

EUTelescope
(for beam tests with Mimosa26 sensors)
final status

Igor Rubinskiy

+

Joerg Behr, Antonio Bulgehroni, Ingrid Gregor,
Tatjana Klimkovich, Slava Libov,
Philipp Roloff, Filip Zarnecki,
and the Telescope users (APIX PPS+3D, ALFA, DEPFET)

EUDET-JRA1 annual meeting
DESY/Hamburg

29 September 2010

Contents

- EUTelescope releases
 - Documentation
 - Installation
- (EU)Telescope users in 2010
- EUTelescope Framework
 - Telescope sensors (Mimosa 26)
 - Easy integration of a new DUT into the framework and DUT analysis
- Summary

- EUTelescope releases
 - [Pro] Version v00-04-01
 - 1 year ago - [Old] v00-02-00, a.k.a. Better User Integration (BUI)
 - » Python based submission scripts introduced
 - 8 intermediate releases in this year
 - The EUTelescope analysis framework
 - » mostly final,
 - » but there are still things to add
 - significant performance improvements
 - » CPU
 - » Memory
 - » human intervention reduced to minimum (almost none)
 - documentation is kept up-to-date with every release
 - » How to run the EUTelescope step by step with python scripts:
<http://projects.hepforge.org/eudaq/Eutelescope/pythonScripts.html>
 - » It is as easy to run analysis on GRID:
<http://projects.hepforge.org/eudaq/Eutelescope/gridtools.html>

EUTelescope installation

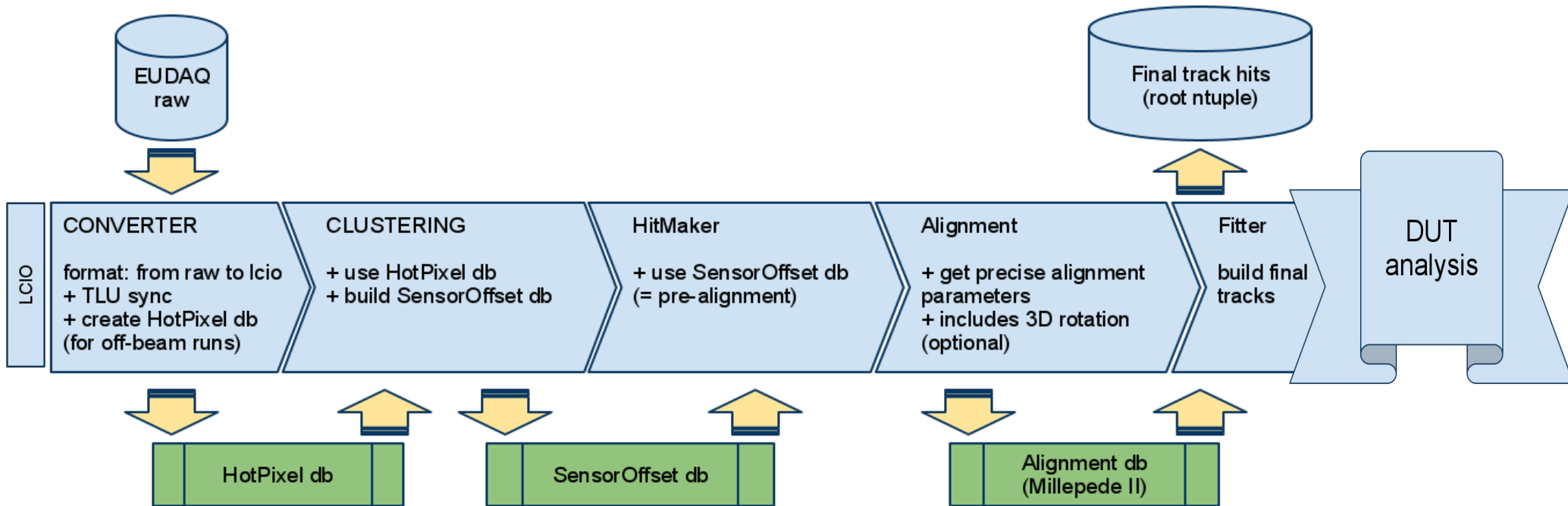
- Step by step (copy-paste style) instructions at <http://projects.hepforge.org/eudaq/Eutelescope/ilcinstall.html>
- installation on SL4/SL5 goes without problem (other OS problematic, e.d. Ubuntu)
- Can be a bit tricky
 - Due to mutual EUDAQ-EUTelescope dependencies
 - Follow strictly the instructions
 - Install EUDAQ
 - » Compile with LCIO=0 and EUTELESCOPE=0 (Makefile flags)
 - Install full ilcsoft
 - » With EUTelescope depending on the EUDAQ library
 - » Recompile EUDAQ with LCIO=1 and EUTELESCOPE=1
 - » Recompile the EUTelescope against the new libeudaq.so
 - Install Millepede II (the latest one from the svn)
 - Would be good to hide this all inside of the ilcinstall scripts (!)
- Before the EUTELESCOPE analysis can be started (every new terminal session) the environment must be loaded
 - %> source \$EUTELESCOPE/build_env.csh
- Now can analyse beam test data
 - Go from pixels in RAW (EUDAQ format) to track parameters in LCIO (or ROOT)
 - Analyse the DUT features

