

Study of Coherent instabilities due to electron cloud at CsrTA and KEKB

K. Ohmi, J. Flanagan

Coherent instabilities due to electron cloud

Single bunch instability

- Threshold is determined by balance with Landau damping due slippage (momentum compaction) factor.
- Dependent on emittance
- Depend only on local electron cloud density

Coupled bunch instability

- Threshold is determined by balance with other damping effects.
- Independent on emittance.
- Independent on momentum compaction.
- Depend on electron cloud density, distribution and motion.

Threshold of the strong head-tail instability (Balance of growth and Landau damping)

- Stability condition for $\omega_e \sigma_z / c > 1$

$$\omega_e = \sqrt{\frac{\lambda_p r_e c^2}{\sigma_y (\sigma_x + \sigma_y)}}$$

$$U = \frac{\sqrt{3} \lambda_p r_0 \beta}{v_s \gamma \omega_e \sigma_z / c} \frac{|Z_{\perp}(\omega_e)|}{Z_0} = \frac{\sqrt{3} \lambda_p r_0 \beta}{v_s \gamma \omega_e \sigma_z / c} \frac{KQ}{4\pi} \frac{\lambda_e}{\lambda_p} \frac{L}{\sigma_y (\sigma_x + \sigma_y)} = 1$$

- Since $\rho_e = \lambda_e / 2\pi \sigma_x \sigma_y$,

$$\rho_{e,th} = \frac{2\gamma v_s \omega_e \sigma_z / c}{\sqrt{3} KQ r_0 \beta L}$$

Origin of Landau damping is momentum compaction

$$v_s \sigma_z = \alpha \sigma_{\delta} L$$

- $Q = \min(Q_{nl}, \omega_e \sigma_z / c)$
 $Q_{nl} = 5-10?$, depending on the nonlinear interaction.
- K characterizes cloud size effect and pinching.
- $\omega_e \sigma_z / c \sim 12-15$ for damping rings.
- We use $K = \omega_e \sigma_z / c$ and $Q_{nl} = 7$ for analytical estimation.

Threshold for various rings

	KEKB	KEKB	KEKB-DRt	CESR chess	CesrTA	ILC-OCS	PEPII
L	3016	3016	3016	768.44	768.44	6695	2200
gamma	6849	6849	4501	10372	3914	9785	6067
Np	3.30E+10	7.60E+10	2.00E+10	1.12E+11	2.00E+10	2.00E+10	8.00E+10
ex	1.80E-08	1.80E-08	1.50E-09	1.11E-07	2.30E-09	5.60E-10	4.80E-08
bx	10	10	10	10	10	30	10
ey	2.16E-10	2.16E-10	6.00E-12	1.11E-09	1.50E-12	2.00E-12	1.50E-09
by	10	10	10	10	10	30	10
sigx	4.24E-04	4.24E-04	1.22E-04	1.05E-03	1.52E-04	1.30E-04	6.93E-04
sigy	4.65E-05	4.65E-05	7.75E-06	1.05E-04	3.87E-06	7.75E-06	1.22E-04
sigz	0.006	0.007	0.009	0.0173	0.009	0.006	0.012
nus	0.024	0.024	0.011	0.0487	0.098	0.067	0.025
Q	3.6	5.9	7	4.7	7	7	3.7
omegae	1.79E+11	2.51E+11	5.29E+11	8.20E+10	6.84E+11	6.31E+11	9.20E+10
phasee	3.6	5.9	15.9	4.7	20.5	12.6	3.7
K	3.6	5.9	12.5	4.7	20.5	12.6	3.7
rhoeth	6.25E+11	3.81E+11	1.22E+11	5.73E+12	2.92E+12	1.91E+11	7.67E+11

Tune shift

- Single bunch instability depends on local electron cloud density near the beam.
- Incoherent tune shift can be an indicator of the single instability.
- Tune shift should linearly increase for every bunch passage, because a certain numbers of photoelectrons are supplied by every bunch.
- Tune shift saturates after several 10 bunches passage, because of space charge limit or dynamic balance of creation and absorption of electrons.

