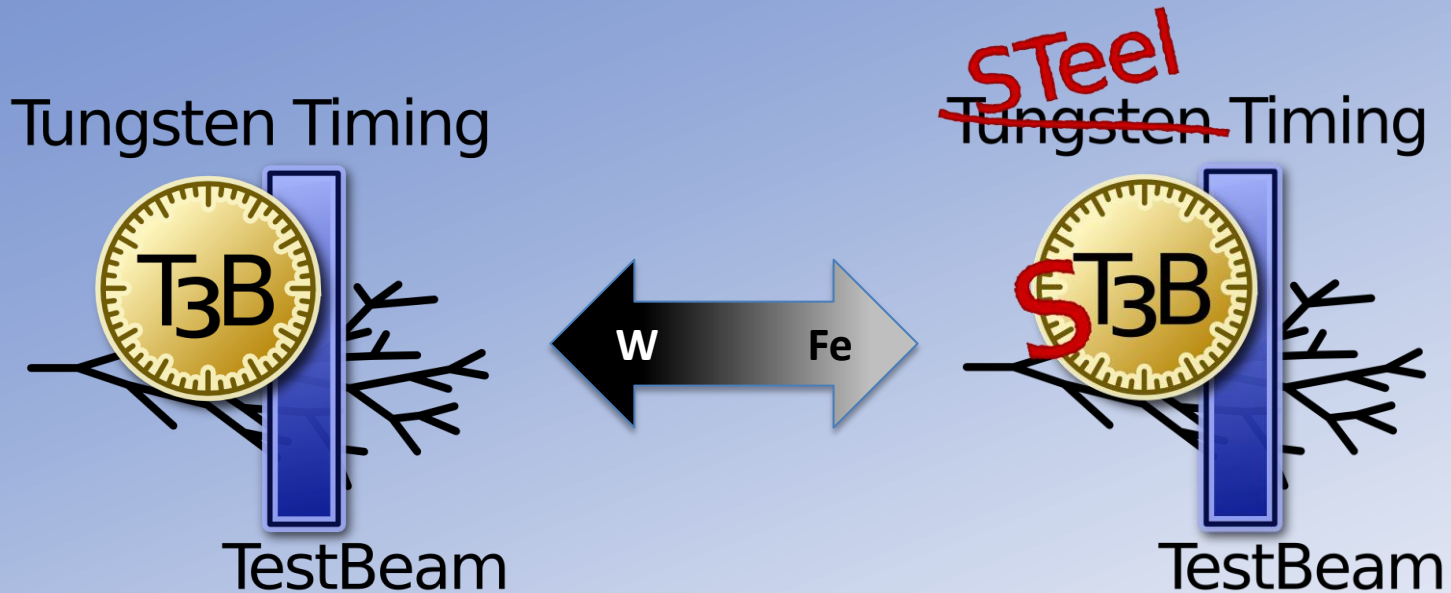


T3B Standalone Analysis

Calibration, Time Profile, Steel/Tungsten Comparison



Calice collaboration meeting – March 6th 2012 – Shinshu, Japan



Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

Christian Soldner
Max-Planck-Institute for Physics





Outline



- Introduction: CALICE and T3B
- Calibration to the MIP scale
 - Time of First Hit Analysis
- Afterpulsing Study
 - Time of Hit Analysis
- Summary and Outlook

THE CALICE CALORIMETER AND **THE T3B EXPERIMENT**



The T3B Experiment



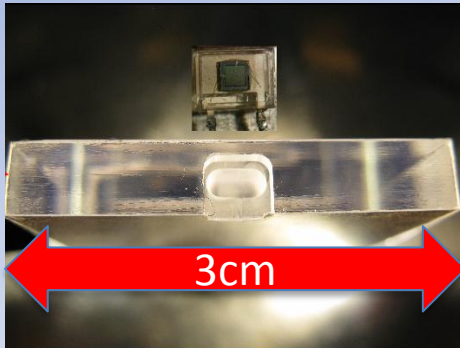
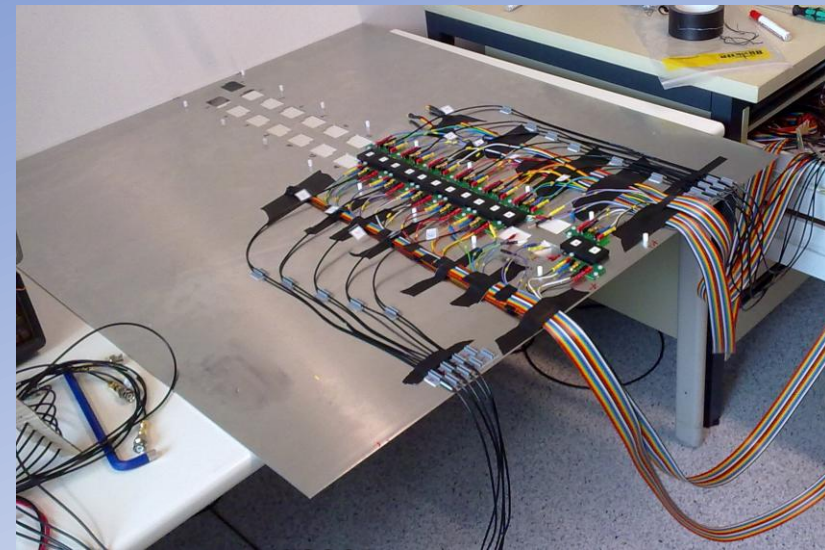
What is T3B?

- One row of 15 scintillator tiles
- Tile dimensions: $3 \times 3 \times 0.5 \text{ cm}^3$
- Light Readout by SiPMs: MPPC-50P
- Data Acquisition: 4 fast USB Oscilloscopes

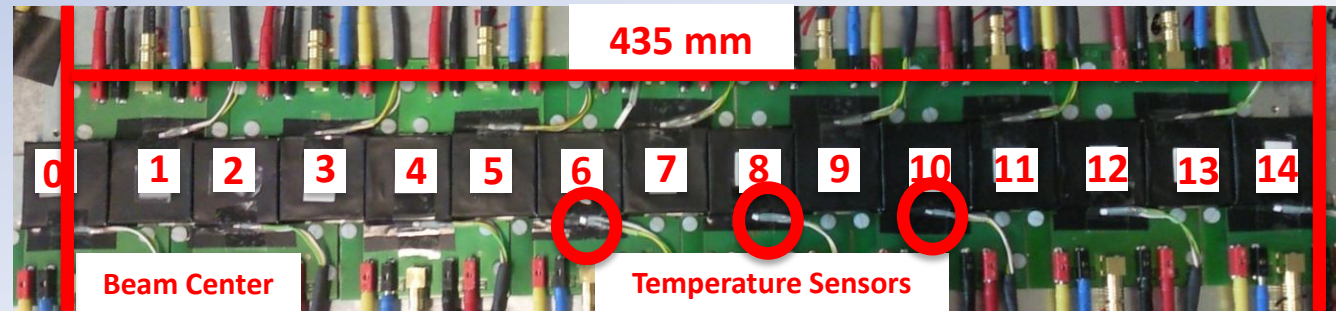
Setup optimized to observe the time development of hadron showers

CALICE:

- + 3D reconstruction of hadronic shower shapes
- No timing information on the showers \rightarrow (s)T3B



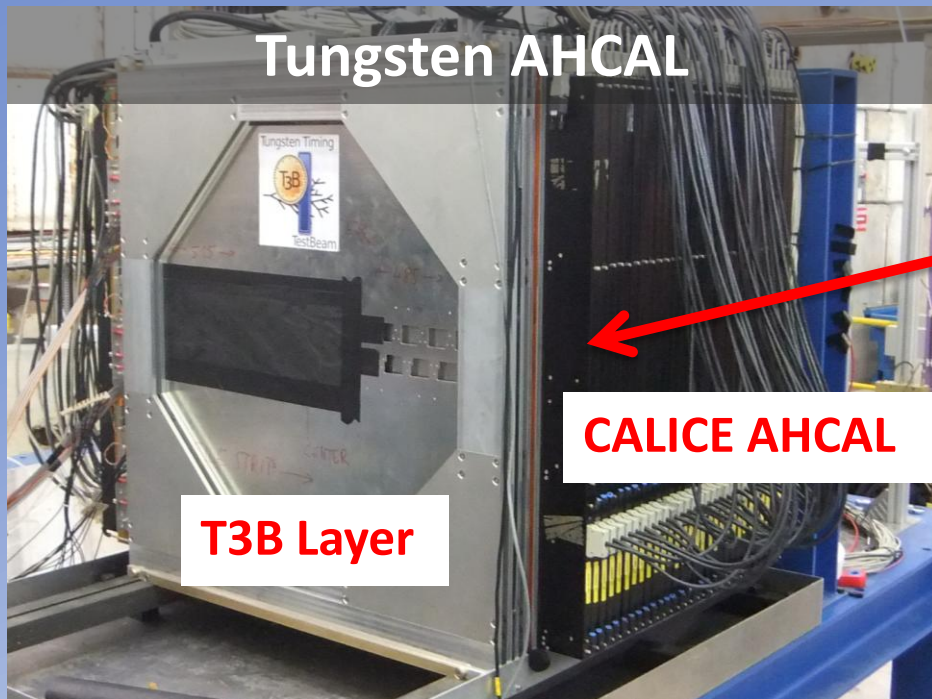
Tile geometry optimized for direct coupling



1 Temperature Sensor PT1000 for each T3B cell



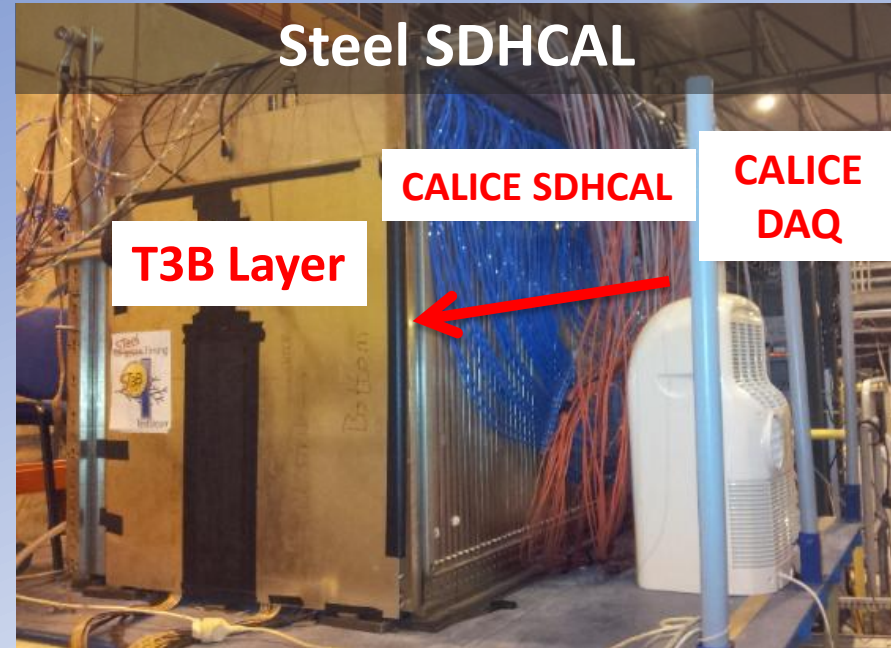
The T3B Experiment within the CALICE Calorimeters



Tungsten AHCAL

CALICE AHCAL

T3B Layer



Steel SDHCAL

CALICE SDHCAL

CALICE DAQ

T3B Layer

Run Periods: PS: Nov 2010
SPS: June/July/Sept 2011

Energy Range: 2-300GeV

Trigger: CALICE Synchronous

Shower Depth: $\sim 3\lambda_1$ (PS), $\sim 5\lambda_1$ (SPS)

Total Had. Events: 27 Million

Run Periods: SPS: October 2011

Energy Range: 40-180GeV

Trigger: T3B Standalone

Shower Depth: $\sim 6\lambda_1$

Total Had. Events: 5 Million

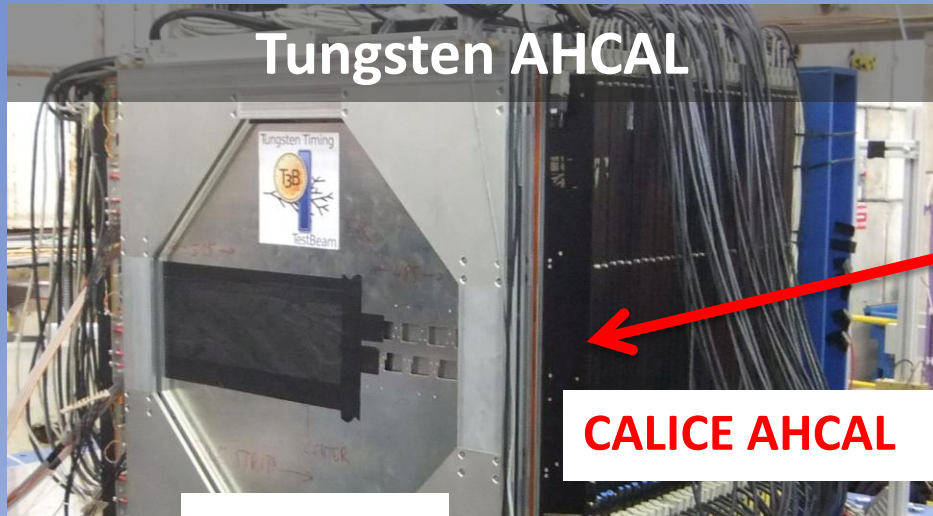




The T3B Experiment within the CALICE Calorimeters



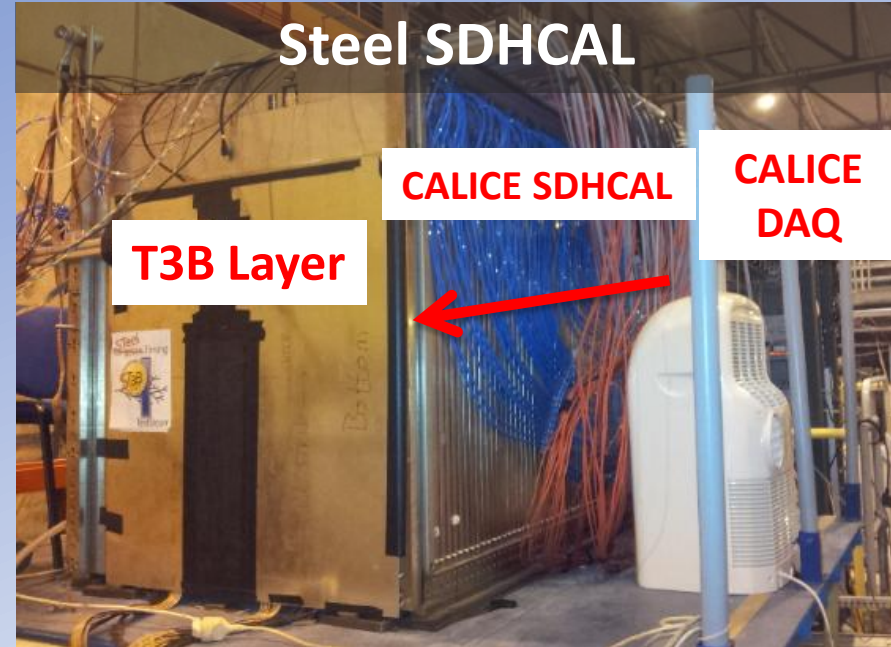
Tungsten AHCAL



CALICE AHCAL

Energy [GeV]	Calice-Sync + [MEv]	Calice-Sync - [MEv]
6	1,2	1,7
8	1,5	1,5
10		4,6
40	2,0	
50		1,7
60	4,1	
80	4,5	
150		1,2
180	0.9	0.7

