

Muon Collider Detector Studies in ILCroot

C. Gatto

On behalf of ILCroot/MARS15 simulation group:

N. Mokhov, S. Striganov

and ILCroot simulation group:

V. Di Benedetto, C. Gatto,

F. Ignatov, N. Terentiev

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Main Detector Challenges

- If we can build a Muon Collider, it will be a precision machine!
- One of the most serious technical issues in the design of a Muon Collider experiment is the background (see all previous talks)
- The major source (from muon decays) is overwhelming: for 750 GeV muon beam with $2 \cdot 10^{12}$ muons/bunch expect $\sim 4.3 \cdot 10^5$ decays/m
- Large background is expected in the detector: $\sim 10^8$ particles/bunch
Xing
- The backgrounds and/or adequate shielding can spoil the physics program
- A Muon Collider Physics&Detector program has been established by Fermilab to address such issues and to guide toward the choice of technology and detector parameters optimization.

MARS and ILCroot Frameworks

MARS – the framework for simulation of particle transport and interactions in accelerator, detector and shielding components.

- New release of MARS15 available since February 2011 at Fermilab (N. Mokhov, S. Striganov, see www-ap.fnal.gov/MARS)
- Among new features:
 - Refined MDI (Machine Detector Interface) with a 10° nozzle
 - Significant reduction of particle statistical weight variation
 - Background is provided at the surface of MDI (10° nozzle + walls)

ILCroot - Software architecture based on ROOT, VMC & Aliroot

- All ROOT tools are available (I/O, graphics, PROOF, data structure, etc)
- Extremely large community of ROOT users/developers
- It is a simulation framework and an offline system:
 - Single framework, from generation to reconstruction and analysis!!
 - Six MDC have proven robustness, reliability and portability
 - VMC allows to select G3, G4 or Fluka at run time (no change of user code)
- Widely adopted within HEP community (Opera, CMB, Panda, 4th Concept, LHeC, T1015)
- It is publicly available at FNAL on ILCSIM since 2006

ILCroot: essential add-ons to Aliroot

1. Interface to external files from Event Generators in various format (STDHEP, text, MARS, etc.)
2. Standalone VTX track fitter
3. Pattern recognition from VTX (for Si central trackers)
4. Track fitters for different trackers technologies (Si Pixels, Si Strips, Drift Chambers, Straw Tubes, TPC's) and a ombination of them
5. Full simulation of Dual Readout calorimeters
6. Parametric beam background (# integrated bunch crossing chosen at run time)

Very important for detector and Physics studies of New Projects

The Virtual Montecarlo (VMC) Concept

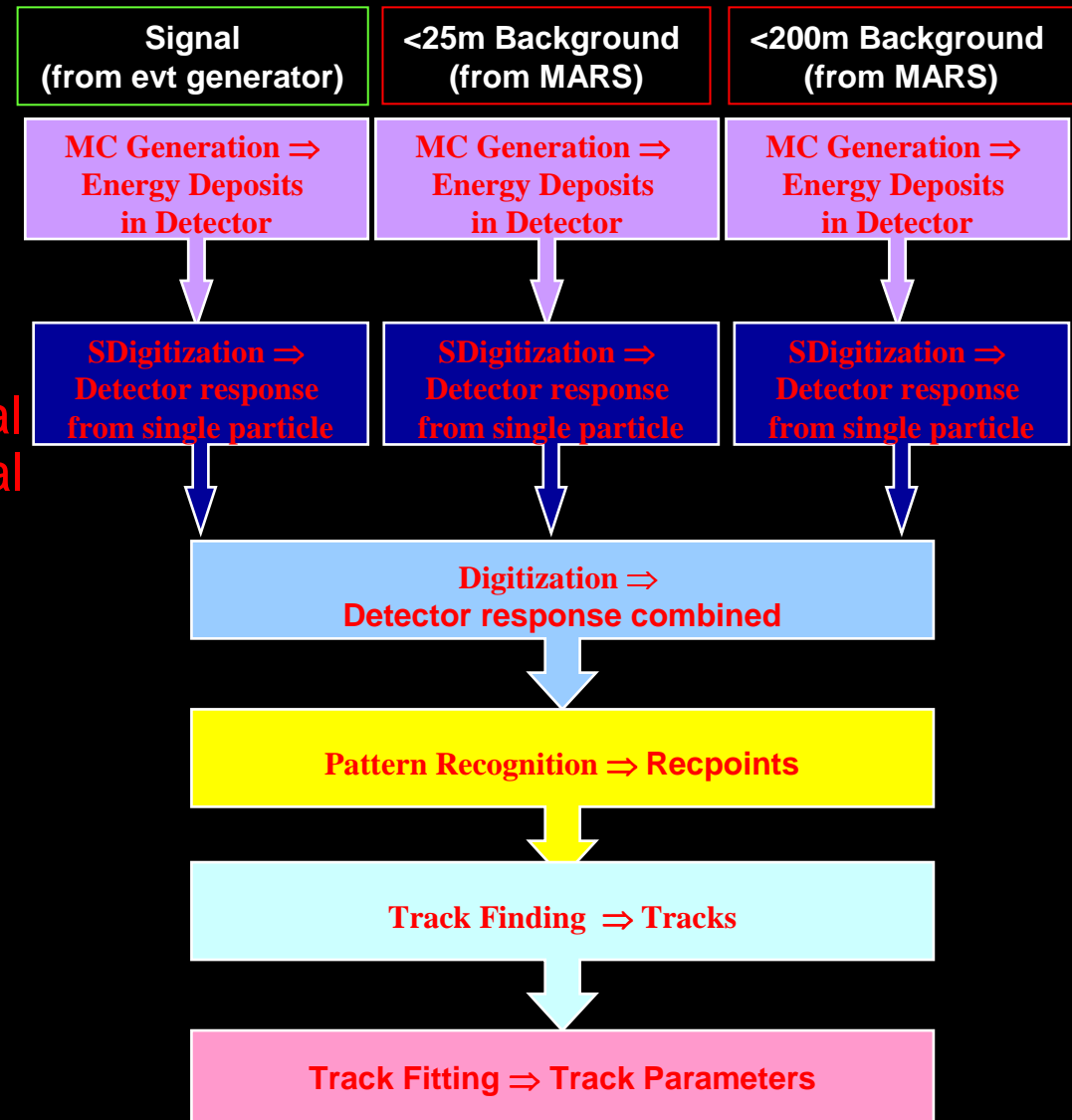
- Virtual MC provides a **virtual interface** to Monte Carlo
- It allows to run the same user application with all supported Montecarlo's
- The real Monte Carlo (**Geant3, Geant4, Fluka**) is selected and loaded at run time

Simulating 1 MARS Event @ $E_{cm} = 1.5 \text{ TeV}$ with 2×10^{12} particles

- About 1×10^8 particles, almost all originating $< 25\text{m}$ from IP
 - Muon beam decay from beam line components and accelerator tunnel is major source of background in the detector
 - 4.3×10^5 decays/m/bunch Xing.
 - Incoherent 3×10^4 e^+e^- pair production per bunch Xing
- Background is split into two sources:
 - Near ($< 25\text{m}$)
 - Far ($25\text{m} < < 200\text{m}$)
- Particle at MDI interface
 - $w \sim 20$
- Particles from beamline
 - $W \ll 1$: need proper statistical treatment
- At large radii also:
 - Bethe-Heitler and beam-halo induced background
- Background with current shielding configuration is reduced by $\mathcal{O}(10^3)$

Processing Flow of Full Simulation: detector hits + digitization + reconstruction

- Hits: produced by MC (G3,G4,Fluka)
- SDigits: simulate detector response for each hit
- Digits: merge digit from several files of SDigits (example Signal + Beam Bkgnd)
- Recpoints: Clusterize nearby Digits
- Pattern recognition + track fit through full Parallel Kalman Filter
- Or Calorimetry shower reco + jet-finder



12 Detectors in ILCroot + 12 from Alice

Detector	Layouts	Digit./Cluster.
VXD (SIDMay06)	1 (parametric)	Full
FTD (SiLC)	1	Full
TPC (Hybrid readout)	1	Gauss. Smear.
Si-Tracker (SID01-Polyhedra)	1+1	Full
μ Collider/CLIC Tracker	1	Full
Hadron Calorimeter	2	Full
THACAL	1	Fast
ADRIANO Calorimeter	1	Full
EM Calorimeter	2	Full
Muon Spectrometer (straw tubes)	1	Gauss. Smear.

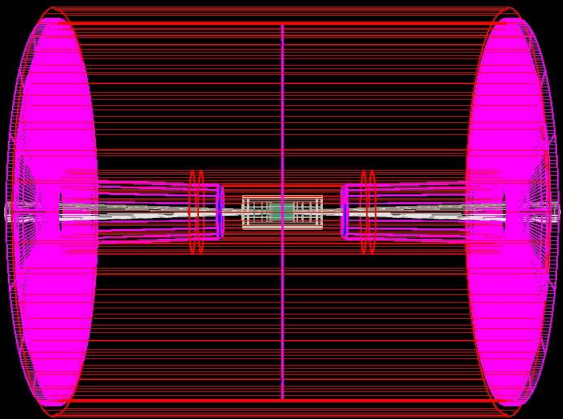
NEW

MARS + ILCroot (Dedicated ILCroot framework for MUX Physics and background studies)

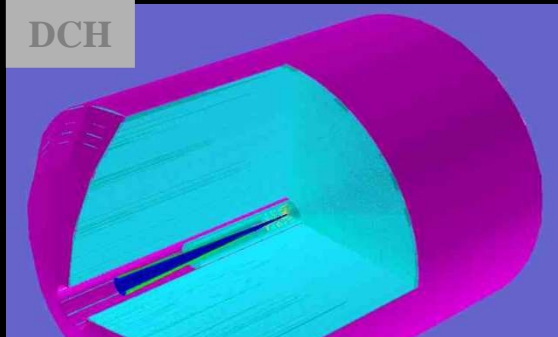
- **The ingredients:**
 - MDI description in MARS & ILCroot
 - Detector description in ILCroot
 - MARS-to-ILCroot interface (**Vito Di Benedetto**)
- **How it works**
 - The interface (**ILCGenReaderMARS**) is a *TGenerator* in ILCroot
 - MARS output is used as a config file
 - **ILCGenReaderMARS** creates a STDHEP file with a list of particles entering the detector area at $z = 7.5\text{m}$
- **ILCGenReaderMARS feeds the Montecarlo with:**
 - 1 particle with corresponding weight
 - OK for calorimetric studies
 - W particles smeared according to their origin
 - OK for detailed tracking occupancy studies
 - ...but very slow and time consuming
 - A mix of the above
 - OK for most tracking studies

Detectors in ILCr

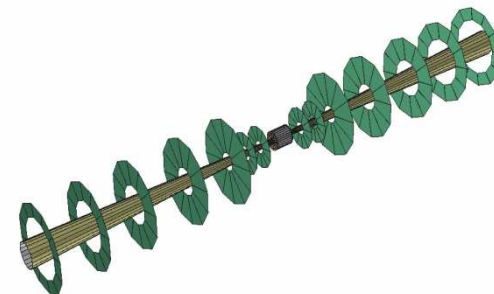
TPC



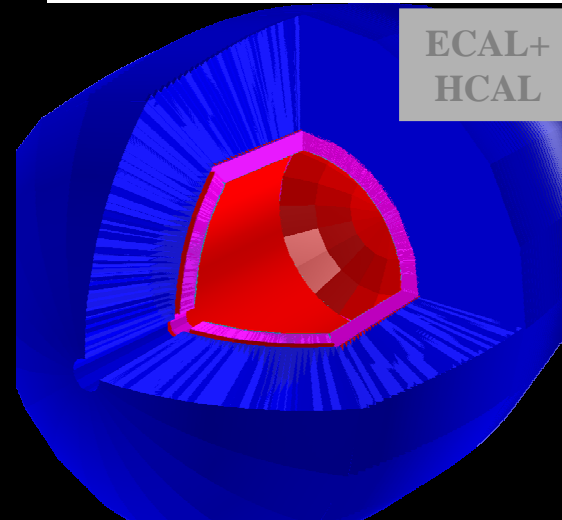
DCH



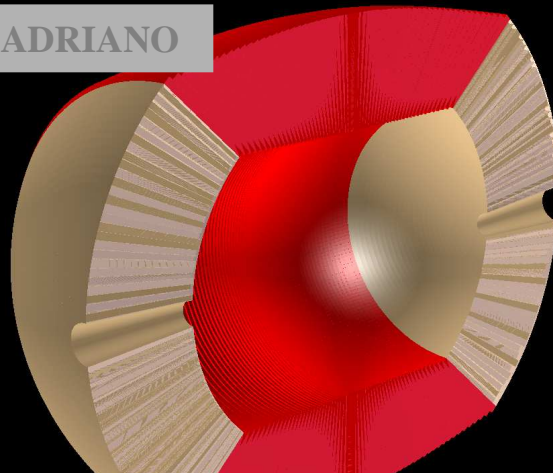
FTD



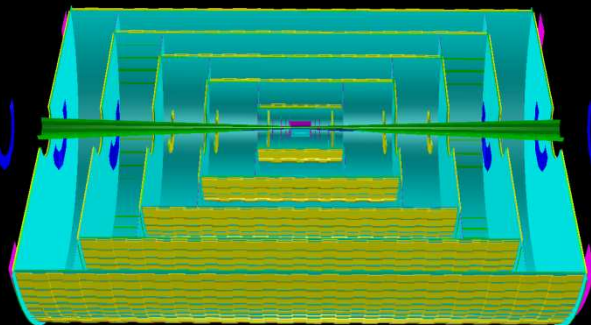
ECAL+
HCAL



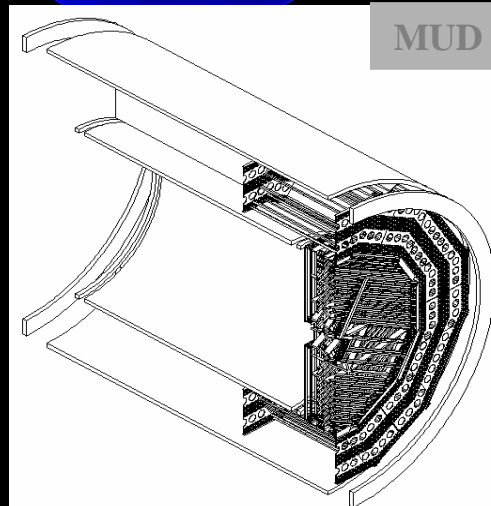
ADRIANO



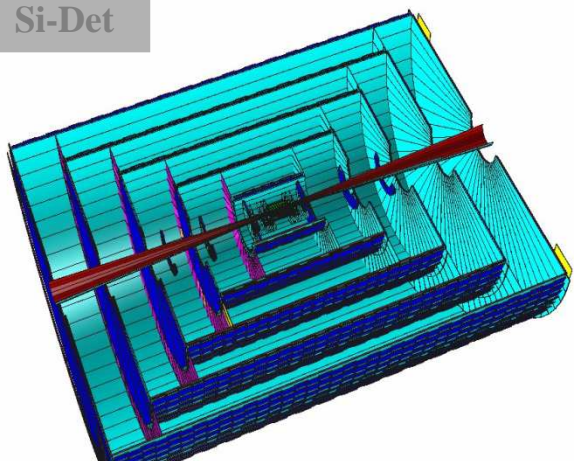
μ C/CLIC



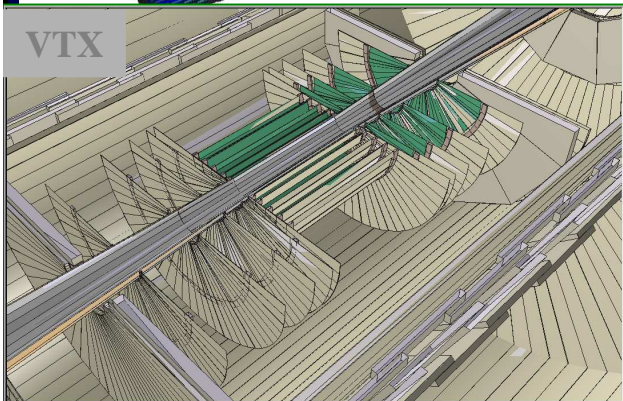
MUD



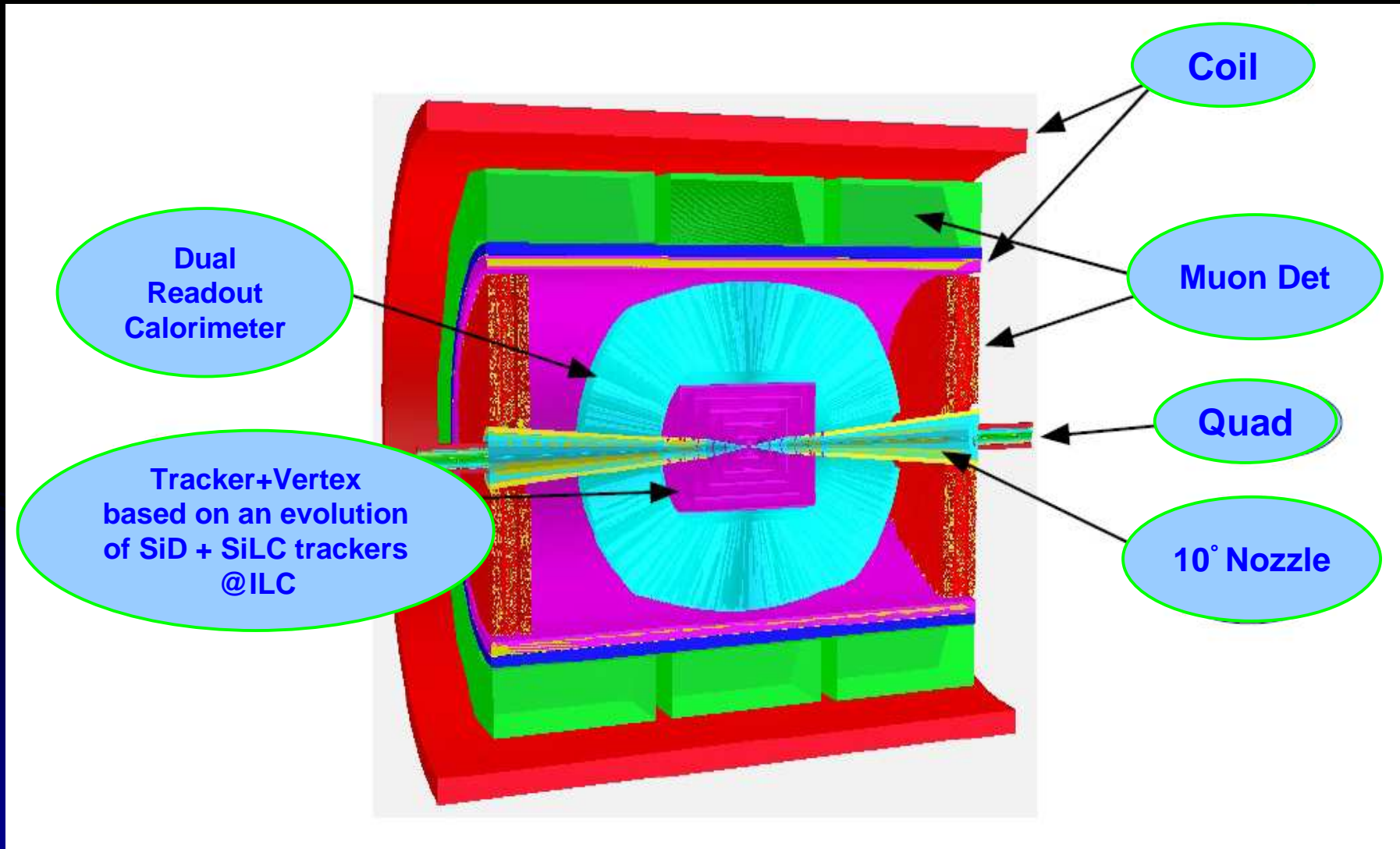
Si-Det



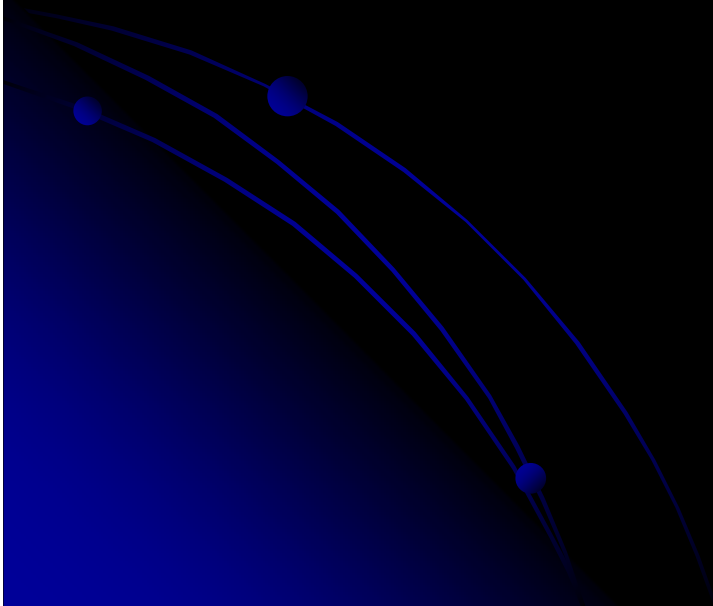
VTX



Baseline Detector for Muon Collider Studies



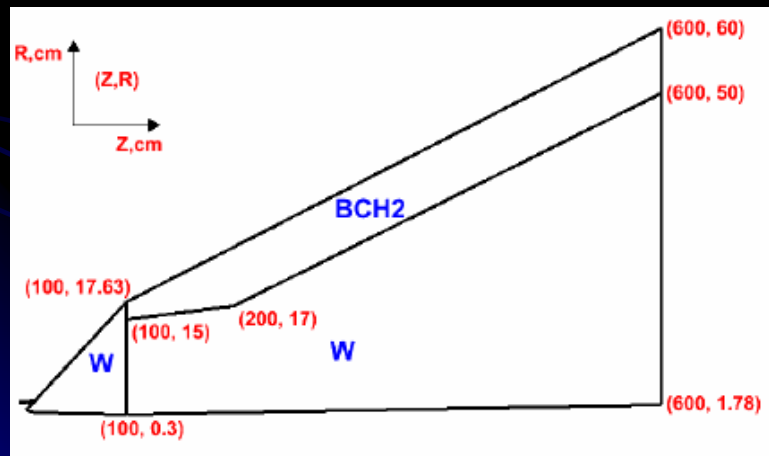
Tracking Studies



Vertex Detector (VXD) 10° Nozzle and Beam Pipe

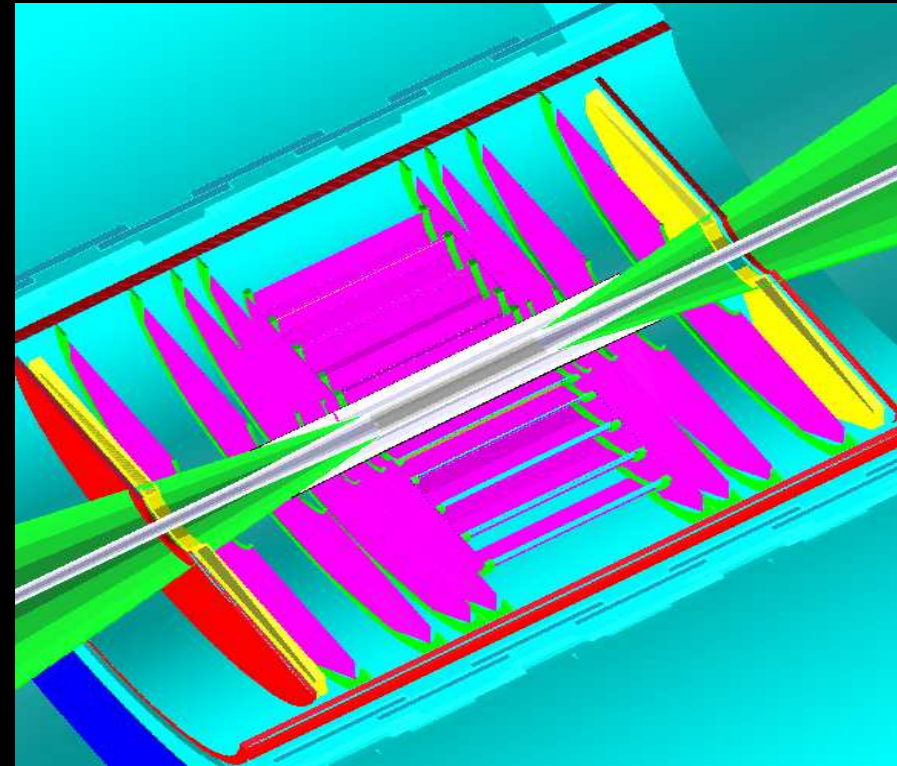
VXD

- Modified SiD design
- 100 μm thick Si layers
- 20 μm x 20 μm Si pixel
- Barrel : 5 layers subdivided in 12-30 ladders
- $R_{\text{min}} \sim 3 \text{ cm}$ $R_{\text{max}} \sim 13 \text{ cm}$ $L \sim 13 \text{ cm}$
- Endcap : 4 + 4 disks subdivided in 12 ladders
- Total length 42 cm



NOZZLE

- Tungsten core
- Borated Polyethylene (BCH2) jacket



PIPE

- Be – Beryllium 400 μm thick
- 12 cm between the nozzles

Silicon Tracker (SiT) and Forward Tracker Detector (FTD)

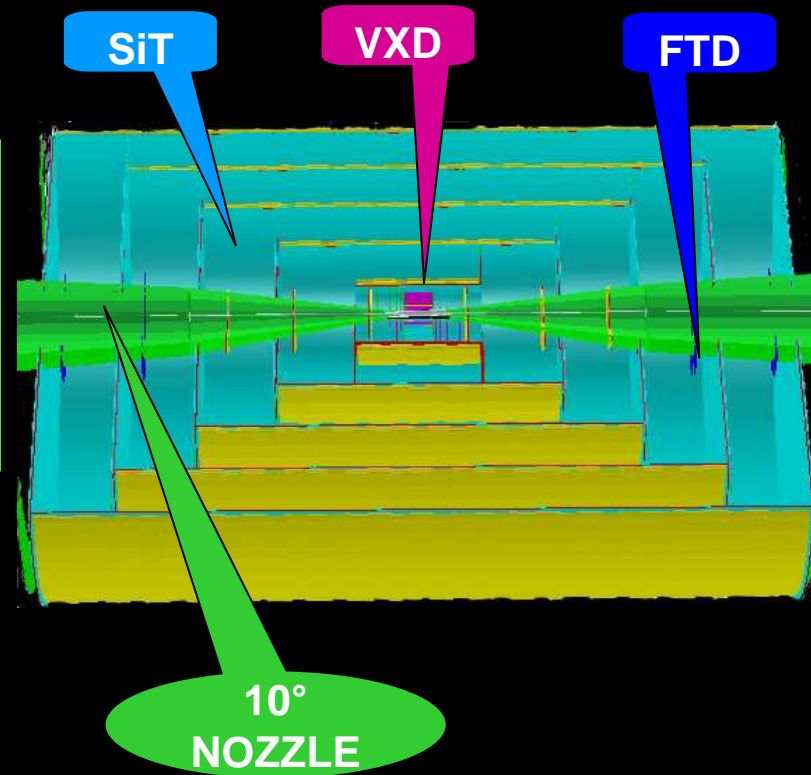
SiT

- original SiD design
- 100 μm thick Si layers
- 50 μm x 50 μm Si pixel (or Si strips or double Si strips available)
- Barrel : 5 layers subdivided in staggered ladders
- Endcap : (4+3) + (4+3) disks subdivided in ladders
- $R_{\text{min}} \sim 20 \text{ cm}$ $R_{\text{max}} \sim 120 \text{ cm}$ $L \sim 330 \text{ cm}$

FTD

- from previous collaboration with SiLC/IFIC
- 20 μm x 20 μm Si pixel
- Endcap : 3 + 3 disks
- Distance of last disk from IP = 190 cm

- Silicon pixel for precision tracking amid up to 10^5 hits
- Tungsten nozzle to suppress the background

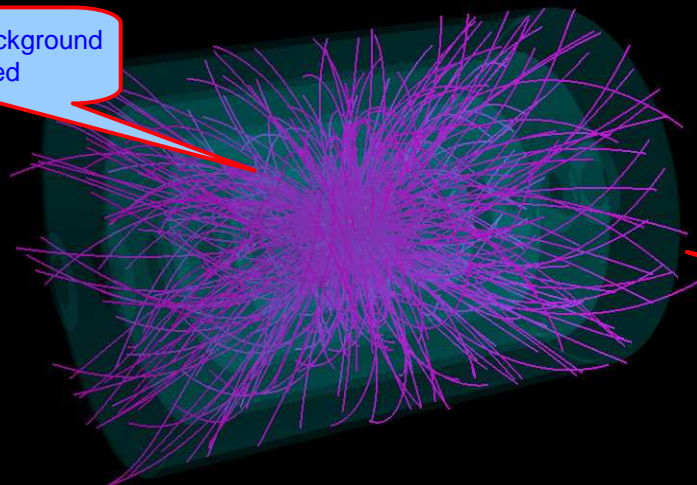


Ingredients for Tracking & Calorimetry

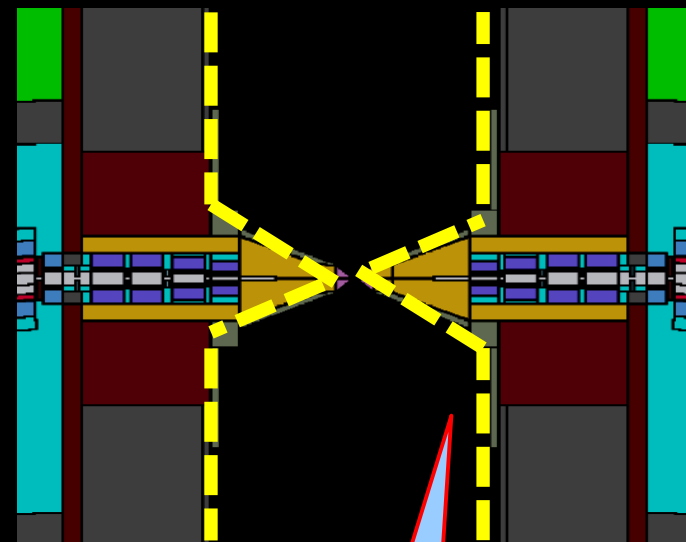
Studies in ILCroot

- MARS background provided at the surface of MDI (10° nozzle + walls)
- GEANT4 /fluka simulated particles in the detector (background + single muons from the I.P.)