Telescope users in 2010

user	data, GB	# runs	# DUTs	location	# events
FORTIS+TPAC	936	1523	1	DESY	~90 mln
TIMEPIX (INGRIDs)	-	-	-	"	-
APIX (3D)/RD42(SPIDER)	534	942	1/1'	CERN	~60 mln
NA62	15	288	0	"	~15 mln
APIX (Diamond)	20	221	1-2'	"	~10 mln
APIX (PPS)	85	908	8	"	~30 mln
ALFA	98	532	0	"	~98 mln
SPIDER	8	72	-	"	~7 mln
SILC
DEPFET

Total: ~300 mln
 Analysis speed:(M26x6) ~10-50 ms/ev
 (raw->Tracks ntuple) [prev. ~1 s/evt]

EUTelescope data flow concept (for Telescope with Mimosa 26)



- All steps are implemented in Python scripts with a set of xml templates
- The transition from a local 'one run test job' to mass production of hundreds of runs on grid is rather easy.

PyConverter

(python script processing of the processors in the template converter-tmp.xml)

- Converts data from RAW to LCIO
 - Converter is closely related to the EUDAQ library
 - All data formats are defined in EUDAQ
 - Data format conversion (raw -> lcio) is done via EUDAQ Plugins
- HotPixelKiller [to be run only on Off-Beam runs!]
 - **Define a “hot” pixel as firing more frequent then 1% of time** per run without beam, it looks like 10K events is enough. Default numbers: 1% and 10K events can be changed via steering template.
 - **Dump the HotPixel Collection into a DB file** with structure identical to a normal data run
 - The HotPixel Collection can be **loaded by the Clustering processor** and the **hot pixels skipped** during clustering
 - It's a good idea to take medium size runs (100K) without beam once in a while to see how the hot pixel distribution changes (if at all).