Steel SDHCAL



T3B Layer

CALICE SDHCAL

CALICE DAQ

Energy [GeV]	T3B Standalone + [MEv]
60	1,6
80	2,0
180	1,2





CALIBRATION TO THE MIP SCALE



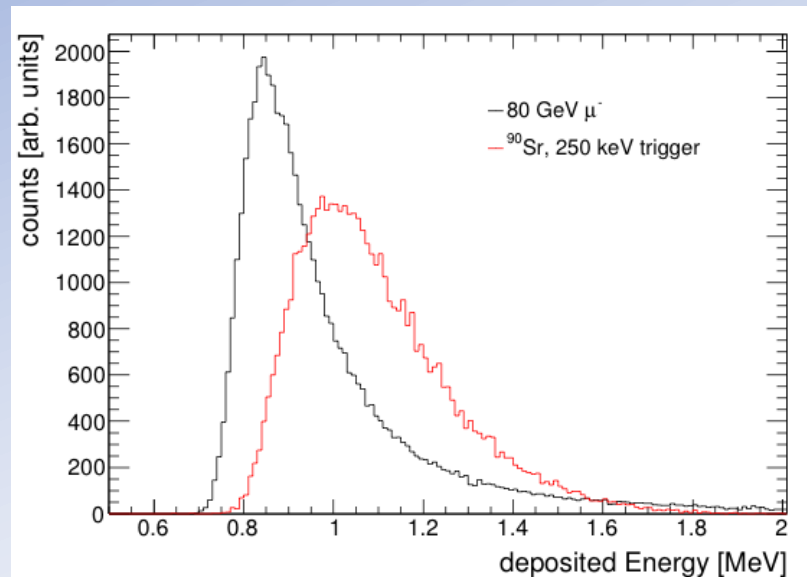
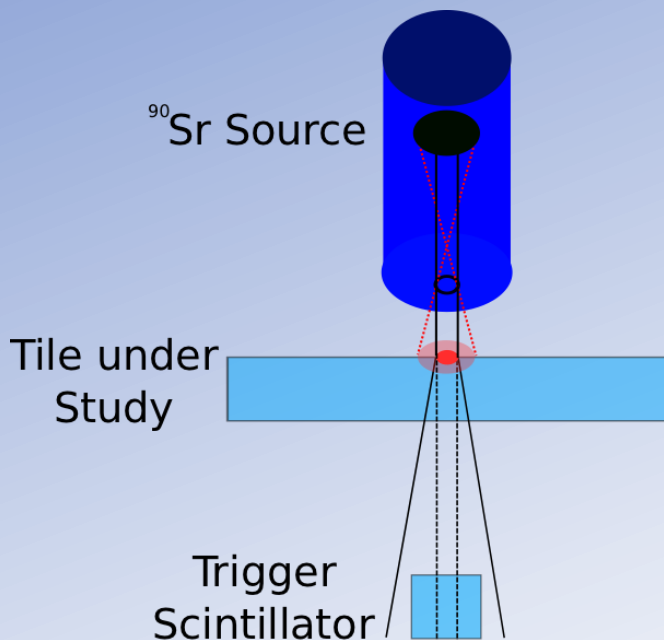
T3B Calibration to the MIP Scale: Sr90 Data



- During the Test Beam T3B monitors the SiPM Gain continuously
→ This data can be used to calibrate energy depositions to the MIP Scale
Assumption: The MIP MPV depends in first order only(!) on the Gain

Offline Calibration Setup:

- Consecutive calibration of all T3B cells individually
→ Use T3B DAQ: Acquire Sr90 and SiPM gain data at the same time

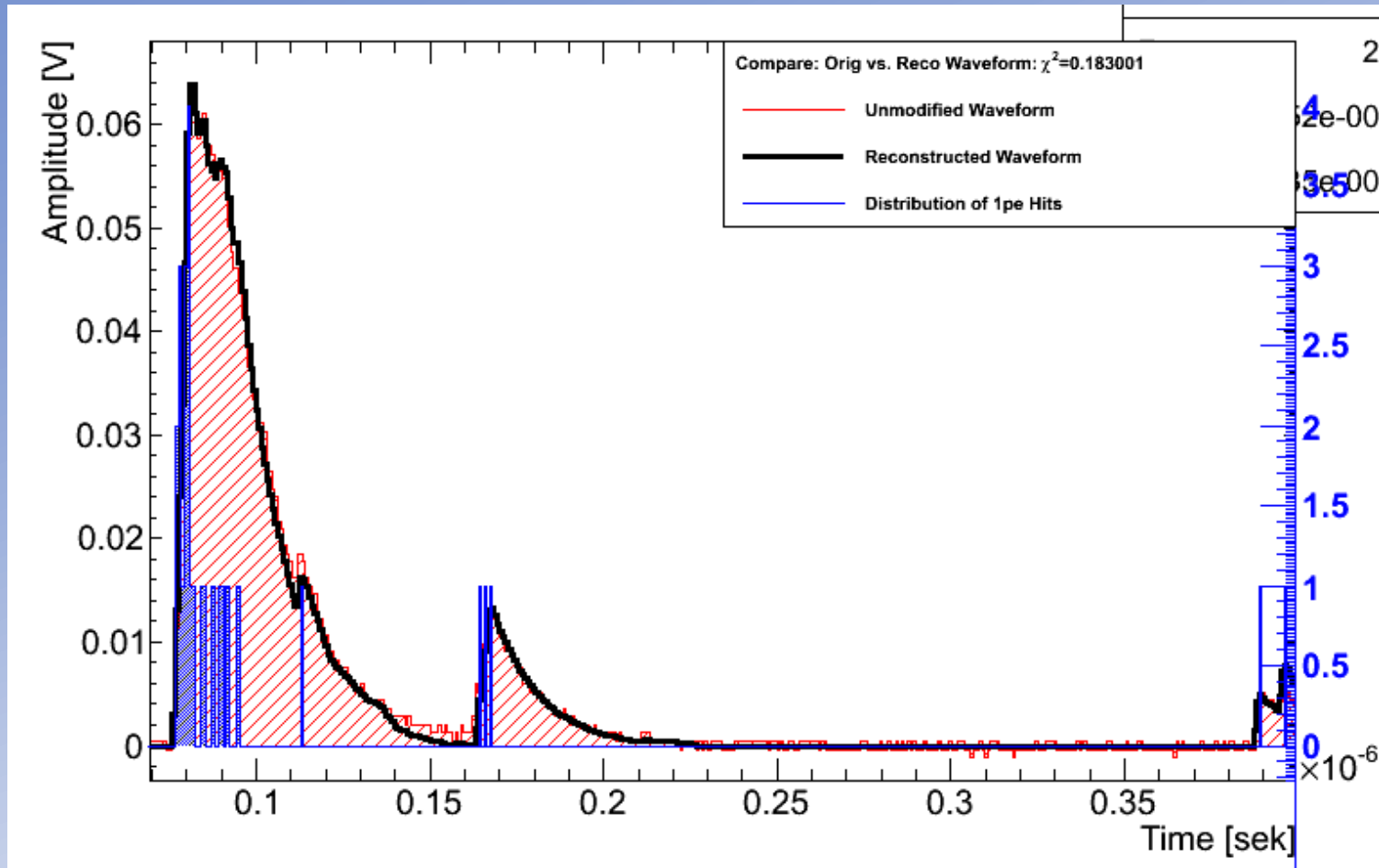


Note:
Electrons are
no perfect
MIPs
→ need scale
factor

GEANT4 Simulation:
 $MPV(\mu) = MPV(e) * 0.82$



The Time Integration Window



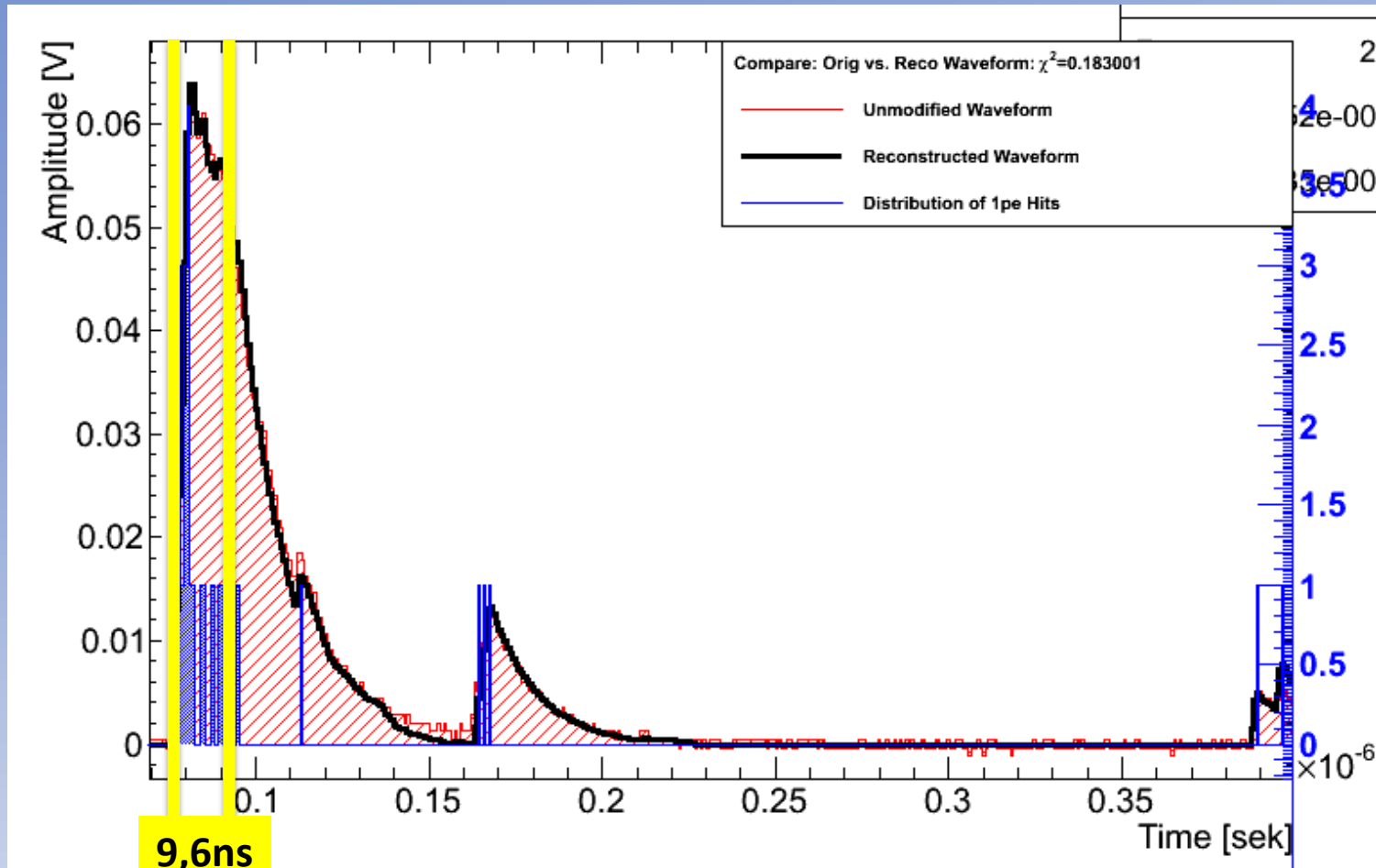
The MPV is very sensitive on the Time Integration Window

→ Dominant effect: SiPM Afterpulsing

- Separate afterpulsing from energy depositions
- Study the effect of afterpulsing