Only 475 background pictured



Reconstructed tracks



MDI surface

- Full digitization of signal and background in VTX, Central tracker, F and ADRIANO
- Fast simulation for TAHCAL
- Reconstruct tracks with a parallel Kalman Filter in a 3.5 T B-field
- Reconstruct jets with A. Mazzacane recursive jet finder

Full Simulation of Si Detectors

VXD SDigitization

- ❑ Follow the track in steps of 1 μm
- ❑ convert the energy deposited into charge
- ❑ spreads the charge asymmetrically (B-field) across several pixels:

$$f(x, z) = \text{Errf}(x_{\text{step}}, z_{\text{step}}, \sigma_x, \sigma_z)$$

$$\sigma_x = \sqrt{T \cdot k / e \cdot \Delta l / \Delta V \cdot \text{step}}$$

$$\Delta l = \text{Si tickness}, \quad \Delta V = \text{bias voltage}, \quad \sigma_x = \sigma_x \cdot \text{fda}$$

- ❑ Add couplig effect between nearby pixels
- ❑ Parameters used:
 - Eccentricity = 0.85 (fda)
 - Bias voltage = 18 V
 - cr = 0% (coupling probability for row)
 - cc = 4.7% (coupling probability for column)
 - T° = 300 K

Charge pile-up is automatically taken into account

VXD Digitization

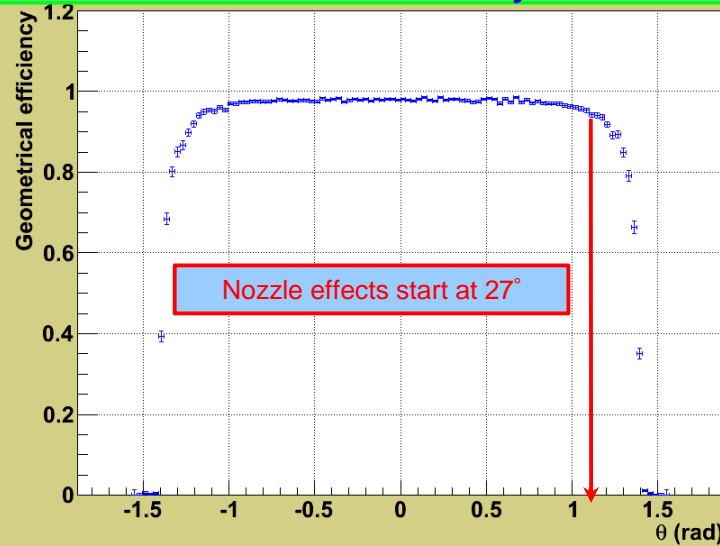
- ❑ Merge signals belonging to the same channel (pixel)
 - ❑ Include non-linearity effects
 - ❑ Add threshold
 - ❑ Add saturation
- ❑ Add electronic noise
- ❑ Save Digits over threshold
 - threshold = 3000 electrons
 - electronic noise = 0 electrons

Pattern recognition

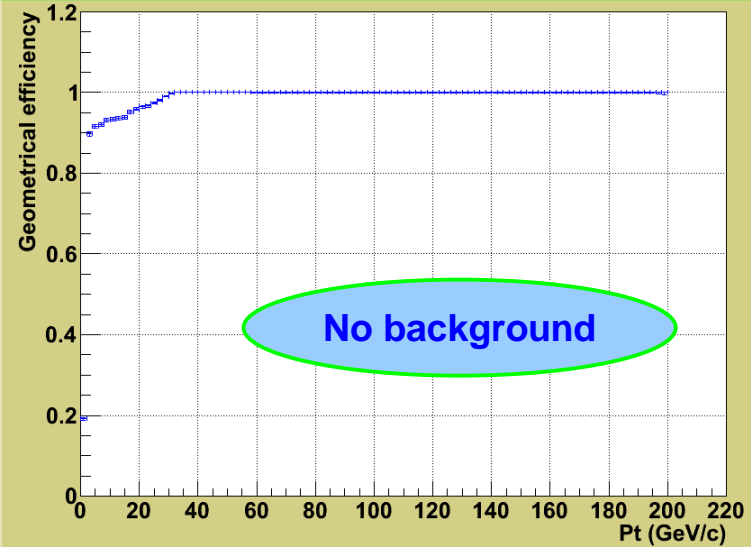
- ❑ Create a initial cluster from adjacent pixels (no for diagonal)
- ❑ Subdivide the previous cluster in smaller NxN clusters
- ❑ Get cluster and error matrix from coordinate average of the cluster
- ❑ Kalman filter picks up the best Recpoints

Reconstruction Efficiency for Single Muons

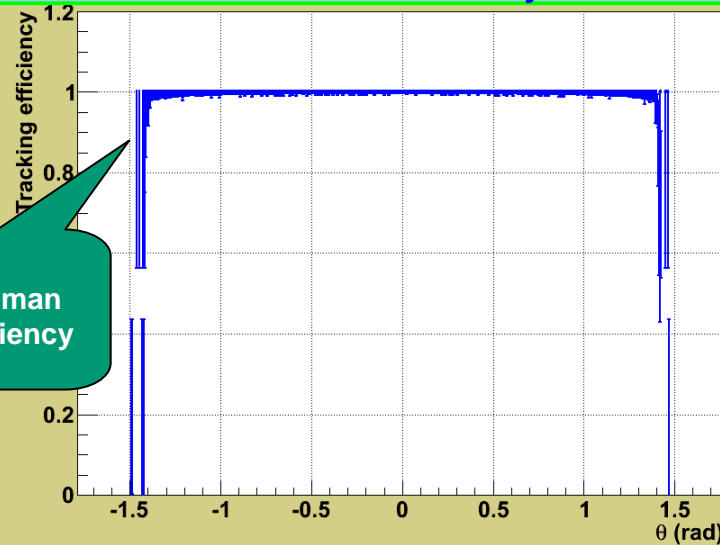
Geometrical Efficiency vs Theta



Geometrical Efficiency vs Pt

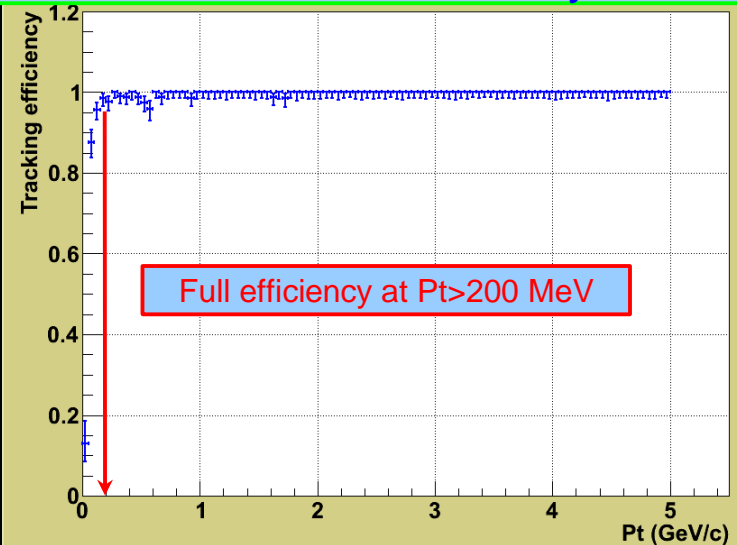


Kalman Filter Efficiency vs Theta



If hits are there, Kalman filter efficiency is ~100%

Kalman Filter Efficiency vs Pt



Effect of the 10° nozzle

ILCroot event display
for 10 muons up to 200 GeV

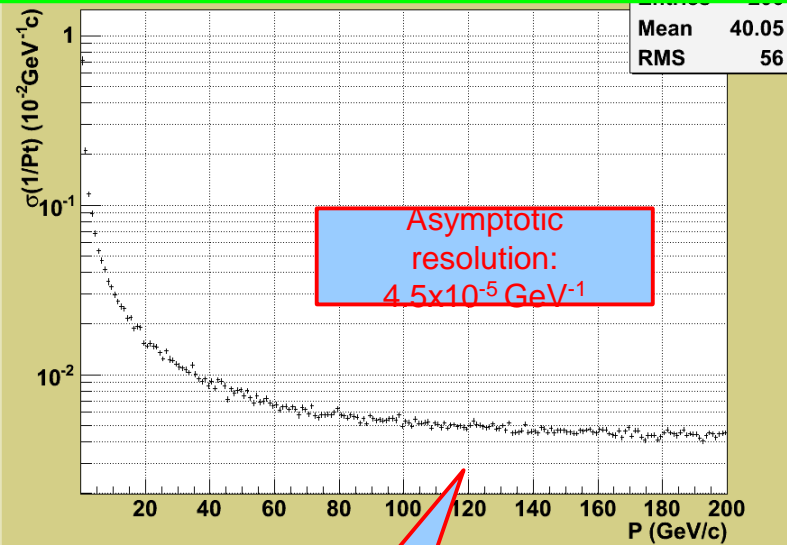
green - hits
purple - reconstructed tracks
red - MC particle

10 generated muons
9 reconstructed tracks

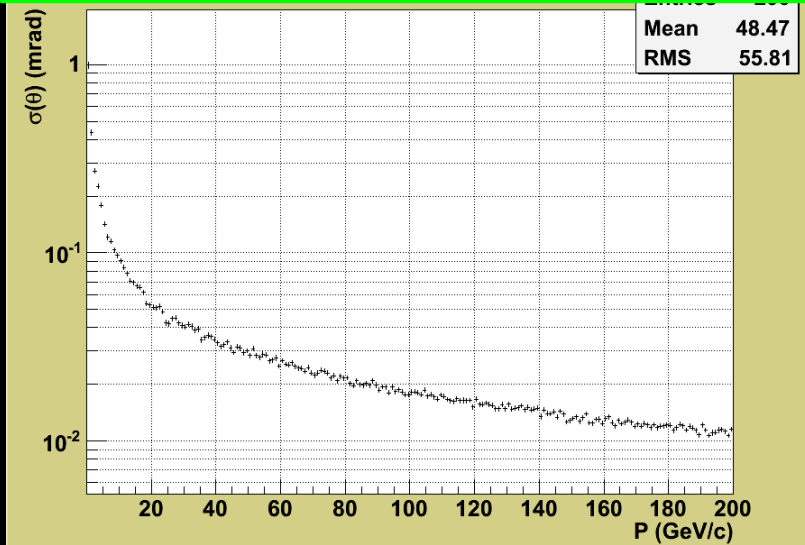
Full simulation of 10 muons – no bkg

Resolutions for single muons

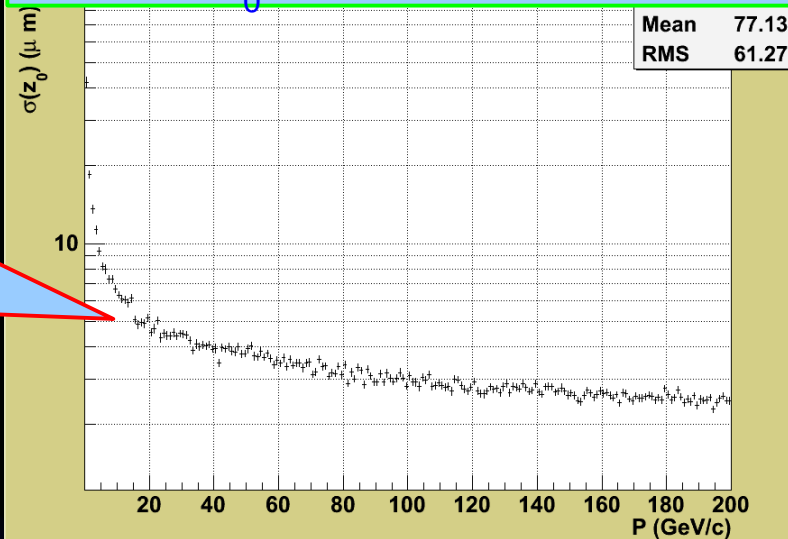
1/Pt Resolution vs P



Theta Resolution vs P



Z_0 Resolution vs P

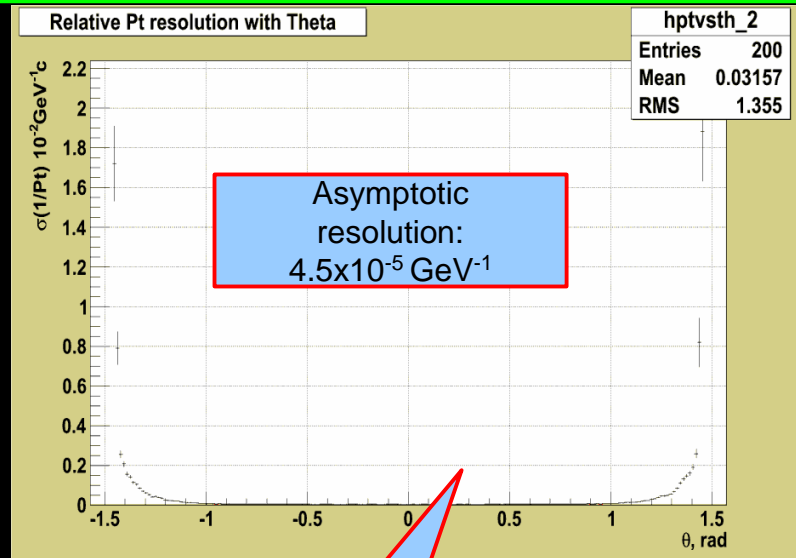


Well within requirements for precision physics

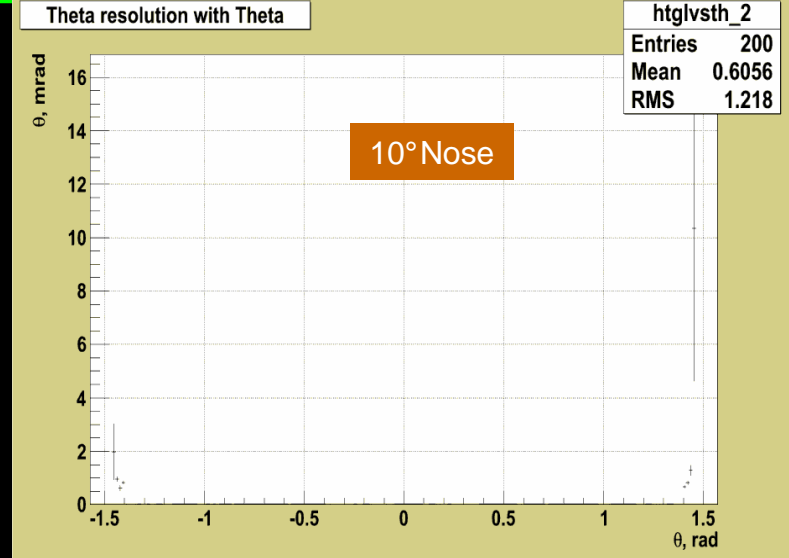
No background

Resolutions vs θ for single tracks

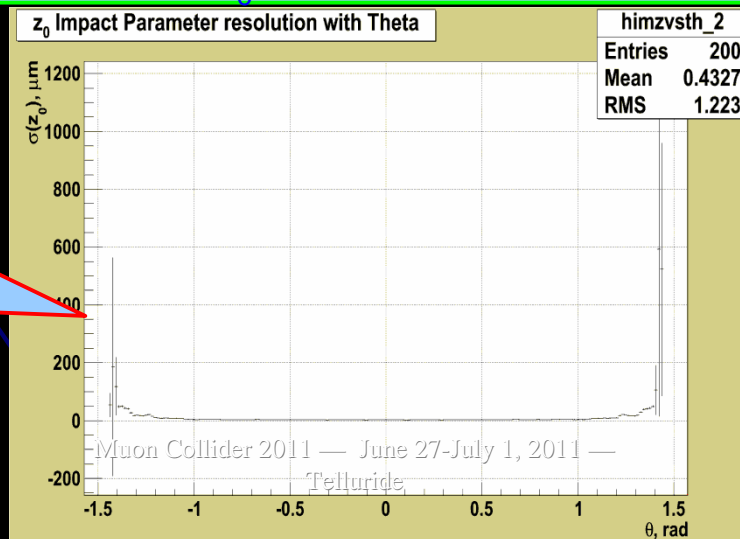
1/Pt Resolution vs θ



Theta Resolution vs θ



Z_0 Resolution vs θ



Well within requirements for precision physics

Beam Background Studies

- We simulated in ILCroot 4 detectors with different timing capabilities:
 - Det. A – No time information (integrates all hits)
 - Det. B – Acquires data in a fixed 7 ns time gate (minimal timing capabilities)
 - Det. C - Acquires data in a adjustable 3 ns time gate tuned to distance from IP (advanced timing capabilities)
 - Det. D - Acquires data in a 1 ns time gate tuned to pixel distance from IP (extreme timing capabilities)

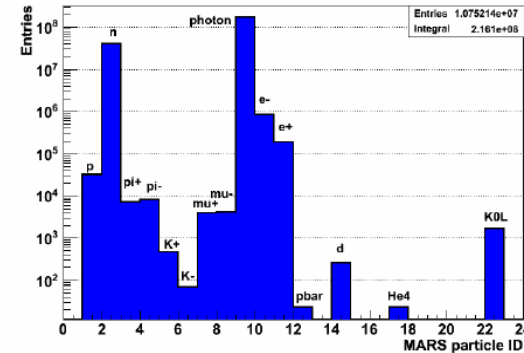
See N. Terentiev 's talk

Add Beam Background

$2 \times 10^5 / \text{cm}^2$
in layer1

Particles
Entering the
Detector region

- MARS background particle ID's yields for 750 GeV 2×10^{12} muons/bunch

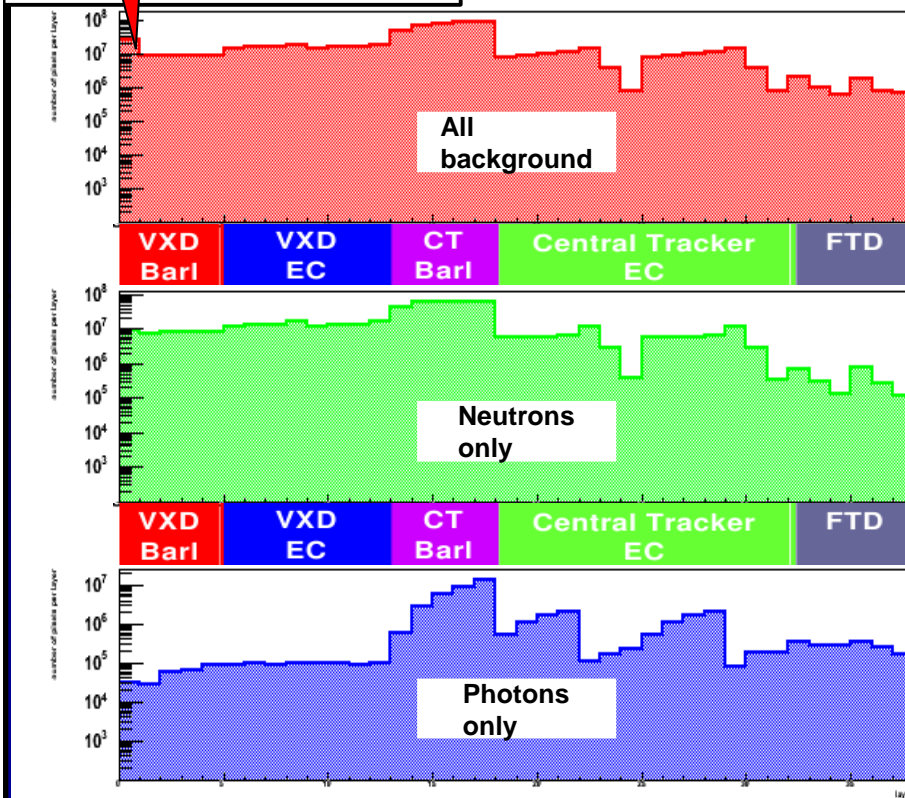


- Background yields/bunch on 10^0 nozzle surface and MARS thresholds

	γ	n	e^{++}	μ^{++}
Yield	1.77e+08	0.40e+08	1.03e+06	0.80e+04
Ethr, MeV	0.2	0.1	0.2	1.0

N. Terentiev (CMU/Fermilab) Muon Collider 2011 27 June – 01 July, 2011

Pixels per layer



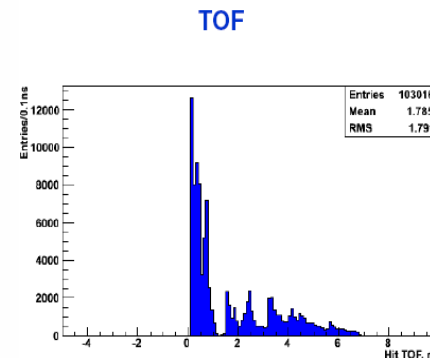
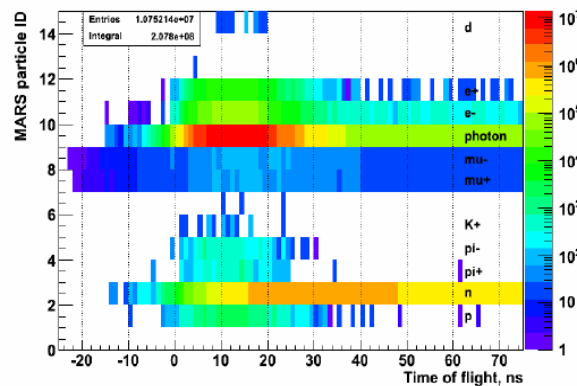
Actual pixels
Turned on

Total clusters found: 4×10^7
Too many for any practical
track reconstruction systems!