Tune shift at CESR

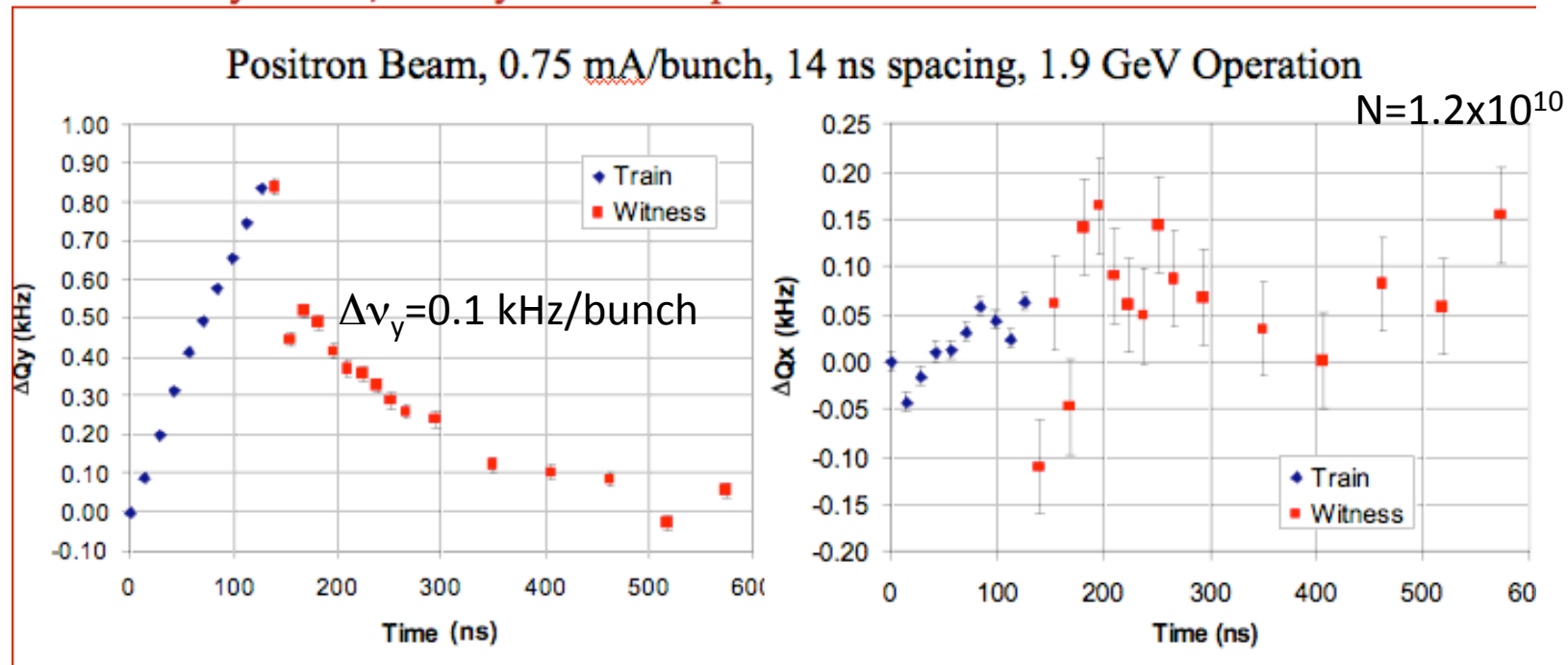
$$\Delta\nu_x + \Delta\nu_y = \frac{r_e}{\gamma} \oint \rho_e \beta ds$$



Cornell University
Laboratory for Elementary-Particle Physics

Witness Bunch Studies – e⁺ Vertical Tune Shift

- Initial train of 10 bunches ⇒ generate EC
- Measure tune shift and beamsize for witness bunches at various spacings
- Bunch-by-bunch, turn-by-turn beam position monitor



Error bars represent scatter observed during a sequence of measurements

1 kHz ⇒ $\Delta\nu = 0.0026$

$\rho_e \sim 1.5 \times 10^{11} \text{ m}^{-3}$

Ohmi, etal, APAC01, p.445

$\beta = 30\text{m}$

Preliminary

Tune shift at KEKB

(T. Ieiri, Proceedings of Ecloud07)

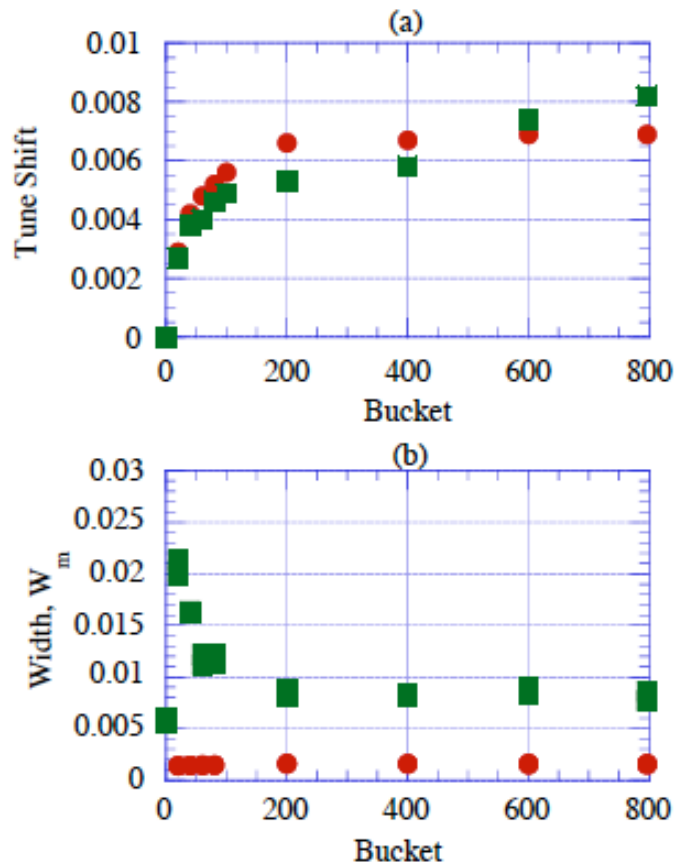


Figure 4: Tune shift (a) and spectrum width (b) along a train. The red dots (horizontal) and green squares (vertical) are measured at a bunch current of 0.5 mA. The tune of the head bunch of the train is used as the reference.

