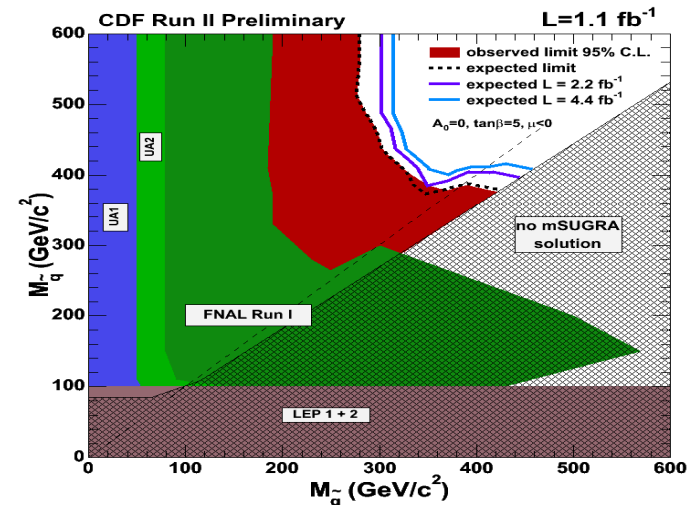
- EWPO : $M_W, \sin^2\theta_{\text{eff}}^1$
 - M_W prefers light stops
- B observables :
 - $\text{Br}(B_s \rightarrow \mu\mu) < 5.8 \times 10^{-8}$ disfavors very large $\tan\beta$ and light H/A
 - b \rightarrow s γ constrain m_0 - $m_{1/2}$, $\tan\beta$ dependent



- $(g-2)_\mu$: prefers light smuons/gauginos

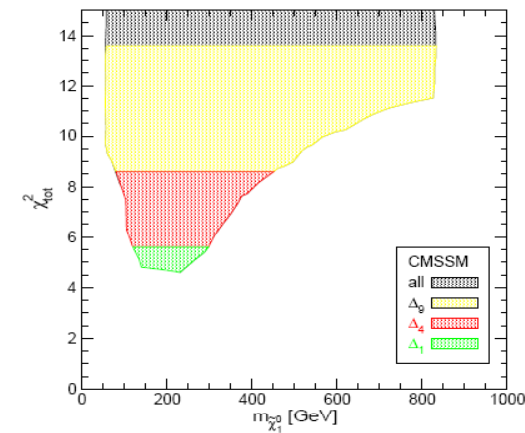
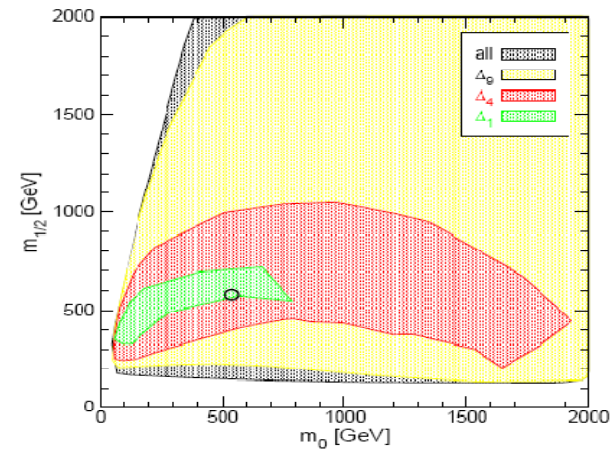
$$a_\mu^{\text{exp}} - a_\mu^{\text{theo}} = (27.5 \pm 8.4) \times 10^{-10},$$

- Tevatron lower limits : squarks/gluino



Constraints on CMSSM

- CMSSM: MSSM with GUT scale universal condition (4 1/2 parameters- $m_0, m_{1/2}, A_0, \tan\beta, \text{sign}(\mu)$) and MFV
- χ^2 fit: Large area of parameter space allowed
- Constraint on mass scale : $\Delta\chi^2 < 4$ -- LSP < 450 GeV, other sparticles could be in reach of ILC500 or 1000