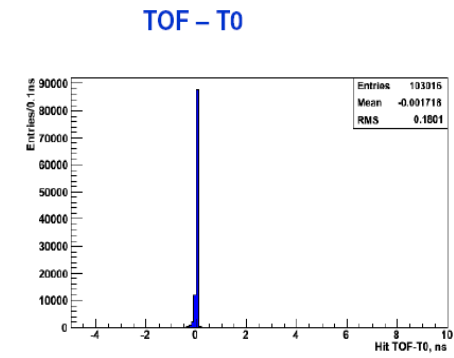
but.. Background is off time

- **MARS particle TOF and their ID (see in backup Ekin, Pt and Z)**
 - Time of flight (TOF) wrt. bunch crossing time, on a surface of the 10^9 nozzle
 - In window $0 \leq \text{TOF} \leq 25$ ns:
 - ~21% of neutrons, ~36% of muons, >94% of other particles
 - $\text{TOF} < 0$ corresponds to the particles making straight path to detector

- **Vertex and tracker timing for IP muons**



RMS ~ 1.8 ns

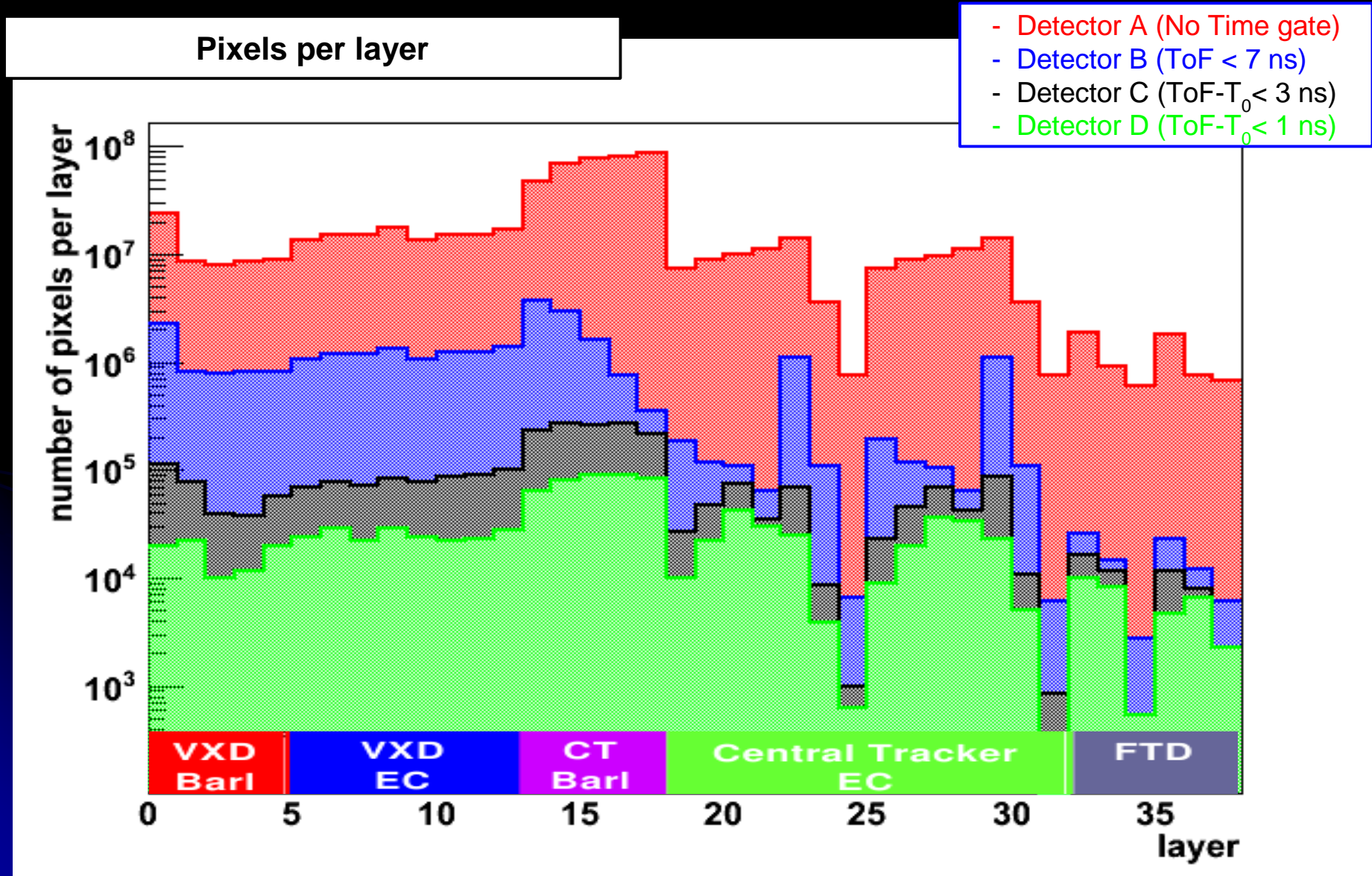


RMS ~ 0.18 ns

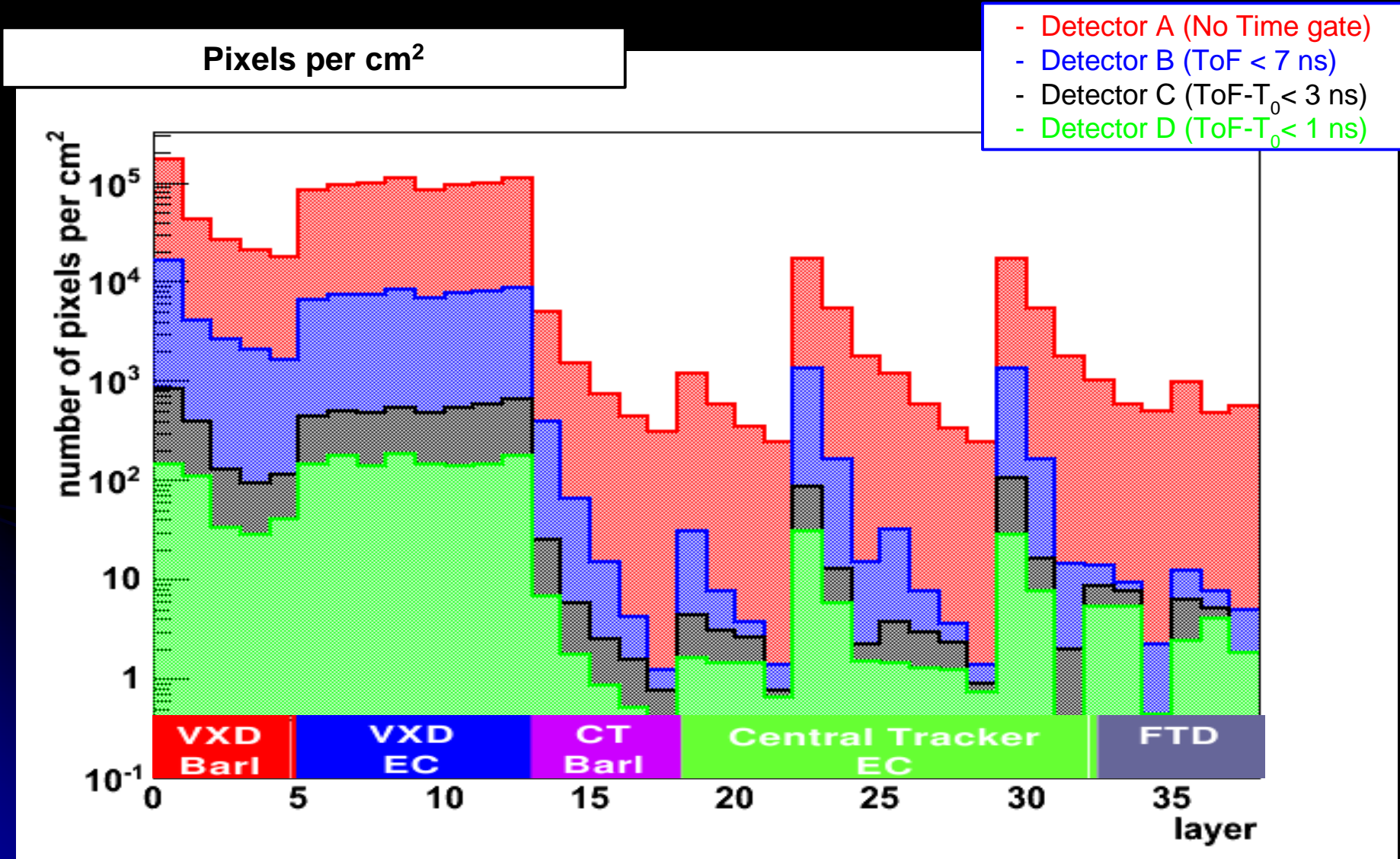
□ We simulated in ILCroot 4 detectors with different timing capabilities:

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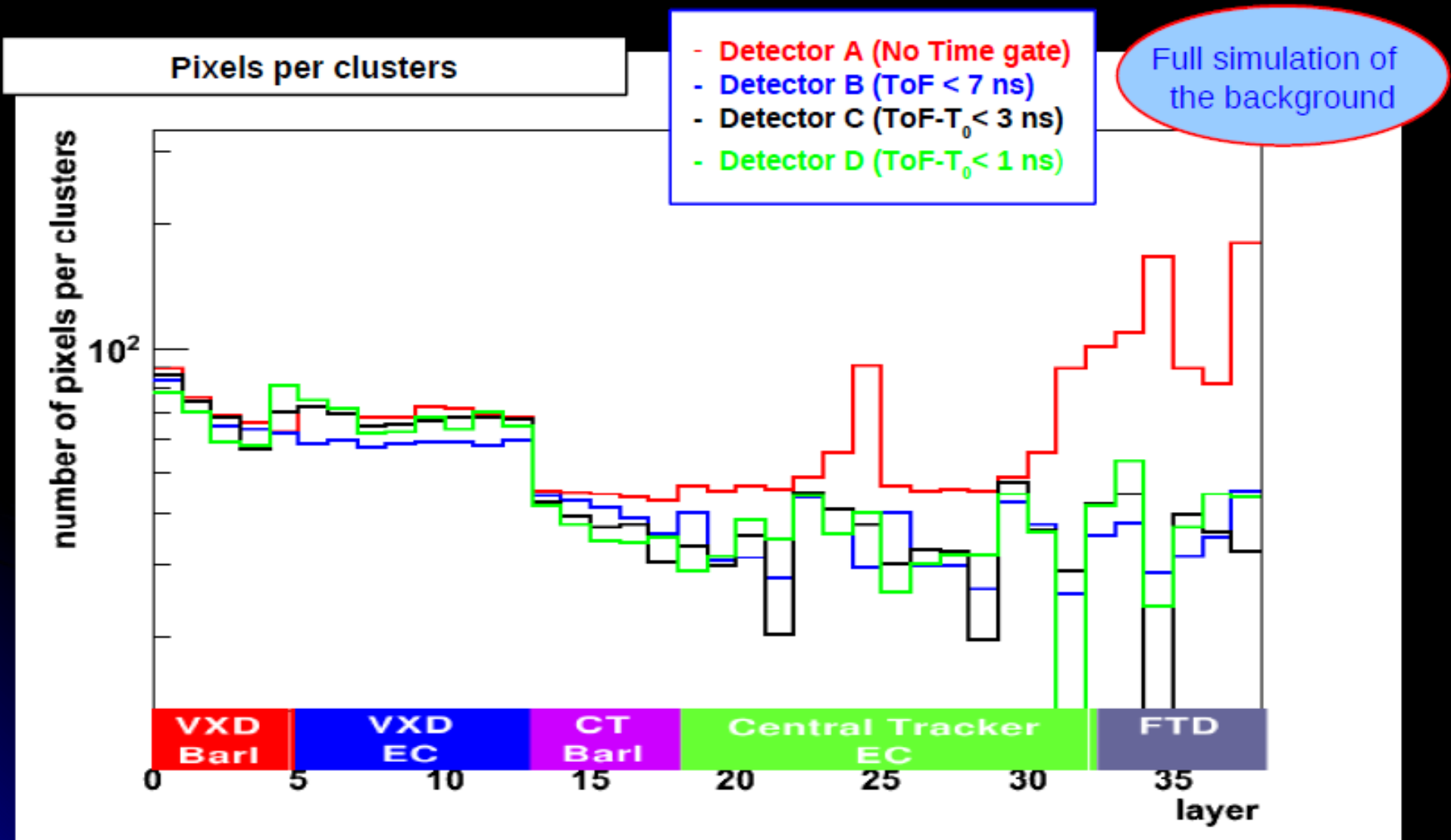
Total Pixels vs Layer for 4 detectors



Pixels occupancy vs Layer for 4 detectors



Pixels Occupancy for 4 detectors



Reconstructed Background Tracks (from Kalman filter)

Detector type	Reconstructed Tracks (full simu)	Reconstructed Tracks (fast simu)
Det. A (no timing)	Cannot calculate	Cannot calculate
Det. B (7 ns fixed gate)	75309	64319
Det. C (3 ns adjustable gate)	6544	4639
Det. D (1 ns adjustable gate)	1459	881

Full reconstruction is paramount when combinatorics is relevant

Reconstructed Background Tracks (from Kalman filter)

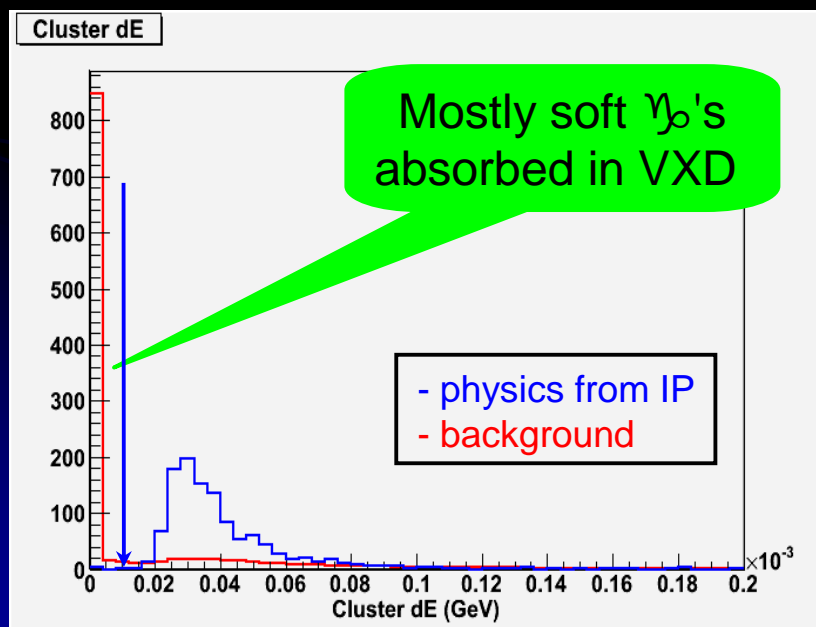
Detector type	Reconstructed Tracks (full simu)	Reconstructed Tracks (fast simu)
Det. A (no timing)	Cannot calculate	Cannot calculate
Det. B (7 ns fixed gate)	75309	64319
Det. C (3 ns adjustable gate)	6544	4639
Det. D (1 ns adjustable gate)	1459	881

STILL TOO MANY
RECONSTRUCTED TRACKS
FROM BACKGROUND

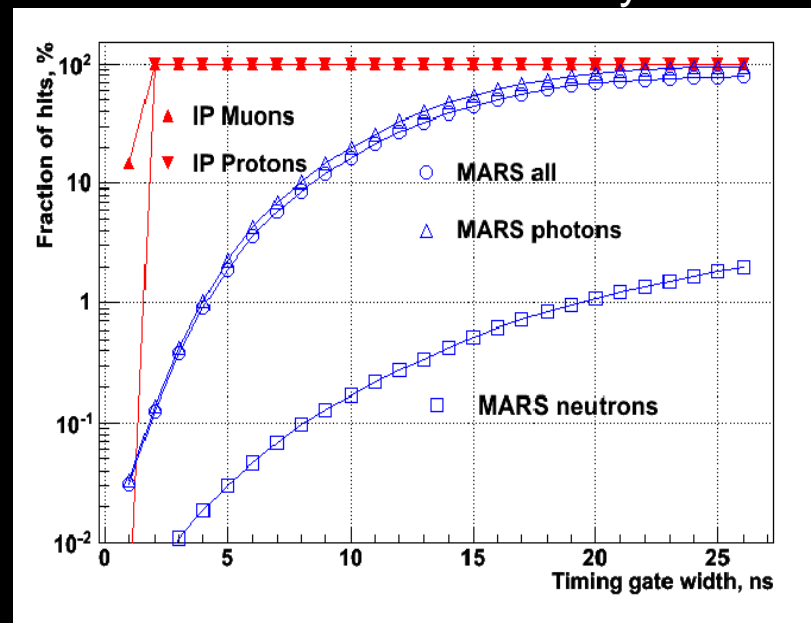
Compare dE/dx and timing for signal vs background

	Kalman Reconstruction	Clusters
Physics: 100 μ (0.2–200) GeV/c	92 (include geom. eff.)	1166
Machine Background	-	4×10^7

N. Terentiev's study



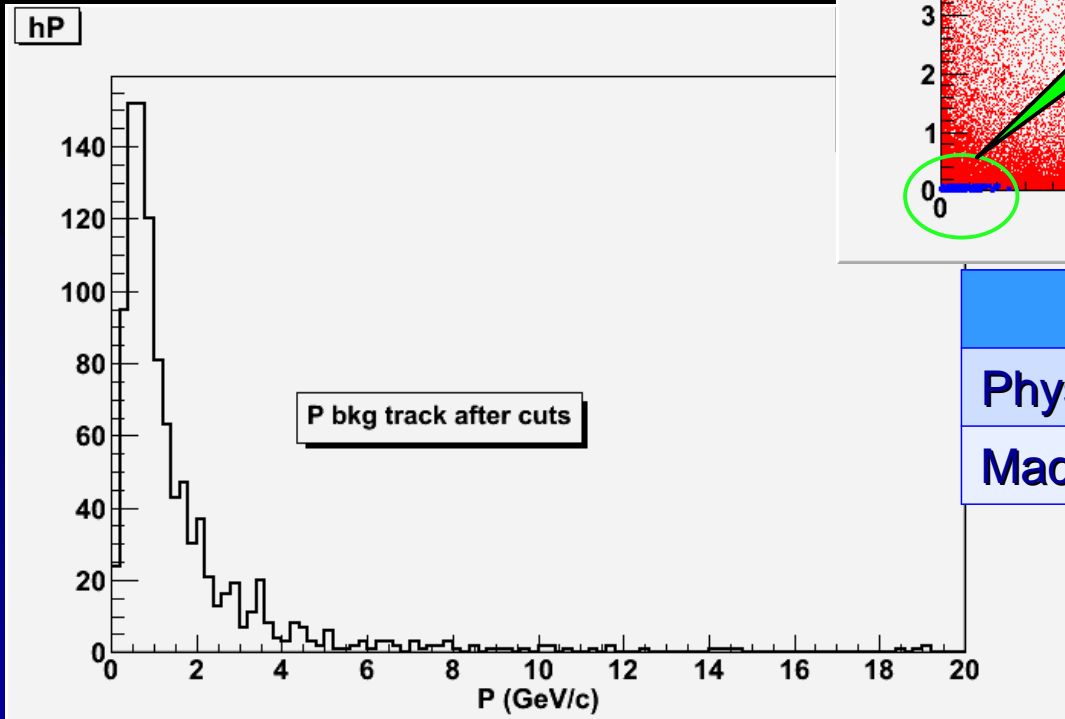
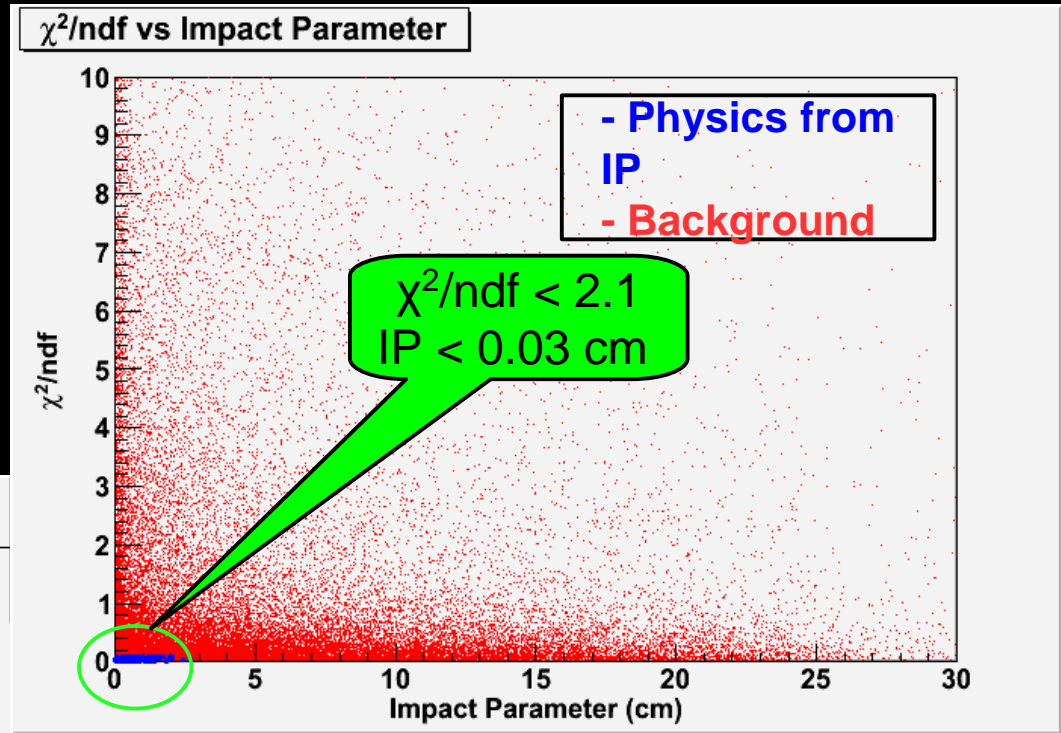
Choose E threshold 10 KeV (2400 e-)



Choose Time Gate Width

Physics vs Background: a strategy to disentangle reconstructed tracks from IP

Studies based on half background

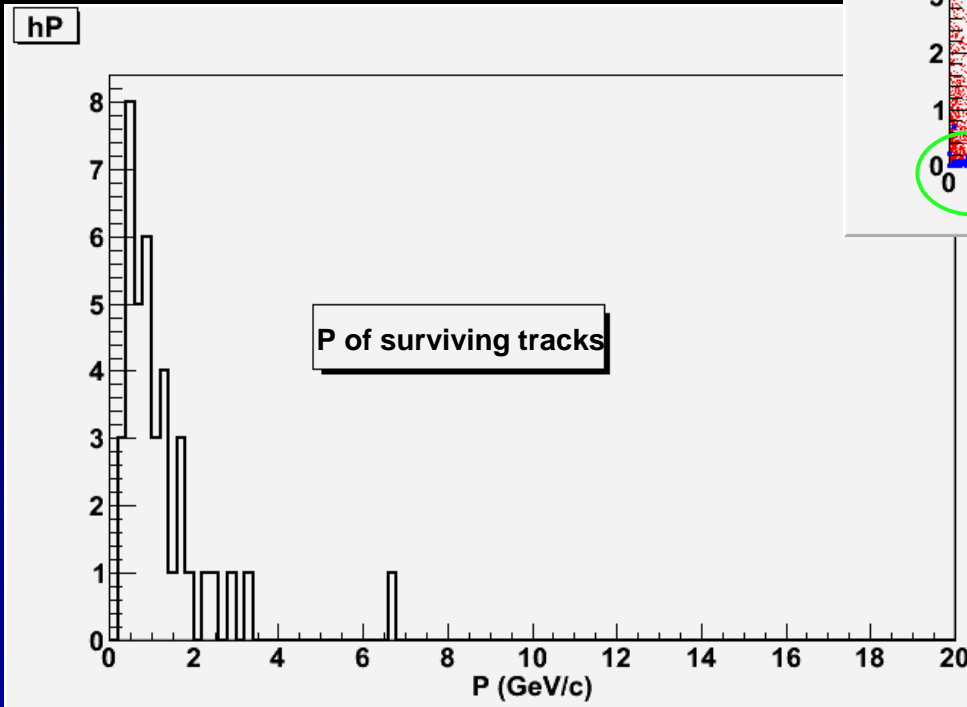
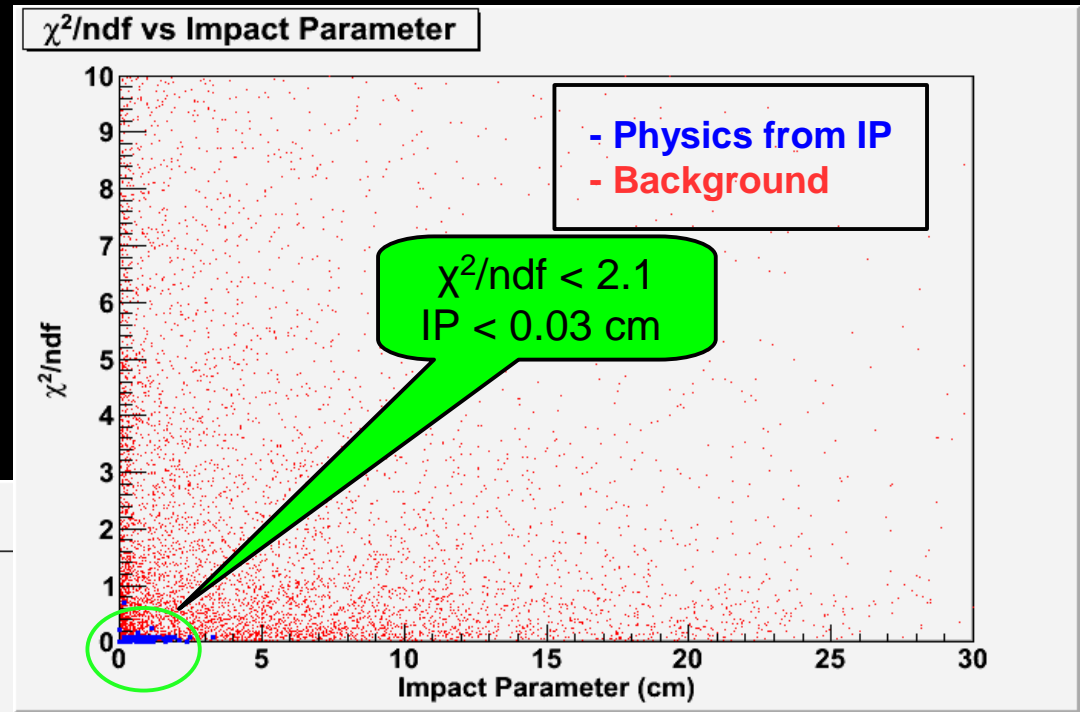


	Surviving tracks
Physics: 100 μ	89
Machine background	2110

3 lost

Physics vs Background in Det. B: a strategy to disentangle reconstructed tracks from IP

Full simulation of physics + bkg

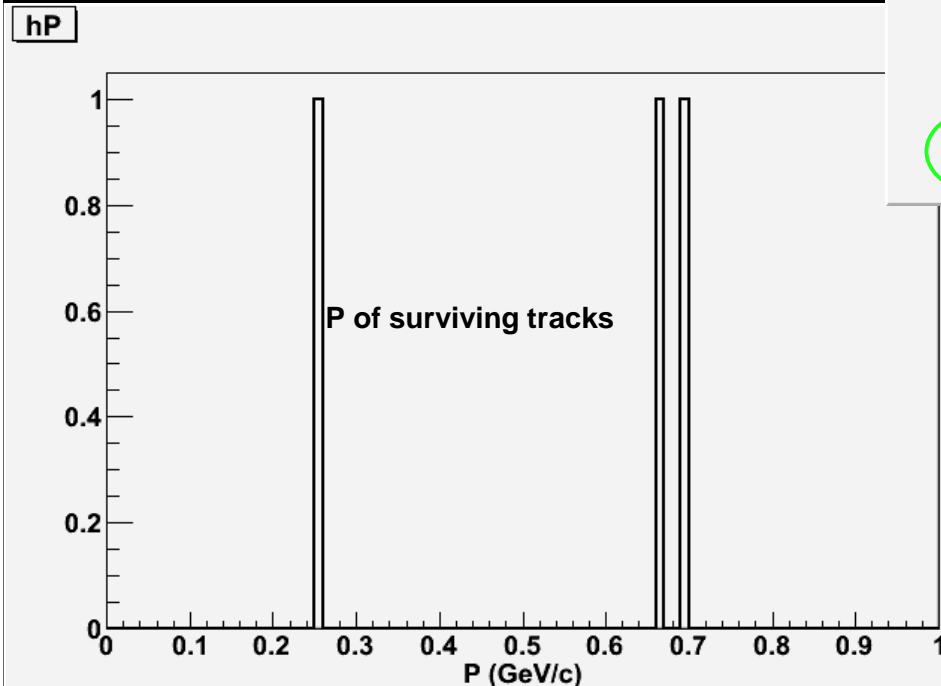
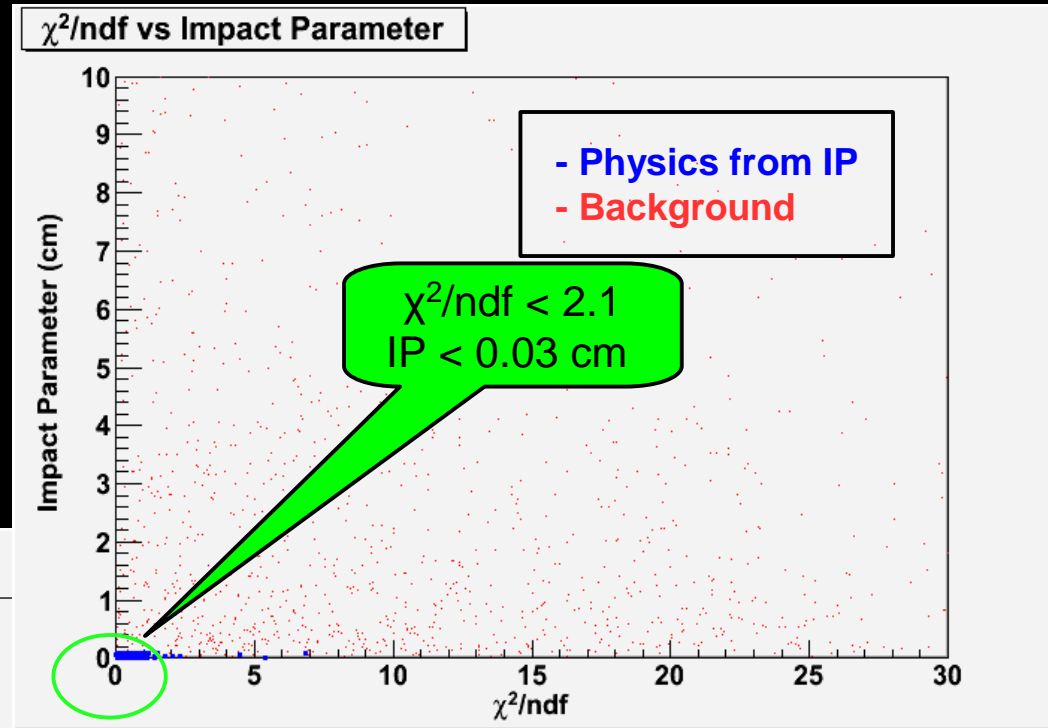