Without solenoid

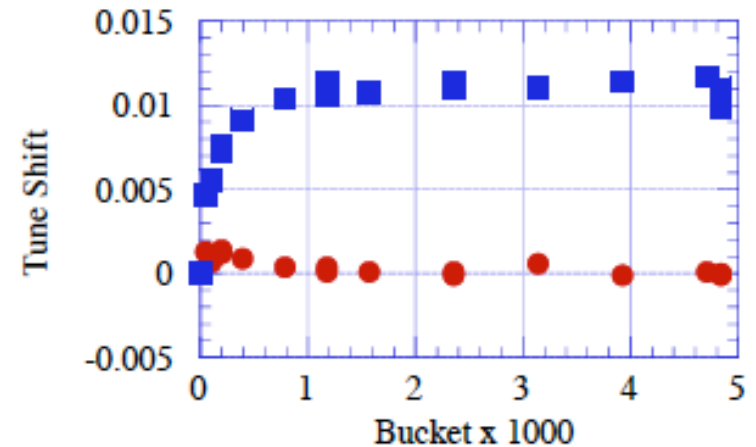


Figure 11: Horizontal (red dots) and vertical (blue squares) tune-shifts along the bunch-train. The bunch current is 1.0 mA with an average spacing of 7 ns.
With solenoid

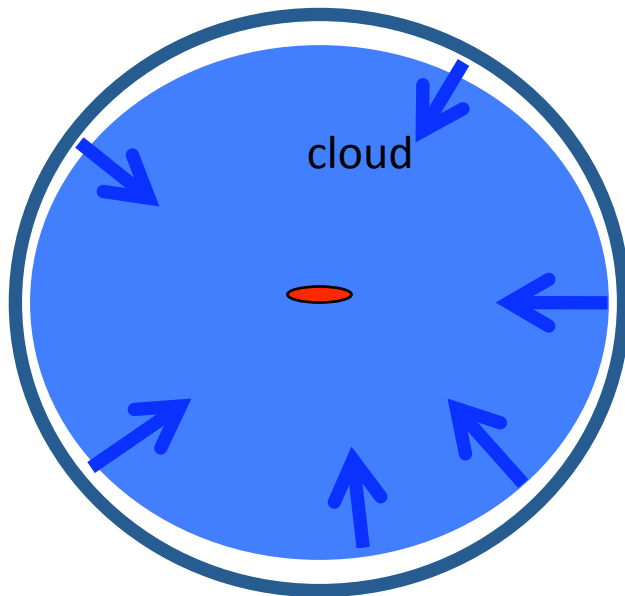
- Both showed similar density because of $v_x + v_y = 0.015$ and 0.012 .
- Round cloud for no solenoid and flat cloud for solenoid. How do we think?

Typical cloud distribution and tune shift

- Tune shift is determined by the electron distribution.
- Electron distribution depends on the initial condition and magnetic field

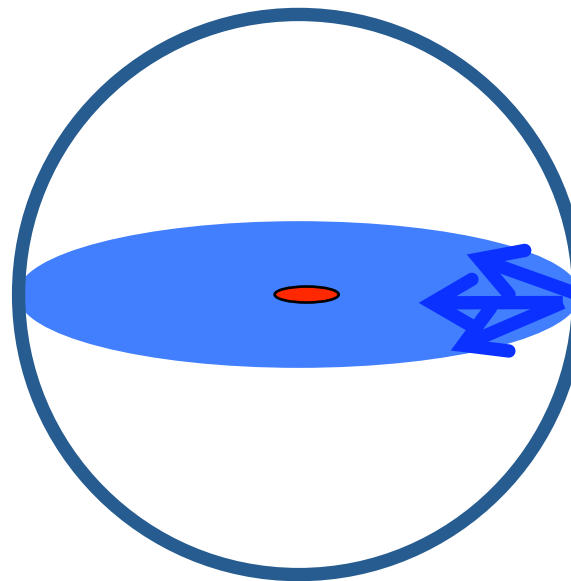
$$\text{div}\mathbf{E} = \frac{\rho_e}{\epsilon_0} \longrightarrow \Delta v_x + \Delta v_y = \frac{r_e}{\gamma} \oint \rho_e \beta ds$$