The Time Integration Window



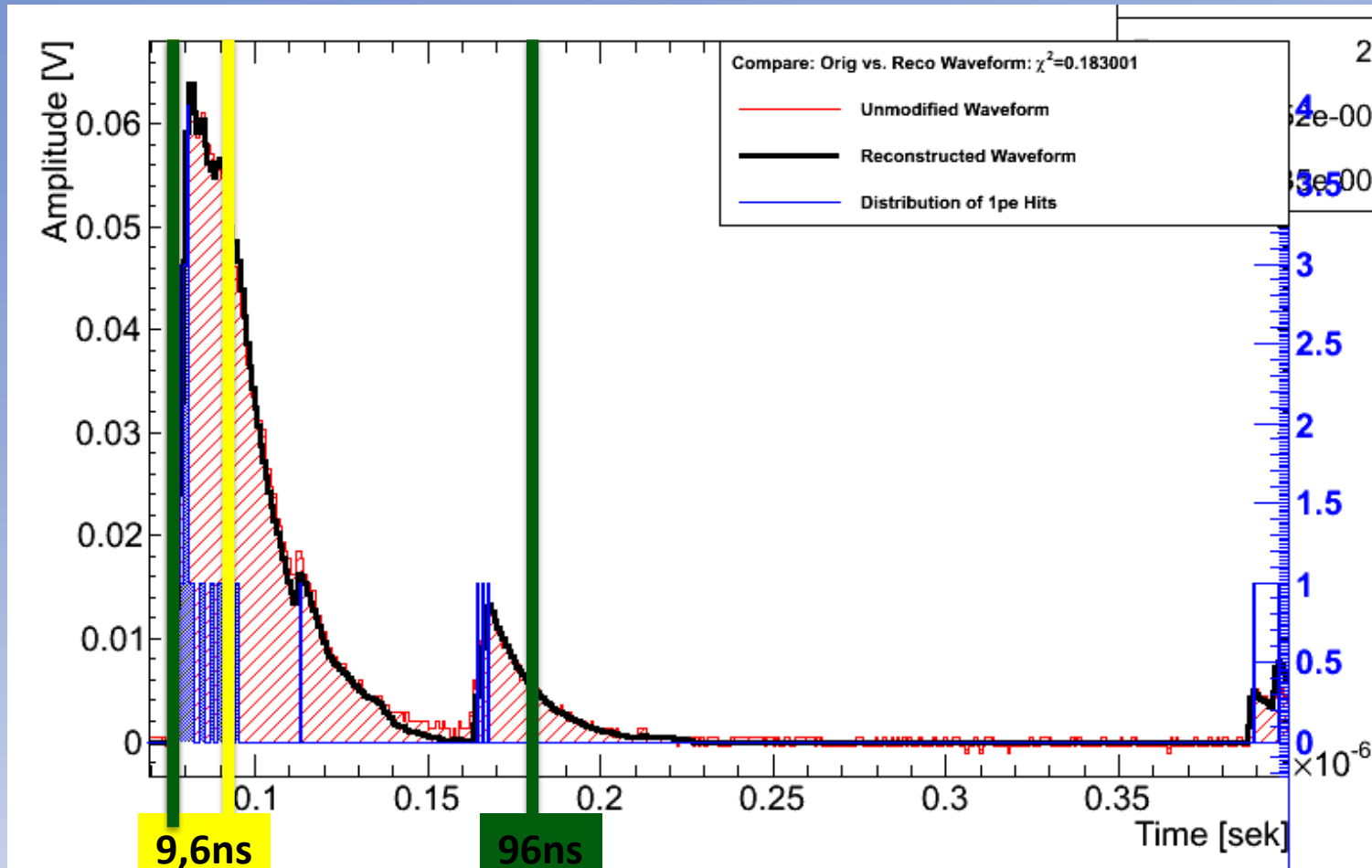
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The Time Integration Window



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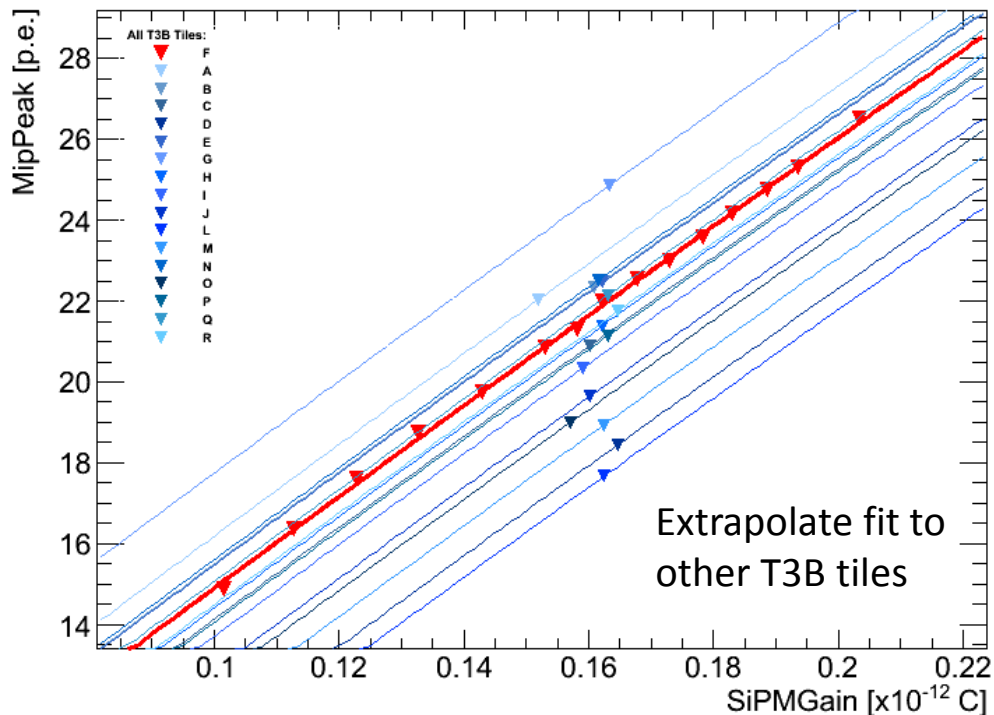


The Effect of Afterpulsing



Bias Voltage Scan for one T3B
“Master Tile”
One Measurement for
all other T3B tiles

Time Window	MPV-Gain dependence
Short (~10 ns)	Linear (no AP)
Long (~100 ns)	Quadratic (with AP)





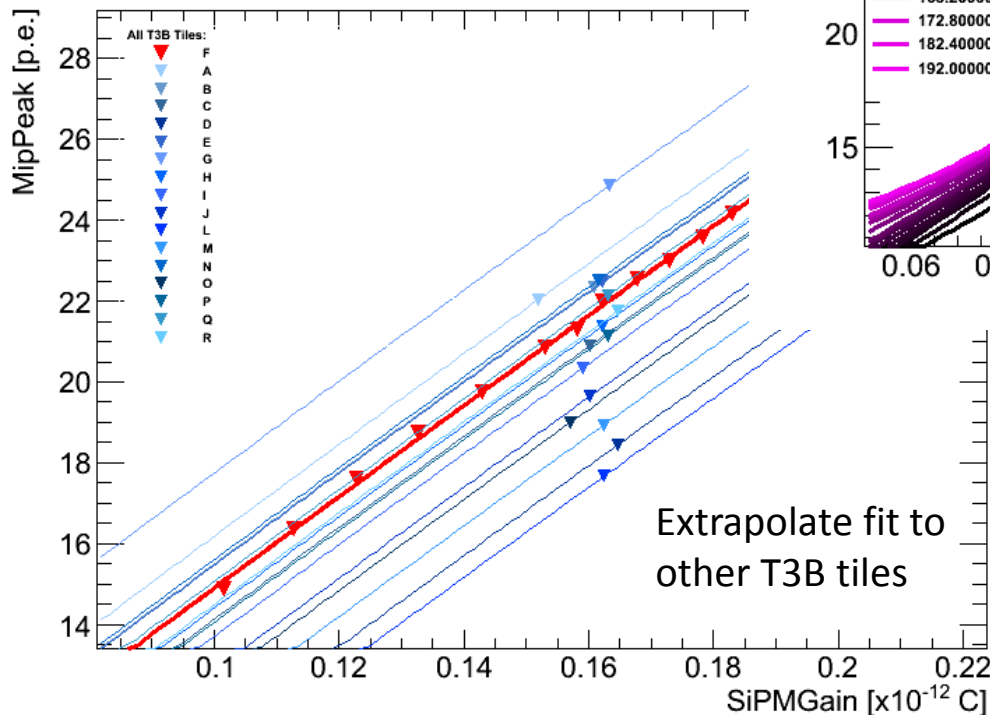
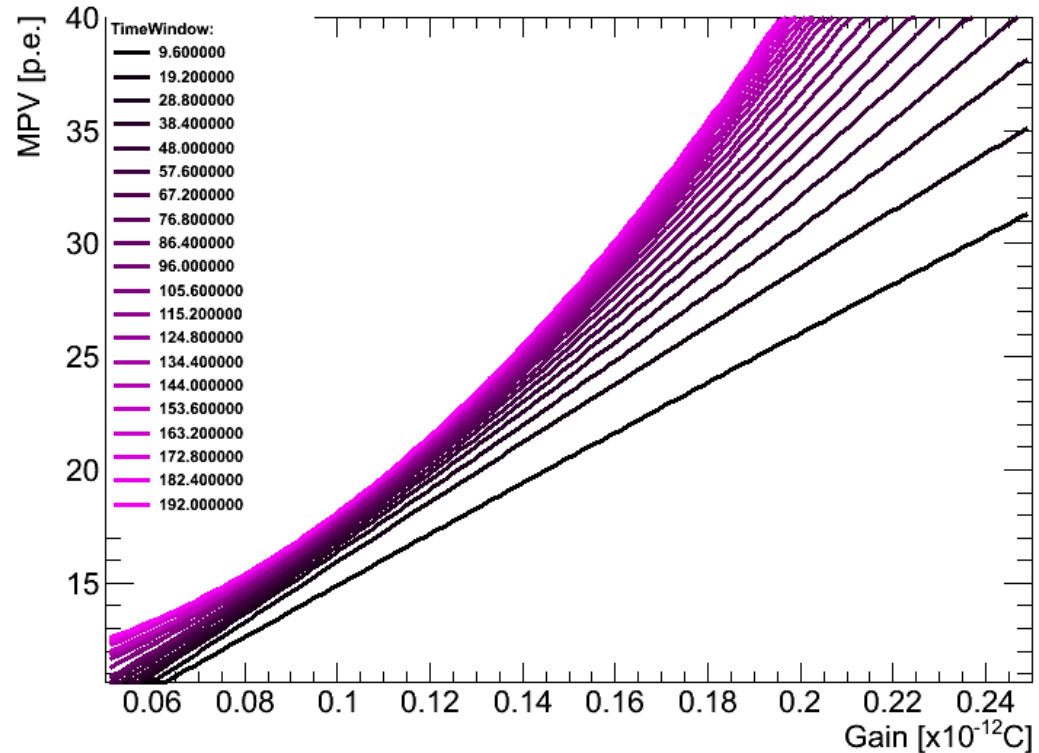
The Effect of Afterpulsing



Bias Voltage Scan for one T3B
"Master Tile"
One Measurement for
all other T3B tiles

Time Window	MPV-Gain dependence
Short (~10 ns)	Linear (no AP)
Long (~100 ns)	Quadratic (with AP)

Sr90 Calibration Data - Langau MPV vs. Gain



Higher Gain

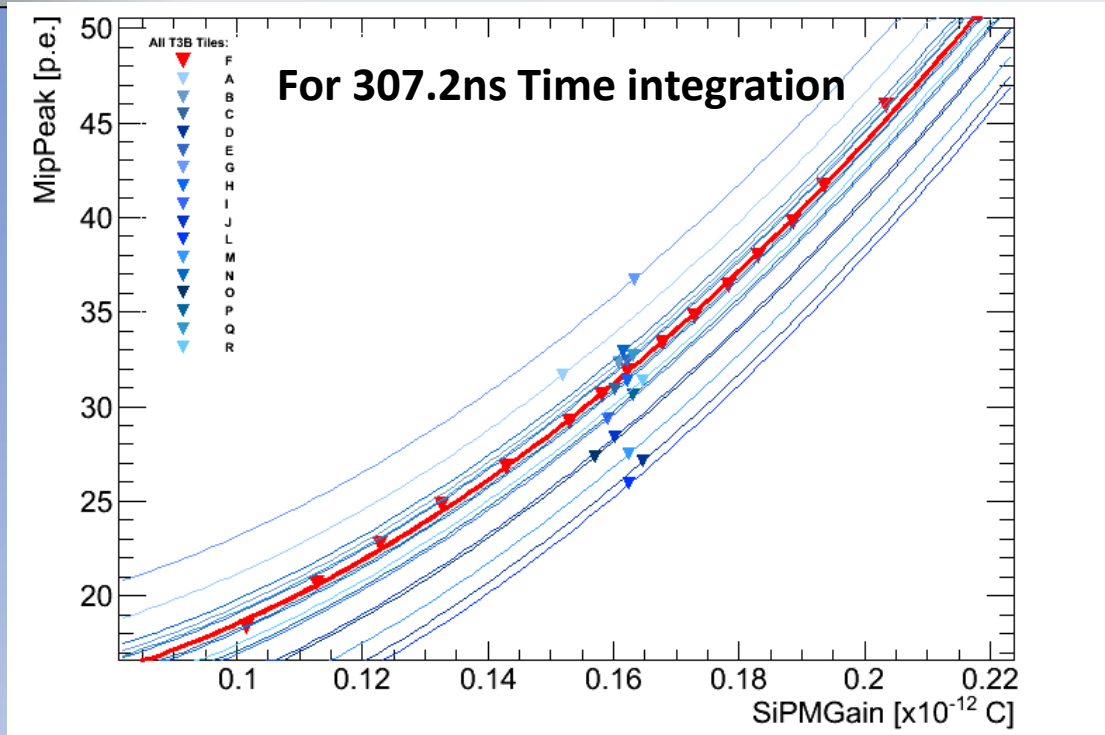
→ Afterpulsing and Crosstalk probability increased

→ Increased MPV dependence for long integration times

Needs to be taken into account in Calib



T3B Calibration to the MIP Scale



Obtain a dictionary:

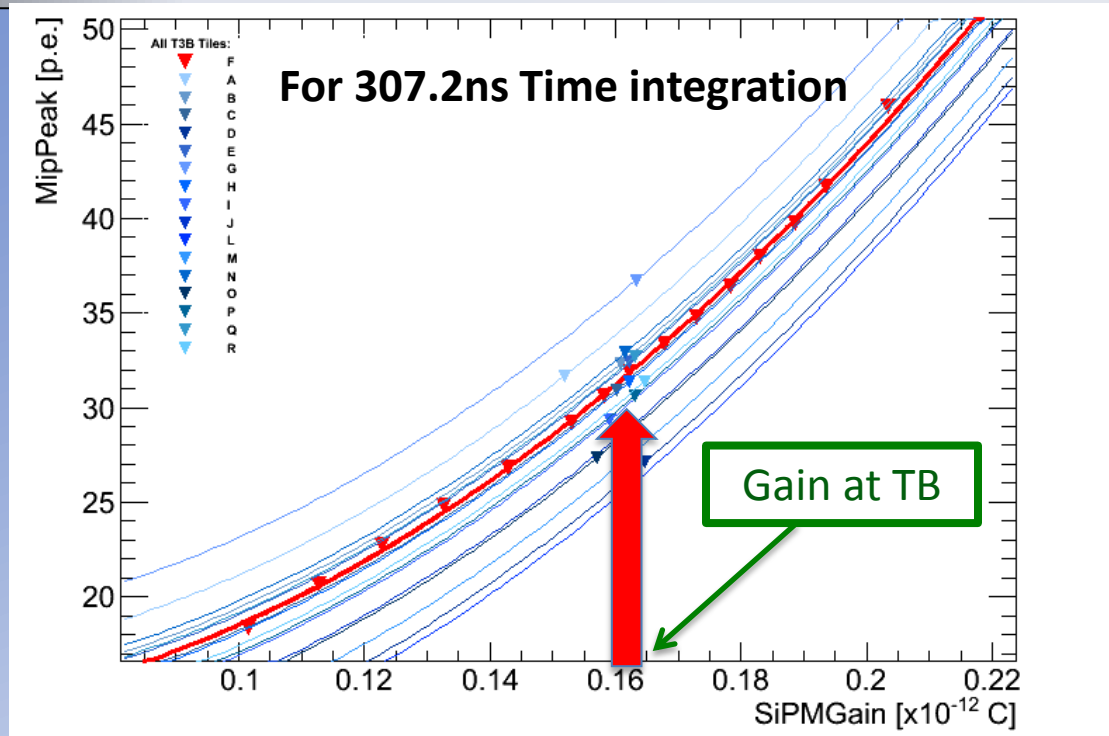
Determine live SiPM Gain from testbeam data

Select MPV-Gain dependence for distinct time integration window

Obtain corresponding MPV of MIP distrib.