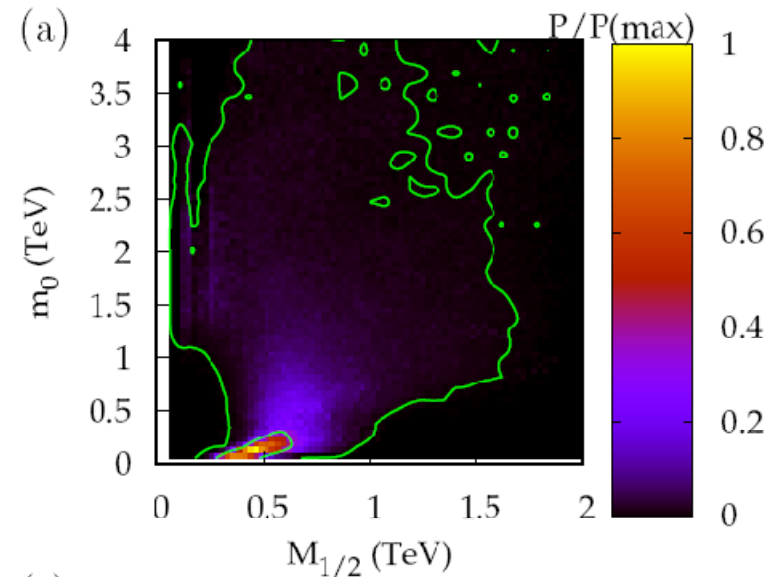
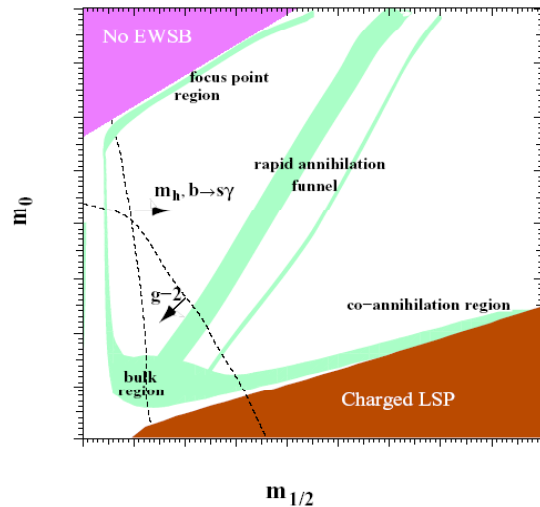
Including DM

- Assume R parity conservation
- Include WMAP/SDSS and DM relic density

$$\Omega h^2 = 0.104^{+0.0073}_{-0.012}$$

- Strong relation $m_0 - m_{1/2}$ for fixed values of A_0 and $\tan\beta$

- MCMC approach+ Bayesian statistics : fits of CMSSM +SM parameters
- 90%CL – no constraint yet on SUSY spectrum – best fit at low $m_0 - m_{1/2}$