PyConverter

(python script processing of the processors in the template converter-tmp.xml)

```
<execute>
  <processor name="UniversalNativeReader"/>
  <processor name="Mimosa26EUTelAutoPedestalNoiseProcessor"/>
  <processor name="HotPixelKiller"/>
</execute>

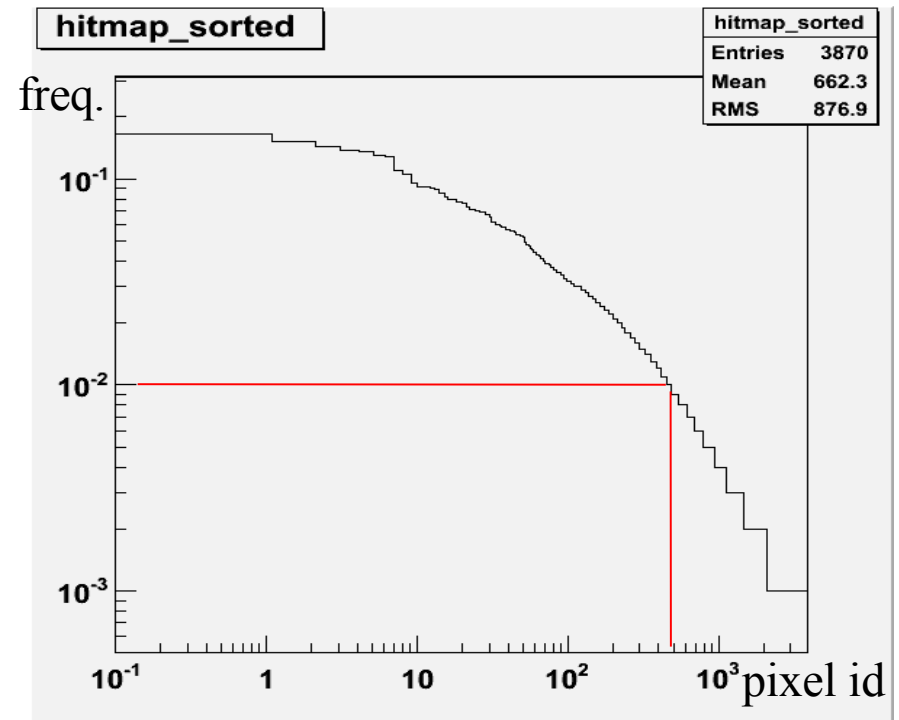
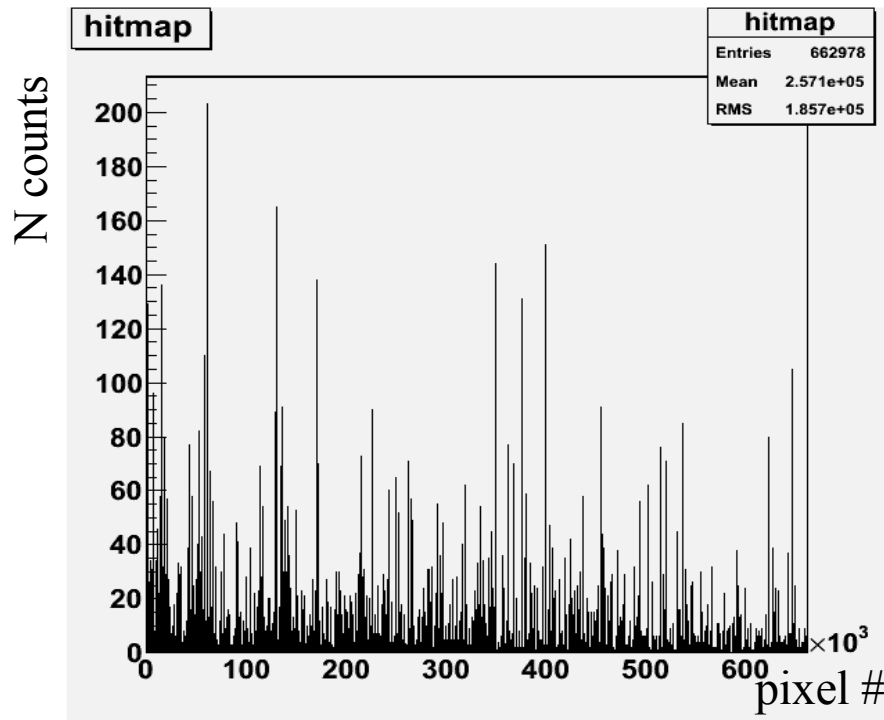
<processor name="UniversalNativeReader" type="EUTelNativeReader">
  <!--Resynchronize the events based on the TLU trigger ID-->
  <parameter name="SyncTriggerID" type="bool" value="false"/>
</processor>

<processor name="HotPixelKiller" type="EUTelHotPixelKiller">
  <parameter name="MaxAllowedFiringFreq" type="float" value="0.01"/>
  <parameter name="NoOfEventPerCycle" type="int" value="10000"/>
  <parameter name="TotalNoOfCycle" type="int" value="0"/>
</processor>
```

new

new

PyConverter :: Hot Pixel Killer



- a run without beam
- ordered pixels
 - 3870 total, spikes are the “hot pixels”
 - define “hot pixels” as firing more often than 1% (600 pixels, 16% of all pixels)

PyClustering

(python script processing of the processors in the template clusearch-tmp.xml)