	Surviving tracks
Physics: 100 μ	92
Machine background	41

no lost

Physics vs Background in Det. D: a strategy to disentangle reconstructed tracks from IP

Full simulation of physics + bkg



	Surviving tracks
Physics: 100 μ	92
Machine background	3

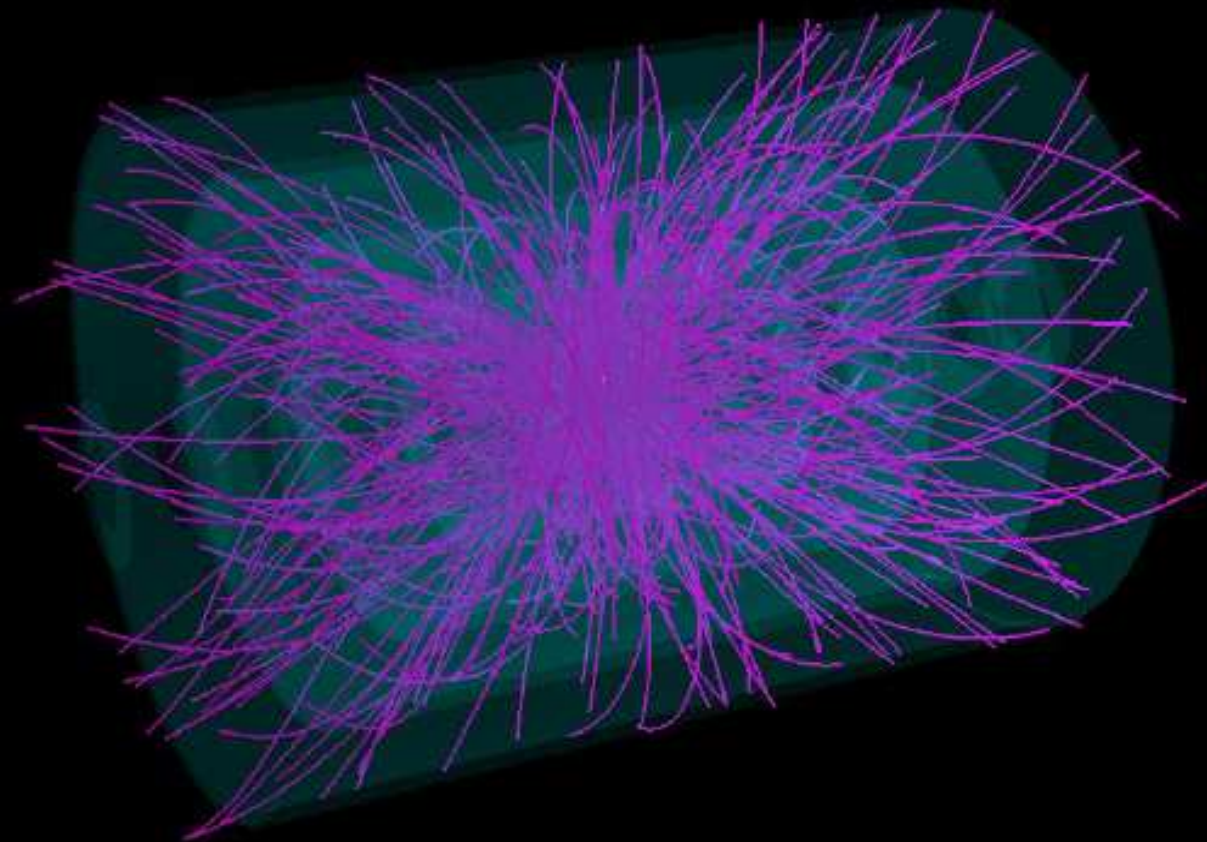
no lost

Reconstructed Background Tracks (from Kalman filter) after χ^2 and Impact Point cuts

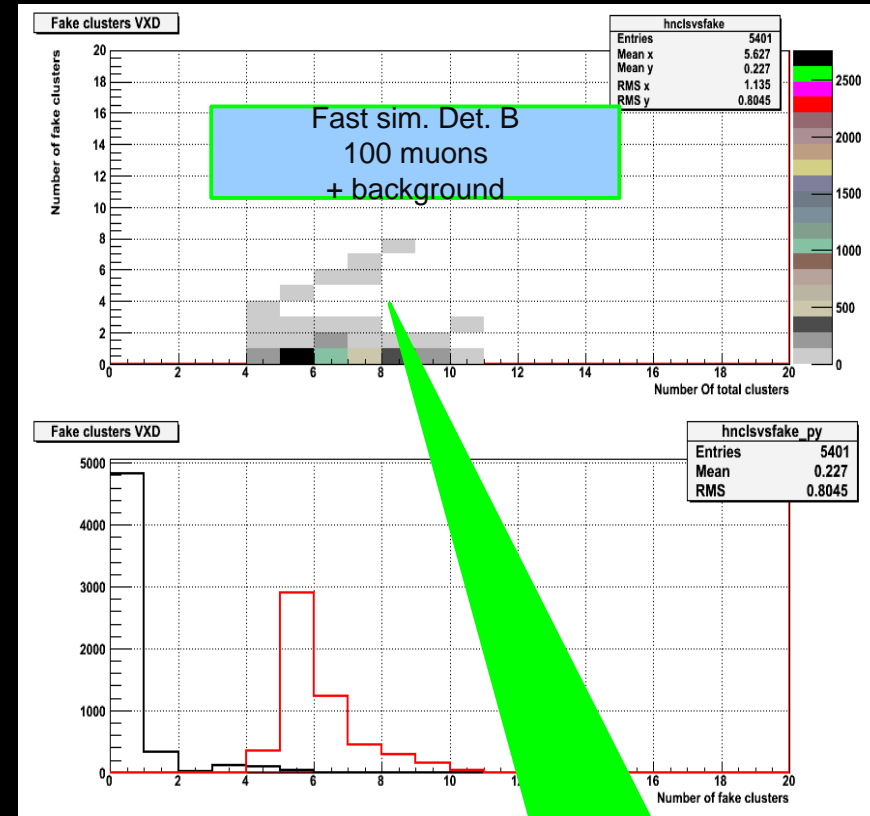
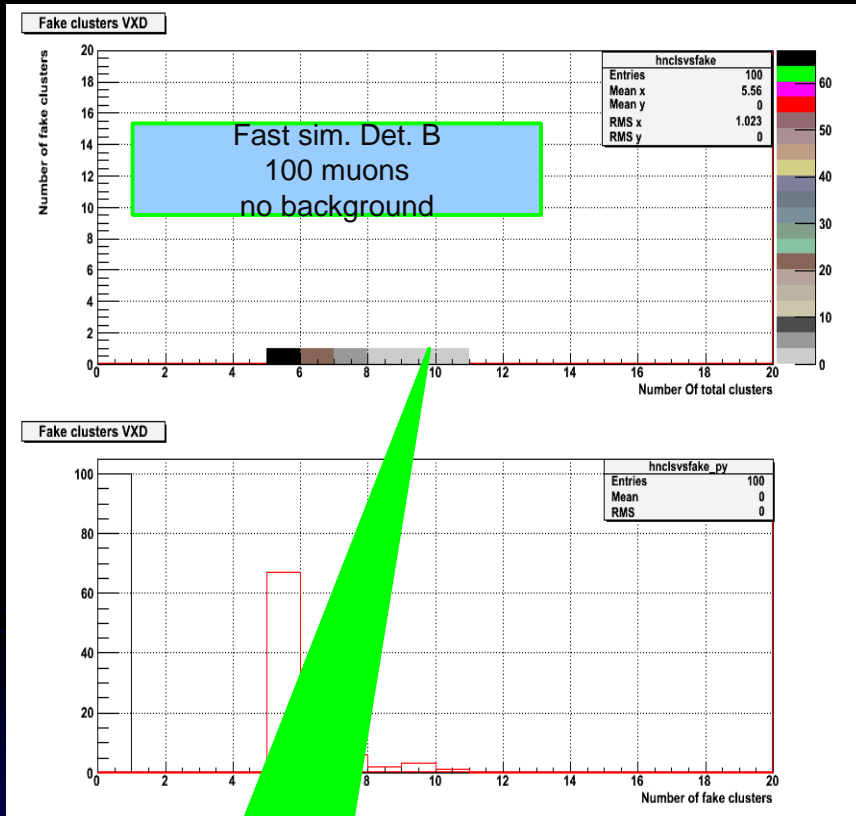
Detector type	Reconstructed Tracks (full simu)	Reconstructed Tracks (fast simu)
Det. A (no timing)	Cannot calculate	Cannot calculate
Det. B (7 ns fixed gate)	475	405
Det. C (3 ns adjustable gate)	11	8
Det. D (1 ns adjustable gate)	3	1

Full reconstruction is paramount when combinatorics is relevant

Event Display of Surviving Background tracks



Effects of background Hits on Physics



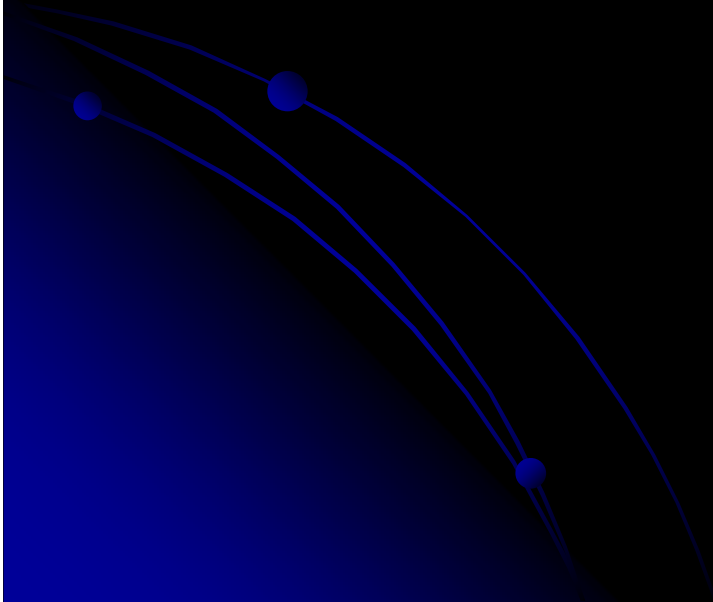
no fake cluster

physics event = 100 muons

< 5% of tracks
have > 1 fake cluster

Effects on track parameter resolution are unaffected by background

Calorimetry Studies



Fundamental issues at a μ Collider

1. Resolution: how good for W/Z separation?

- Run a Toy Montecarlo with several $\sigma(E)$ models
- Assume 10 mrad jet-axis resolution and perfect pattern recognition

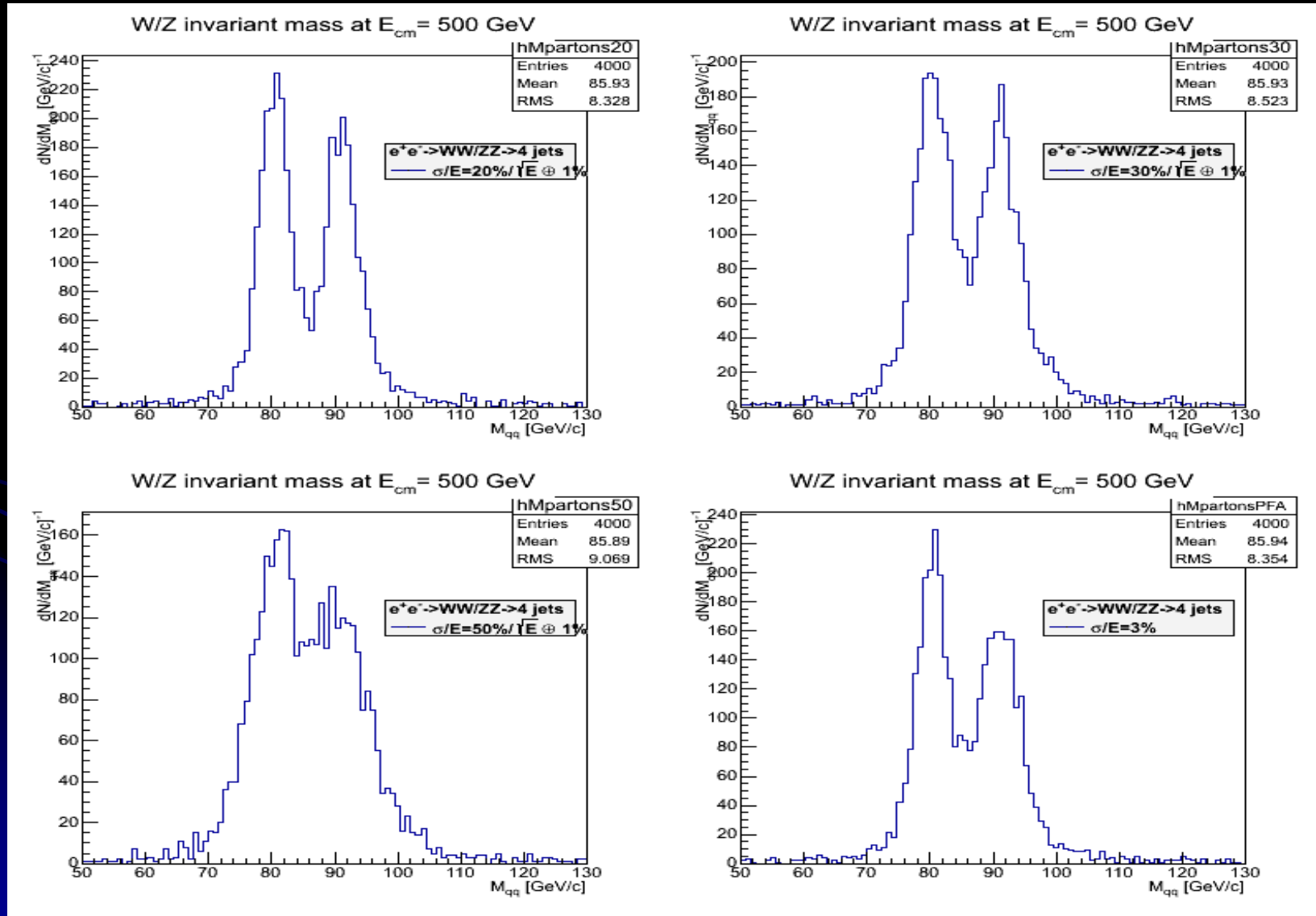
2. Longitudinal segmentation

3. Timing

- Need a full study including instrumentation effects

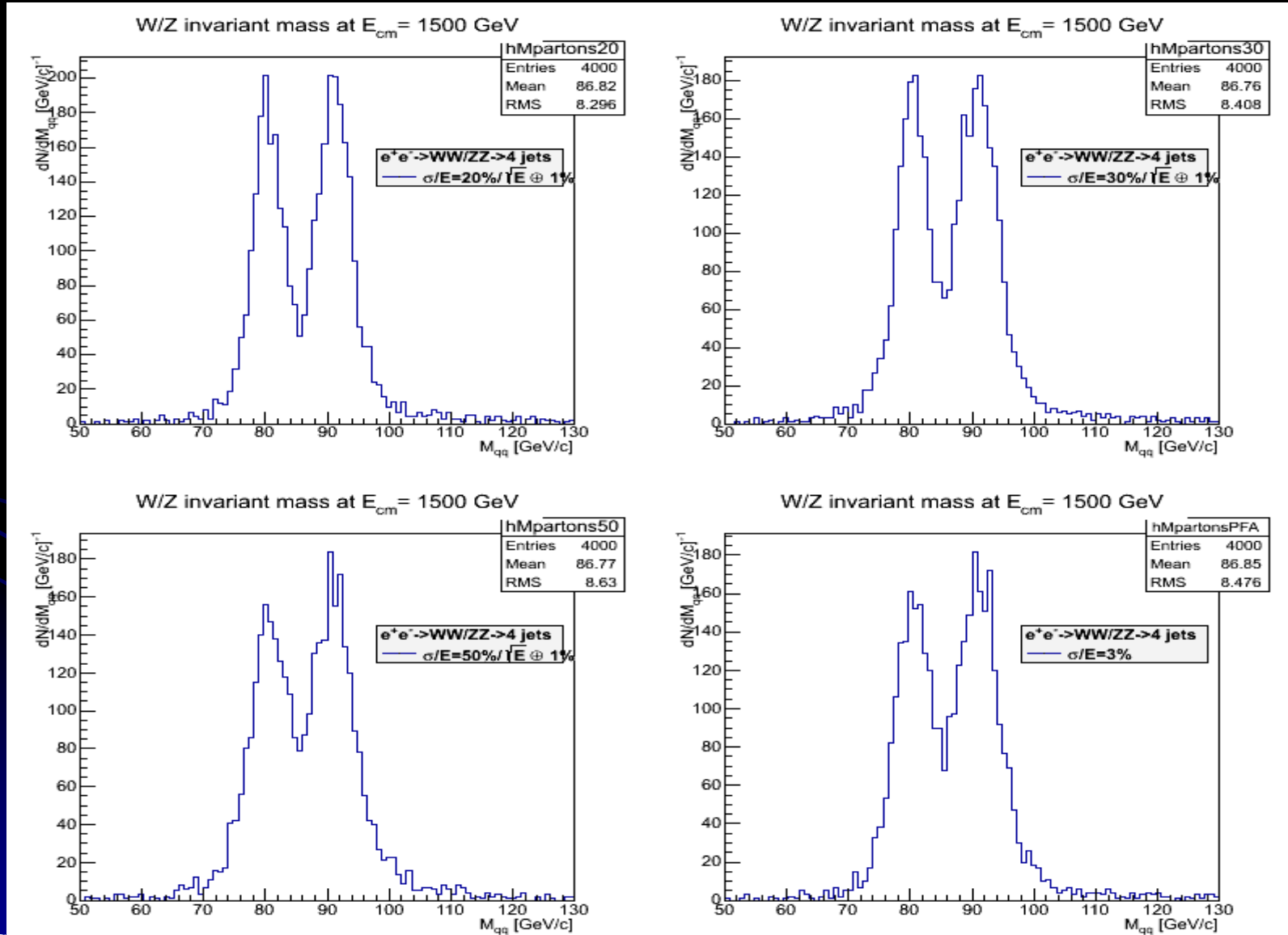
W/Z separation at IL

Toy Montecarlo



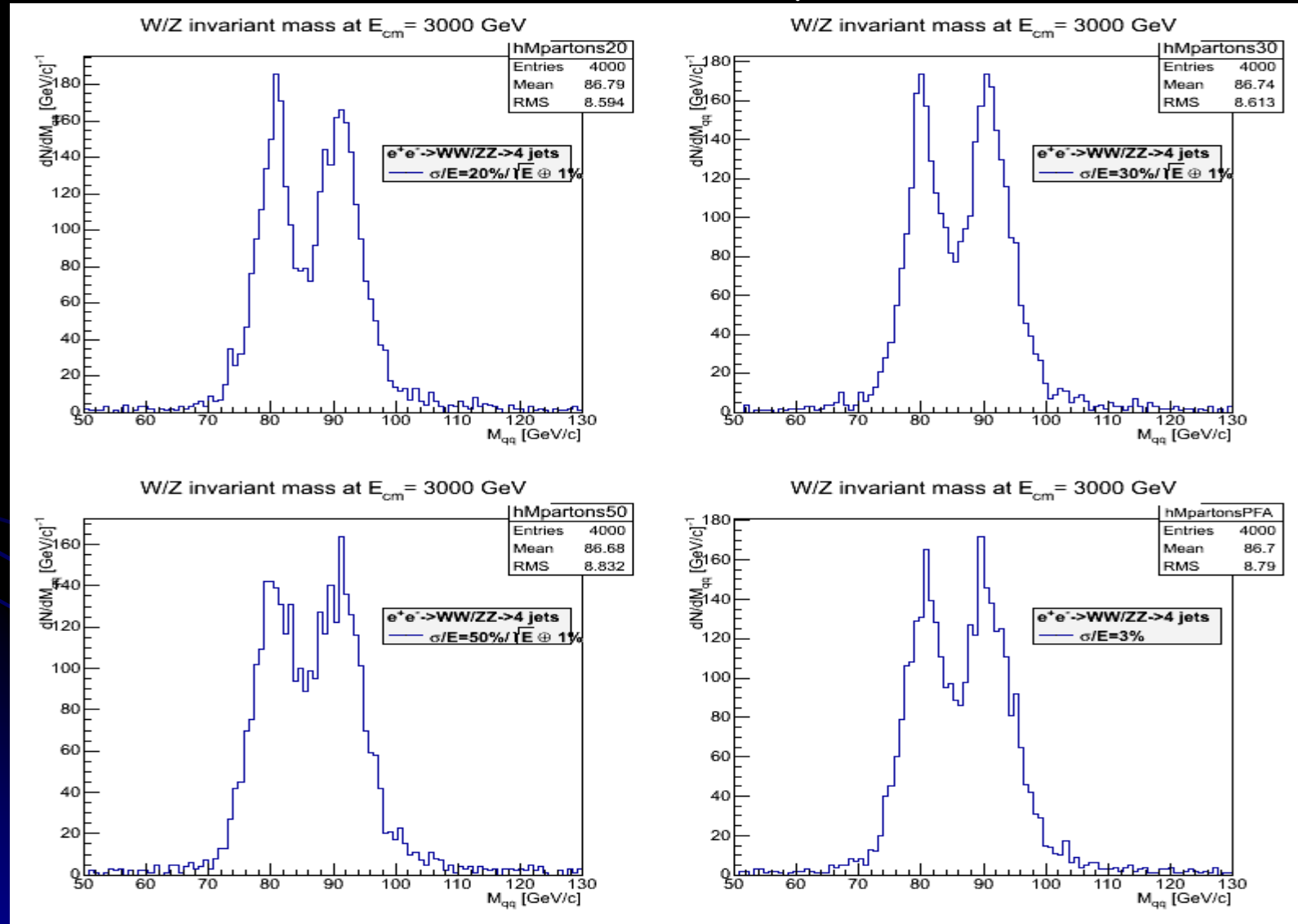
W/Z separation at CLIC/ μ Coll

Toy Montecarlo



W/Z separation at CLIC/ μ Coll

Toy Montecarlo

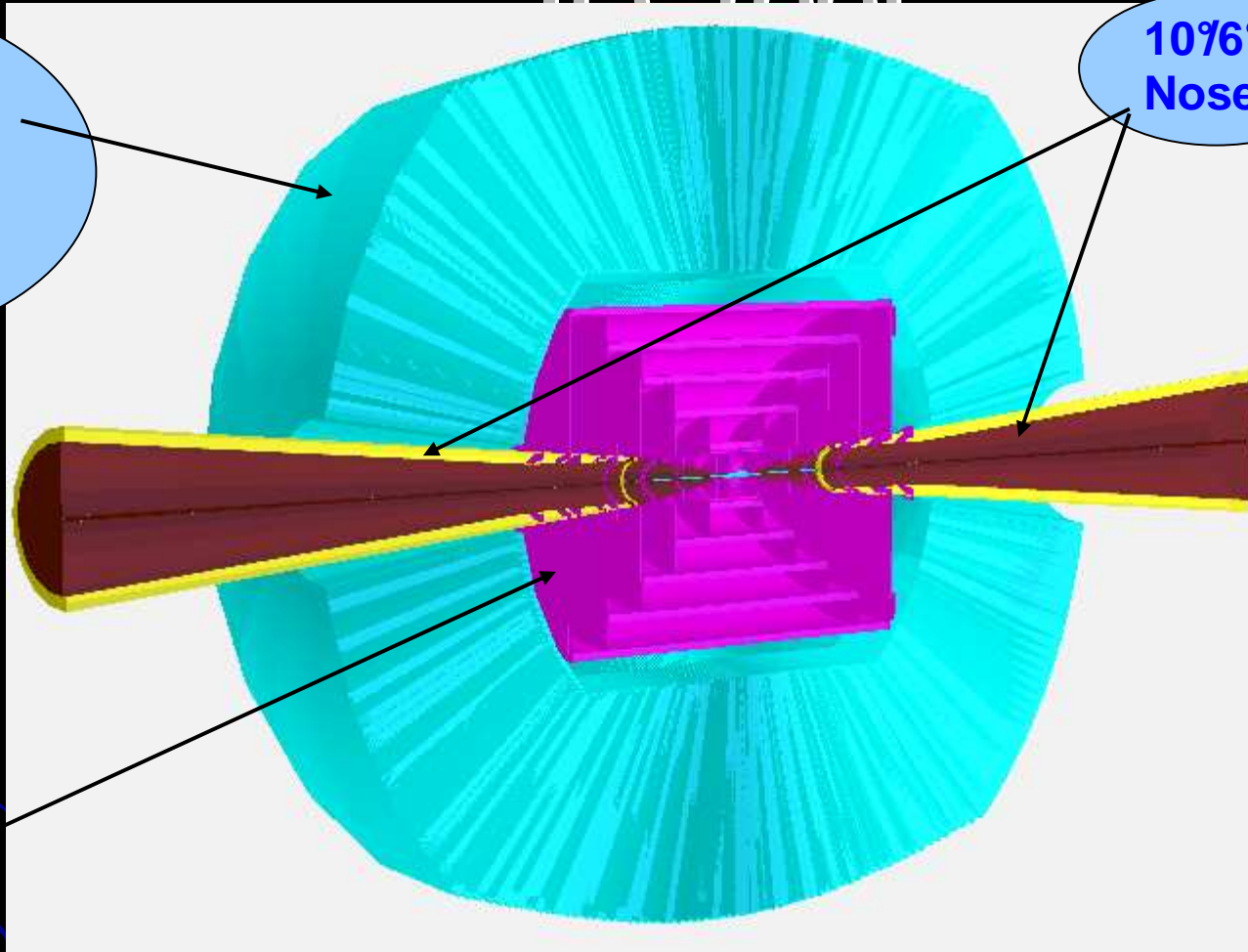


Calorimeter Geometry in IL Croc

Dual Readout

- 1) ADRIANO (full simulation)
- 2) TAHCAL (fast simulation)