T3B Calibration to the MIP Scale



Obtain a dictionary:

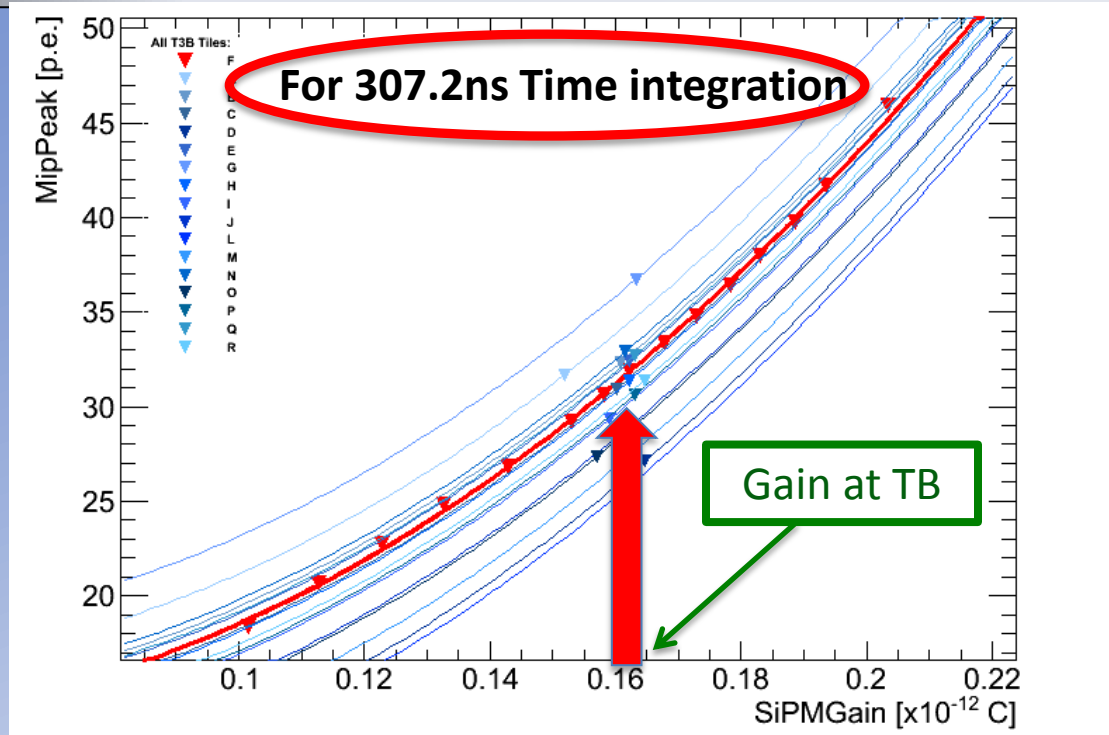
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T3B Calibration to the MIP Scale



Obtain a dictionary:

Determine live SiPM Gain from testbeam data

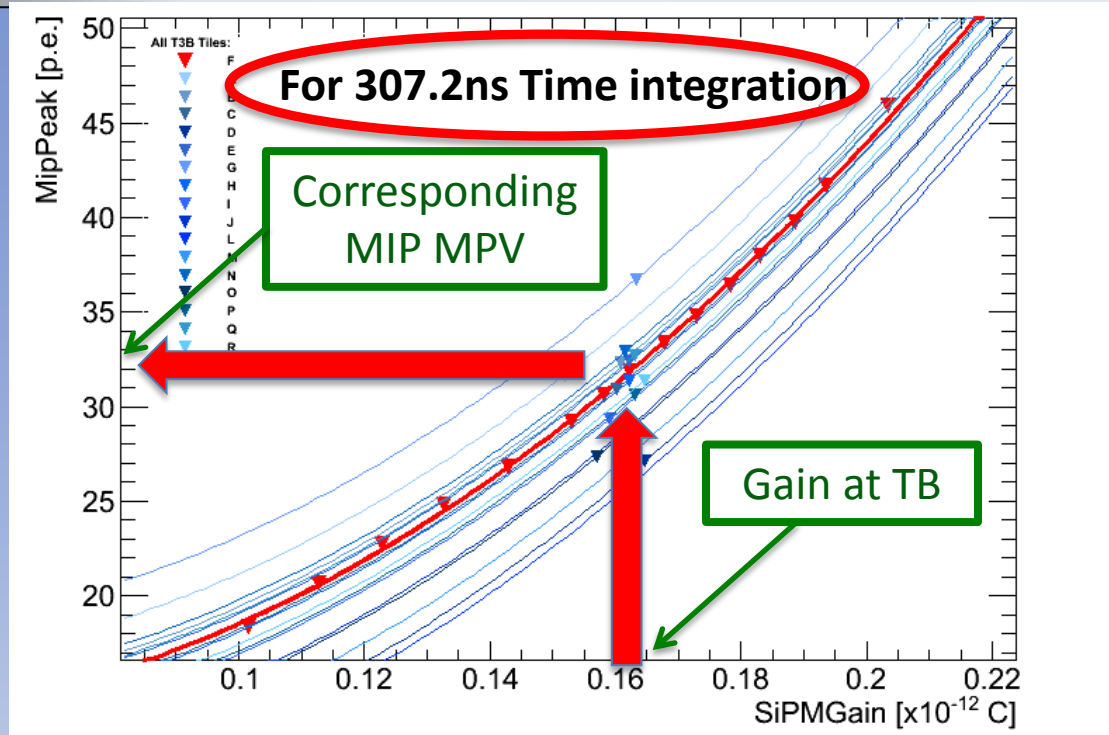


Select MPV-Gain dependence for distinct time integration window

Obtain corresponding MPV of MIP distrib.



T3B Calibration to the MIP Scale



Obtain a dictionary:

Determine live SiPM Gain from testbeam data



Select MPV-Gain dependence for distinct time integration window



Obtain corresponding MPV of MIP distrib.



Verification of the Calibration

Principle: Muon Data

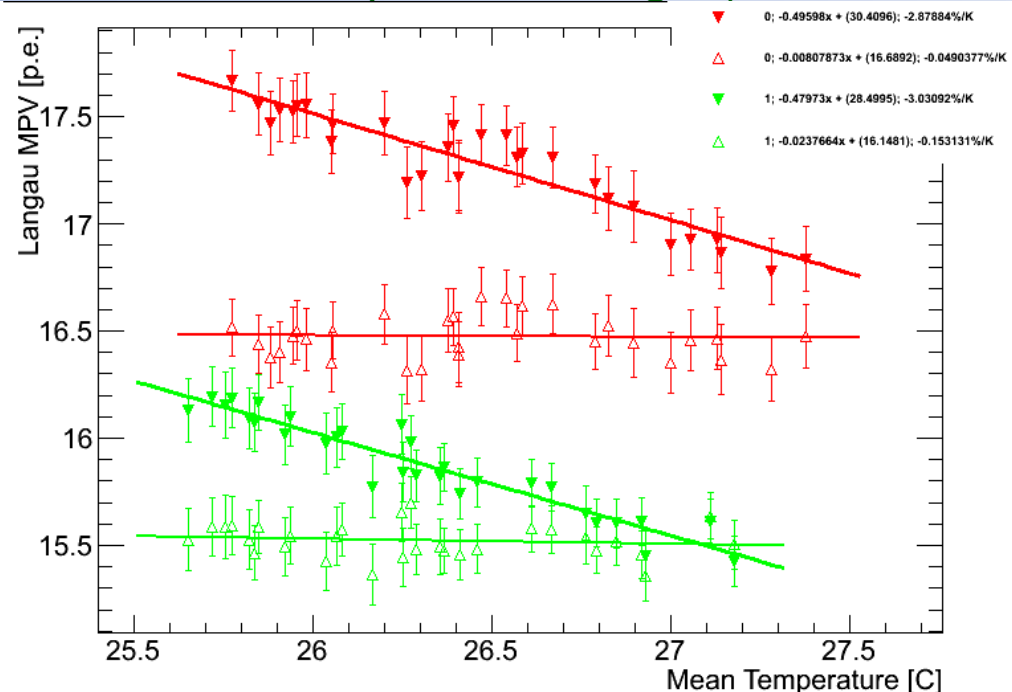


During the commissioning of the SDHCAL we could take an large amount of muon data:

- 14 mio Muon Events
- 40 hours without interruption
- Day-night-cycle Temperature Range: ~25.5C to 27.5C

→ Enough to extract the Mip MPV-Temperature dependence

→ Then: Apply correction factor from Sr90 Data to eliminate the dependence (remember: We assume the MPV depends in first order only on the SiPM gain)





Verification of the Calibration

Principle: Muon Data



Corrected MPV-Temperature dependence

Calibration results in efficient elimination of the dependence

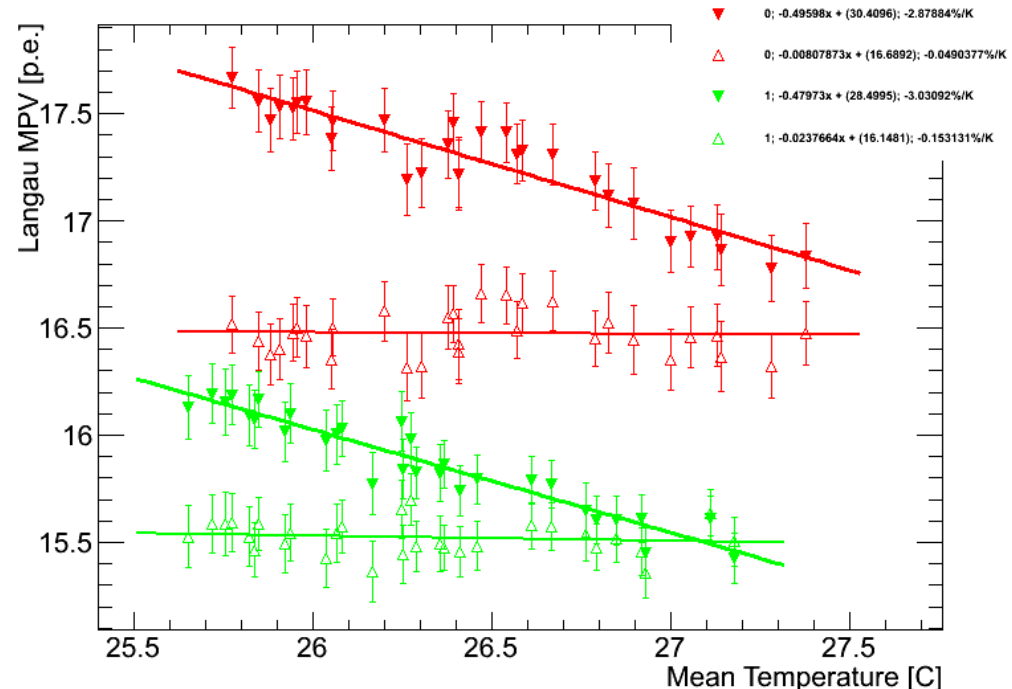
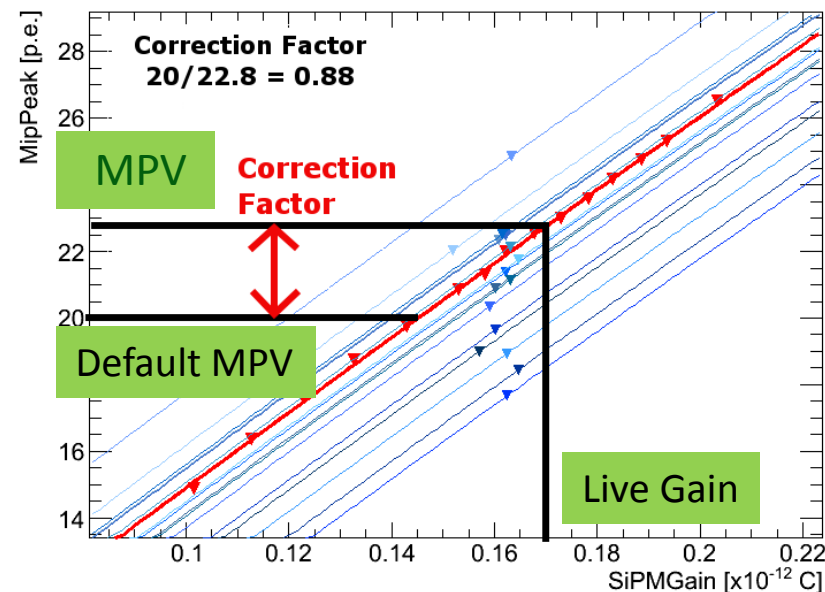
Note (TilePosition 0, analogous for other tiles):

Corrected MPV values at ~ 16.5 p.e., not at the 20 p.e. we corrected to.