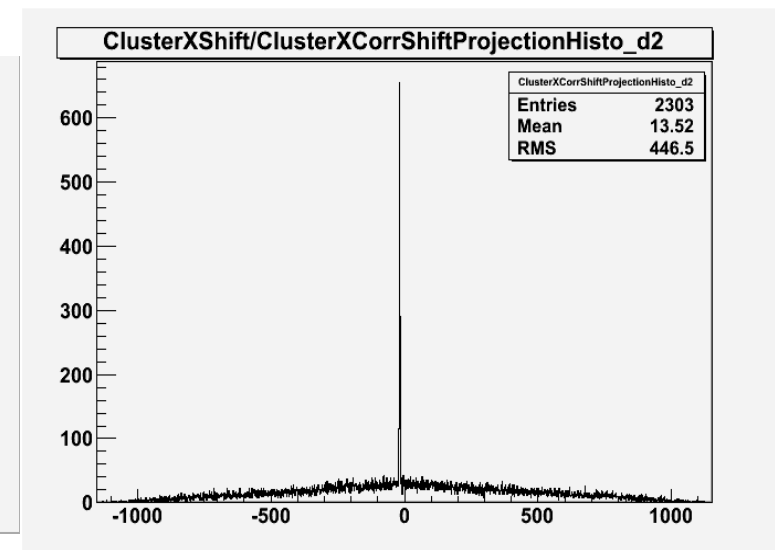
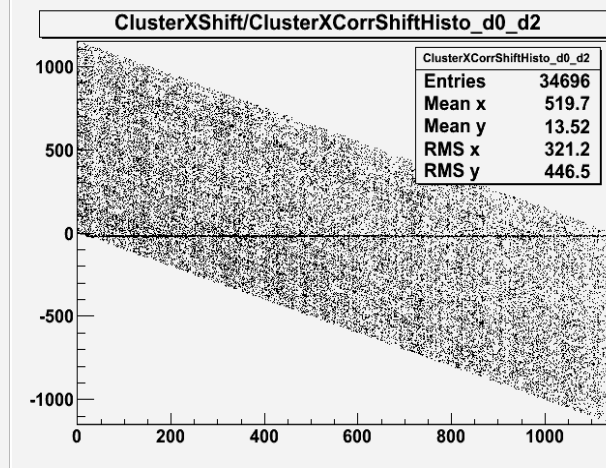
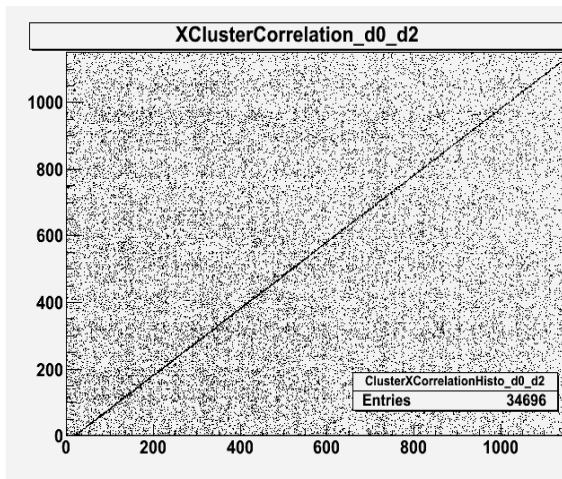
```
<execute>
  <processor name="AIDA"/>
  <processor name="Mimosa26EUTelAutoPedestalNoiseProcessor"/>
  new | <processor name="LoadHotPixelDB"/>
  <processor name="Clustering"/>
  <processor name="Correlator"/>
  <processor name="Save"/>
</execute>

</processor>
  <parameter name="ZSClusteringAlgo" type="string" value="DFixedFrame"/>
  <!--parameter name="ZSClusteringAlgo" type="string" value="SparseCluster2"/-->
</processor>
```

- Optimised clustering algorithms:
 - Digital Fixed Frame (DFF) – a version of the Zero Suppressed Fixed Frame (FF) [J.Behr]
 - Looking for clusters NxM size
 - Sparse Clustering 2 (SP2) - a version of the Sparse Clustering (SP) [A.Bulgheroni]
 - Follows all pixels neighbour-by-neighbour (can be complex structures)
 - In both cases
 - The hot pixel collection is read only once in the beginning of the run
 - If in a run a pixel is found which also exists in the HotPixel DB file, it's being skipped
 - » for hotpixel frequency 0.01 the cluster rate goes down by 50% (speedup ~ x2)
 - Some tests on a (beam) run (10602) of the clustering performance (first 10K events):
 - HotPixel freq 0.01:
 - » DFF(3x3): 238K (sensor 0), timing (for all): 31 ms/evt.
 - » SP2: 228K (sensor 0), timing (for all): 18 ms/evt.
 - SP2:
 - » HotPixel freq -0.05: 307/23, -0.10: 334/25, -0.15: 342/26 [K cluster/ ms/evt]
 - DFF(3x3):
 - » HotPixel freq -0.05: 317/36, -0.10: 344/28, -0.15: 352/39 [K cluster/ ms/evt]

PyClustering:: Correlator

- Correlator processor
 - Build 2D correlations between sensors (the first one #0 and all others)
 - sensor0(X).vs.sensor1(X), sensor0(Y).vs.sensor1(Y), etc.
 - If there was beam, one must see a clear diagonal line (“correlation band”)
 - By the shift of the “correlation band” from the real diagonal (from 0:0 to 1156:1156) one can calculate the relative sensor shift -> preAlignment
 - Build 2D “biased” correlation plot
 - sensor0(X).vs.[sensor1(X)-sensor0(X)], etc. [must take into account sensor flip!]
 - In this case the “correlation band” goes horizontal
 - Make a 1D projection and the peak position gives the sensor offset value
 - Dump them into a db file with structure identical to the existing alignment-constants db file
 - Apply this preAlignment constants already at the Hitmaker level