Straight section

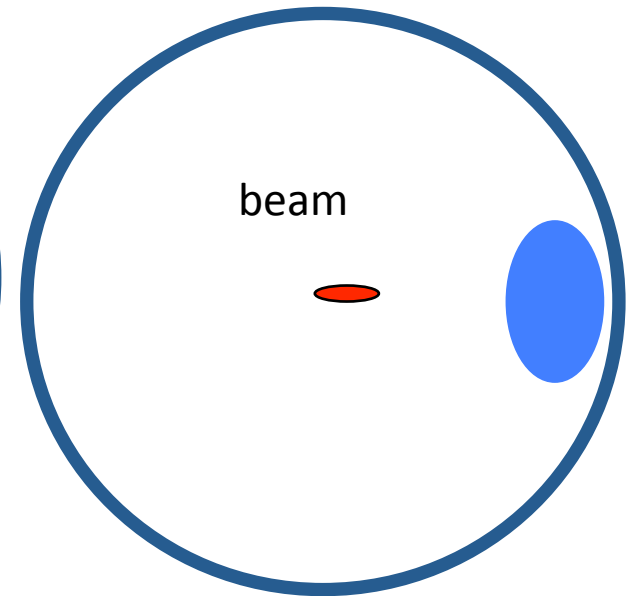


$$\Delta v_y \sim \Delta v_x$$

Magnetic field section

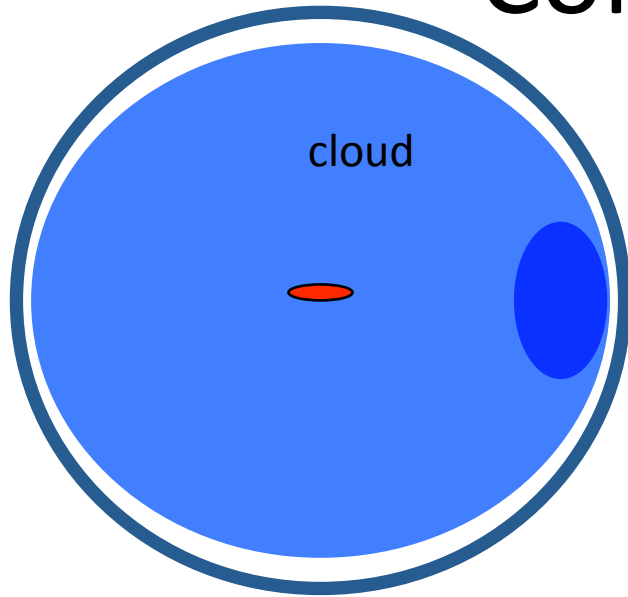


$$\Delta v_y > 0 \quad \Delta v_x = 0$$

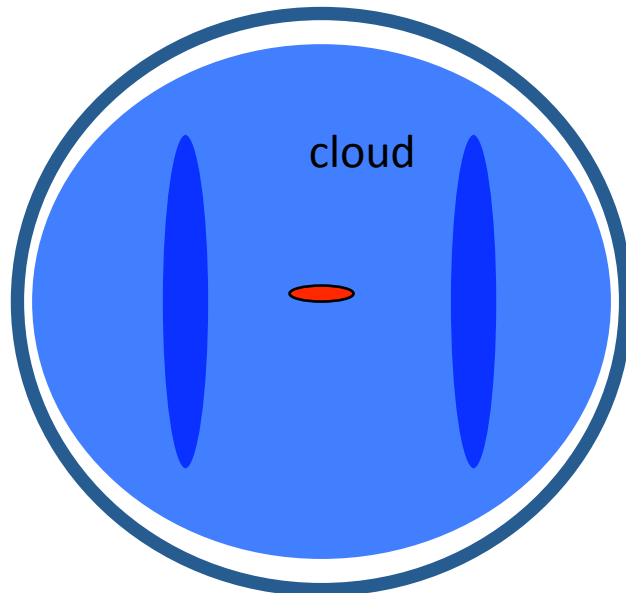


$$\Delta v_y \sim -\Delta v_x > 0$$

Combination



- $\Delta v_y > 0$ $\Delta v_x \sim 0$ can be realized, if Δv_x is cancelled in two distributions.



Number of produced electrons

Number of photon emitted by a positron per unit bending angle.

$$\frac{dY_{pe}}{d\theta} = \frac{5}{2\sqrt{3}} \alpha \gamma \times 0.1 (\text{/rad}) \quad \text{Quantum eff.}=0.1$$

- ◆ CCSR 5GeV $\gamma=10000 \rightarrow Y_{pe}=0.086/\text{m}, E_c=3 \text{ keV}$
- ◆ Ccsr-TA 2GeV (arc) $=4000 \rightarrow Y_{pe}=0.034/\text{m}, E_c=100 \text{ eV}$
- ◆ KEKB 3.5 GeV $=7000 \rightarrow Y_{pe}=0.015/\text{m},$
- Bunch population
 $N_p=1.2 \times 10^{10}$ (0.75mA) 3.3×10^{10} (KEKB)
- electrons created by a bunch passage in a meter
 $N_p \times Y_{pe}=1.0 \times 10^9$ (5GeV) 4.0×10^8 (2GeV) 4.9×10^8 (KEKB)
- Increase of volume density per bunch ($\Delta\rho[\text{m}^{-3}\text{bunch}^{-1}]$)
 2.0×10^{11} (5GeV) 8.1×10^{10} (2GeV) 6.2×10^{10} (KEKB)
- Tune shift per bunch
 0.00045 (5GeV) 0.00045 (2GeV) 0.00077 (KEKB)
- Beam line density $N_p/4.2=2.9 \times 10^9$ (Ccsr) 1.4×10^{10} (KEKB)