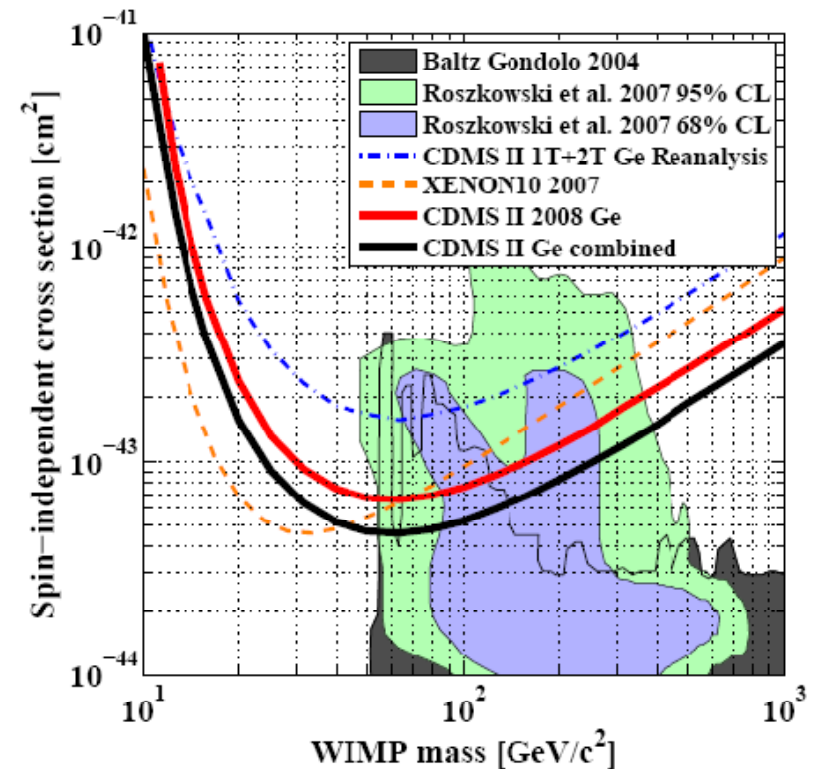


Allanach et al. 0705.0487

Dark matter – Searches

- Direct detection – search for WIMPS interaction with nuclei in large detectors -- experiments are improving : new limits
- Xenon and CDMS – $\sigma^{\text{SI}} > 4 \times 10^{-8} \text{pb}$
- Probe parameter space of MSSM
- Goal 10^{-10}pb in < 10 years

- SD limits also improving-
COUPP and Kims $\sim \sigma^{\text{SD}} \sim 10^{-2} \text{pb}$ --
Not yet probing the MSSM



SUSY at ILC

- Search for sparticles hard to see at LHC
- Determination of parameters :
 - Masses can be measured at per-mil level : threshold scans/endpoint measurements
 - Spin of particles
 - Determination of soft SUSY parameters
 - LHC/ILC : Phases from stop production and decay (G. Moortgat-Pich)
 - Underlying model, SUSY breaking mechanism..
 - DM properties

SUSY at ILC

- **Determination of parameters :**

- Sneutrinos – (T. Robens)

- Mass measurement mass even when decay is invisible (e.g. SPS1A)

Freytas'05 $e^+ e^- \longrightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^- \longrightarrow (\tilde{\nu}_e \tilde{\nu}_\mu e^- \mu^+ \longrightarrow) \tilde{\chi}_1^0 \tilde{\chi}_1^0 e^- \mu^+ \nu_\mu \bar{\nu}_e$

signal: $e^- \mu^+ + \mathbf{E}_{\text{miss}}$

- Full matrix element for both signal and background (interference effects)
- Chargino mass from threshold scan, sneutrino mass from lepton energy spectrum $\sim 1\%$ precision
- study is ongoing : improved cuts etc...

Precision

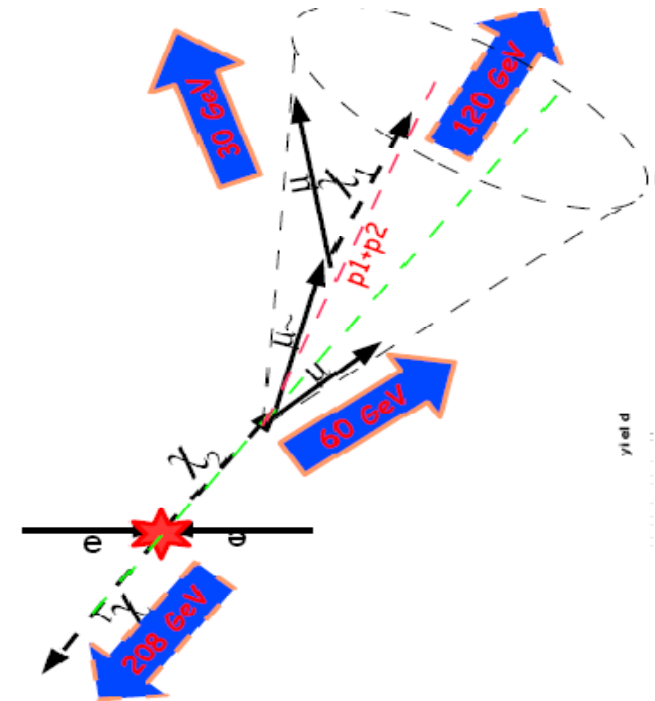
- To take advantage of high precision at ILC – must make theoretical predictions at one-loop level
 - Many processes computed in the last few years
 - Chargino pair production with CP violation (K. Rolbiecki)
 - Asymmetries induced at one-loop only – sensitive to phases, e.g. φ_t could be a few %

$$A_{12} = \frac{\sigma^{\text{loop}}(e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_2^-) - \sigma^{\text{loop}}(e^+e^- \rightarrow \tilde{\chi}_2^+ \tilde{\chi}_1^-)}{\sigma^{\text{tree}}(e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_2^-) + \sigma^{\text{tree}}(e^+e^- \rightarrow \tilde{\chi}_2^+ \tilde{\chi}_1^-)}$$