PyHitmaker

(python script processing of the processors in the template hitmaker-tmp.xml)

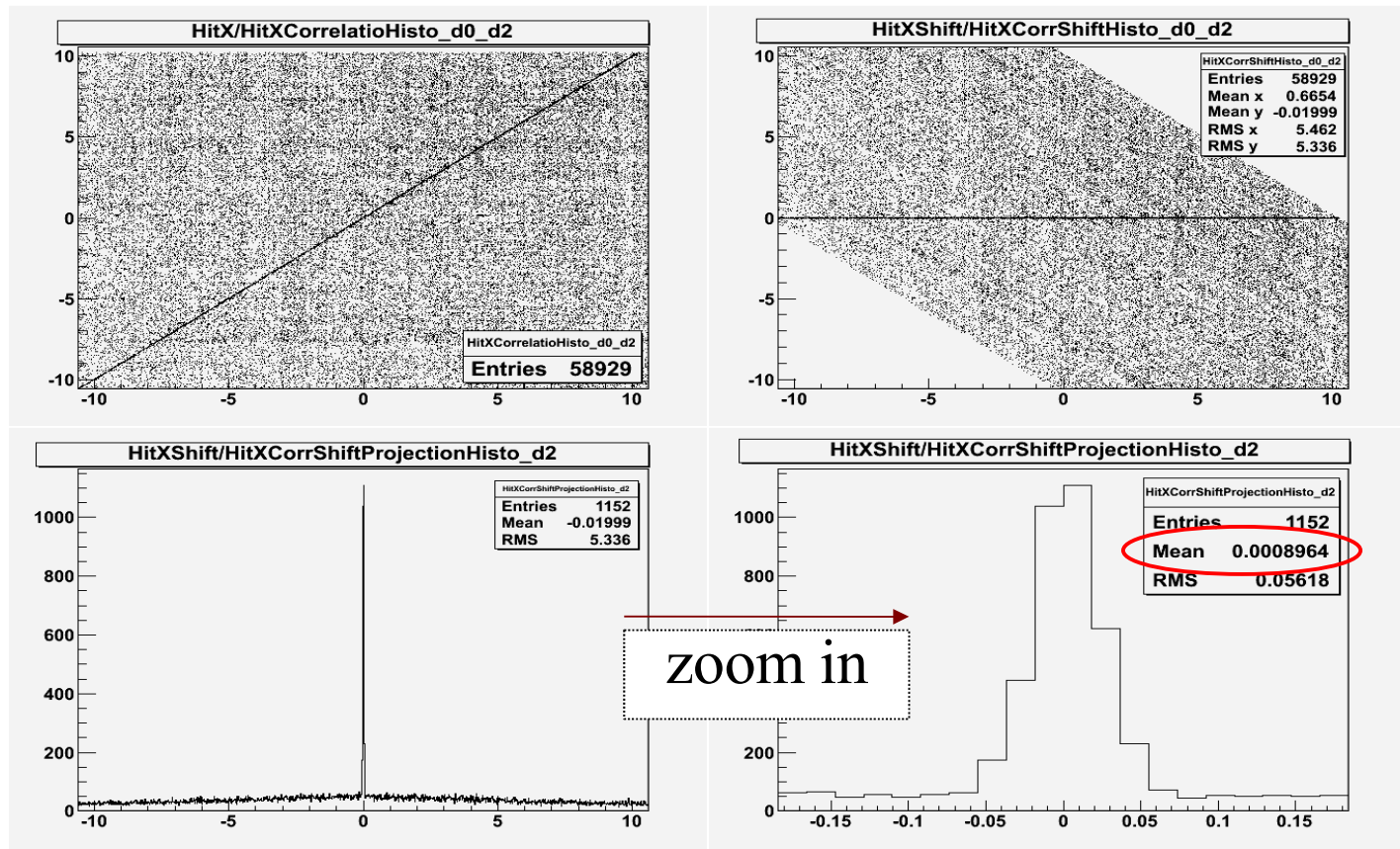
```
<execute>
  <processor name="AIDA"/>
  <processor name="LoadOffsetDB"/>
  <processor name="HitMaker"/>
  <processor name="Correlator"/>
  <processor name="Save"/>
</execute>

<processor name="LoadOffsetDB" type="ConditionsProcessor">
  <parameter name="SimpleFileHandler" type="StringVec"> preAlignment offset-db.slcio preAlignment </parameter>
</processor>

<processor name="HitMaker" type="EUTelHitMaker">
  <parameter name="CoGAlgorithm" type="string" value="FULL"/>
</processor>
```