Tune shift at the space charge limit

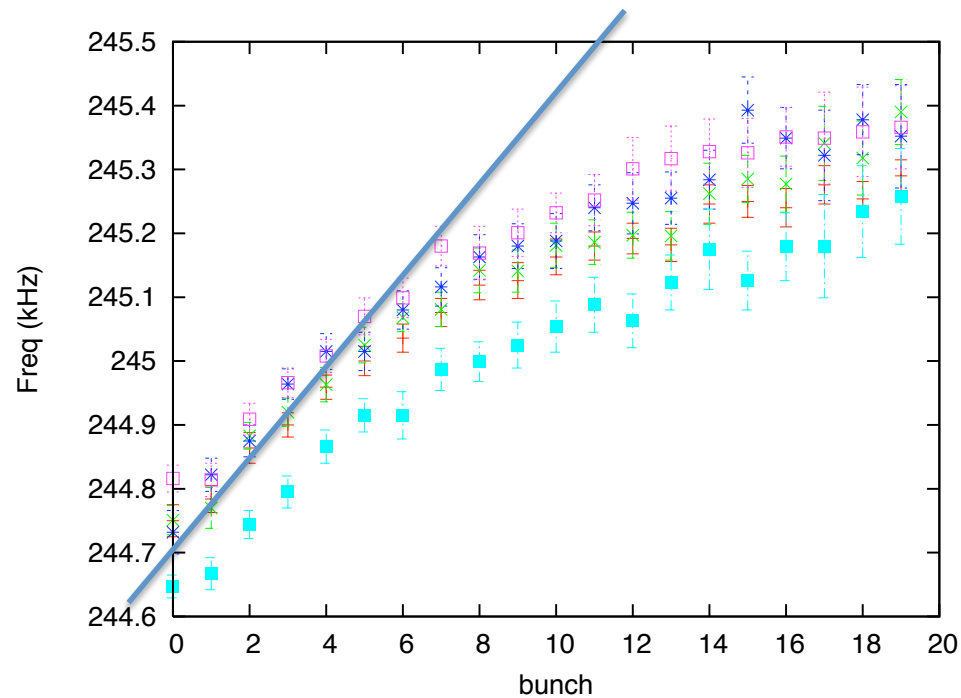
		Cesr 14 ns	Cesr 14 ns	KEKB 8ns
Bunch popu.	N_p	1.2e10	2.0e10	3.3e10
Spacing	L_{sp} (m)	4.2	4.2	2.4
Line density	λ_p (m ⁻¹)	2.9e9	4.8e9	1.4e10
Neutralized density	ρ_e (m ⁻³)	5.7e11	9.5e11	1.7e12
Tune shift	$\Delta\nu$	0.0032	0.0053	0.021

Tune shift at the threshold

	KEKB	KEKB	KEKB-DRt	Cesr chess	CesrTA	ILC-OCS	PEPII
L	3016	3016	3016	768.44	768.44	6695	2200
gamma	6849	6849	4501	10372	3914	9785	6067
Np	3.30E+10	7.60E+10	2.00E+10	1.12E+11	2.00E+10	2.00E+10	8.00E+10
ρ_{eth}	6.25E+11	3.81E+11	1.22E+11	5.73E+12	2.92E+12	1.91E+11	7.67E+11
$\Delta v_{x+y}@th$	0.0078	0.0047	0.0023	0.0120	0.0162	0.0111	0.0078
DampT-xy	40	40	75	22	56.4	26	40
DampR-xy	2.51E-04	2.51E-04	1.34E-04	1.16E-4	4.54E-05	8.58E-04	1.83E-04

Tune shift at Cesr Chess

- 0.75mAx20bunches, 14ns
- $\Delta\nu_y=0.0002/\text{bunch}(=0.07\text{-}0.08\text{kHz}/\text{bunch})$,
saturate 10-20 bunches, $\Delta\nu_y=0.002=0.7\text{kHz}$.



Threshold of cloud density as a function of bunch population (Cesr Chess)

$$\rho_{e,th} = \frac{2\gamma v_s \omega_e \sigma_z / c}{\sqrt{3} K Q r_0 \beta L} = \frac{2\gamma v_s}{\sqrt{3} r_0 \beta L \omega_e \sigma_z / c} = \frac{2\gamma v_s}{\sqrt{3} r_0 \beta L} \sqrt{\frac{2\pi \sigma_x \sigma_y}{N_p r_e \sigma_z}}$$

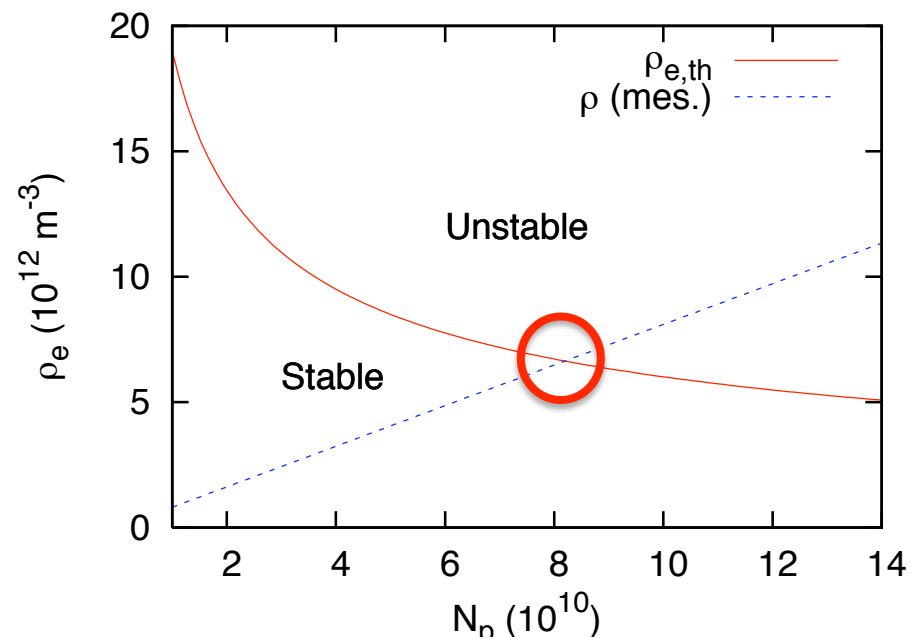
$$\sqrt{N_p} \rho_{e,th} = 1.9 \times 10^{18} \text{ m}^{-3}$$