10%⁶
Nose



Tracker

ADRIANO Calorimeter (FNAL-INFN Collaboration)
is used for the studies presented here

A Dual Readout Integrally Active Non-segmented Option (T1015 Collaboration)

Lead glass + scintillating fibers

$\sim 1.4^\circ$ tower aperture angle

180 cm depth

$\sim 7.5 \lambda_{\text{int}}$ depth

$> 100 X_0$ depth

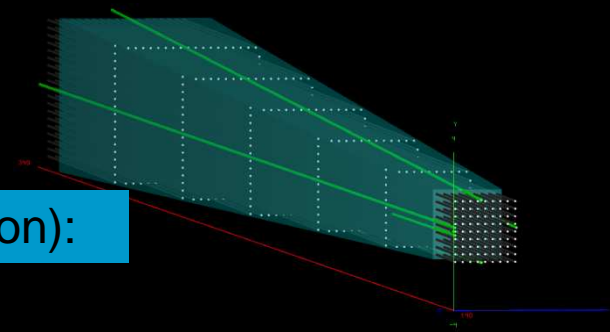
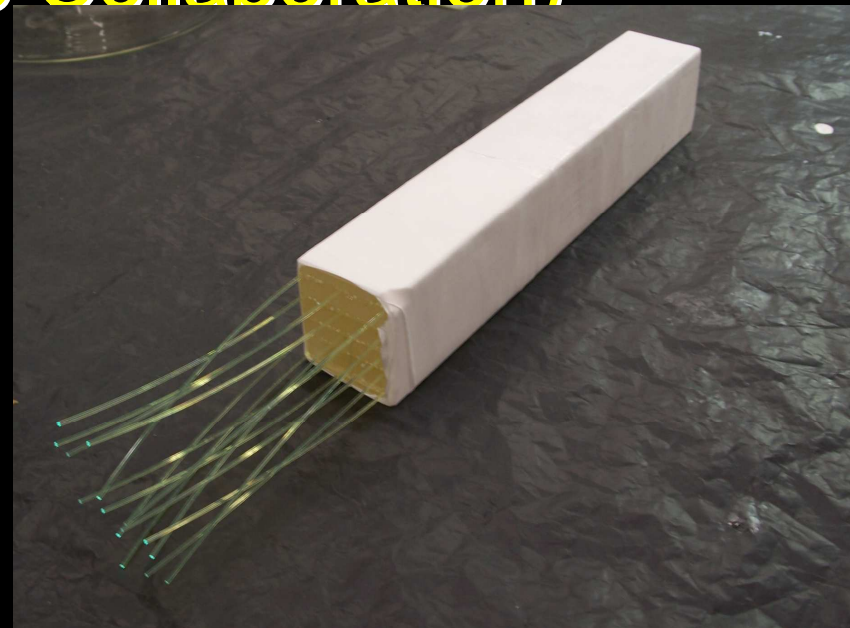
Fully projective geometry

Azimuth coverage

- down to $\sim 8.4^\circ$ (Nozzle)
- Barrel: 16384 towers
- Endcaps: 7222 towers

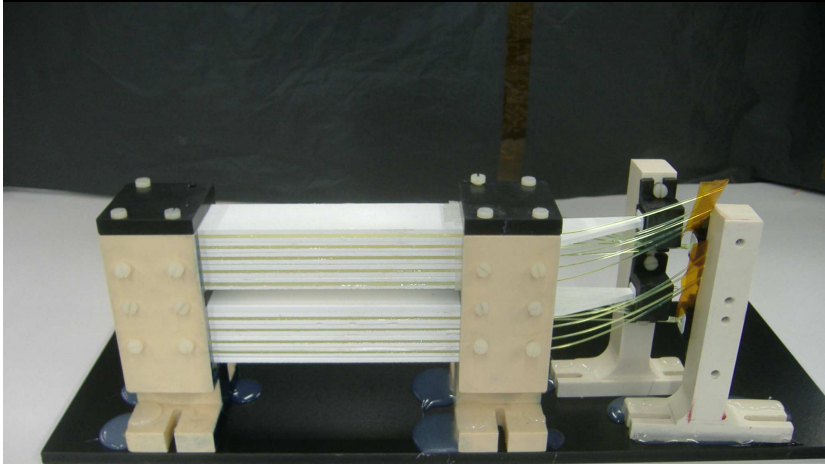
Expected resolution (see my talk at calorimetry session):

$$\sigma_E / E = 30\% / \sqrt{E}$$



ADRIANO Simulation

Details



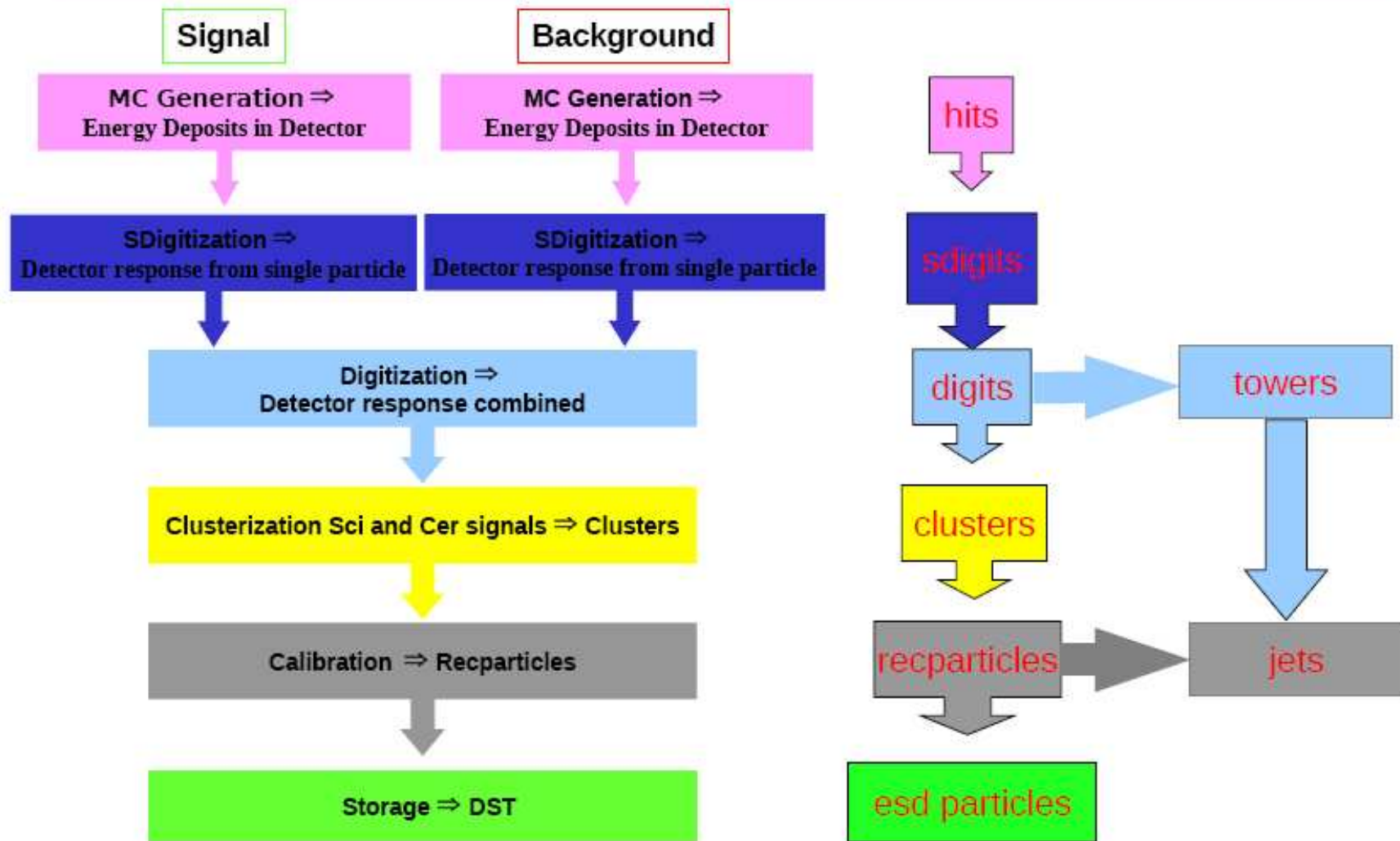
- WLS's collect Cerenkov photons from lead glass (front and back readout)
- Generate and transport scintillating photons (front and back readout for fibers in the core of the tower; only back readout for WLS and Q20)

Simulations include:

- $\tau_{\text{scifi}} = 2.4$ nsec (Kuraray SCSF81)
- $\tau_{\text{WLS}} = 2.7$ nsec (Bicron BCF92)
- SiPM with ENF=1.016
- Fiber non-uniformity response = 0.8% (scaled from CHORUS)
- Threshold = 3 p.e. (SiPM dark current < 50 kHz)
- ADC with 14 bits
- Gaussian noise with $\sigma = 1$ p.e.

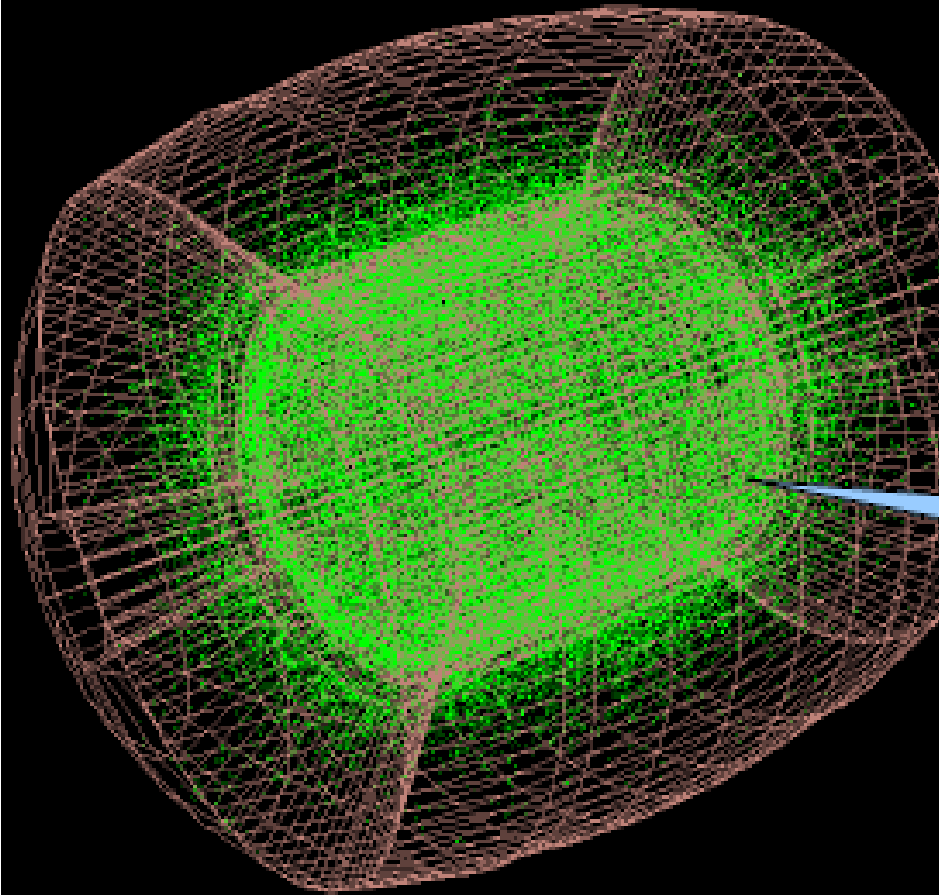
Already underwent two test-beams at FNAL

LCroot: for Full Calorimetry

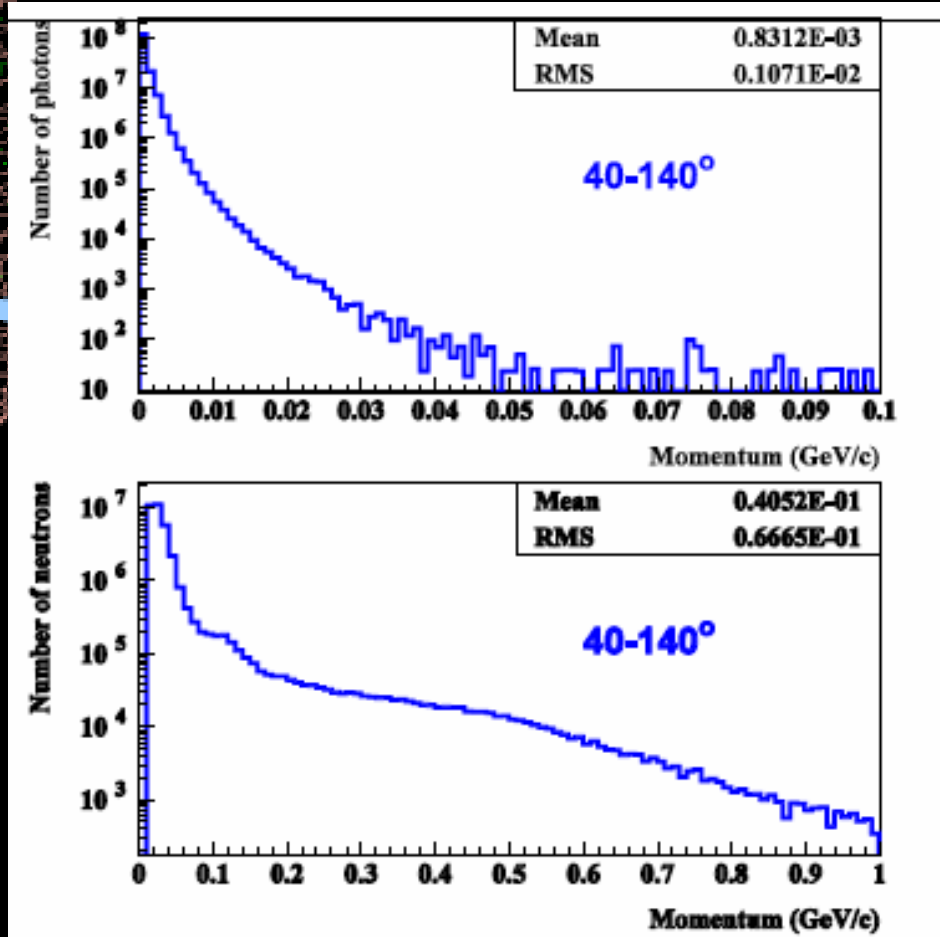


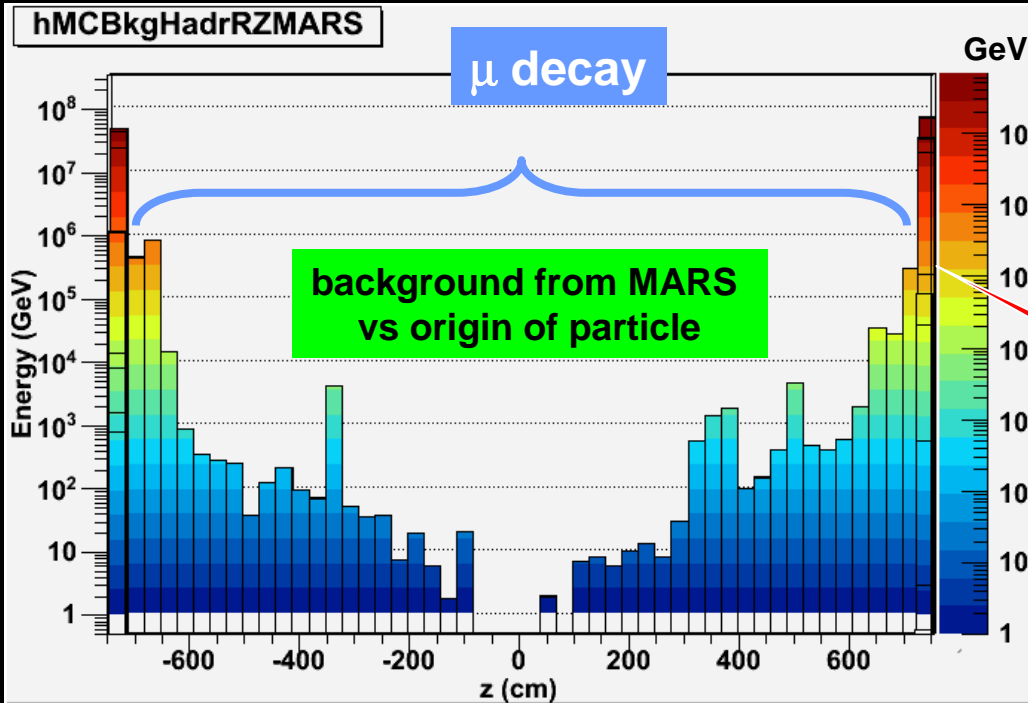
Background Entering the Calorimeter

γ and neutrons entering the calorimeter



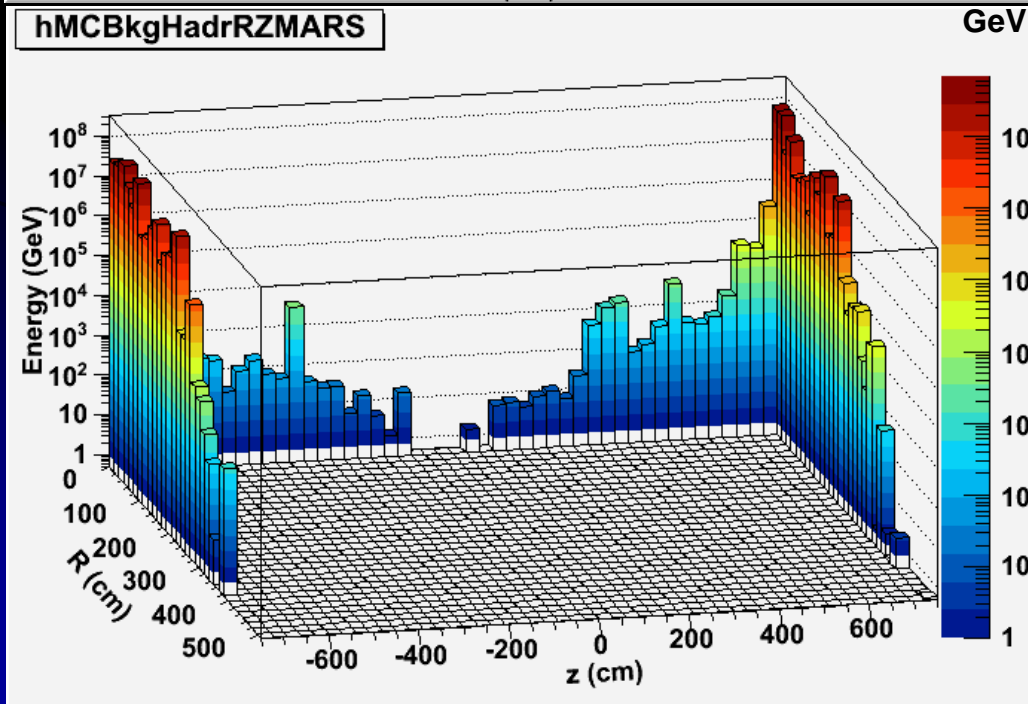
4% of a background event depicted





Start with a plain vanilla calorimeter (non segmented & no timing)

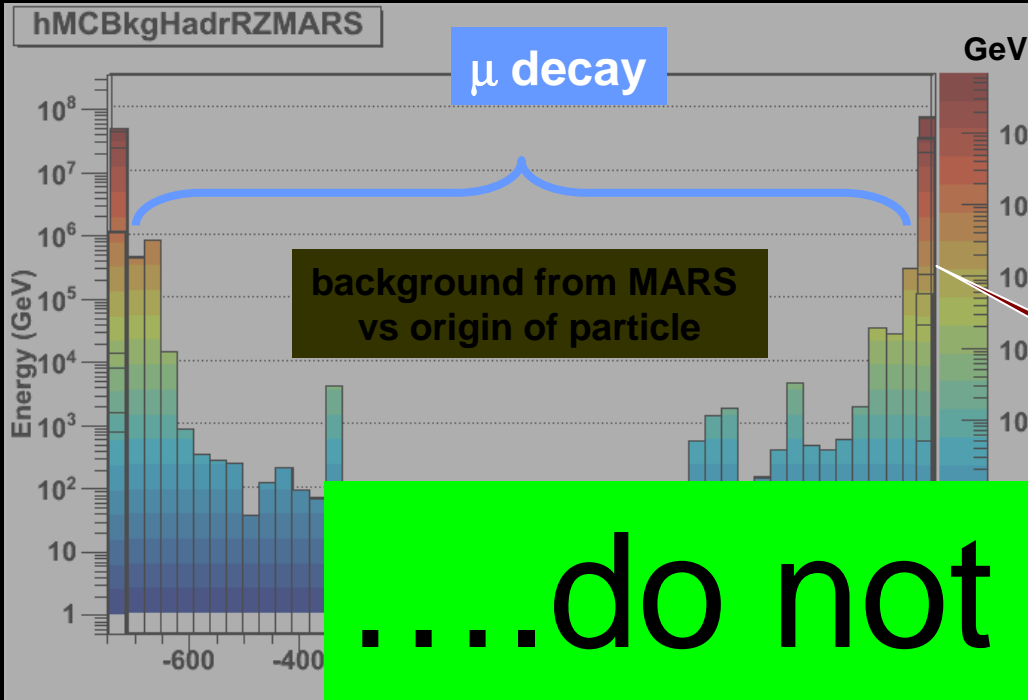
Entering detector area at 7.5 m



$R = r_{xy}$ of particle origin (1bin= 30cm)

Z=7.5 means that the particle originated outside the MDI separation plane (1bin=16cm)

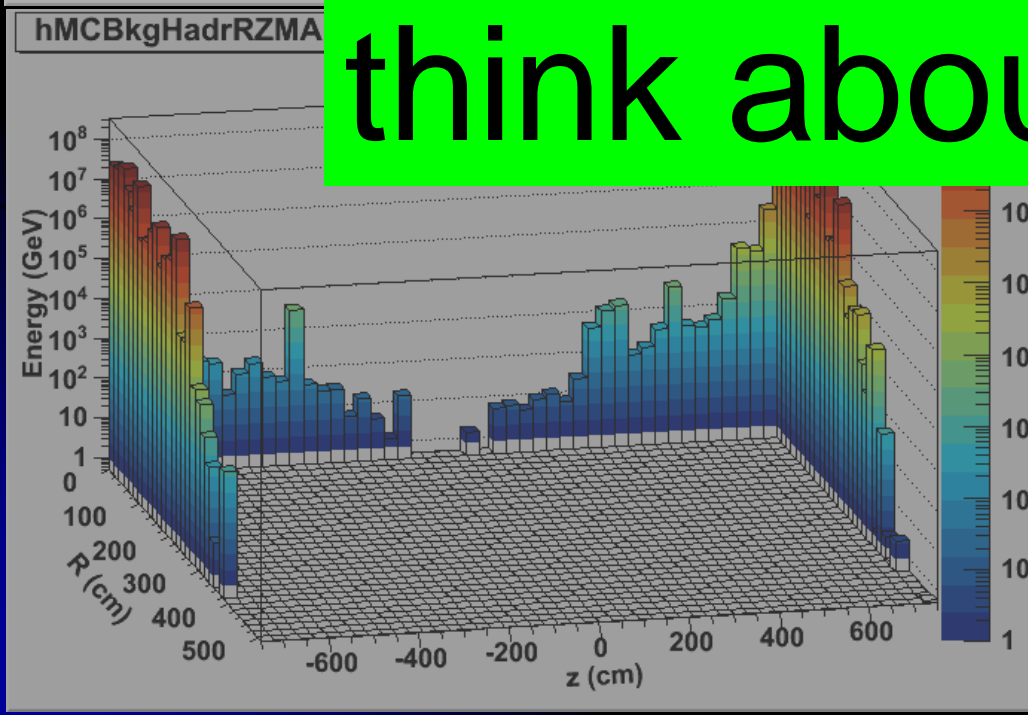
6°Nose



Start with a plain vanilla calorimeter (non segmented & no timing)

....do not even think about....