- Experimental level : studies with full simulation
 - Di-muon decay of neutralino2 (N. D'Ascenzo)
 - Smuon pair production (Chen)

Di-muon decay of neutralino 2

- Muon channel has small BR (2.5% for SPS1a) but di-muon signature is clean and easy to detect at ILC
- Signal is small, background is huge (more than 2 orders of magnitude)
- Optimisation, extended Likelihood method for enhancing signal vs background for small signals
- Full background + full simulation on the way



- Two isolated muons with :
 - Missing energy > 300 GeV
 - Transversal momentum > 40 GeV
 - Acoplanarity $> 0.5 \pi$

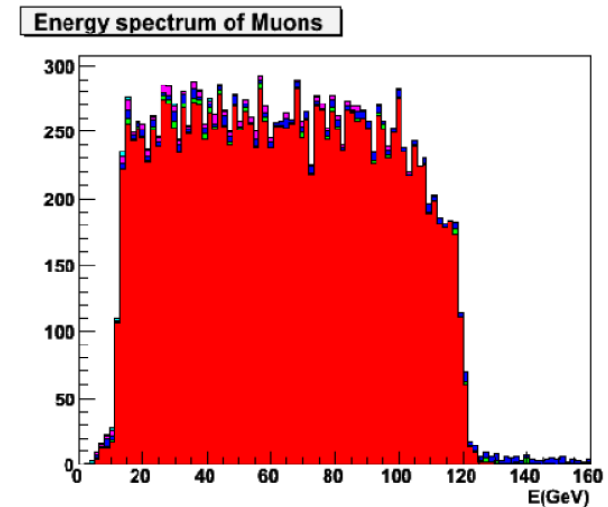
RH- Smuon pair production

- Determine mass and spin of smuon
- Full simulation
- Signal: 2 muons +missing energy
- Background small after cuts
- Masses from endpoint energy :
precision <1%
- Determination of spin
- Angular distribution

$$\frac{d\sigma}{d\cos\theta} \propto \sin^2\theta$$

- Shows that smuon has spin zero

$$e_L^+ e_R^- \rightarrow \tilde{\mu}_R \tilde{\mu}_R \rightarrow \mu^+ \mu^- \tilde{\chi}_1^0 \tilde{\chi}_1^0$$



Calculated: $m_{\tilde{\mu}_R} = 142.89 \pm 0.33 \text{ GeV}$
 $m_{\tilde{\chi}_1^0} = 97.63 \pm 0.22 \text{ GeV}$

Dark Matter

- Evidence for BSM physics?
- Many candidates for DM (SUSY and non-SUSY) - neutralino
- MSSM, CPVMSSM
- CMSSM, GMSB, AMSB, NUHM, VCMSSM (GUT scale conditions and symmetry breaking mechanism)
- NMSSM, nMSSM, MNSSM, CNMSSM, VCNMSSM, USSM – μ problem
- SO10SSM, E6SSM...

Models can have special DM properties and face different challenges for measurement of sparticles at colliders/ILC

- **MNSSM** (S. Hesselbach)
 - Extra neutralino (singlino) can be light
 - Extra scalars possibly light
- **USSM** (J. Roberts)
 - Extra neutralinos, Z'