$$\rho_e = \frac{\gamma}{r_0 \beta L} (\Delta v_x + \Delta v_y)$$

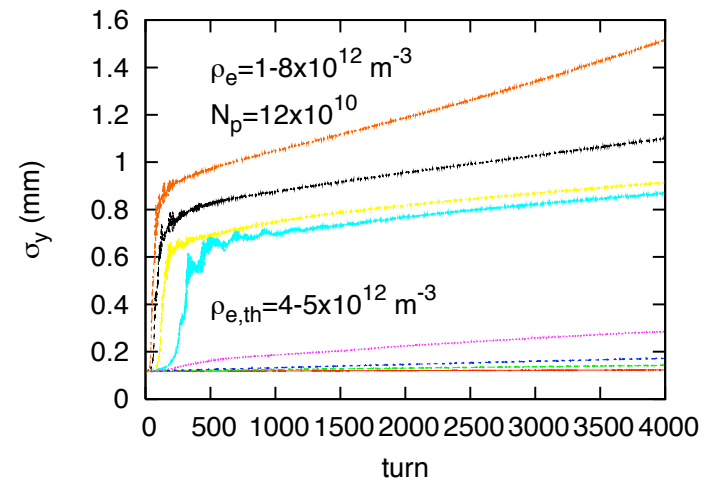
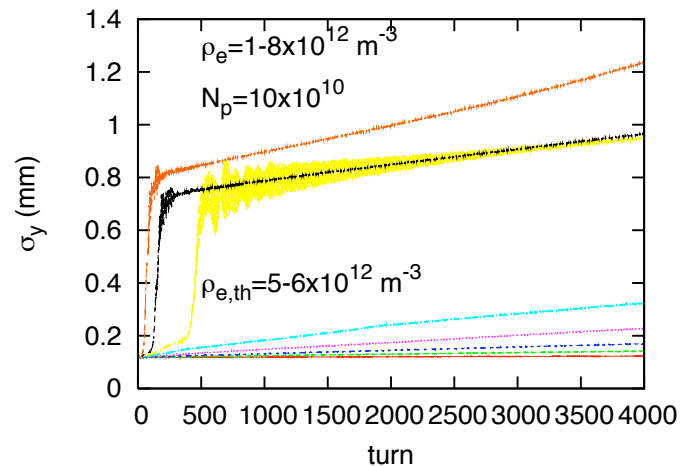
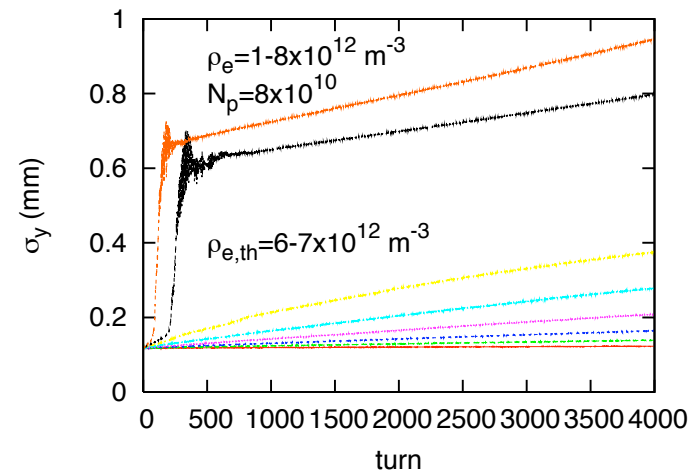
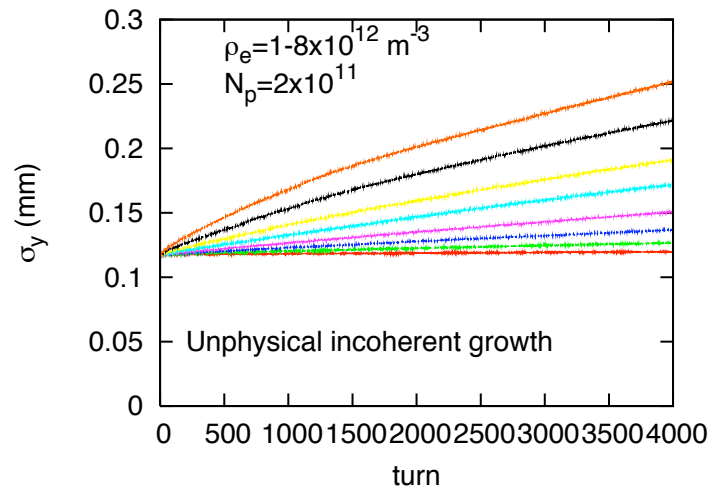
$$\Delta v = 0.0027 I_b \text{ (mA)}$$

$$\rho_e [\text{m}^{-3}] = 81 N_p$$

Fast head-tail instability should be observed at the bunch population more than 8×10^{10} .