PyHitmaker :: HitMaker + Correlator

- main improvement – preAlignment!
- without preAlignment (previously) the sensors relative offset could be 0.1-2 mm, which make the real alignment (Millepede II) a bit problematic
- Now all sensors are preAligned at 0, with precision of better then pixel (half) pitch



PyAlignment

(python script processing of the processors in the template align-tmp.xml)

```
<execute>
  <processor name="AIDA"/>
  <processor name="Align"/>
</execute>
```

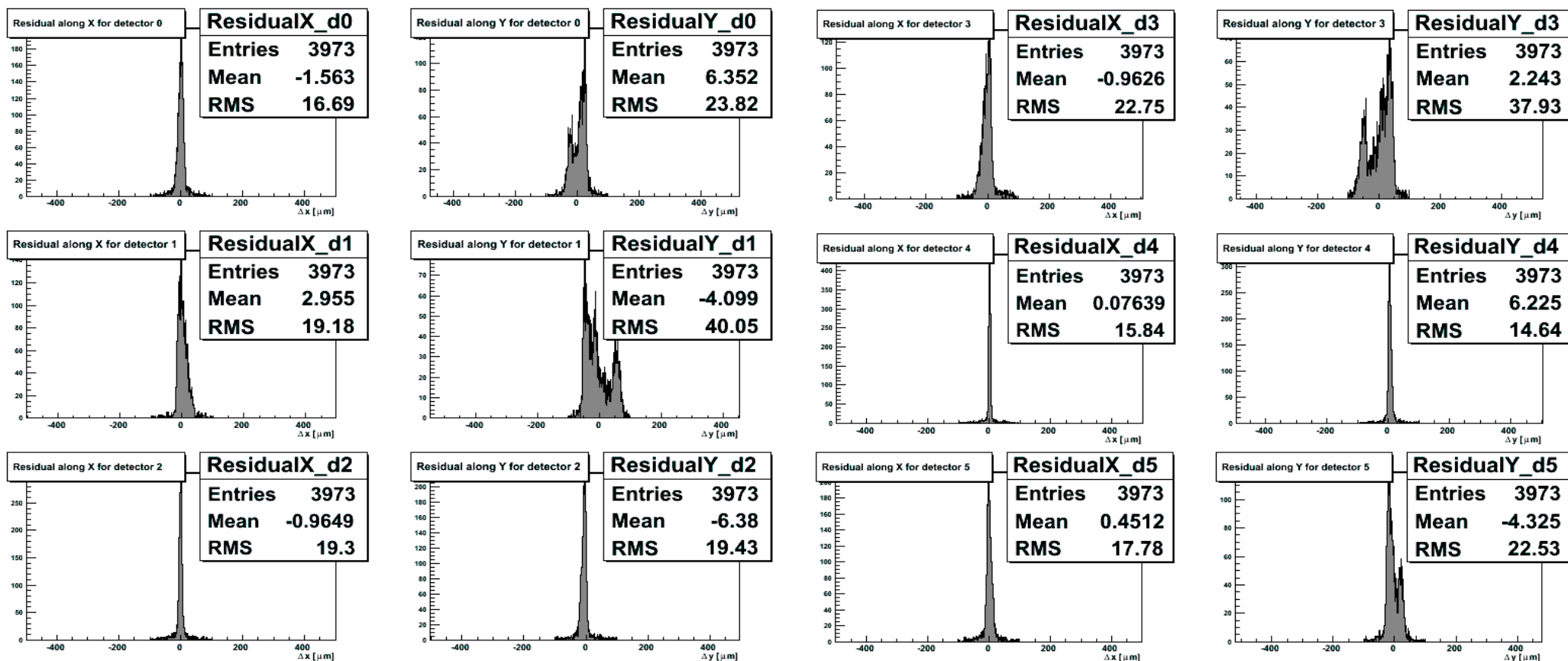
- The sensor X/Y offset values are calculated and applied in the PyHitmaker step, so we can **set new default Residual cut values** for the PyAlign step in the configuration file
- This step required a lot of manual work previously, now the demand of babysitting significantly reduced

[AlignOptions]

ResidualXMin	=	-100	-100	-100	-100	-100	-100
ResidualXMax	=	100	100	100	100	100	100
ResidualYMin	=	-100	-100	-100	-100	-100	-100
ResidualYMax	=	100	100	100	100	100	100
DistanceMax	=	2000					

