Physics with calibration data

Gudrid Moortgat-Pick

- Idea
- Thoughts
- Physics
- Questions



- Triggered by assumption that one calibrates data on the Z-pole after each detector change (push-pull scenario):
 - → Z-pole luminosity: 7-8 x 10³²cm⁻² s⁻¹
 - factor 10 more zees / day than in 100 days of SLD data taking
- Even if no Z-run per pull was needed:
 - using the calibration data for physics may be advantageous
- Questions:
 - can we use these data already for physics ew measurements?
 - which physics gain could we achieve?
 - what do we need for 'using' them? (polarimeter?)

Thoughts

- GigaZ option allows ew measurement with unprecedented precision
 - solves existing discrepancies between AFB and ALR
 - important for consistency tests at quantum level already now
 - delivers precise measurement of ALR, but also other observables
 - it is not only ALR that we need!
- GigaZ only discussed as option after the run -- late anyway
 - requires changes in the machine
 - either via bypass or via deceleration or via drift (@half rep. rate)



Accuracy in sin²Θ_{eff}

$$A_{\rm LR} = \frac{2(1 - 4\sin^2\theta_W^{\rm eff})}{1 + (1 - 4\sin^2\theta_W^{\rm eff})^2}$$

woheadrightarrow precision in ALR directly transferred to $\sin^2\Theta_{
m eff}$

- → GigaZ will provide $\Delta \sin^2 \theta_{eff} \sim 1.3 \times 10^{-5}$ (if Blondel scheme)
- → only electron polarization at GigaZ: ~9.5 x 10⁻⁵
- → current value: 16 x 10⁻⁵
- What could we gain with a 'fraction' of GigaZ ?

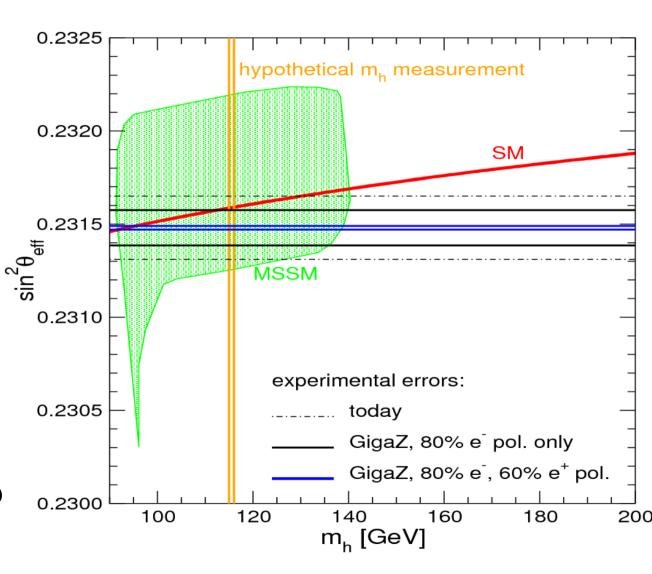
Constraints on the Higgs mass

Distinction between SM and SUSY Higgs

obviously distinction
 between Higgs in
 SM and Susy needs
 GigaZ accuracy

even GigaZ with onlye- Pol. not enough

→ but 3 x 10⁻⁵ from calibration could help



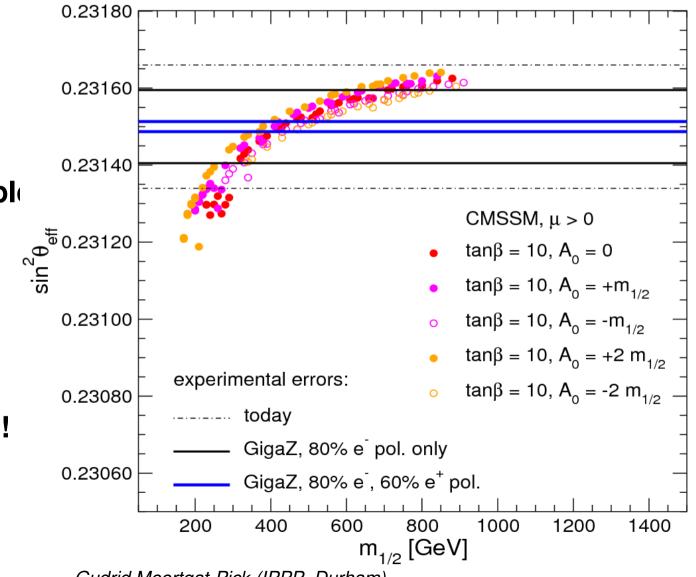
Constraints on SUSY parameters

Prediction of gaugino parameter in CMSSM

every little helps

→ if 3x10⁻⁵ achievable
with 70-100 days
of calibration

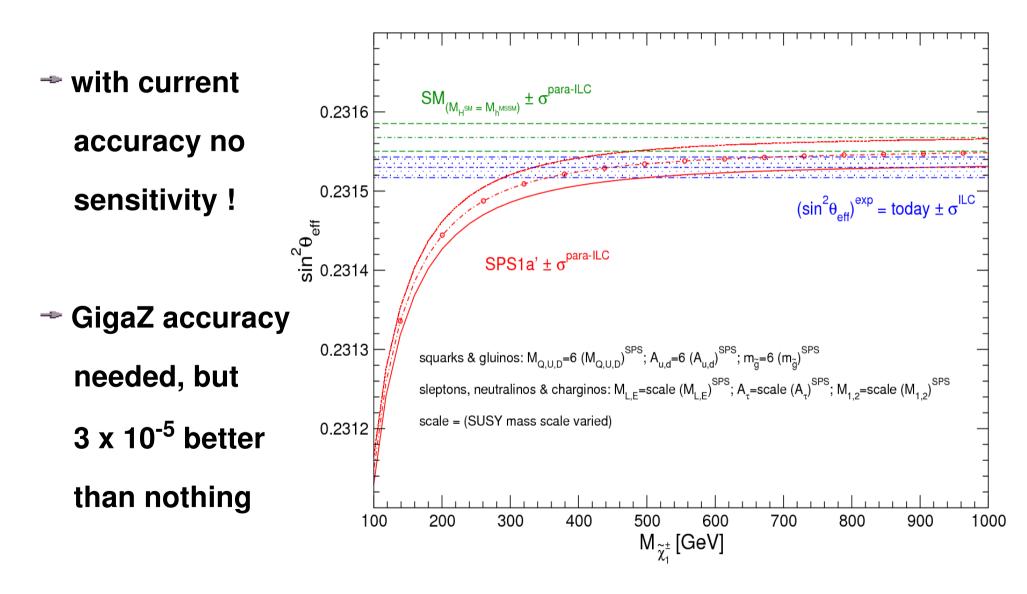
already excellent!



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SUSY constraints in worst scenarios

Case: light Higgs, but squarks in multi-TeV -> nothing@LHC!



Questions

Can we control the systematics if we combine calibration data from many different runs?

- Do we have to flip the polarity?
 - → for e-? for e+?

(remeber: baseline undulator source is already slightly polarized!)

- if flipping: within one calibration run?
- to which accuracy do we have to keep the polarization degree?

Which technical features do we need?