Interpretation: 0.82 is the $\text{Sr90} \leftrightarrow \text{Muon}$ MPV conversion factor

Matches simulations \rightarrow Experimental proof

T3B Tile	MPV Drop	Slope
Center	-2.9 %/K	-0.5 p.e./K
Center corr.	-0.05 %/K	-0.008 p.e./K





TIME OF FIRST HIT

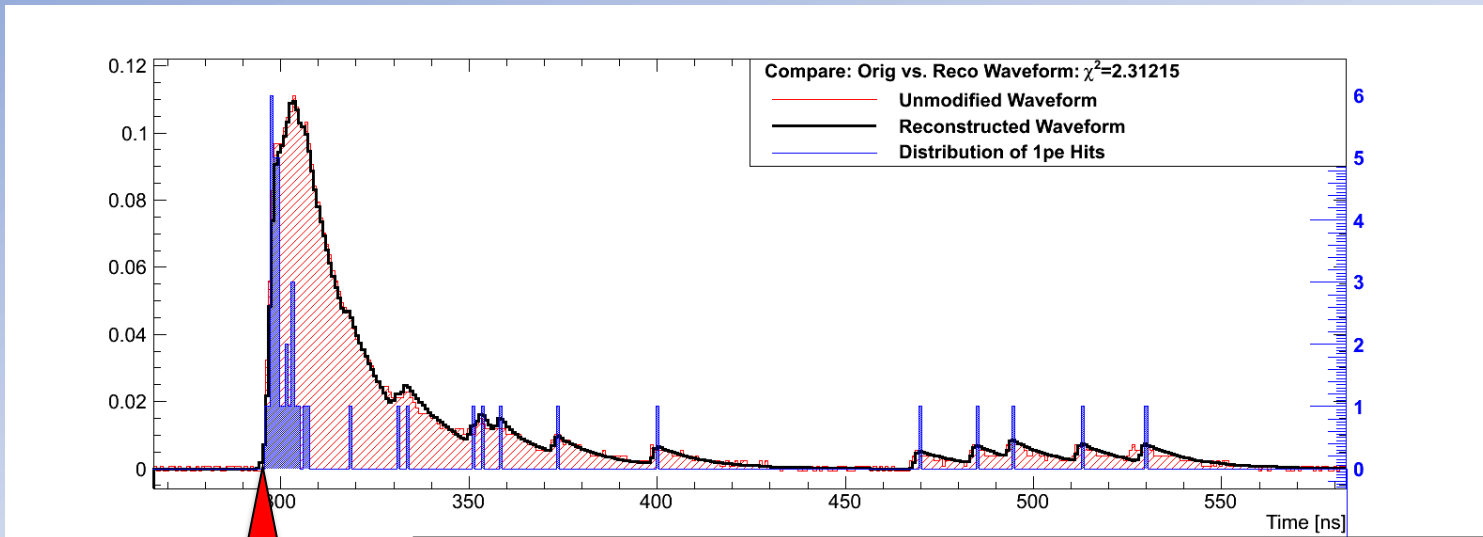


Time of First Hit



Analysis:

- Consider a hit if > 8 p.e. are deposited within 9.6 ns (basically AP free!)
- Take the time of the second fired SiPM Pixel as the ToFH
(reduced bias through thermal darkrate \rightarrow on average 1 pixel per waveform of 2.4 microsec)



Analysis of the time of first energy deposition within the tiles of the T3B layer for hadronic showers



Time of First Hit



Analysis:

- Consider a hit if > 8 p.e. are deposited within 9.6 ns (basically AP free!)
 - Take the time of the second fired SiPM Pixel as the TofH
- (reduced bias through thermal darkrate \rightarrow on average 1 pixel per waveform of 2.4 microsec)

Run selection:

- Choose runs @ 60GeV, 80GeV and 180GeV for Tungsten, Steel and Muon data
 - All runs have > 1 mio Events (Tu @ 60, 80GeV ~ 4 mio Ev, Steel @ 60, 80GeV ~ 2 mio Ev)
 - All runs for one energy from same testbeam period (no mixing)
 - Quality of all runs checked with CALICE and T3B run log
- Note: T3B after 5 lambda for Tu, but after ~ 6 lambda for Steel data
- Slightly increased overall bias voltage for Tu compared to Steel data (200mV)

We present work in progress \rightarrow all plots preliminary



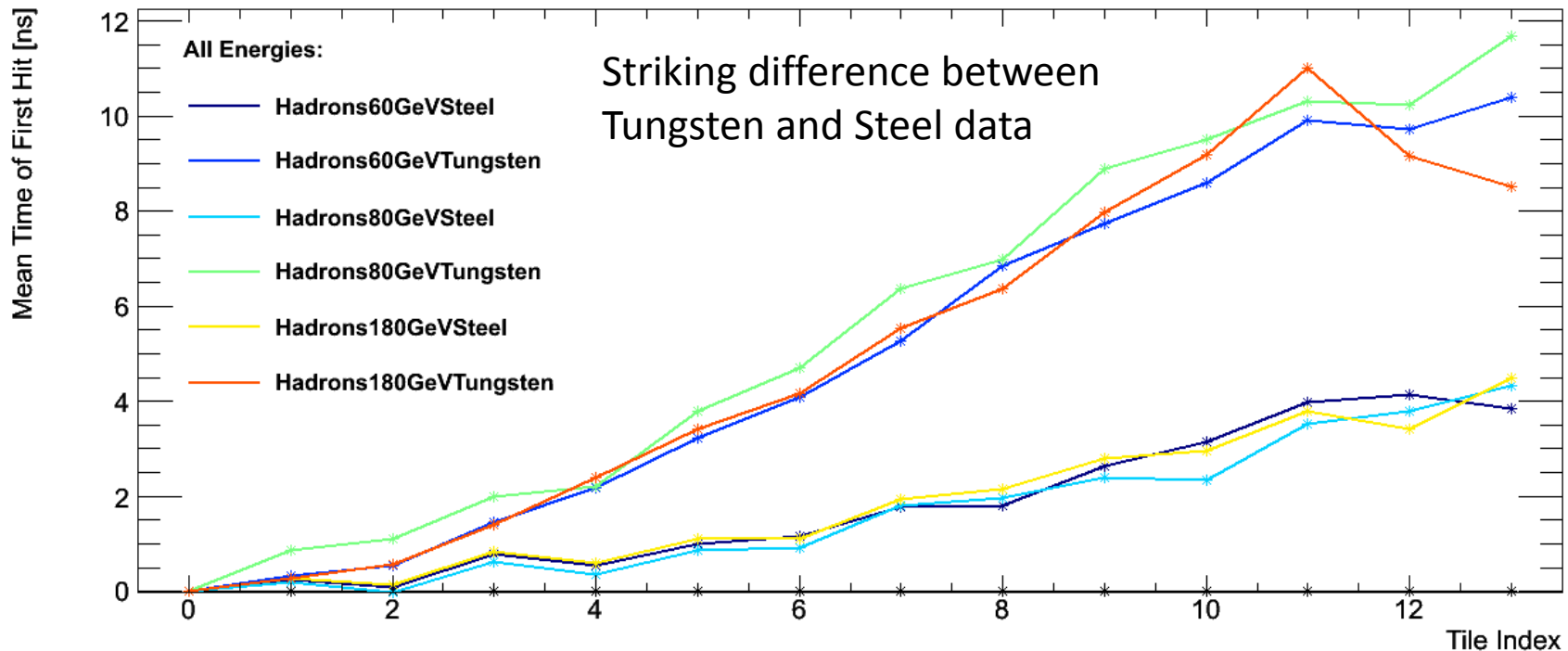
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Mean Time of First Hit



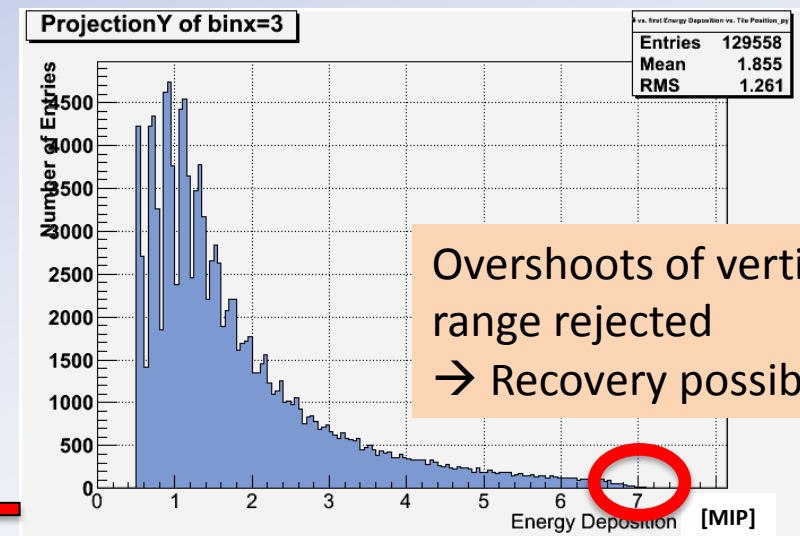
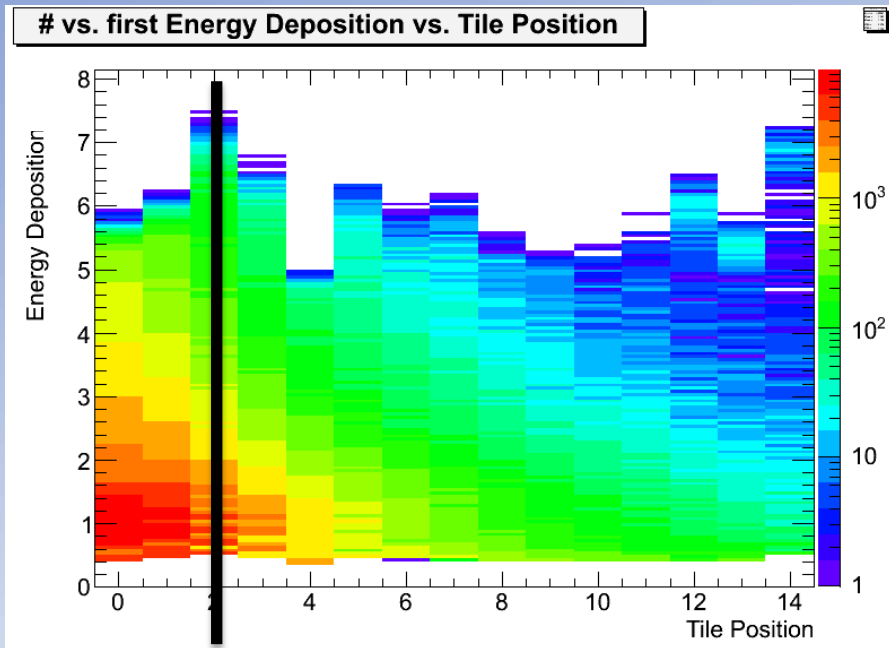


Time of First Hit



Analysis:

- Consider a hit if > 8 p.e. are deposited within 9.6 ns (basically AP free!)
- Take the time of the second fired SiPM Pixel as the ToFH
(reduced bias through thermal darkrate \rightarrow on average 1 pixel per waveform of 2.4 microsec)
- Sum the number of pixels fired within the identified hit
 \rightarrow Calibrate to the MIP scale



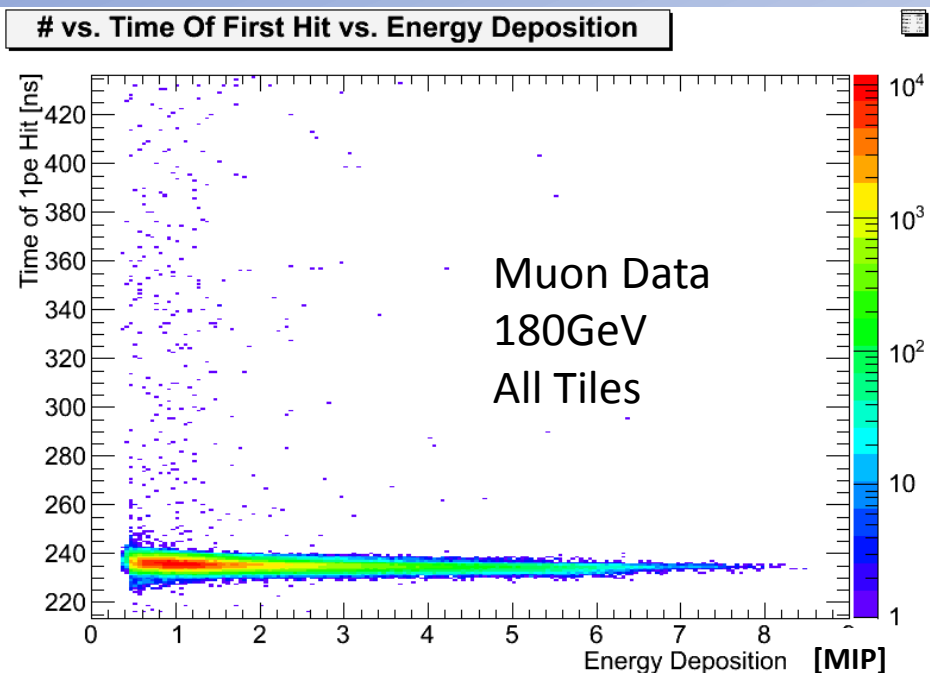


Time of First Hit



Analysis:

- Consider a hit if > 8 p.e. are deposited within 9.6 ns (basically AP free!)
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(reduced bias through thermal darkrate \rightarrow on average 1 pixel per waveform of 2.4 microsec)
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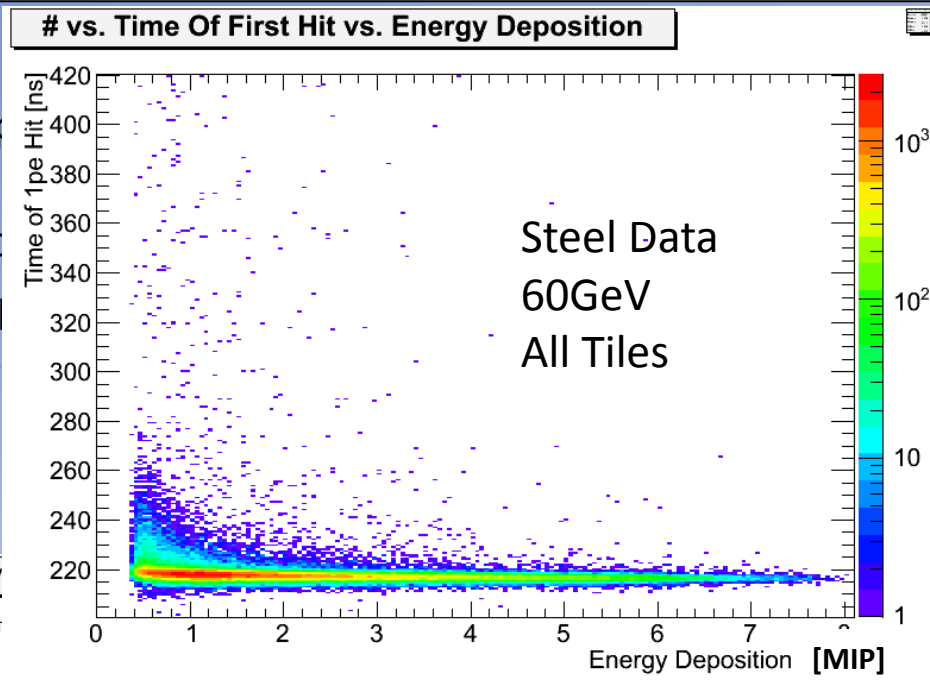