PyAlignment – residual plots

- Main goal achieved – sensors initially aligned around '0'
- now at the final (fine) alignment (Millepede II) we can do um level alignment
- X direction is fine
- Y direction not always fine – artefact (previously seen in other runs)

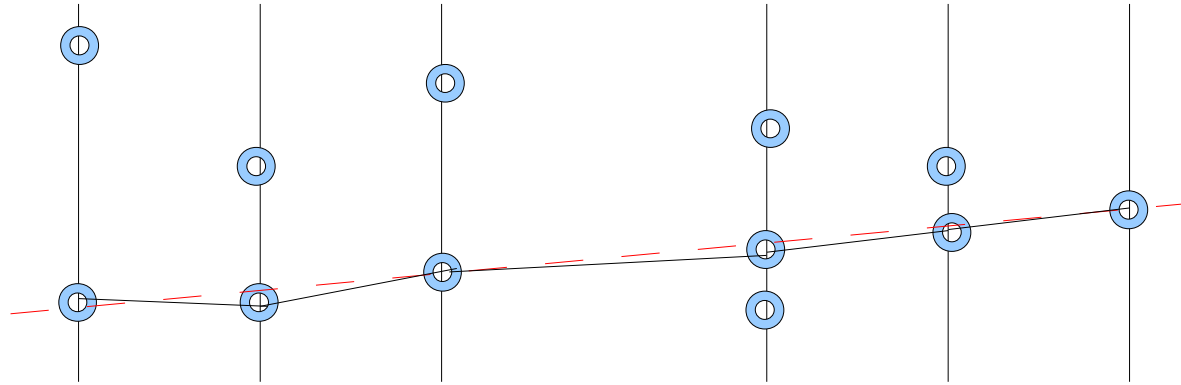


PyFitter

(python script processing of the processors in the template fitter-tmp.xml)

```
<execute>
  <processor name="LoadAlignment"/>
  <processor name="ApplyAlignment"/>
  <processor name="Fitter"/>
</execute>

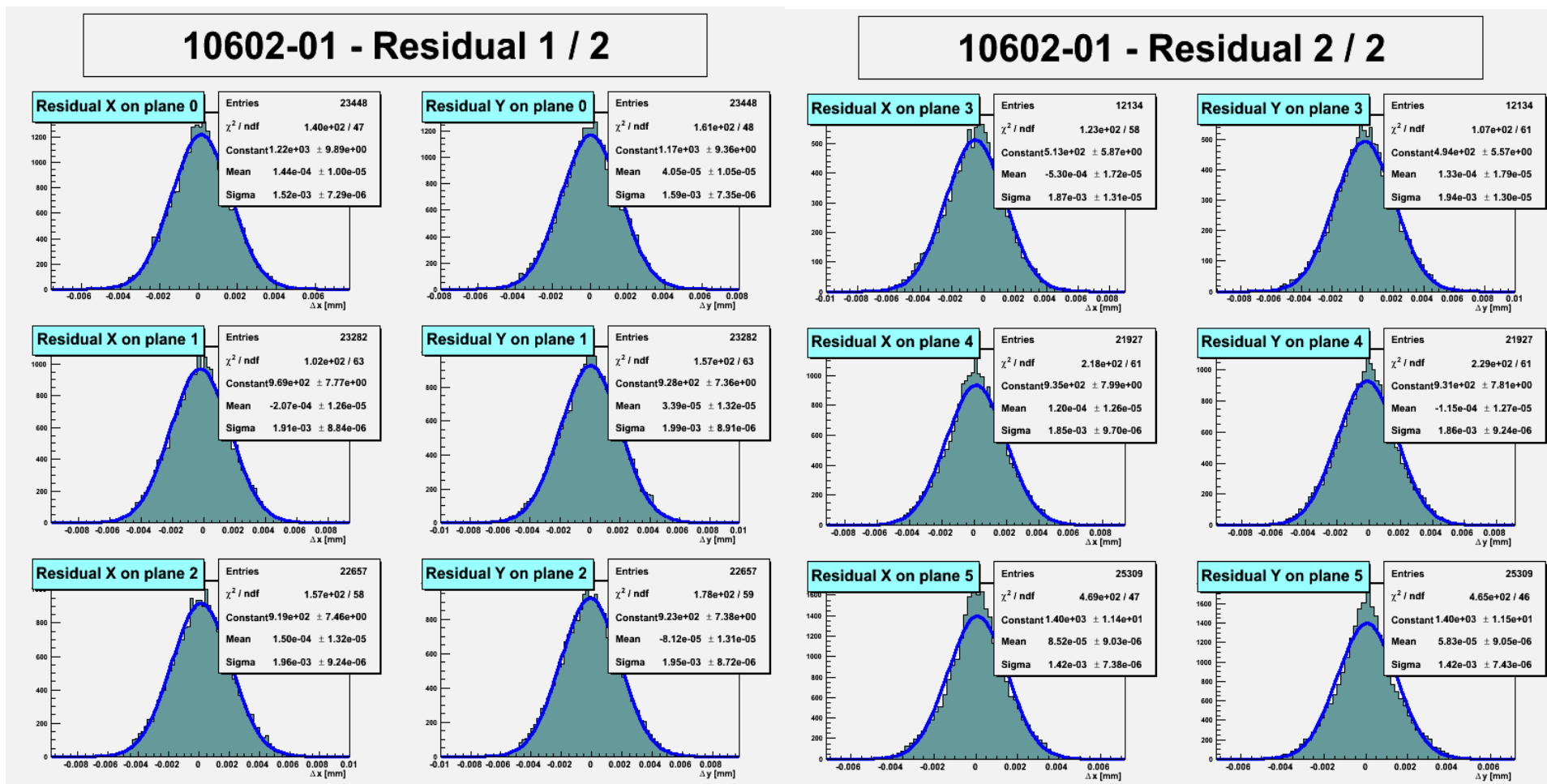
<processor name="Fitter" type="EUTelTestFitter">
  <!--Decide now weather you want to rely on the track candidate slope permanence in X and Y, default=true -->
  <parameter name="UseSlope" type="bool" value="true"/>
  <!--Set the allowed maximum difference of the slope in X (from plane to plane), default = 0.01 -->
  <parameter name="SlopeXLimit" type="float" value="0.0001"/>
  <!--Set the allowed maximum difference of the slope in Y (from plane to plane), default = 0.01 -->
  <parameter name="SlopeYLimit" type="float" value="0.0001"/>
  <!--Maximal allowed (initial) distance between hits in the XY plane between the planes,default = 2. mm -->
  <parameter name="SlopeDistanceMax" type="float" value="@DistanceMax@"/>
</processor>
```