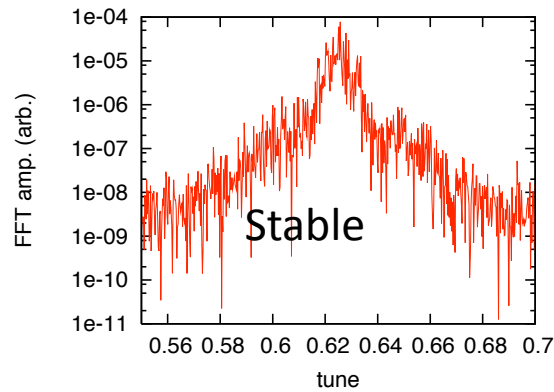
Simulations for instability threshold



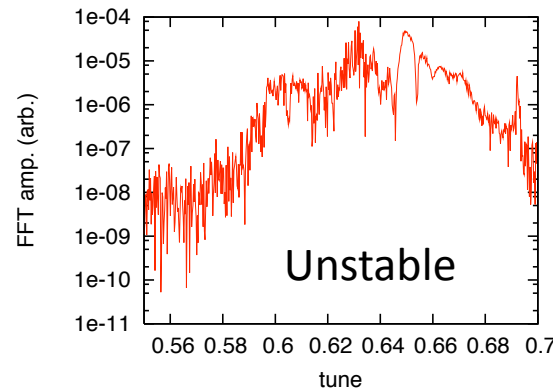
- 8 interactions in one revolution.
- No magnetic field

Unstable mode of the instability

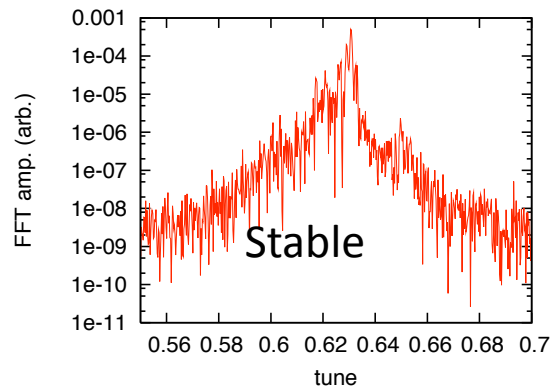
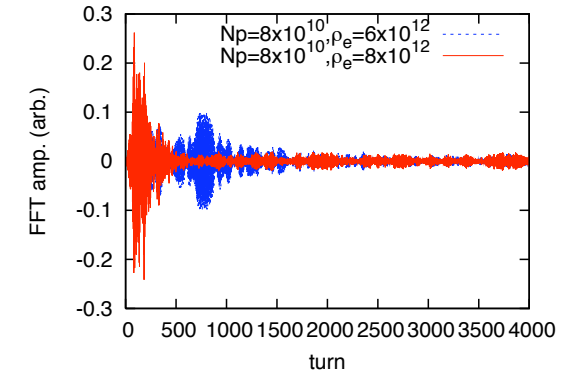
- FFT spectra under and over the threshold
- $v_{y0}=0.6, v_s=0.0486$



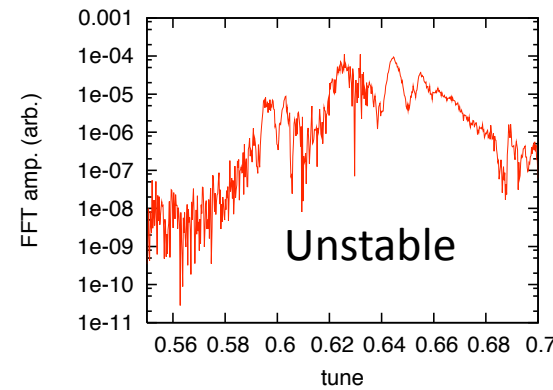
$N_p=6 \times 10^{10}, \rho_e=6 \times 10^{12} \text{ m}^{-3},$



$\rho_e=8 \times 10^{12}$



$N_p=8 \times 10^{10}, \rho_e=6 \times 10^{12} \text{ m}^{-3},$

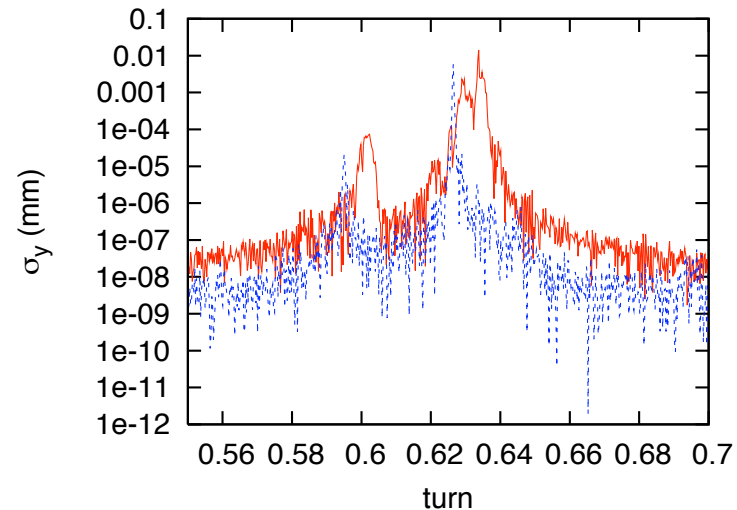
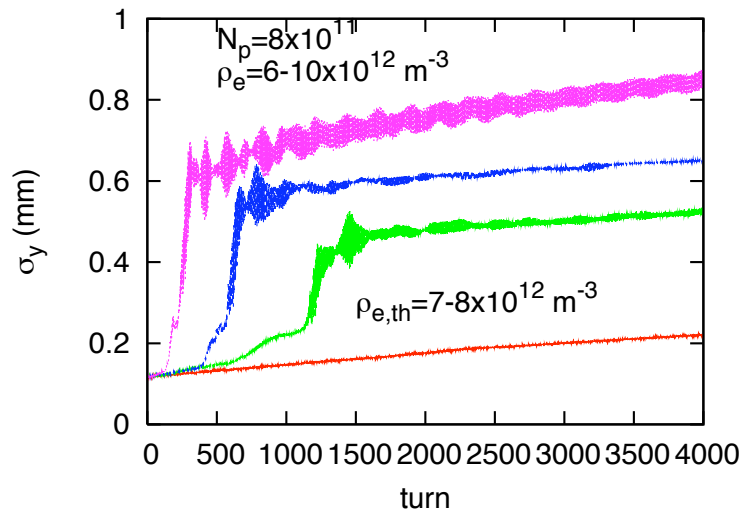