Time of First Hit

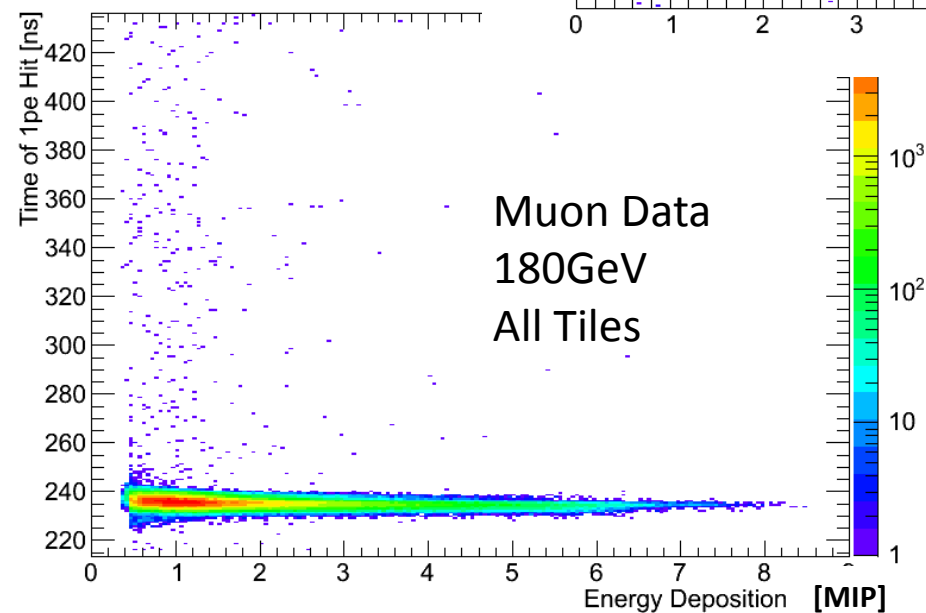
Analysis:

- Consider a hit if > 8 p.e.
- Take the time of the first hit (reduced bias through threshold)
- Sum the number of hits
- Calibrate to the



!)
orm of 2.4 microsec)

vs. Time Of First Hit vs. Energy

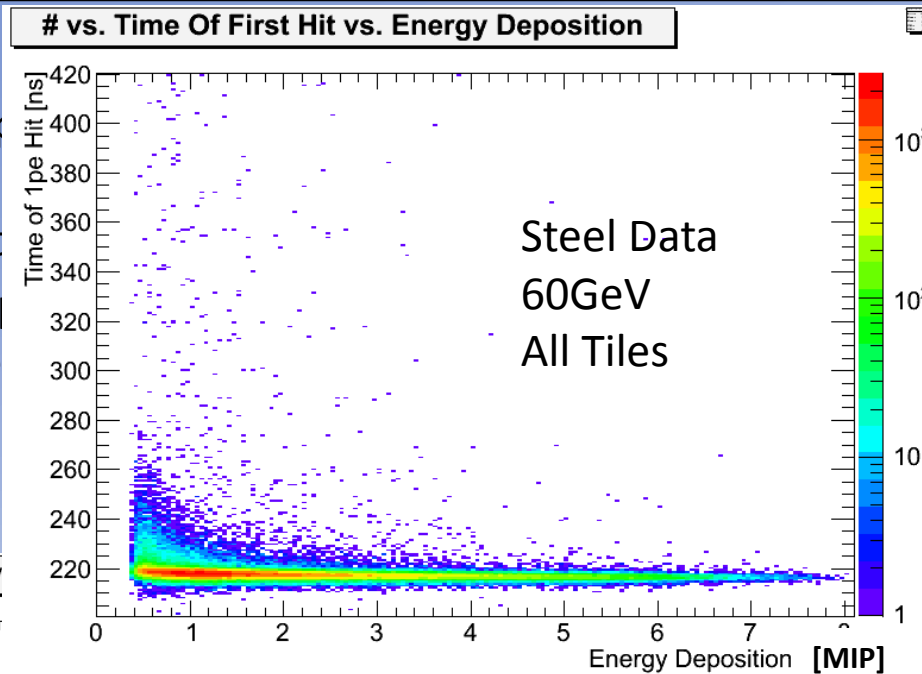




Time of First Hit

Analysis:

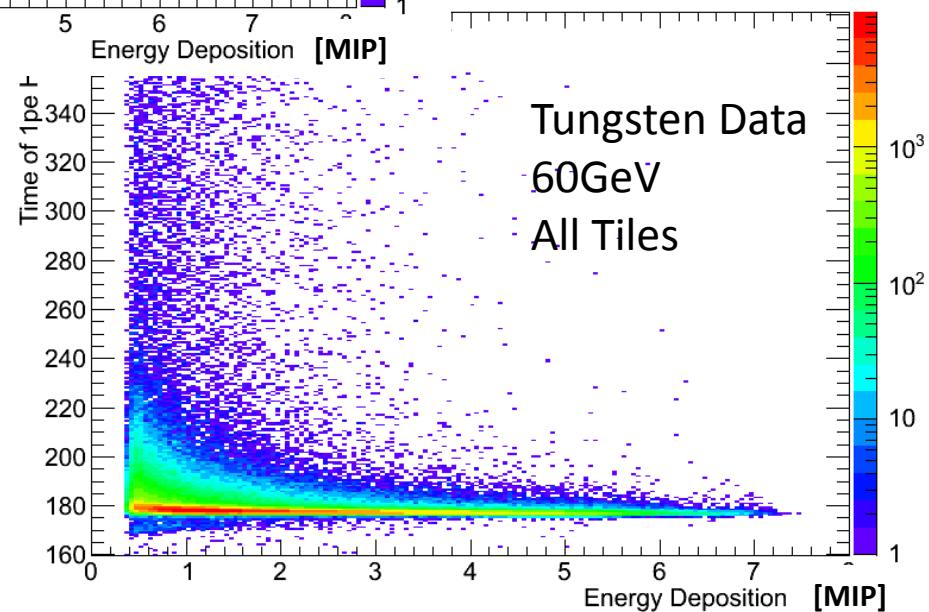
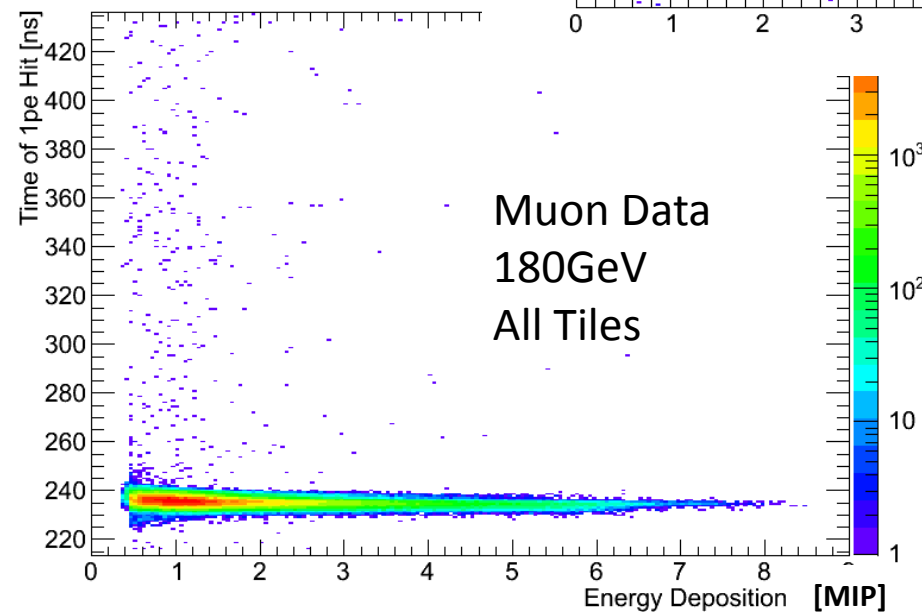
- Consider a hit if > 8 p.e.
- Take the time of the first hit (reduced bias through threshold)
- Sum the number of hits in a bin \rightarrow Calibrate to the



!)
orm of 2.4 microsec)

gy Deposition

vs. Time Of First Hit vs. Energy





Time of First Hit



Analysis

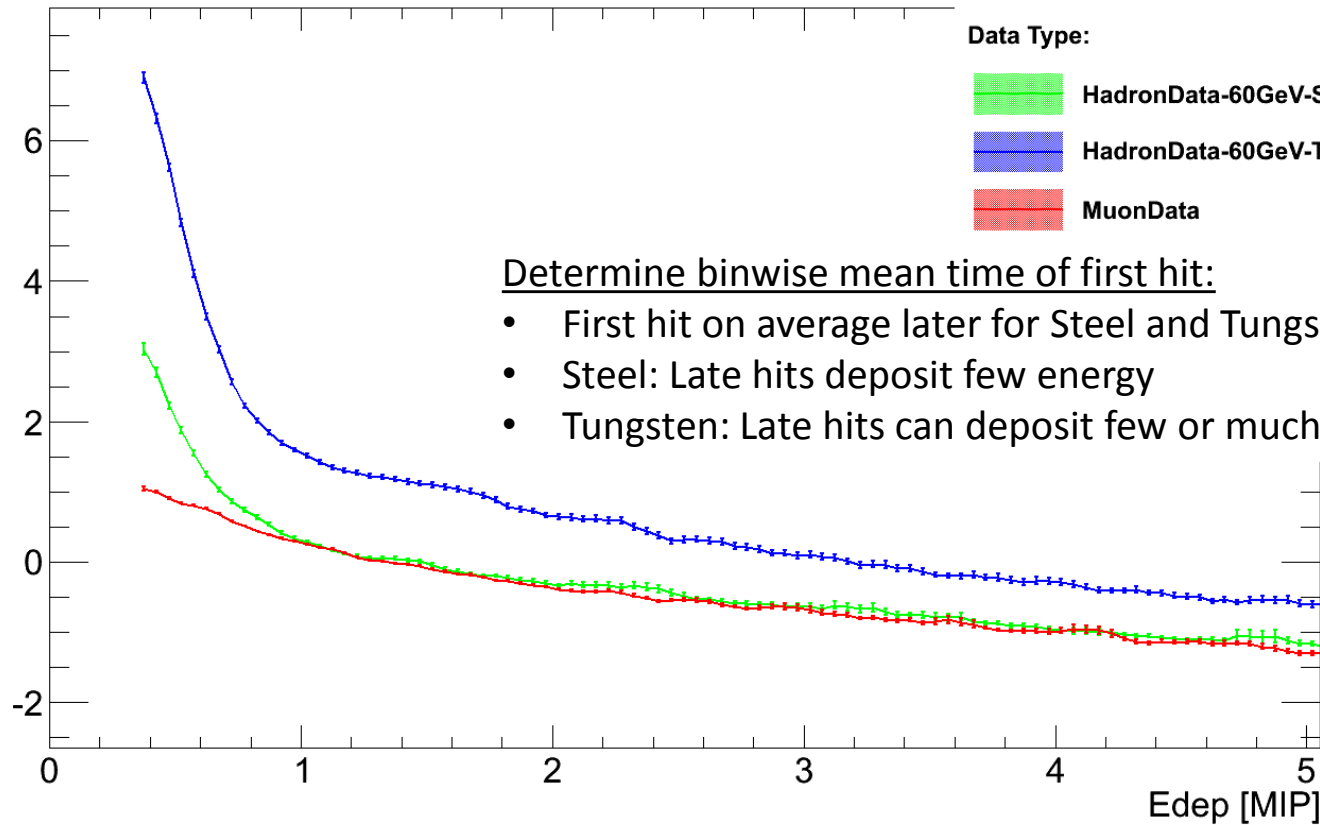
- Const
- Take
- (reduce
- Sum

vs.

8

vs. Time

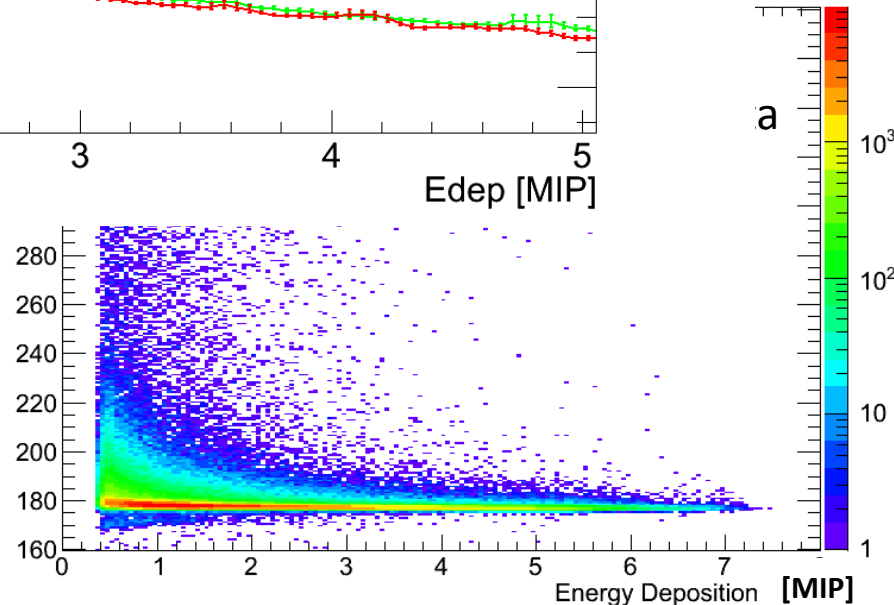
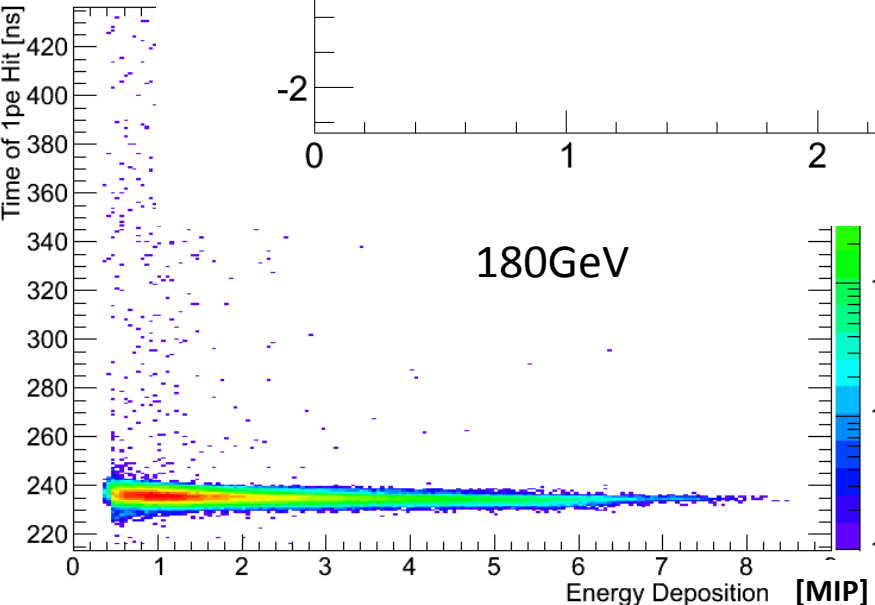
Mean Time of First Hit [ns]