Entering detector area at 7.5 m

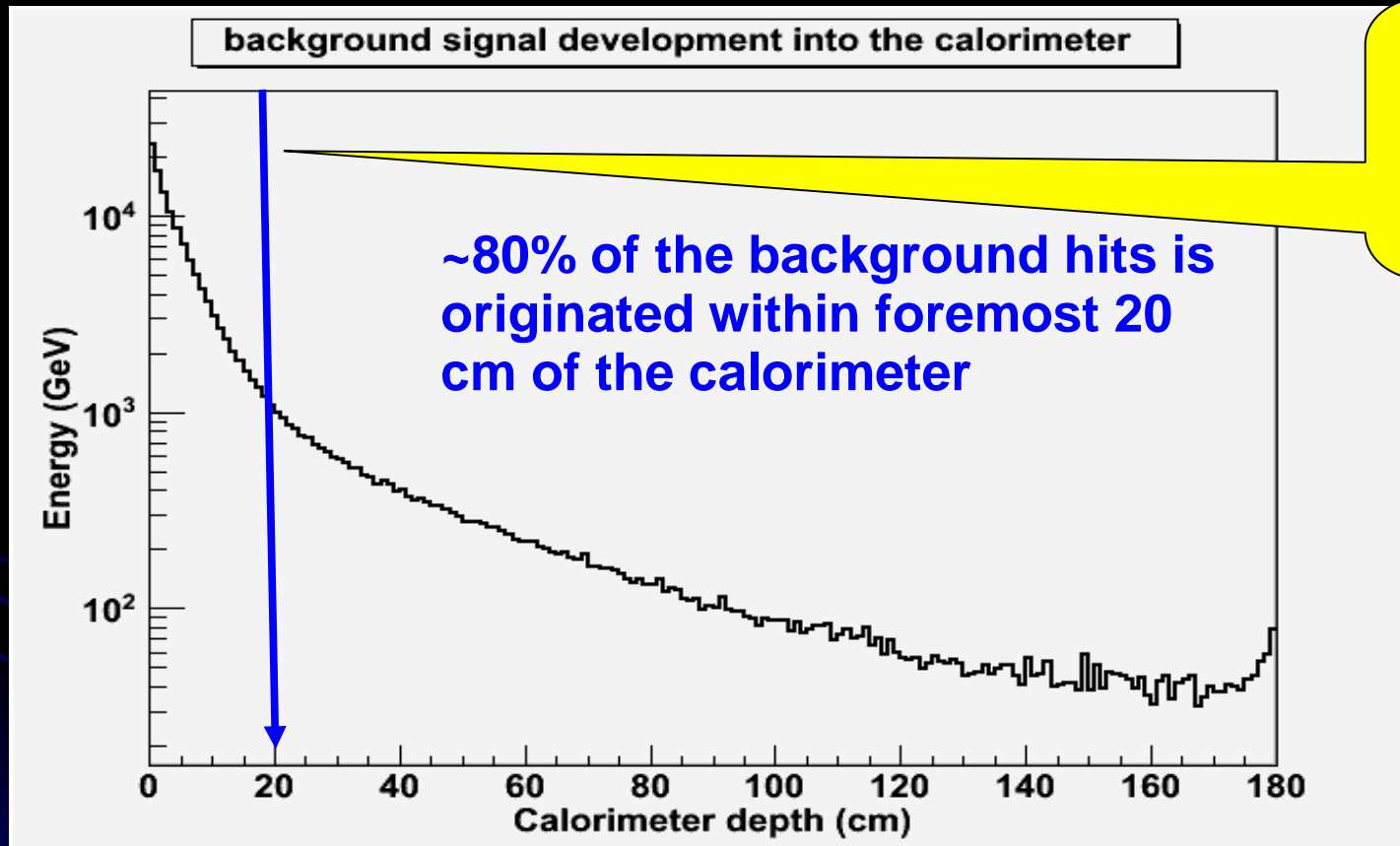


of particle (1bin= 30cm)

Z=7.5 means that the particle originated outside the MDI separation plane (1bin=16cm)

6° Nose

First step: look at the longitudinal energy deposition in the calorimeter produced by 1 background event



Longitudinal segmentation of the calorimeter is mandatory

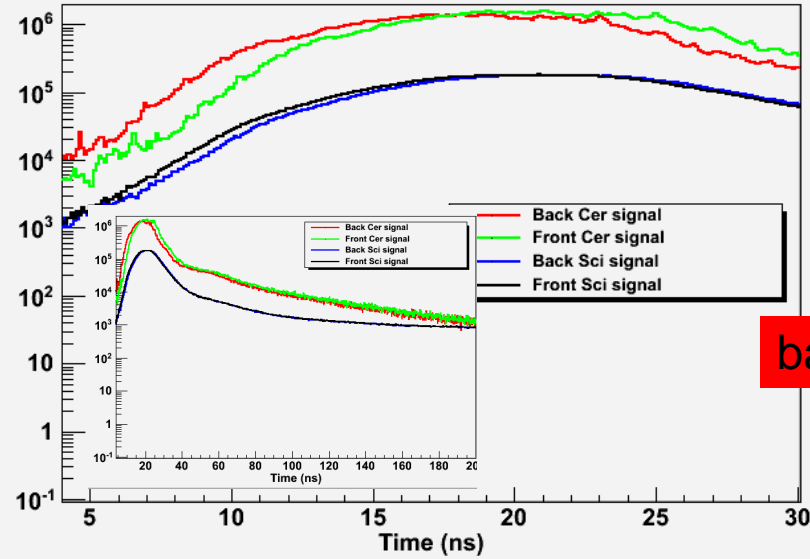
Minimal segmentation at μ Collider requires a 20cm front section to stop the overwhelming EM background

Next step: Time distribution of signal vs background

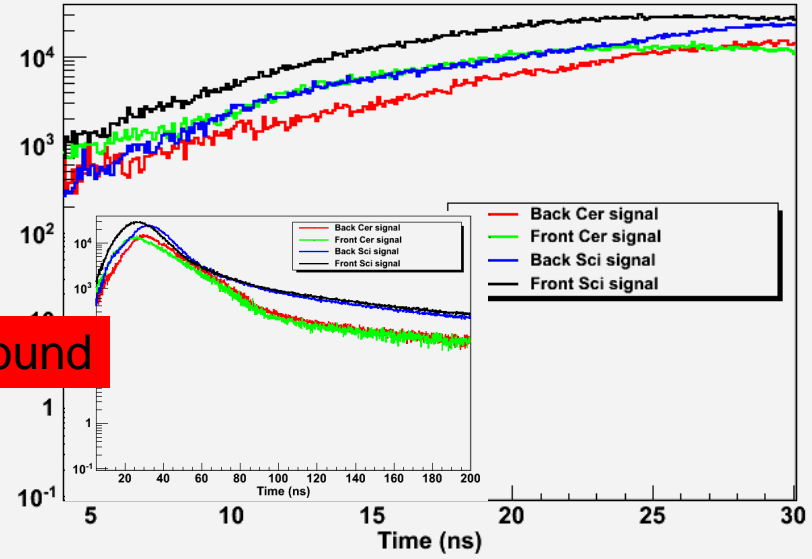
Front calorimeter section

Front calorimeter section

Average time distribution for μC bkg (Front Calorimeter Section)

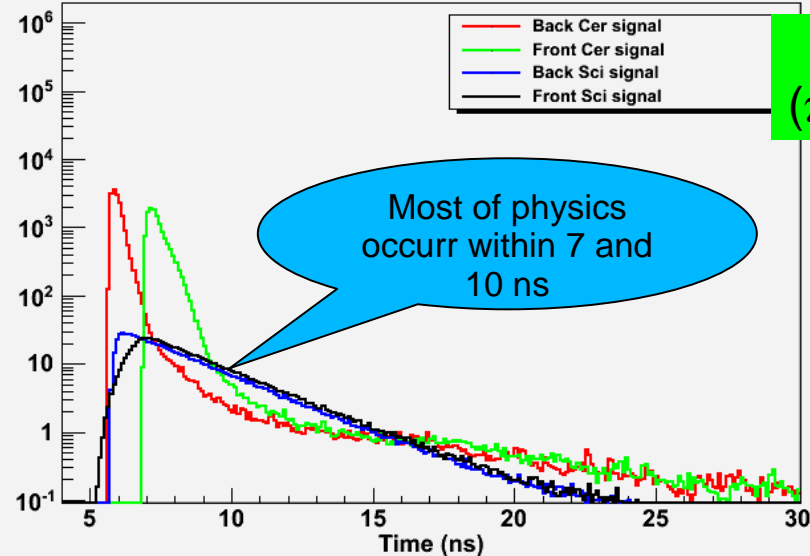


Average time distribution for μC bkg (Rear Calorimeter Section)



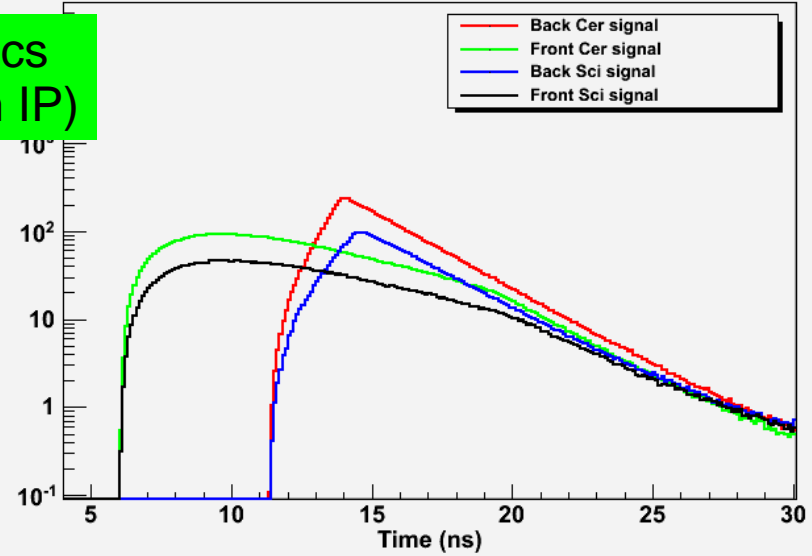
background

Average time distribution for π^- @ 40 GeV (Front Calorimeter Section)



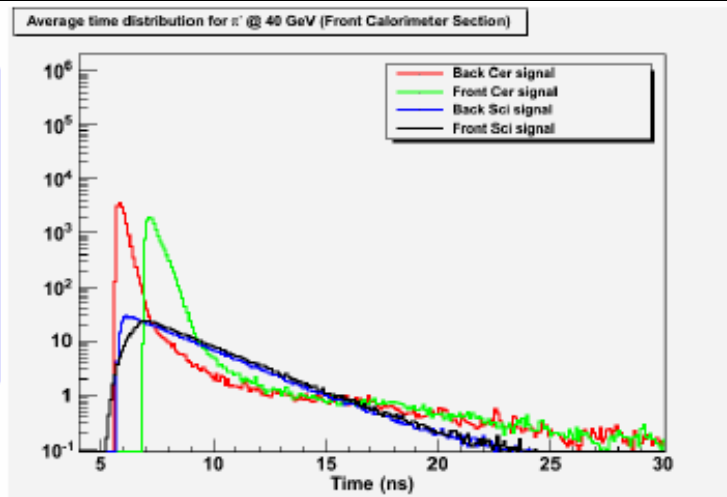
Physics
(π from IP)

Average time distribution for π^- @ 40 GeV (Rear Calorimeter Section)

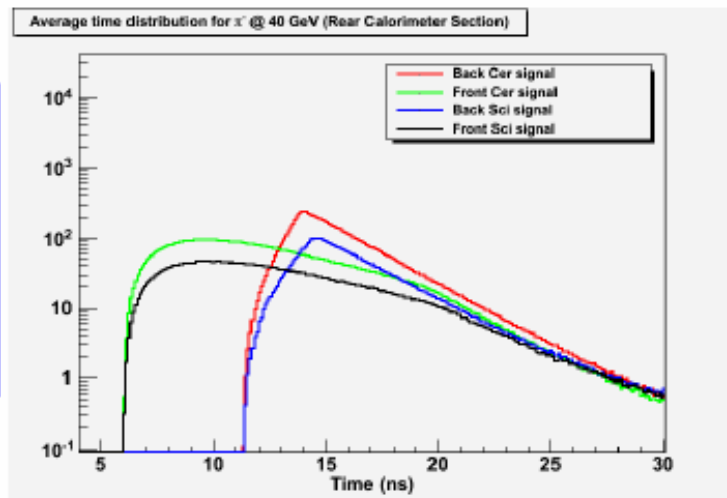


Three Detector Configurations under Study

Front Section



Rear Section



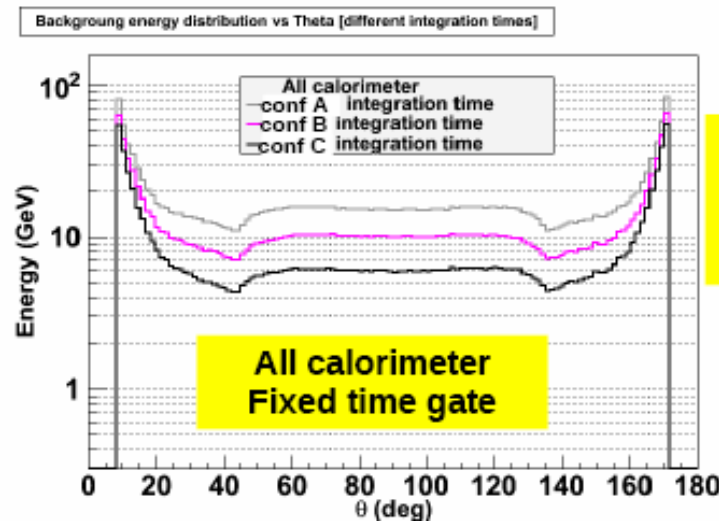
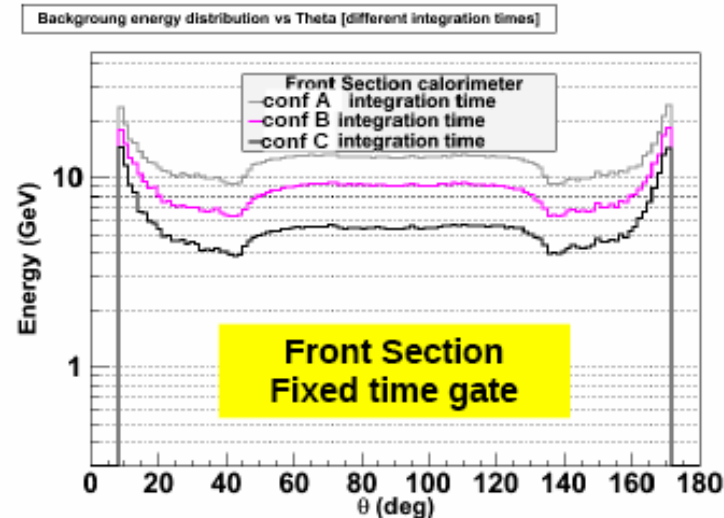
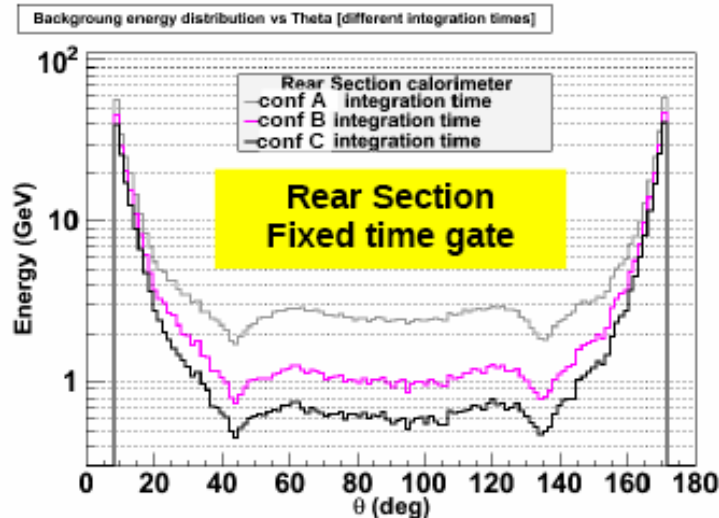
Integration time gate for each section

	Front Section		Rear Section	
	Scint	Cer	Scint	Cer
conf A	100 ns	100 ns	100 ns	100 ns
conf B	20 ns	15 ns	25 ns	25 ns
conf C	15 ns	6 ns	22 ns	22 ns

- In **conf B** 95% of the signal is integrated
- In **conf C** 90% of the signal is integrated

Angular distribution of background for fixed integration times

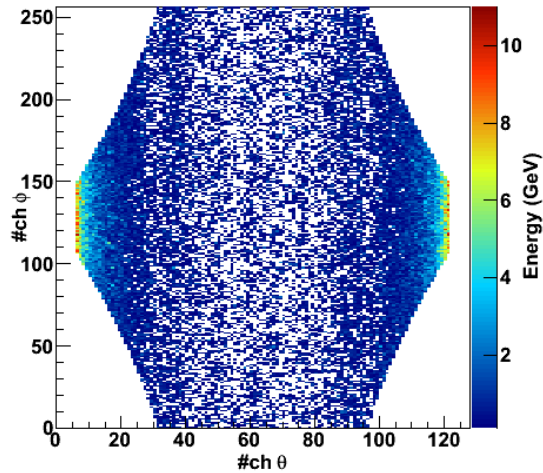
1 entry = <1 tower>



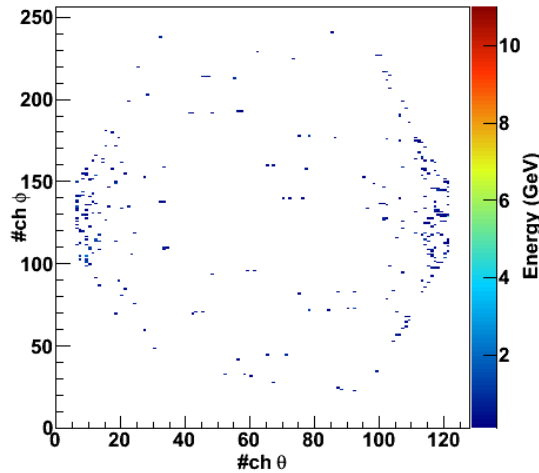
The background reduction is higher going from 25 to 15 ns than going from 100 to 25 ns
Note: the background energy peak is between 20 – 35 ns

Energy distribution per tower. Calorimeter Front Section Integration time gate 15 ns.

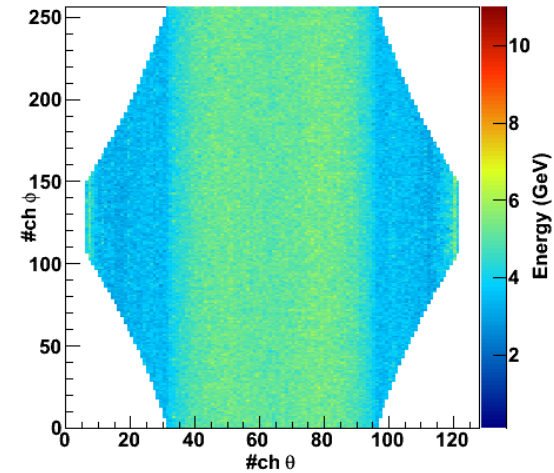
neutrons



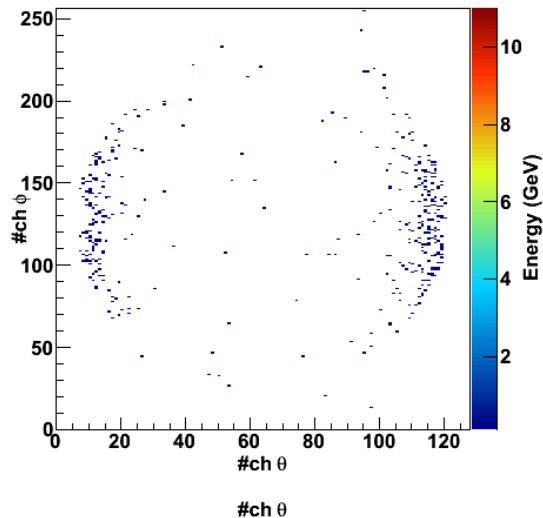
muons



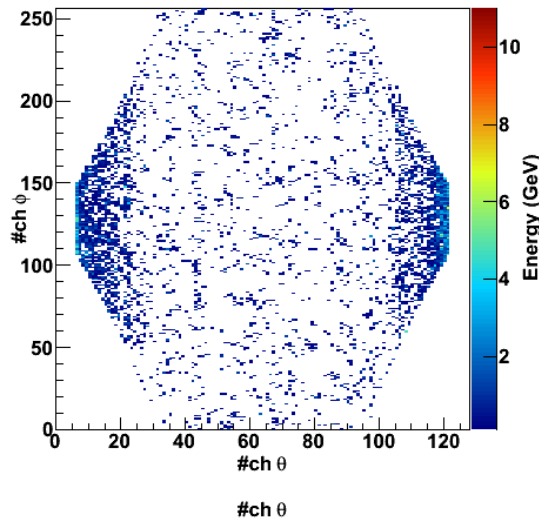
gammas



electrons



others

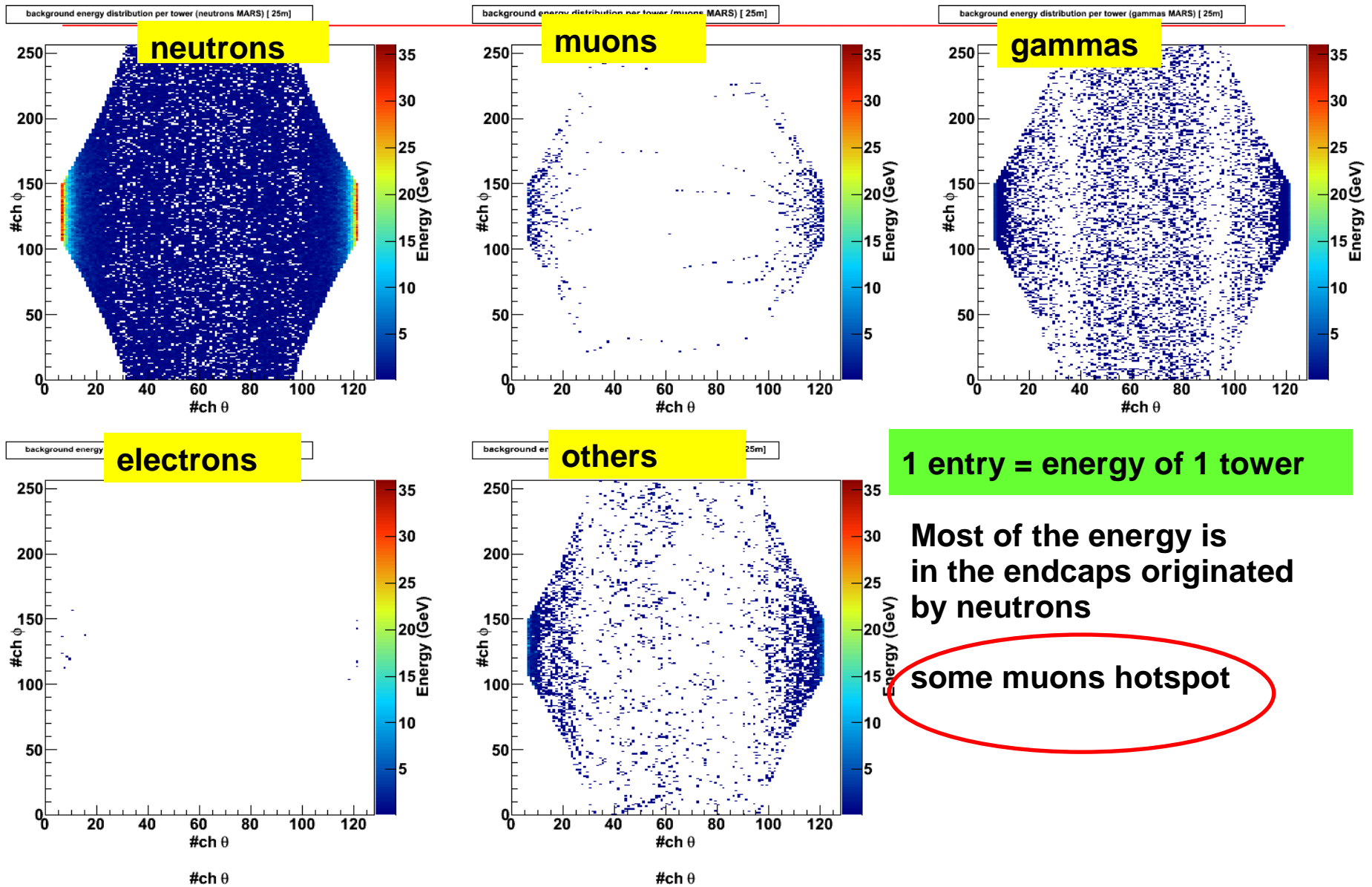


1 entry = energy of 1 tower

Most of the energy is in the endcaps originated by Neutrons and in the barrel originated by gammas

With shorter integration time gate background is reduced

Energy distribution per tower. Calorimeter Rear Section Integration time gate 15 ns.



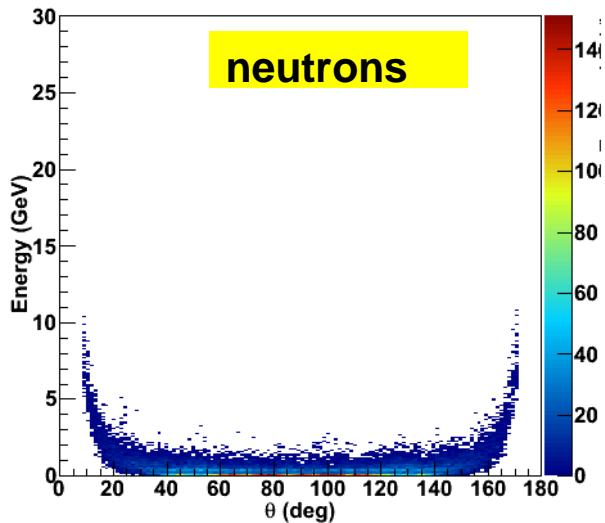
1 entry = energy of 1 tower

Most of the energy is in the endcaps originated by neutrons

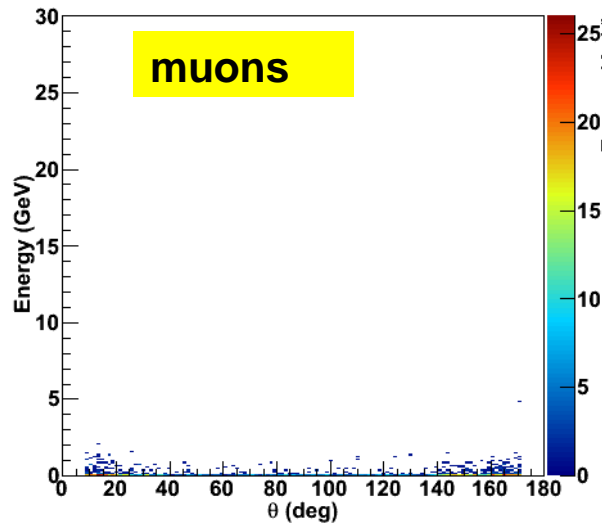
some muons hotspot

Energy distribution per tower vs theta. Calorimeter Front Section; Integration time gate 15 ns

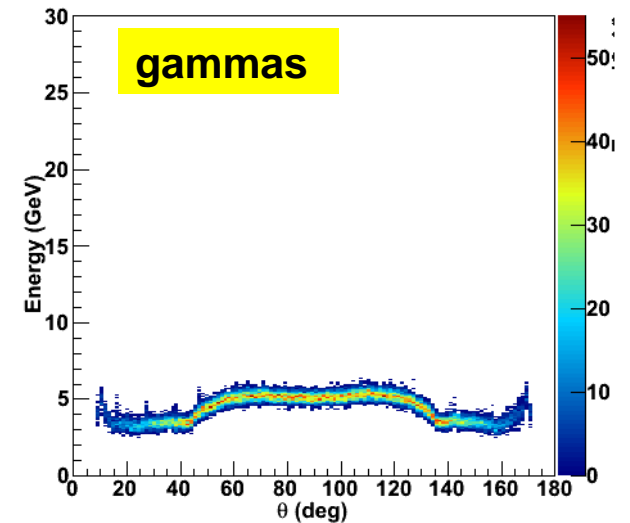
background energy distribution per tower (neutrons MARS) [25m]



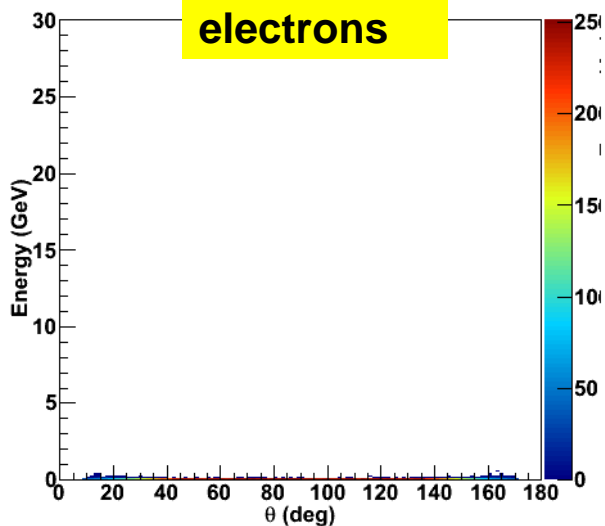
background energy distribution per tower (muons MARS) [25m]



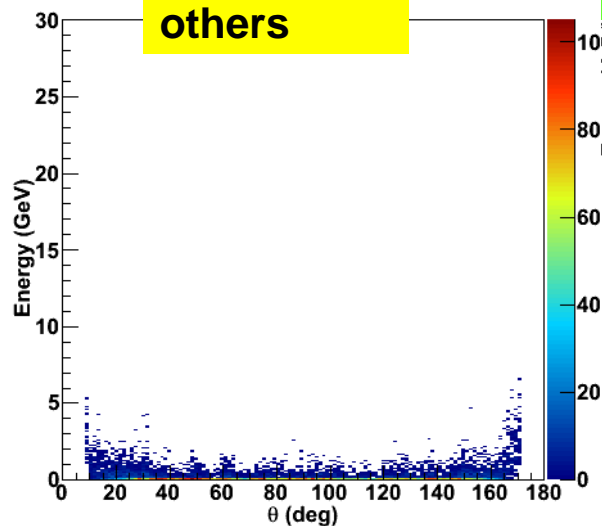
background energy distribution per tower (gammas MARS) [25m]



background energy distribution per tower (electrons MARS) [25m]



background energy distribution per tower (others MARS) [25m]