$\rho_e=8 \times 10^{12}$

Maybe signal near $v_\beta + v_s$ is observed.

Simulations in a strong bending field

- When electrons in bending or wiggler magnets are dominant, the threshold somewhat goes higher, because electron pinch is prevented in horizontal.
- Frequency spectrum is somewhat different from that for free electron motion.



Proposal of experiments II

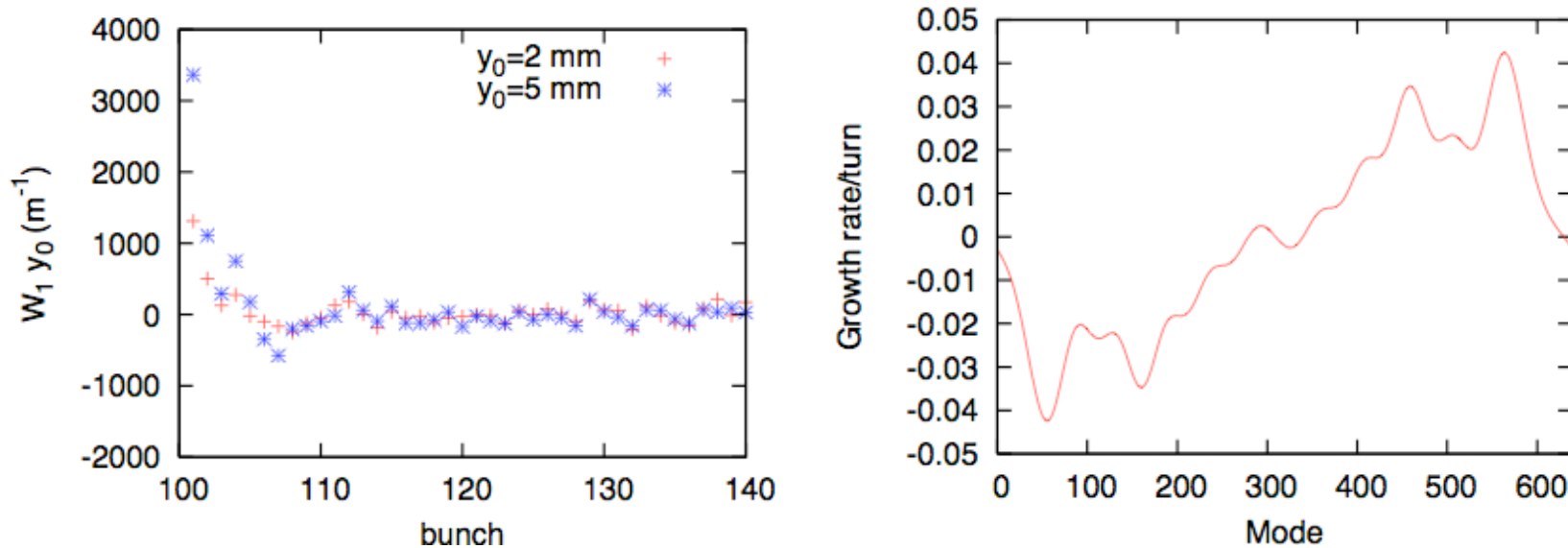
Coherent instability in CEsrTA

Measurement of the fast head-tail instability

- CEsr Chess optics can be used for the present.
- Higher current and longer bunch train. 7mA/bunch, 10 bunch train, 14ns may be satisfied to the instability condition.
- Coupled bunch instability or coherent dipole motion should be controlled in the measurements.
- Measure the bunch by bunch position and beam size turn by turn. Longer bunch train has an advantage for statistics; more bunches are unstable.

Measurement of electron cloud induced Coupled bunch instability

- $N_p=1 \times 10^{10}$, 4 ns spacing uniformly for example. Number of bunch is 640. It is possible to do 14 ns, 90 bunches.
- Cut off the feed back power and measure the positions of all bunches turn by turn.
- Growth time ~ 25 turn, $64 \mu\text{sec}$ for this condition.



This spectrum is given for free electron motion. If bending magnet is dominant, different spectrum is obtained.