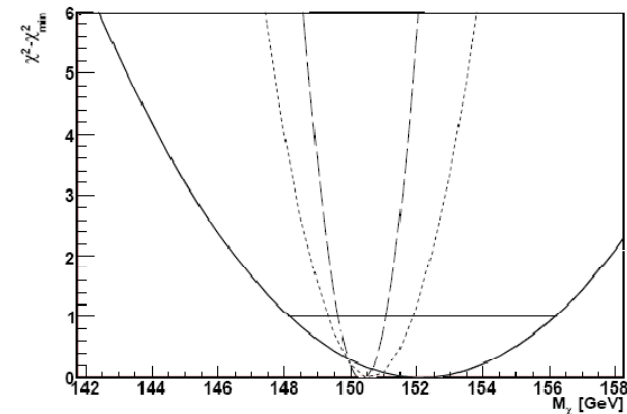
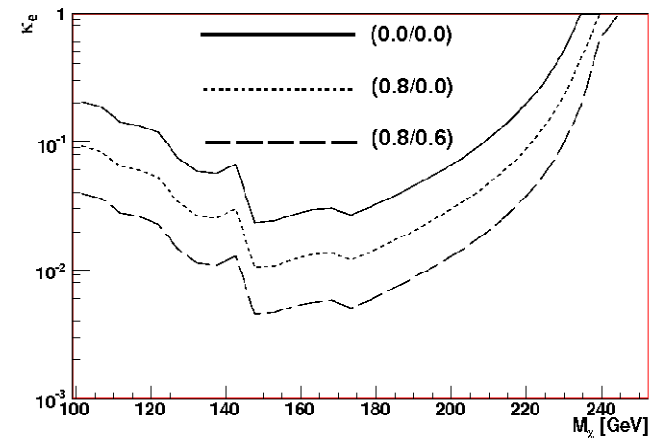
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Model independent approach

Model independent WIMP search

- WIMP pair with single photon (Bartels)
- One of physics analyses of ILD detector concept –optimisation
- Could be only SUSY signal at 500GeV
- Simple and clean signature
- Estimate cross-section from Ωh^2 assuming a certain fraction of annihilation DM in e^+e^- pairs
- Reach in coupling+ mass determination (2007)
- Improved analysis going on

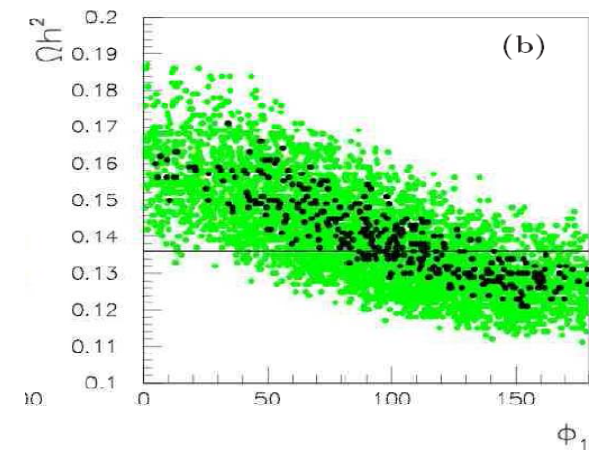


DM at colliders

- WMAP and SDSS gives precise information on the amount of dark matter
- “Doing cosmology at colliders” : discovery of new particles and measurement of their properties →
 - “collider prediction” for the relic density of DM
 - Matches what has been measured in the sky/confront cosmological model
- Precision measurements at colliders are needed
- How difficult strongly depends on the details of the new physics model – which SUSY scenario, what is the dominant DM annihilation process
- Studies exist for both LHC and ILC in CMSSM and MSSM – bulk scenario, stau coannihilation- focus point scenario
 - Polesello, Tovey, Nojiri, Martyn, Bambade et al, Baltz et al,
 - Some scenario, with precision expected at ILC could match the precision of PLANCK

CPVMSSM and ILC

- New study within CPVMSSM in specific scenario with light stau $\tau_1, \tau_2, \chi_1, \chi_2, \chi^+$ accessible at ILC
- Staus are mixed and both contribute to annihilation of LSP into tau pairs
- All signals in $\tau\tau E_{\text{miss}}$ - disentangle sources –
- Use measurement of masses + $\theta_\tau + P_\tau$ to determine model parameters and infer Ωh^2
- $0.116 < \Omega h^2 < 0.19$
- WMAP: $0.094 < \Omega h^2 < 0.136$
- Need to know couplings of LSP not only masses



Conclusion

- ILC can be exploited for precise determination of parameters –
 - Underlying model – supersymmetry breaking mechanism
 - Dark matter properties
- Experimental studies with full simulation of signal and background are going on