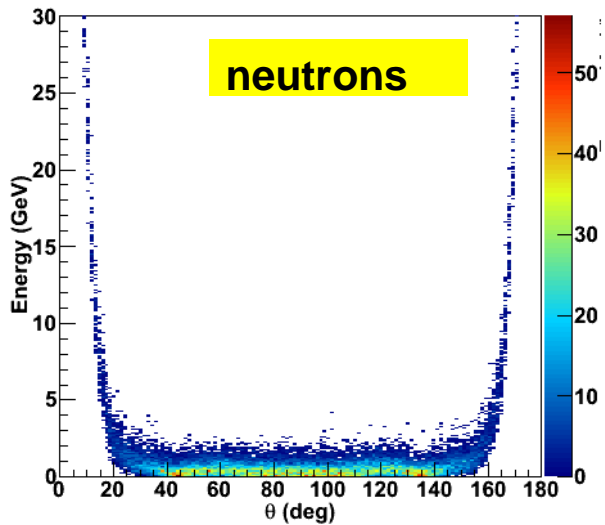


1 entry = energy of 1 tower
averaged over ϕ

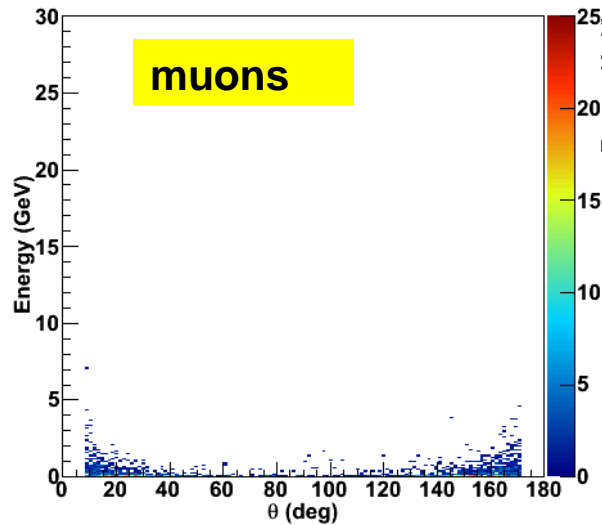
With background particles
within 25 m mostly neutrons
and gammas contribute to
energy into the calorimeter

Energy distribution per tower vs theta. Calorimeter Rear Section; Integration time gate 15 ns

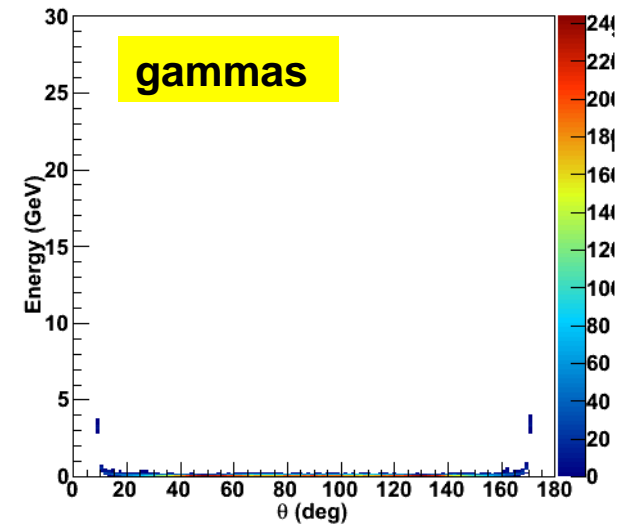
background energy distribution per tower (neutrons MARS) [25m]



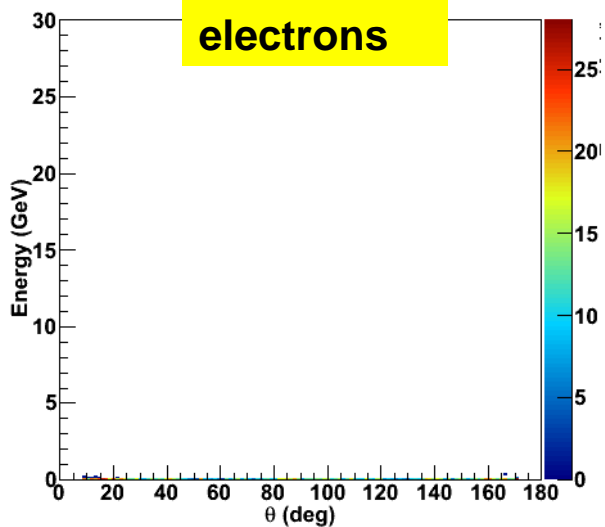
background energy distribution per tower (muons MARS) [25m]



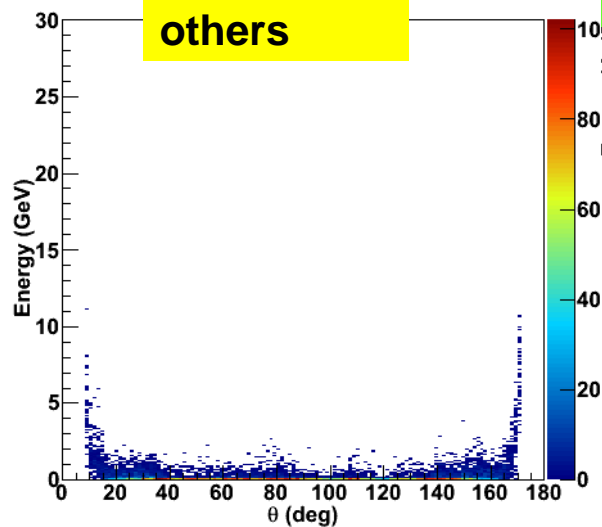
background energy distribution per tower (gammas MARS) [25m]



background energy distribution per tower (electrons MARS) [25m]



background energy distribution per tower (others MARS) [25m]

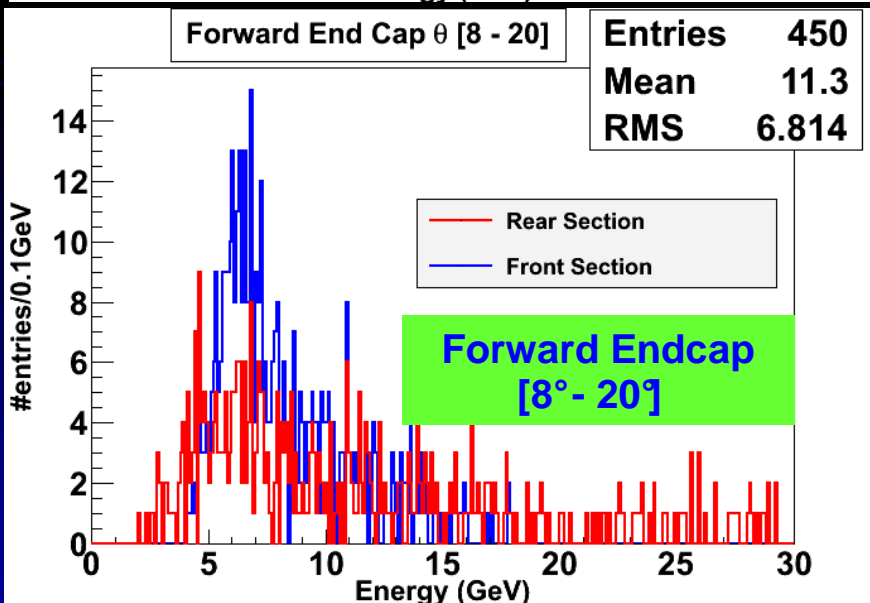
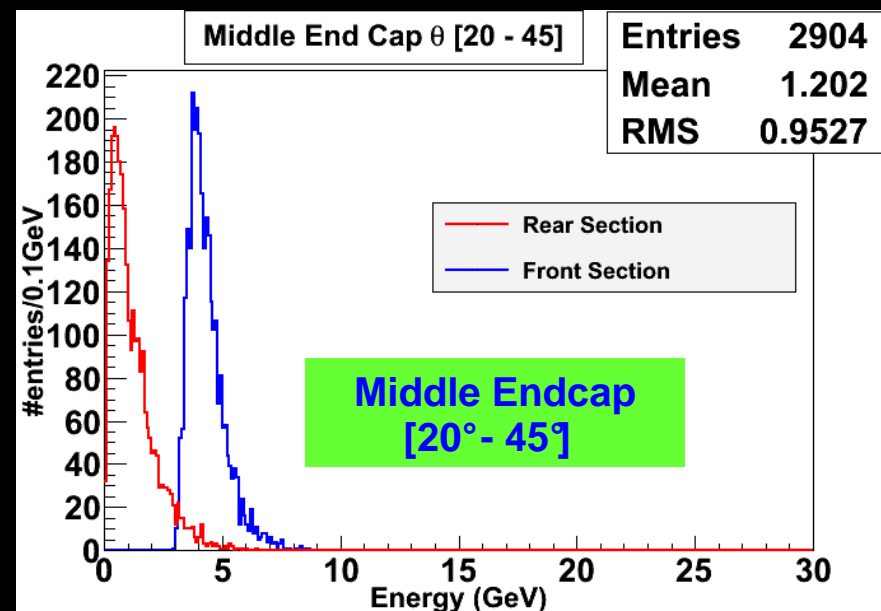
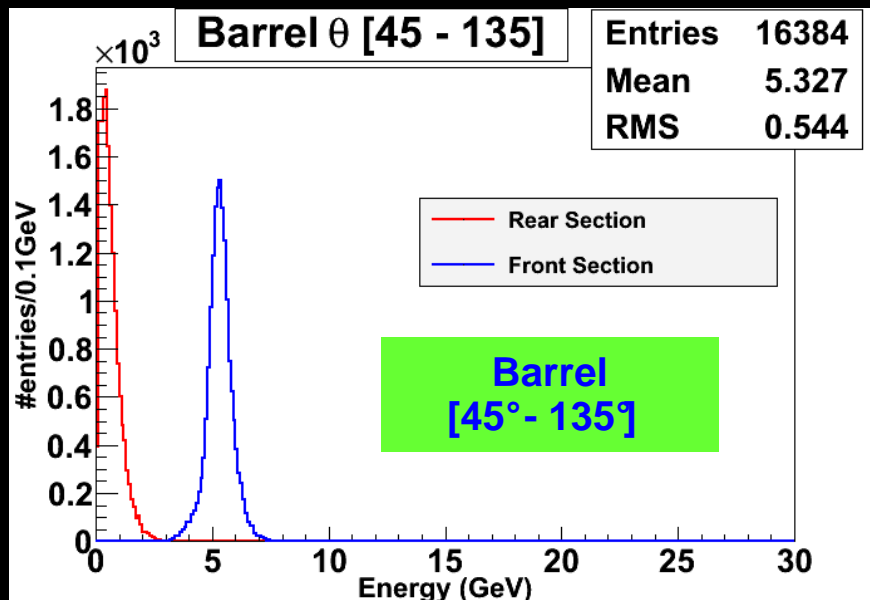


1 entry = energy of 1 tower
Averaged over ϕ

With background particles
within 25 m mostly **neutrons**
contribute to energy into the
calorimeter

SUMMARY: Energy per tower in Front and Rear Section

Integration time 15 ns



In the front section most of contribution is from gammas and neutrons

Background energy fluctuation		
Energy (GeV)	Front	Rear
Barrel	5.33±0.54	0.63±0.43
Mid ECap	4.33±0.79	1.20±0.95
Forward ECap	8.31±2.94	11.3±6.8

Effect on a typical 150 GeV jet

Assume 16-25 towers interested

	RMS(E)	Contribution to resolution
Barrel [45° - 135°]	2.5-3 GeV	12%/√E
Middle Endcap [20° - 45°]	5-6 GeV	25%/√E
Fwd Endcap [8° - 20°]	20-25 GeV	100%/√E

Some Considerations on Calorimetry for a Muon Collider

A $30\%/ \sqrt{E}$ calorimeter (stochastic) or RMS90 = 3% detector (non stochastic) have adequate energy resolution for a lepton collider with $E_{cm} = 500$ GeV or above

However, the higher E_{cm} the more important jet axis angular resolution becomes over energy resolution -> high granularity

Two-sections ADRIANO calorimeter is adequate for physics and resilient to background at $\theta > 20^\circ$. But:

- Must use fast (3 nsec) fibers

- Must have short (15 nsec) integration times and sophisticated time gate

Below $\theta = 20^\circ$ the present solution should be improved (more segmentation?).

The very problem are fluctuations in $E_{background}$

T1004 and T1015 at Fermilab are also looking for total active technique with crystals and heavy scintillating glasses.

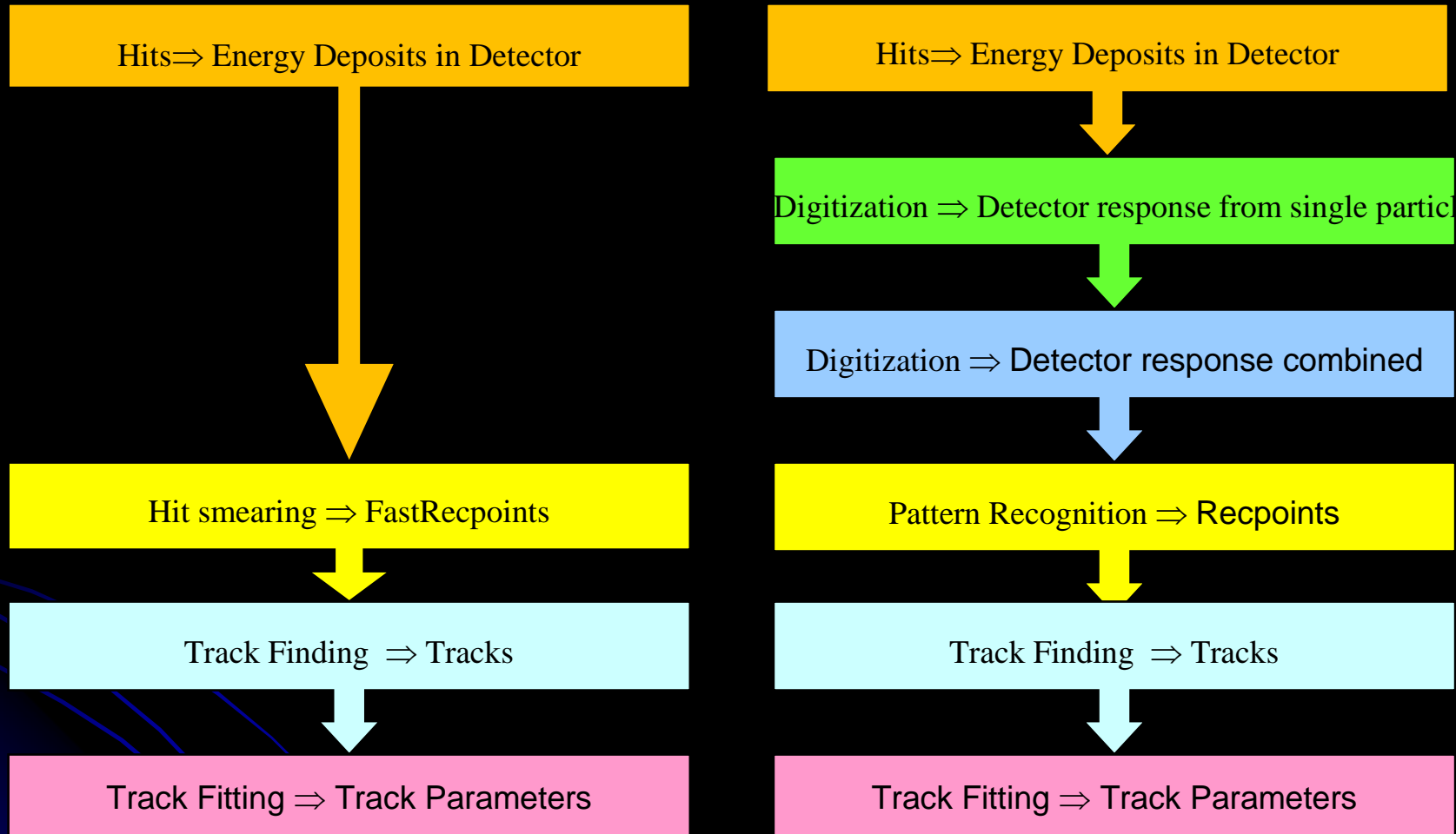
However, both such materials are intrinsically slow ($t > 30$ nsec) and probably inadequate at a muon collider.

Conclusions

- A ILCroot-based simulation effort is ongoing at Fermilab: the goal is understanding the Physics capabilities and optimal technology choices
- Background is very nasty even with 10° tungsten nozzle, but fully understood
- Tracking systems requires very fast technologies (sub nsec) to reject out of time background
- Several solutions for tracking are being considered:
 - 3-D Si-tracker with precision timing
 - 4-D Kalman filter
- Calorimetry (being an integrating device) is far more subject to background.
- What really matters are fluctuations around a pedestal.
 - A two-section calorimeter with appropriate time gate has been considered
 - It works fine down to 20 degrees
 - More effort is needed to fight background at smaller angles

Backup slides

Fast vs Full Simulation



Fast vs Full Simulation

Hits \Rightarrow Energy Deposits in Detector



Hit smearing \Rightarrow FastRecpoints



Track Finding \Rightarrow Tracks



Track Fitting \Rightarrow Track Parameters

Same as a detector with perfect pattern recognition

Hits \Rightarrow Energy Deposits in Detector



Digitization \Rightarrow Detector response from single particles



Digitization \Rightarrow Detector response combined



Pattern Recognition \Rightarrow Recpoints



Track Finding \Rightarrow Tracks

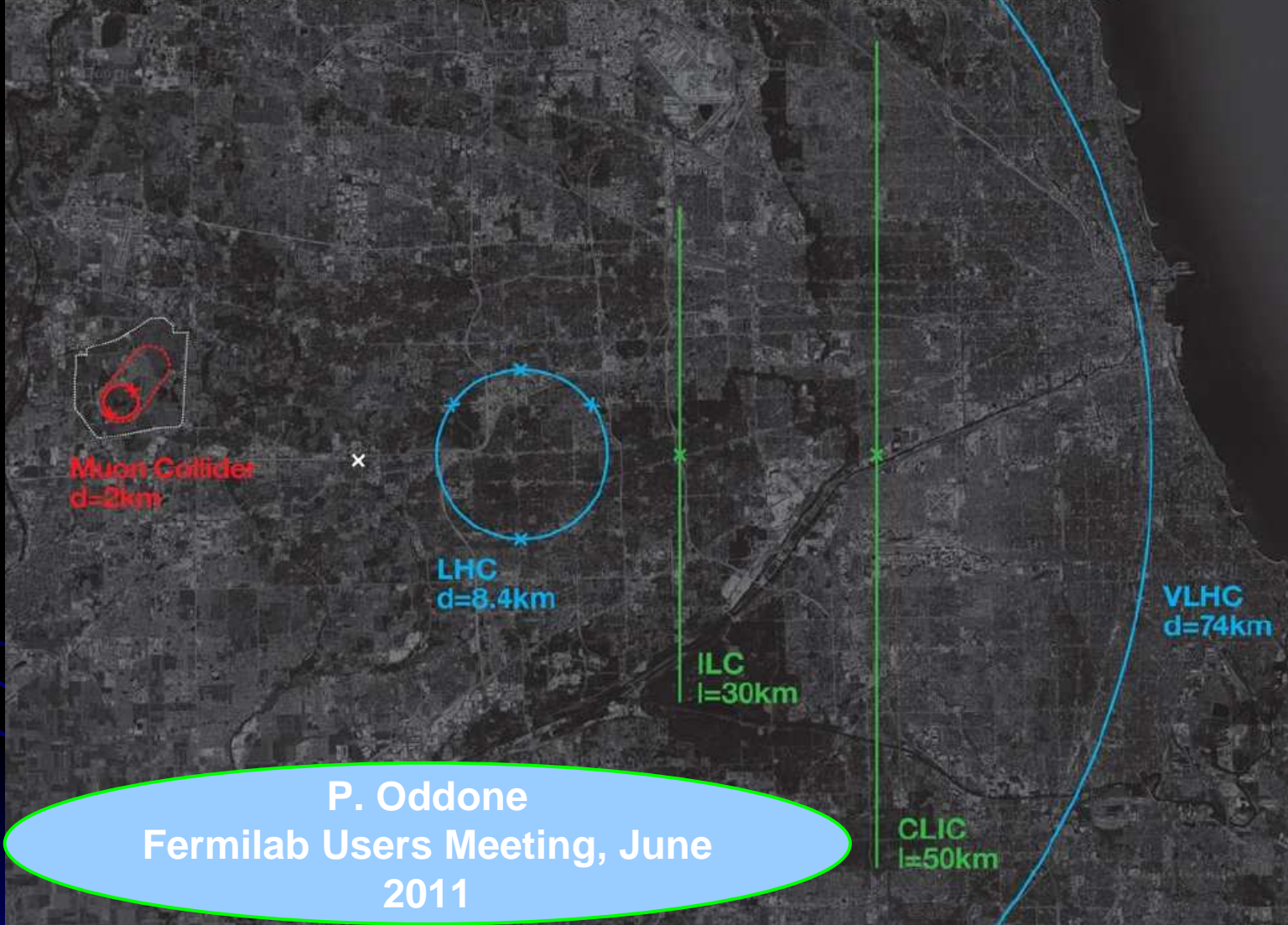


Track Fitting \Rightarrow Track Parameters

Used for most studies in this talk

Comparison of Particle Colliders

To reach higher and higher collision energies, scientists have built and proposed larger and larger machines.



P. Oddone
Fermilab Users Meeting, June
2011

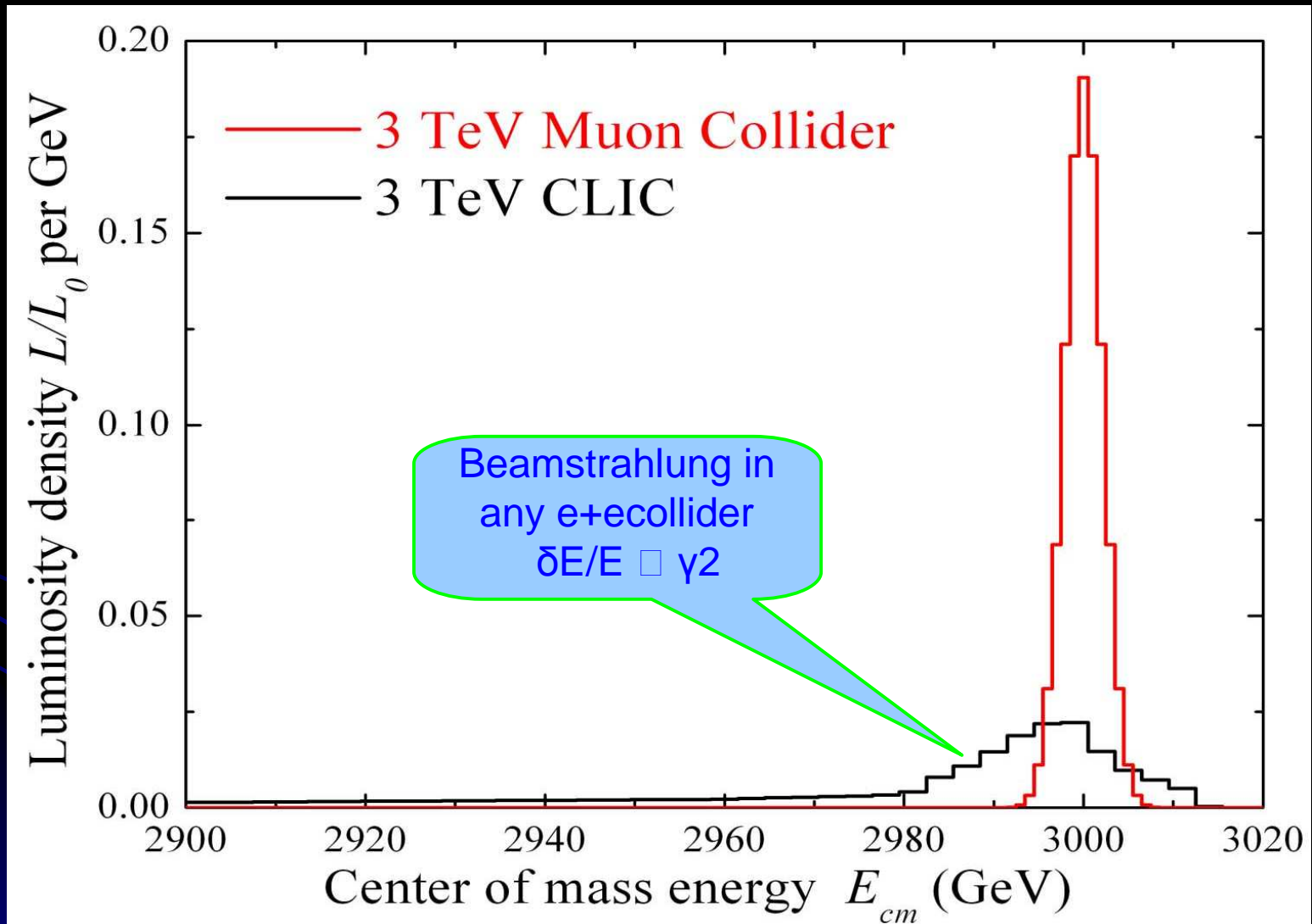
MUON COLLIDER MOTIVATION

If we can build a muon collider, it is an attractive multi-TeV lepton collider option because muons don't radiate as readily as electrons ($m_\mu / m_e \sim 207$):

- **COMPACT**
Fits on laboratory site
- **MULTI-PASS ACC**
Cost Effective operation & construction
- **MULTIPASS COLLISIONS IN A RING** (~1000 turns)
Relaxed emittance requirements & hence relaxed tolerances
- **NARROW ENERGY SPREAD**
Precision scans, kinematic constraints
- **TWO DETECTORS (2 IPs)**
- $\Delta T_{\text{bunch}} \sim 10 \mu\text{s} \dots$ (e.g. 4 TeV collider)
Lots of time for readout
Backgrounds don't pile up
- $(m_\mu/m_e)^2 = \sim 40000$
Enhanced s-channel rates for Higgs-like particles

S. Geer- Accelerator
Seminar
SLAC 2011

Energy Spread



Challenges

- Muons are produced as tertiary particles.

To make enough of them we must start with a MW scale proton source & target facility.

- Muons decay

Everything must be done fast and we must deal with the decay electrons (& neutrinos for CM energies above ~ 3 TeV).

- Muons are born within a large 6D phase-space.

For a MC we must cool them by $O(10^6)$ before they decay □ New cooling technique (ionization cooling) must be demonstrated, and it requires components with demanding performance (NCRF in magnetic channel, high field solenoids.)

- After cooling, beams still have relatively large emittance.

S. Geer- Accelerator
Seminar
SLAC 2011

Fast simulation and/or fast digitization also available in ILCroot for tracking system

- Fast Simulation = hit smearing
- Fast Digitization = full digitization with fast algorithms
- Do we need fast simulation in tracking studies?