- UseSlope control card - reduces the hit selection combinatorics
- Each subsequent pair of hits is required to have only a very small difference in the slope,
 - default value is 0.01 for the (X and Y) difference of $(x_1 - x_0)/(z_1 - z_0)$
 - Since the slope has a meaning of the track angle it does not depend on the z-distance between the sensors

PyFitter – biased residuals

- Performance improvement – 1-2 orders of magnitude
- Y-Artefacts from the alignment level are almost not seen (track $\chi^2 < 10$)



- The benchmarking was done on run 10602 (July 2010 APIX data)
 - Using run 10494 without beam to prepare the HotPixel DB
 - Define a hot pixel as one with firing freq =1%
 - Only 6 Mimosa 26 sensors are considered (no DUT)
- Presently the slowest step is (still) the Clustering
 - Converter: **1 us/evt** [tot: 0.5 ms/evt, or 8 ms/evt HotPixelDB]
 - Clustering: **18 ms/evt** [tot: 20 ms/evt, excl. Correlator which runs on first 10K]
 - HitMaker: **4ms/evt** [tot: 5 ms/evt, also excl. Correlator]
 - Alignment: **1.3 ms/evt** [tot: 1.3 ms/evt]
 - Fitter: **14 ms/evt** [tot: 19 ms/evt], allowing 2 hits to be skipped (slow)
 - if none of the hits are to be skipped, the Fitter is ~x10 faster.
- Total 12 min per 10K events, but **scales not linearly with # events**

Summary

- The latest improvements to the EUTelescope library are discussed
 - Main focus in the last year developments have been given to performance improvements
 - Lots of effort put into it
 - Net estimation of the performance gain is at the level of at least factor x10 (some tests show even higher gain up to 2 orders of magnitude)
- Any Mimosa 26 sensor can be treated as DUT and visa versa (!)
 - any DUT can be treated as part of the Telescope
 - So full DUT analysis can be fully performed in the EUTelescope
- The EUTelescope library is in a very good shape now
- Still few things to finish:
 - Work on the sensor Geometry description in Gear to allow tilted sensors beam tests
 - 6D alignment with Millepede II have been already implemented
 - very urgent (!)
 - Continue on the B.U.I. (Better User Integration)
 - Can we do the installation fully automatic?