Yes!

- Calorimetry related studies do not need full simulation/digitization for tracking
- Faster computation for quick answer to response of several detector layouts/shielding

- Do we need full simulation in tracking studies?

Yes!

- Fancy detector and reconstruction needed to be able to separate hits from signal and background

Technologies Implemented

- 3 detector species:
 - Silicon pixels
 - Silicon Strips
 - Silicon Drift
- Pixel can have non constant size in different layers
- Strips can also be stereo and on both sides
- Dead regions are taken into account
- Algorithms are parametric: almost all available technologies are easily accomodated (MAPS, 3D, DEPFET, etc.)

Used for VXD SiT and FTD
in present studies

Track Fitting in ILCRoot

Track finding and fitting is a global task: individual detectors collaborate

It is performed after each detector has completed its local tasks (simulation, digitization, clusterization)

It occurs in three phases:

- Seeding in SiT and fitting in VXD+SiT+MUD
- Standalone seeding and fitting in VXD
- Standalone seeding and fitting in MUD

• Two different seedings:

- Primary seeding with vertex constraint
- Secondary seeding without vertex constraint

Not yet implemented

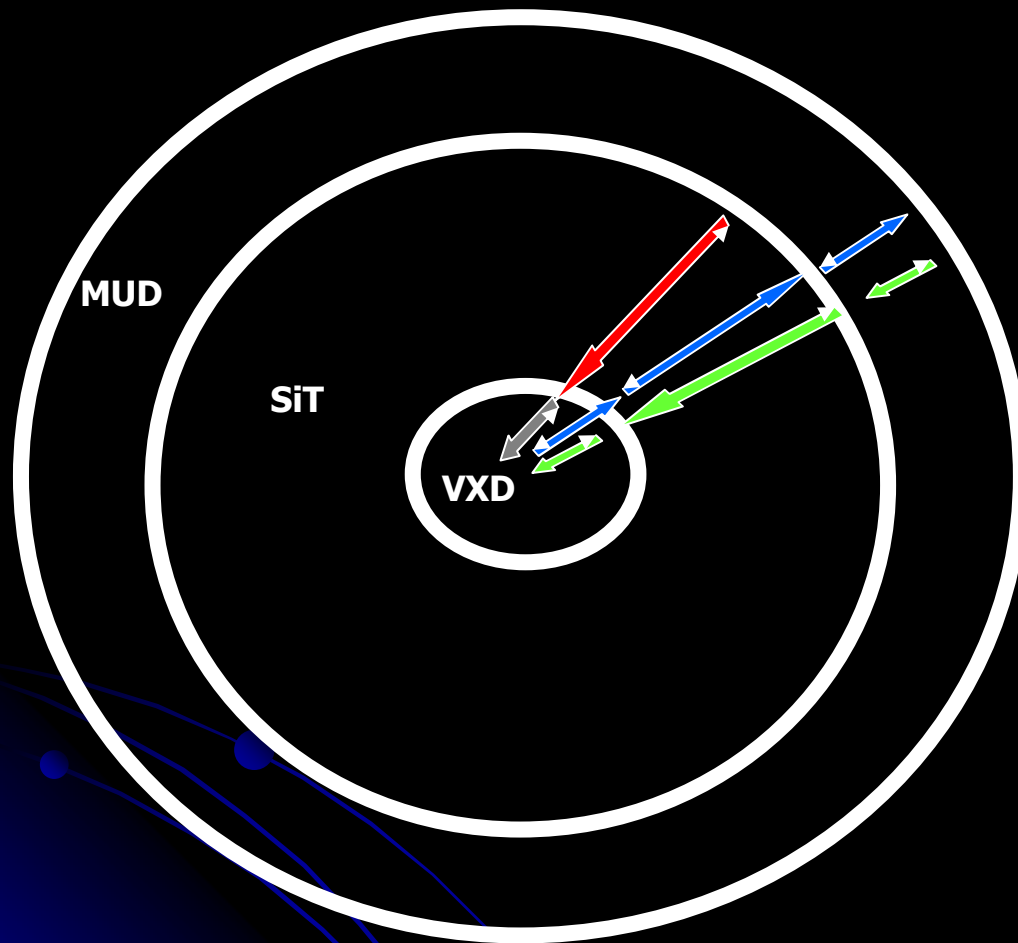
Kalman Filter (classic)

- Recursive least-squares estimation.
- Equivalent to global least-squares method including all correlations between measurements due to multiple scattering.
- Suitable for combined track finding and fitting
- Provides a natural way:
 - to take into account multiple scattering, magnetic field inhomogeneity
 - possibility to take into account mean energy losses
 - to extrapolate tracks from one sub-detector to another

Parallel Kalman Filter

- Seedings with constraint + seedings without constraint at different radii (necessary for kinks and V0) from outer to inner
- Tracking
 - Find for each track the prolongation to the next layer
 - Estimate the errors
 - Update track according current cluster parameters
 - (Possible refine clusters parameters with current track)
- Track several track-hypothesis in parallel
 - Allow cluster sharing between different track
- Remove-Overlap
- **Kinks and V0** fitted during the Kalman filtering

Tracking Strategy – Primary Tracks



- Iterative process
 - **Seeding in SiT**
 - Forward propagation towards to the vertex
 $\text{SiT} \rightarrow \text{VXD}$
 - Back propagation towards to the MUD
 $\text{VXD} \rightarrow \text{SiT} \rightarrow \text{MUD}$
 - Refit inward
 $\text{MUD} \rightarrow \text{SiT} \rightarrow \text{VXD}$
- Continuous seeding –track segment finding in all detectors

VXD Standalone Tracking

- Uses Clusters leftover in the VXD by Parallel Kalman Filter
- **Requires at least 4 hits to build a track**
- Seeding in VXD in two steps
 - Step 1: look for 3 Clusters in a narrow row or 2 Clusters + IP constraint
 - Step 2: prolongate to next layers each helix constructed from a seed
- After finding Clusters, all different combination of clusters are refitted with the Kalman Filter and the tracks with lowest χ^2 are selected
- Finally, the process is repeated attempting to find tracks on an enlarged row constructed looping on the first point on different layers and all the subsequent layers
- In 3.5 Tesla B-field $P_t > 20$ MeV tracks reconstructable

Event Display

ILCroot event display
for 10 muons up to 200 GeV

green - hits
purple - reconstructed tracks
red - MC particle

10 generated muons
9 reconstructed tracks

Effects on Track Resolution

Background in the calorimeter for different particle species originating within 25 m from IP

Background in the calorimeter for different particle species originating in [25-200] m from IP

Future Prospects

Conclusions

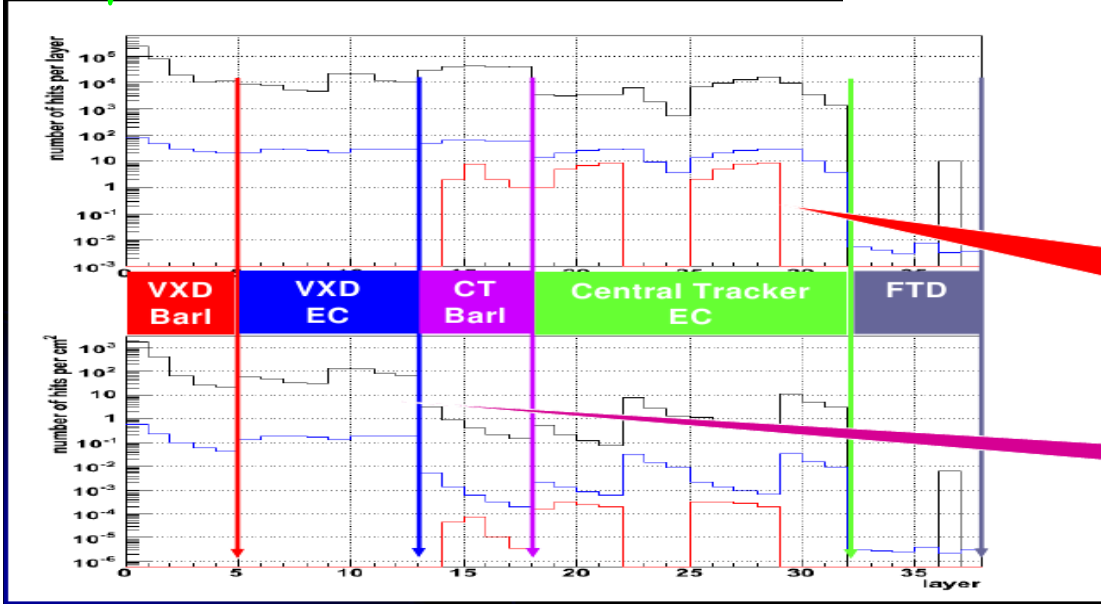
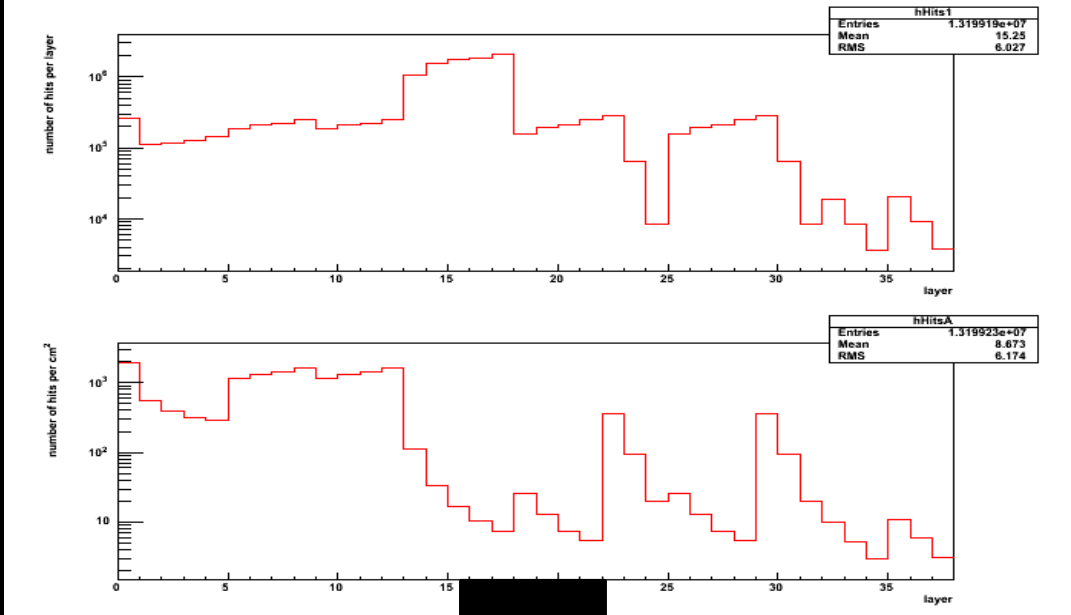
Backup slides

Occupancy in the Tracking System

NEW

Nozzle 10°

Nozzle 6°



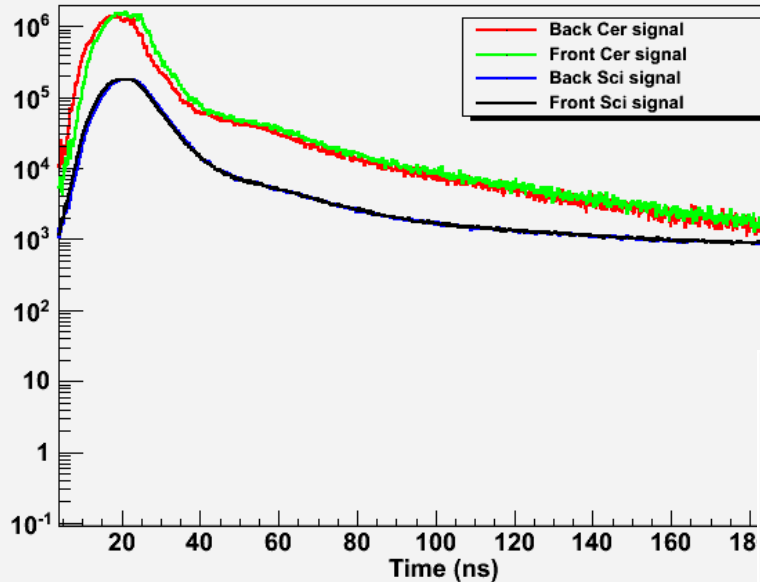
Legenda
 -WWnunu
 -Beam bkg except muons
 -muons

About 10 muons per BX (rejected easily by μ spectrometer)

5 hits/cm² at R=20 cm

Next step: look at the time distribution of the background

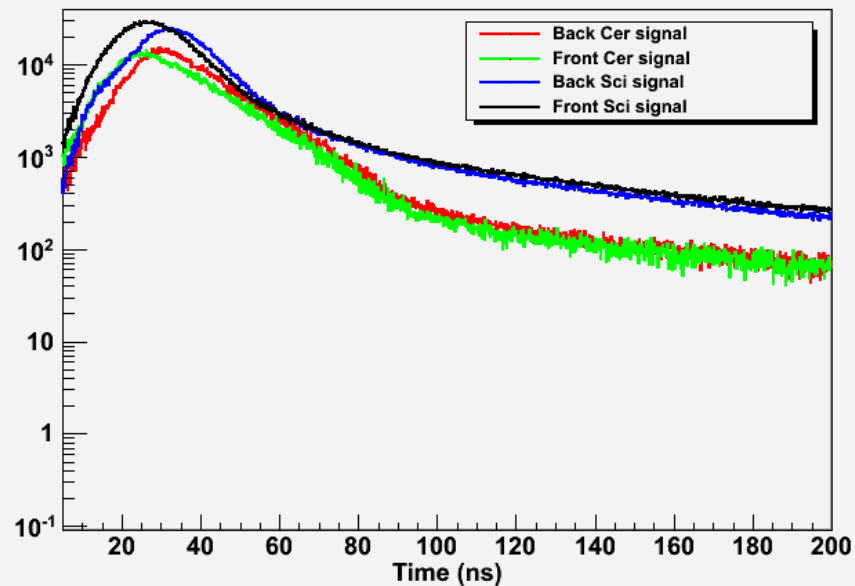
Average time distribution for μC bkg (Front Calorimeter Section)

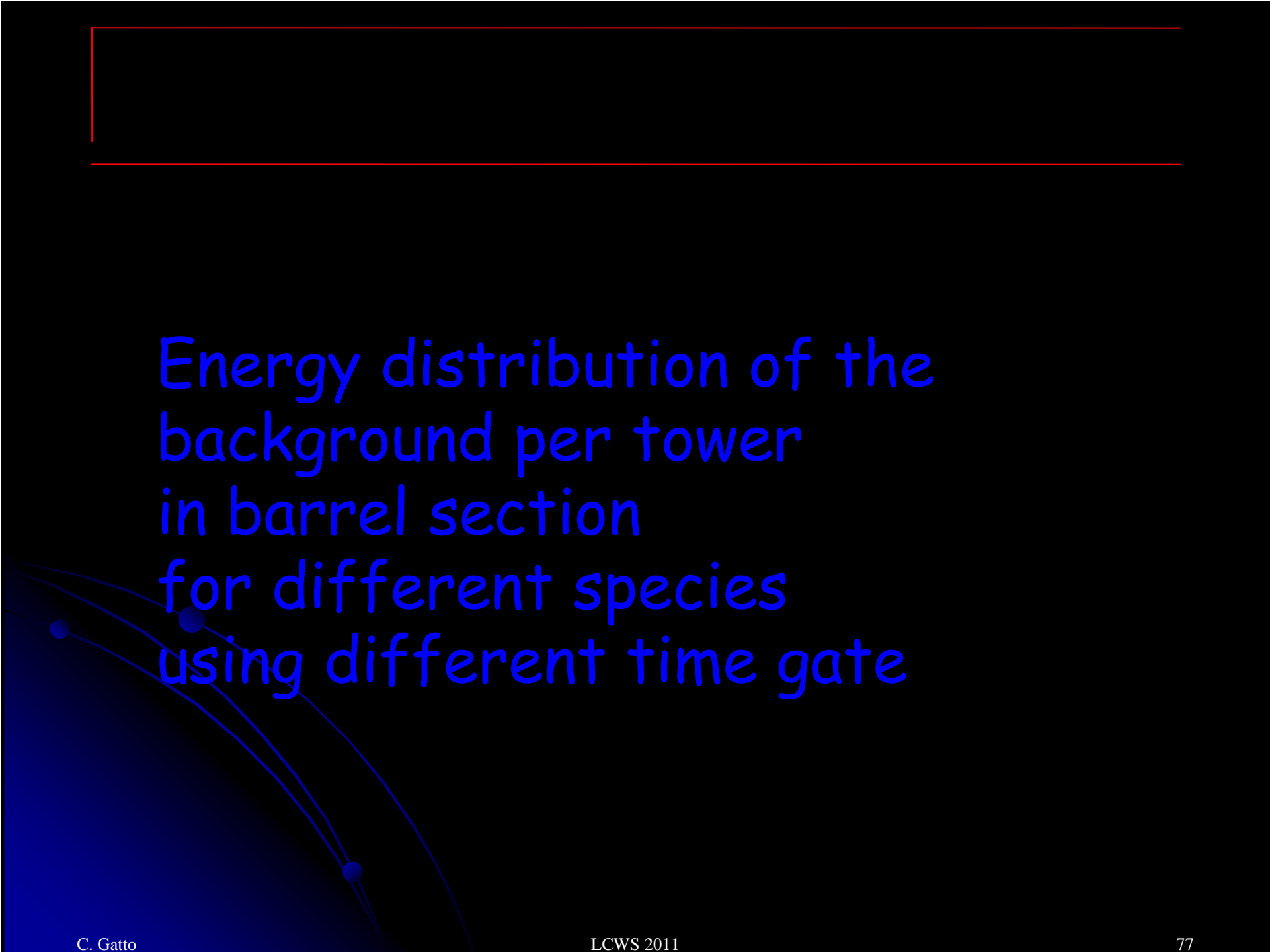


Calorimeter is now split in a forward (20cm) and rear (160 cm) section

Light propagation in fibers and lead glass is implemented in ILCroot

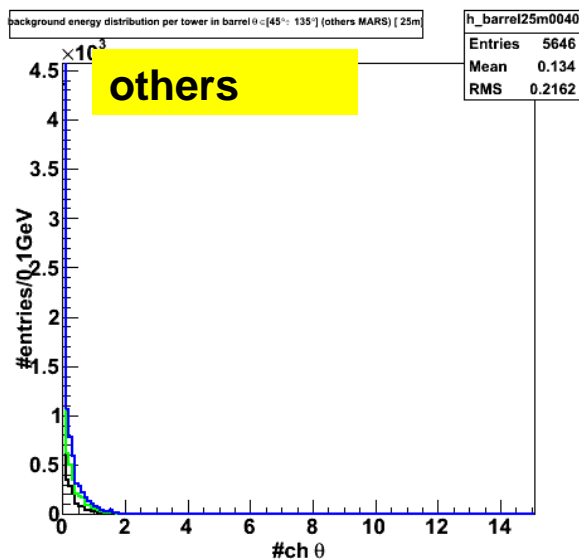
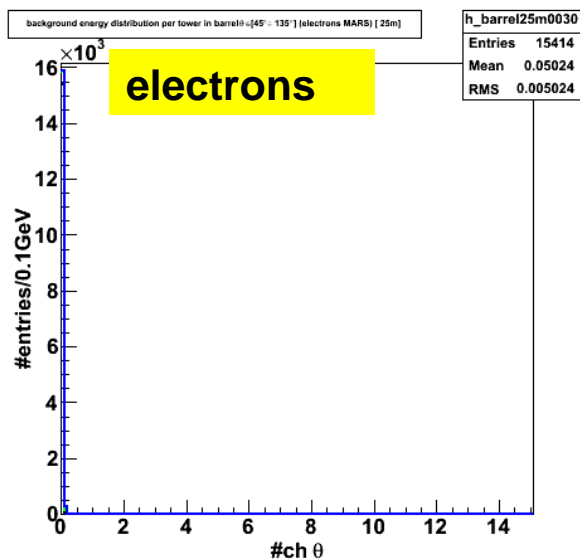
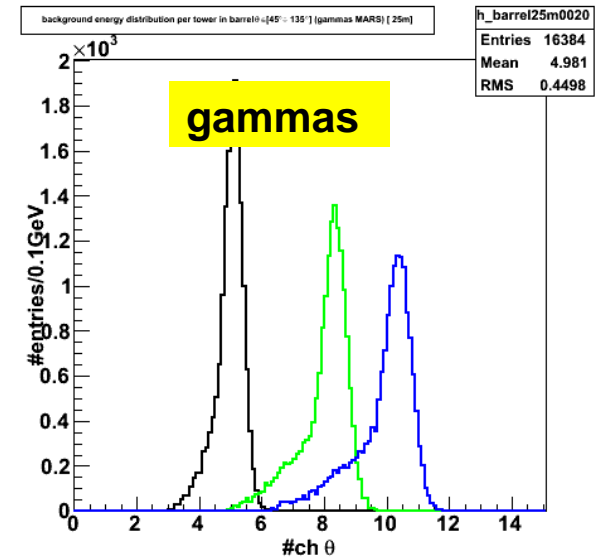
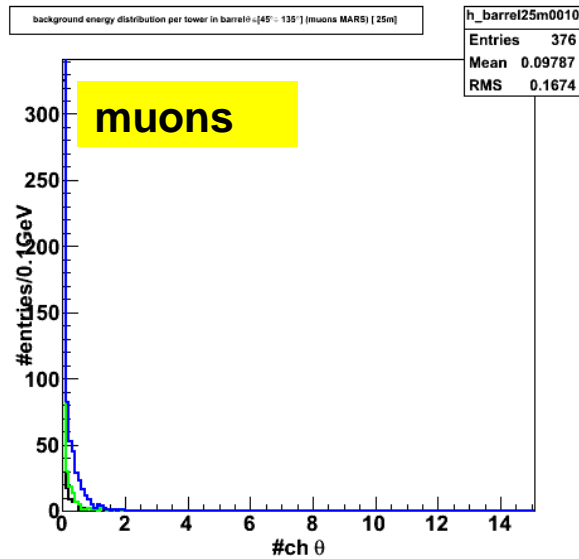
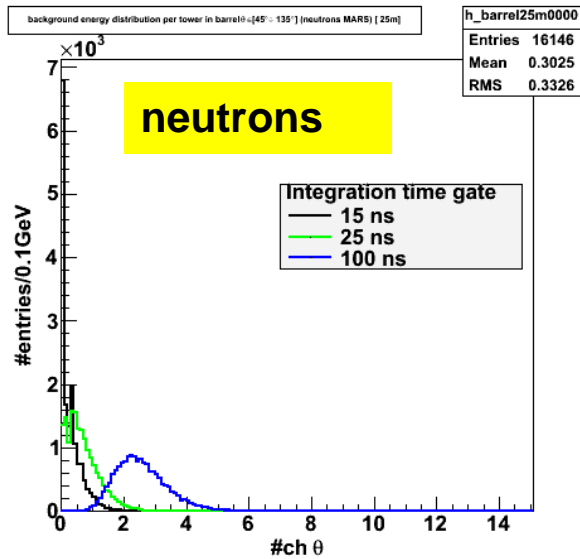
Average time distribution for μC bkg (Rear Calorimeter Section)





Energy distribution of the
background per tower
in barrel section
for different species
using different time gate

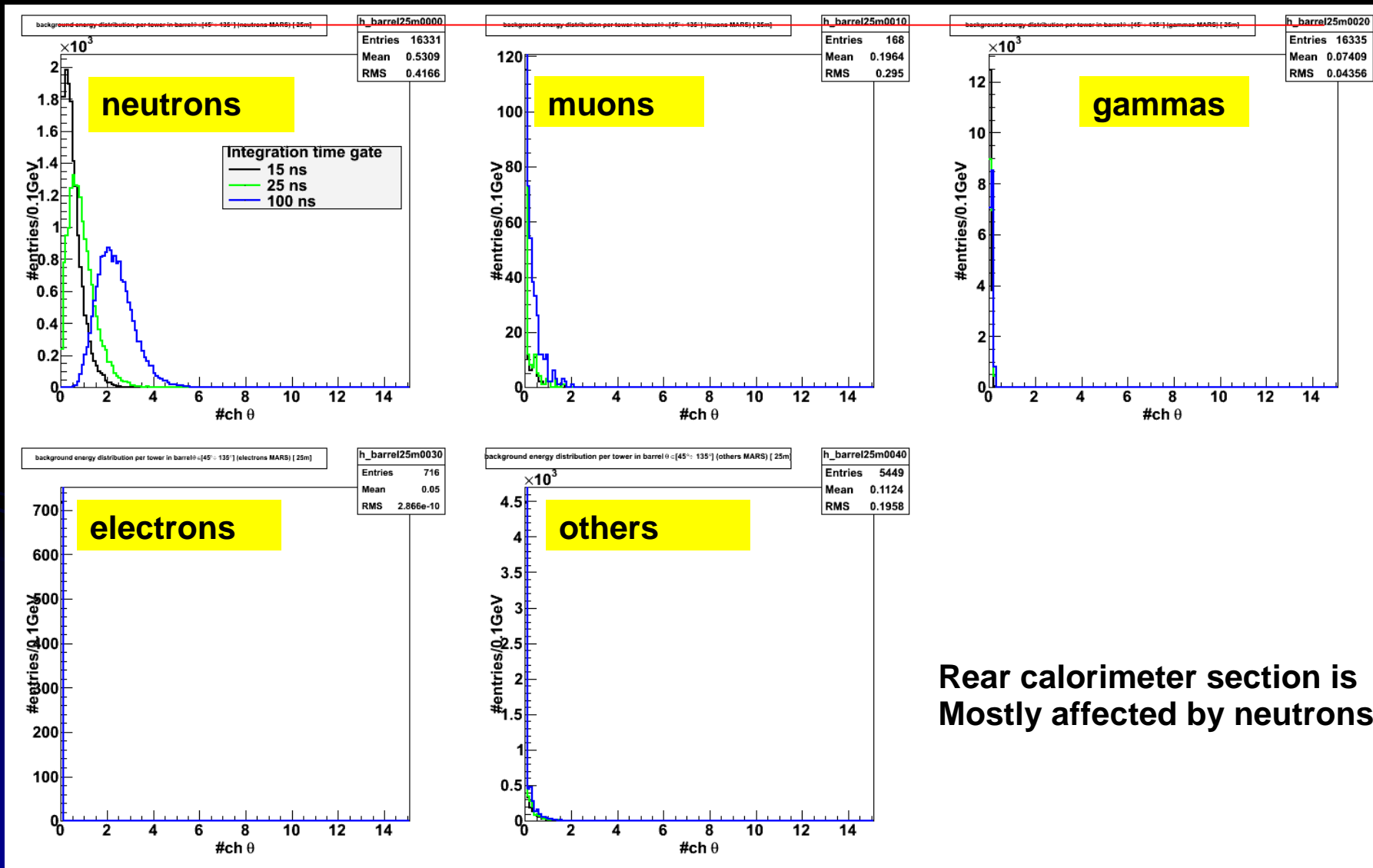
Energy distribution per tower in barrel Calorimeter Front Section for different species



With background particles within 25 m mostly neutrons and gammas contribute to energy into the calorimeter

Substantial background Reduction is obtained with Short time gates

Energy distribution per tower in barrel Calorimeter Rear Section for different species



Rear calorimeter section is Mostly affected by neutrons