Determine binwise mean time of first hit:

- First hit on average later for Steel and Tungsten
- Steel: Late hits deposit few energy
- Tungsten: Late hits can deposit few or much energy

sec)





Time of First Hit



Analysis

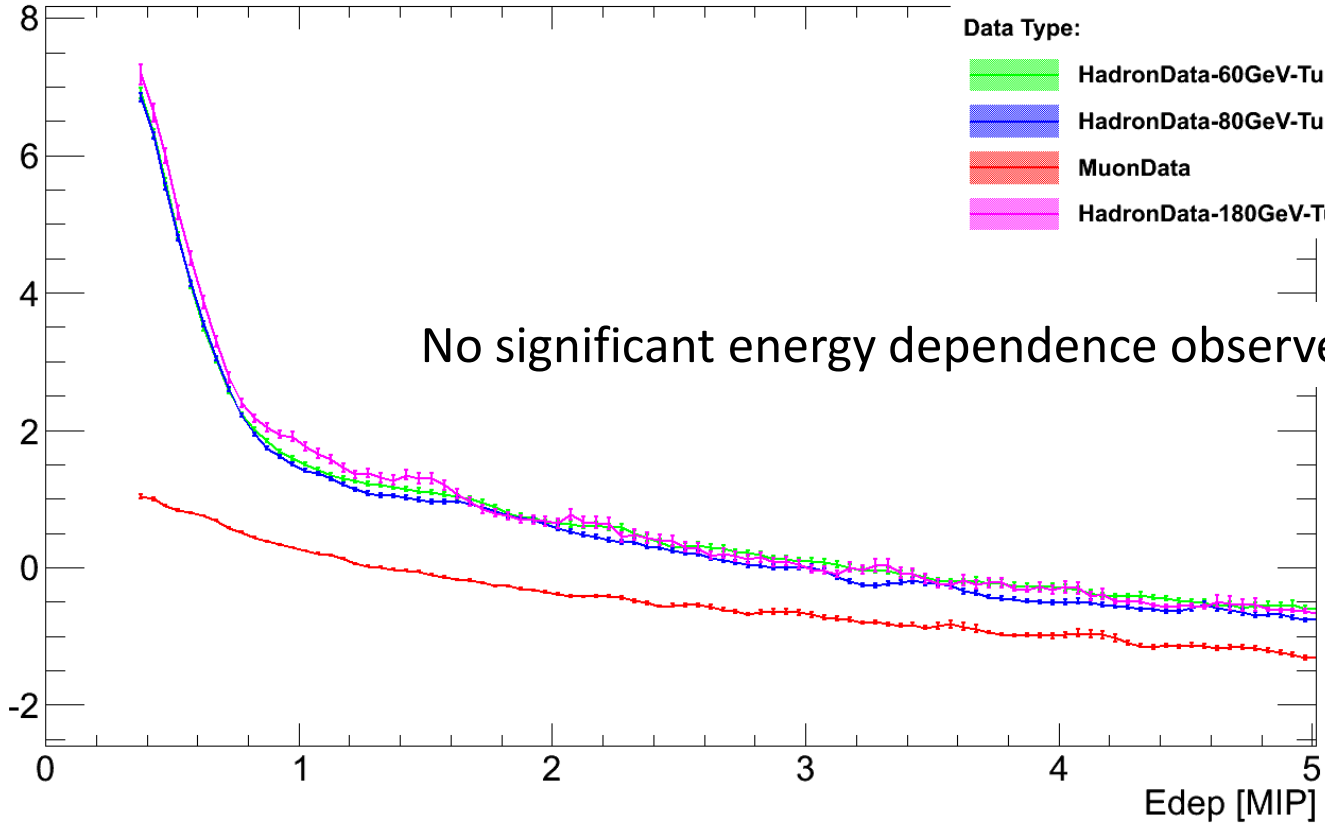
- Const
 - Take
 - (reduce
 - Sum
-

vs.

8

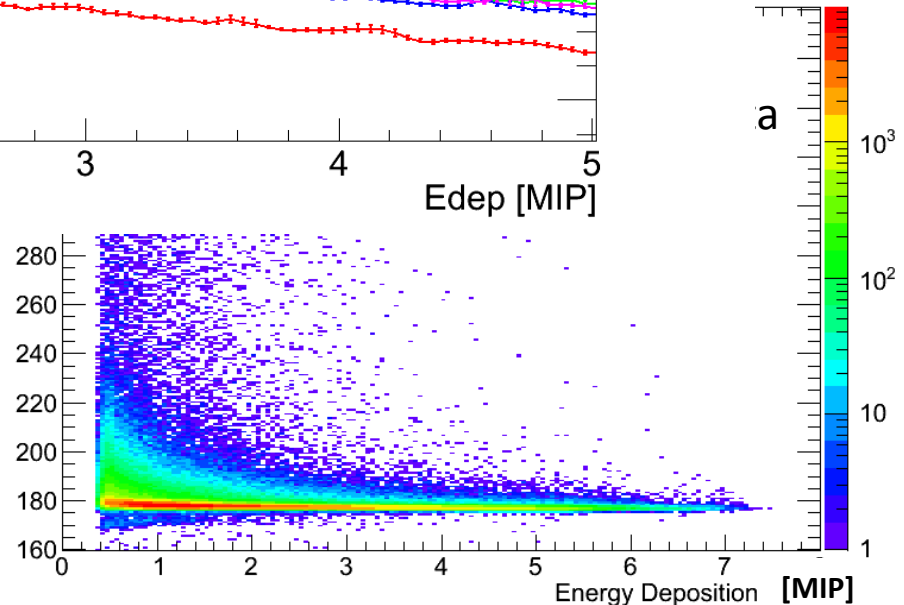
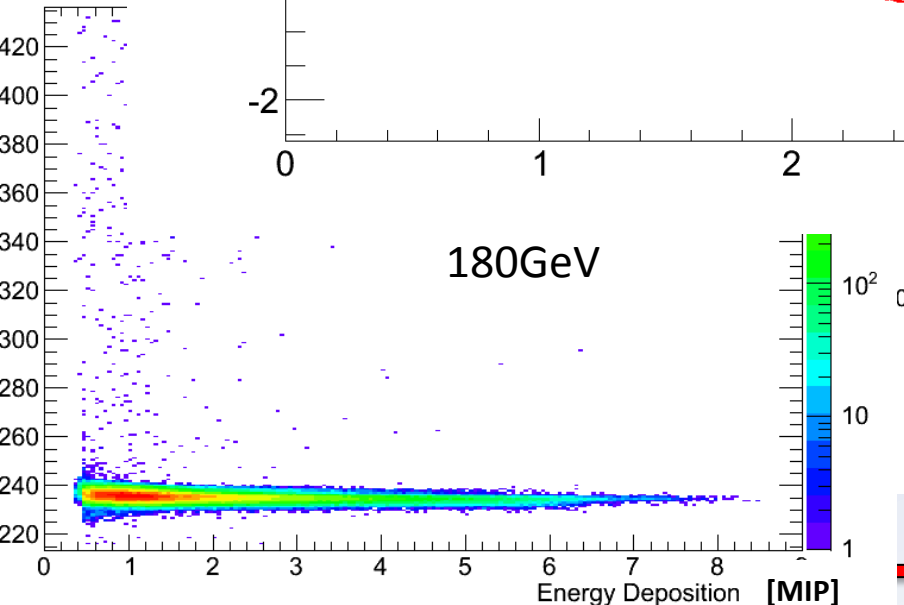
vs. Time

Mean Time of First Hit [ns]



sec)

Time of 1pe Hit [ns]





AFTERPULSING ANALYSIS



Afterpulsing Analysis



Dedicated Afterpulsing Runs:

- Taken in May 2011 with all 15 T3B tiles → No SPS beam available at that time
- Record only SiPM Darkrate → higher pixel counts due to crosstalk (higher counts less likely)
- Threshold scan: 5 runs with 400k Events @ 0.5, 1.5, 2.5, 3.5 and 4.5 p.e. threshold
+ 1 random trigger run (for pedestal subtraction)



Afterpulsing Analysis



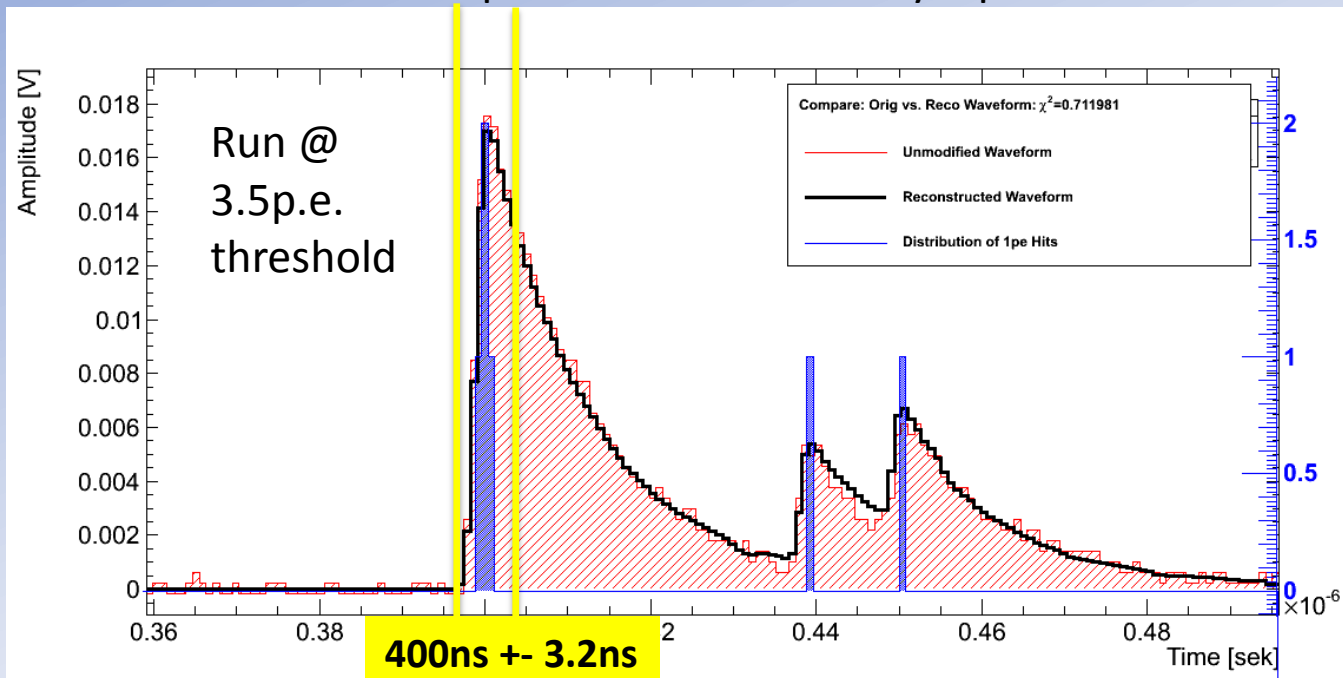
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- Threshold scan: 5 runs with 400k Events @ 0.5, 1.5, 2.5, 3.5 and 4.5 p.e. threshold + 1 random trigger run (for pedestal subtraction)

– Criterion to accept waveform for analysis:

Demand exactly N p.e. within the acceptance range of $\pm 3.2\text{ns}$ around the trigger time for the N p.e. threshold run.

e.g.: Run with threshold at 3.5 p.e. → demand exactly 4 p.e. within $400\text{ns} \pm 3.2\text{ns}$





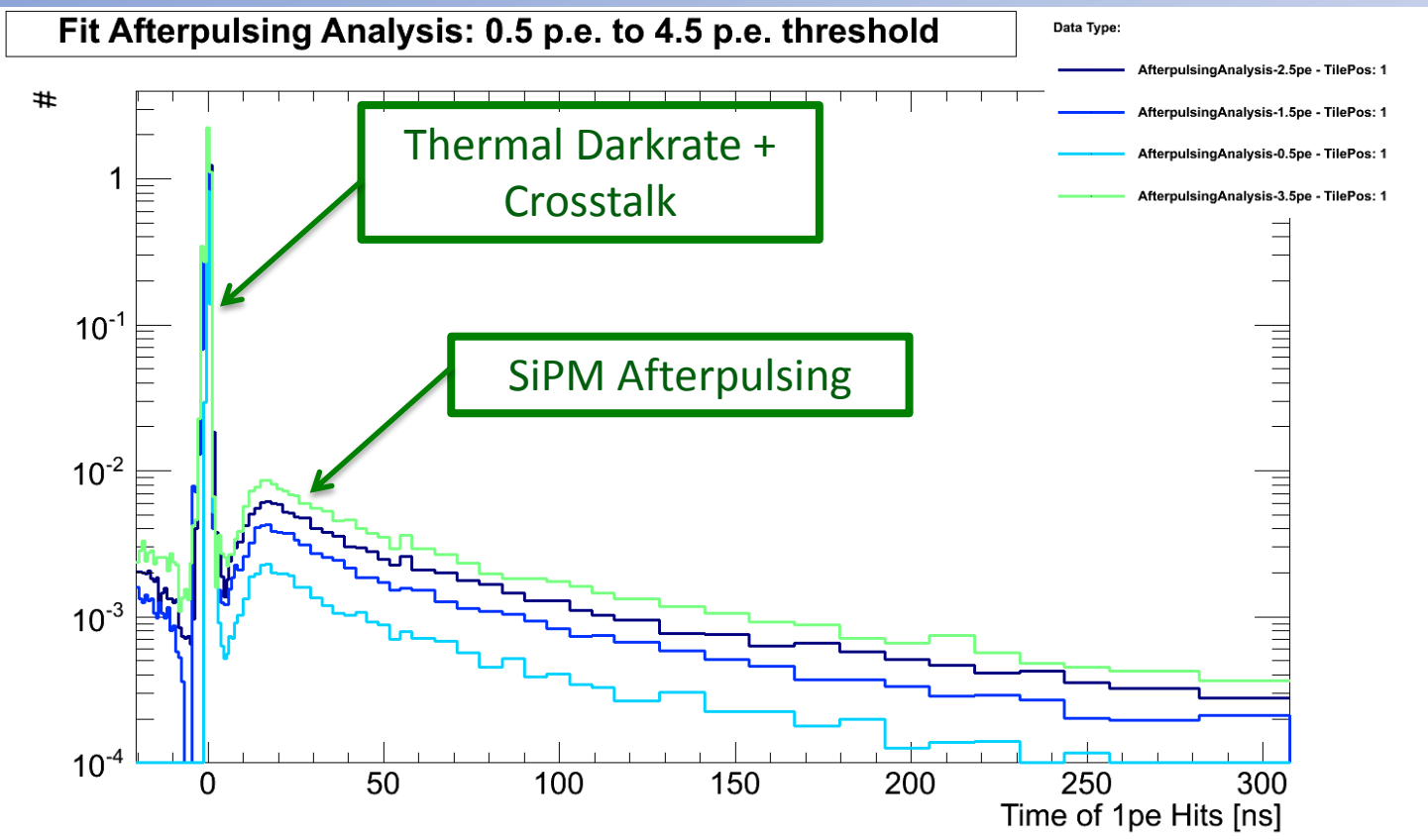
Afterpulsing Analysis



Procedure to extract the average SiPM afterpulsing per fired pixel:

Fill the time of each firing pixel into one histogram for all selected waveforms

→ Normalize to the number of selected waveforms





Afterpulsing Analysis

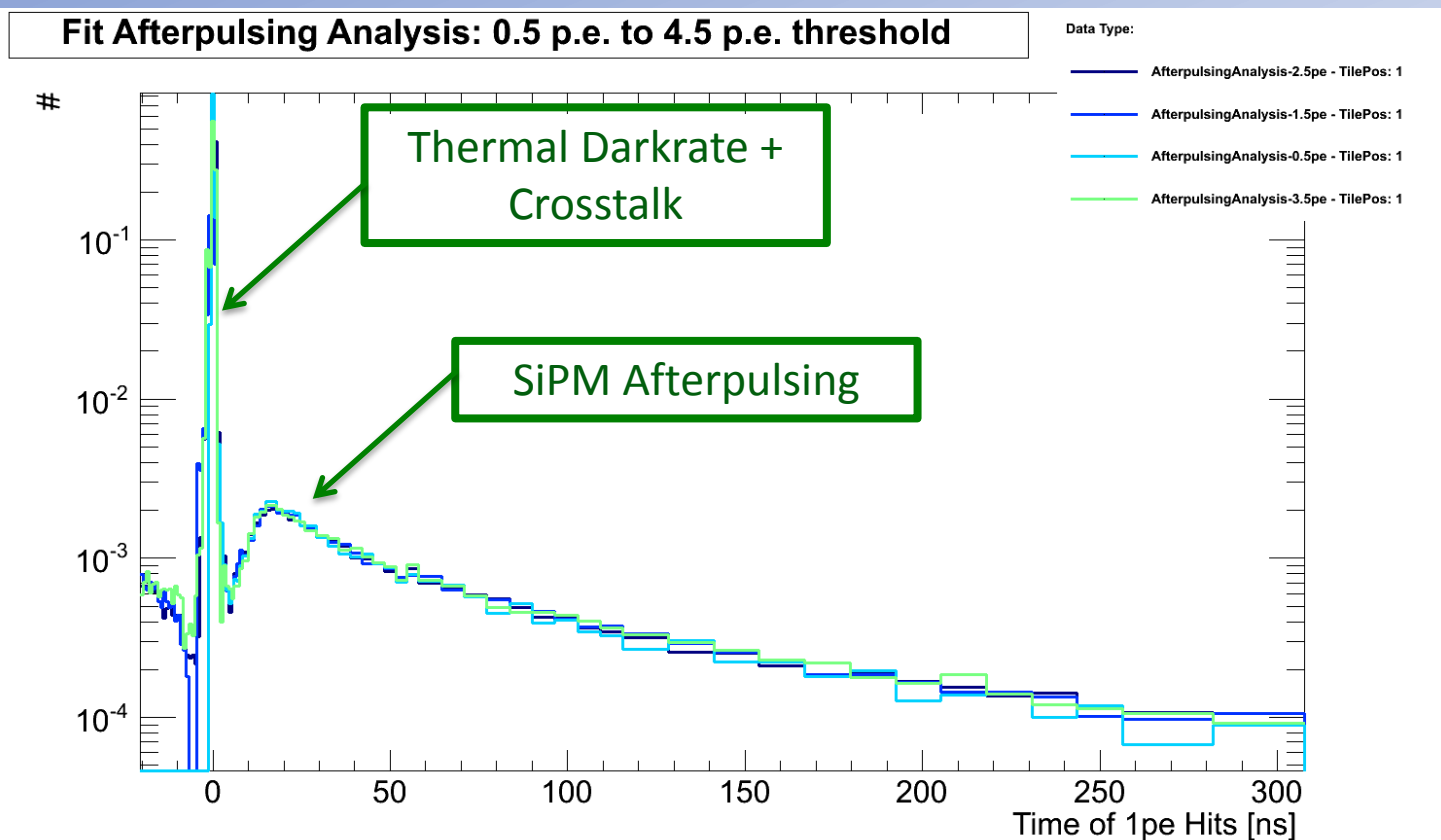


Procedure to extract the average SiPM afterpulsing per fired pixel:

Fill the time of each firing pixel into one histogram for all selected waveforms

→ Normalize to the number of selected waveforms

→ AP distribution independent of the number of fired pixels → scale





Afterpulsing Analysis



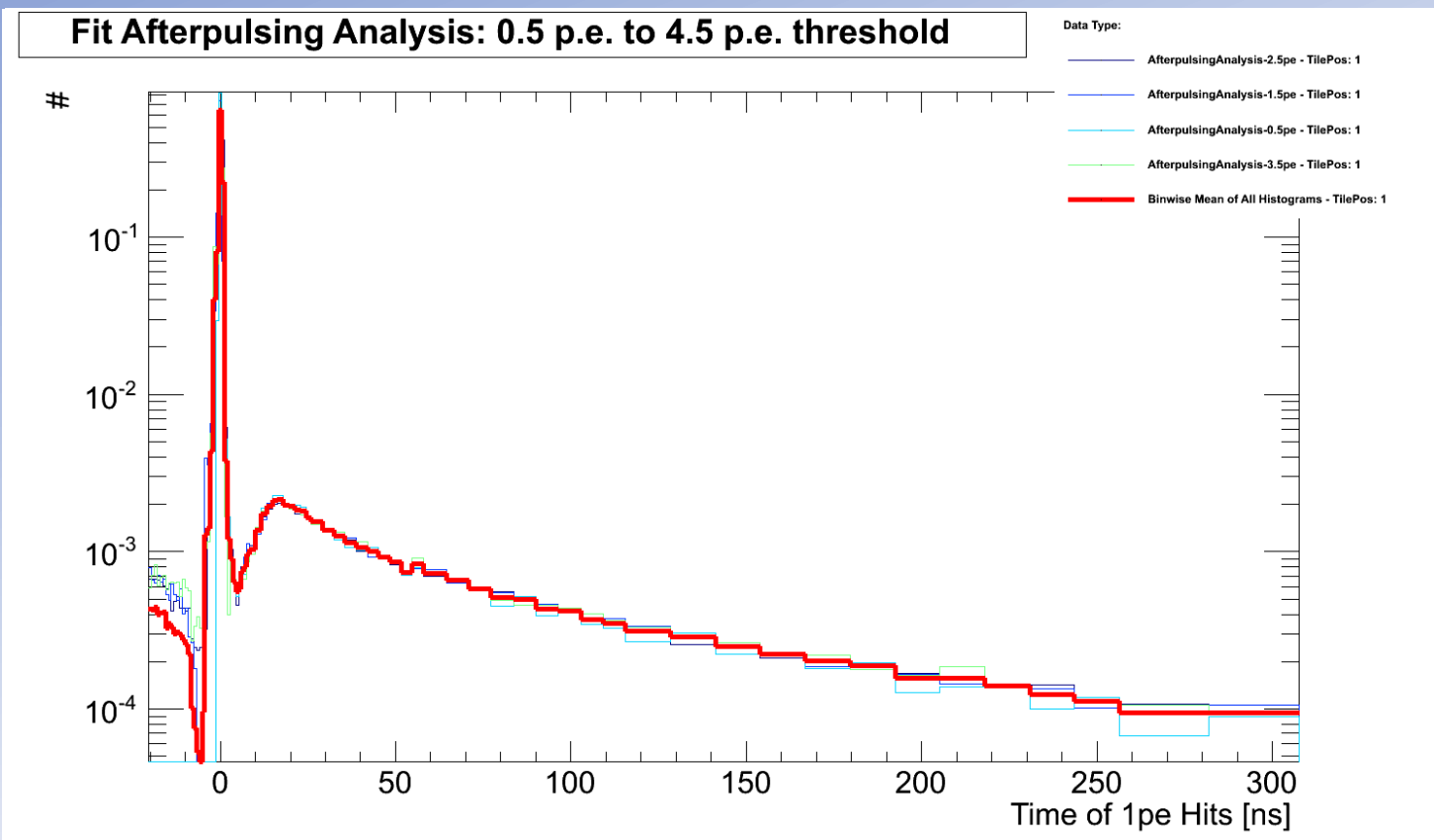
Procedure to extract the average SiPM afterpulsing per fired pixel:

Fill the time of each firing pixel into one histogram for all selected waveforms

→ Normalize to the number of selected waveforms

→ AP distribution independent of the number of fired pixels → scale

→ Increase the statistics by averaging binwise





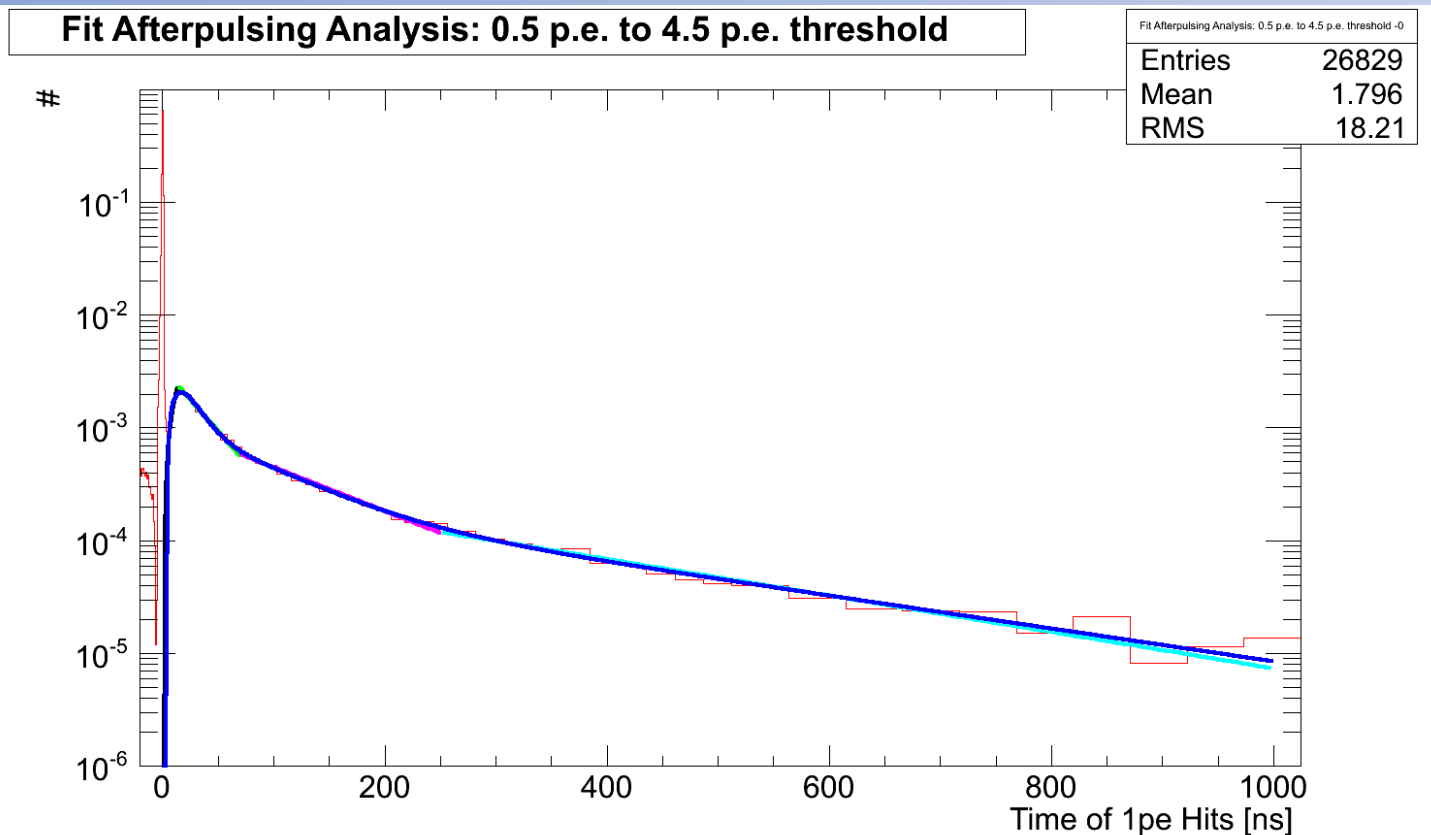
Afterpulsing Analysis



Procedure to extract the average SiPM afterpulsing per fired pixel:

AP distribution is best fit by function:

$$f(t)_{AP} = (1 - e^{-\frac{(t-t_0)}{\tau_{rec}}}) \cdot (a \cdot e^{-\frac{t}{\tau_1}} + b \cdot e^{-\frac{t}{\tau_2}} + c \cdot e^{-\frac{t}{\tau_3}})$$





Afterpulsing Analysis

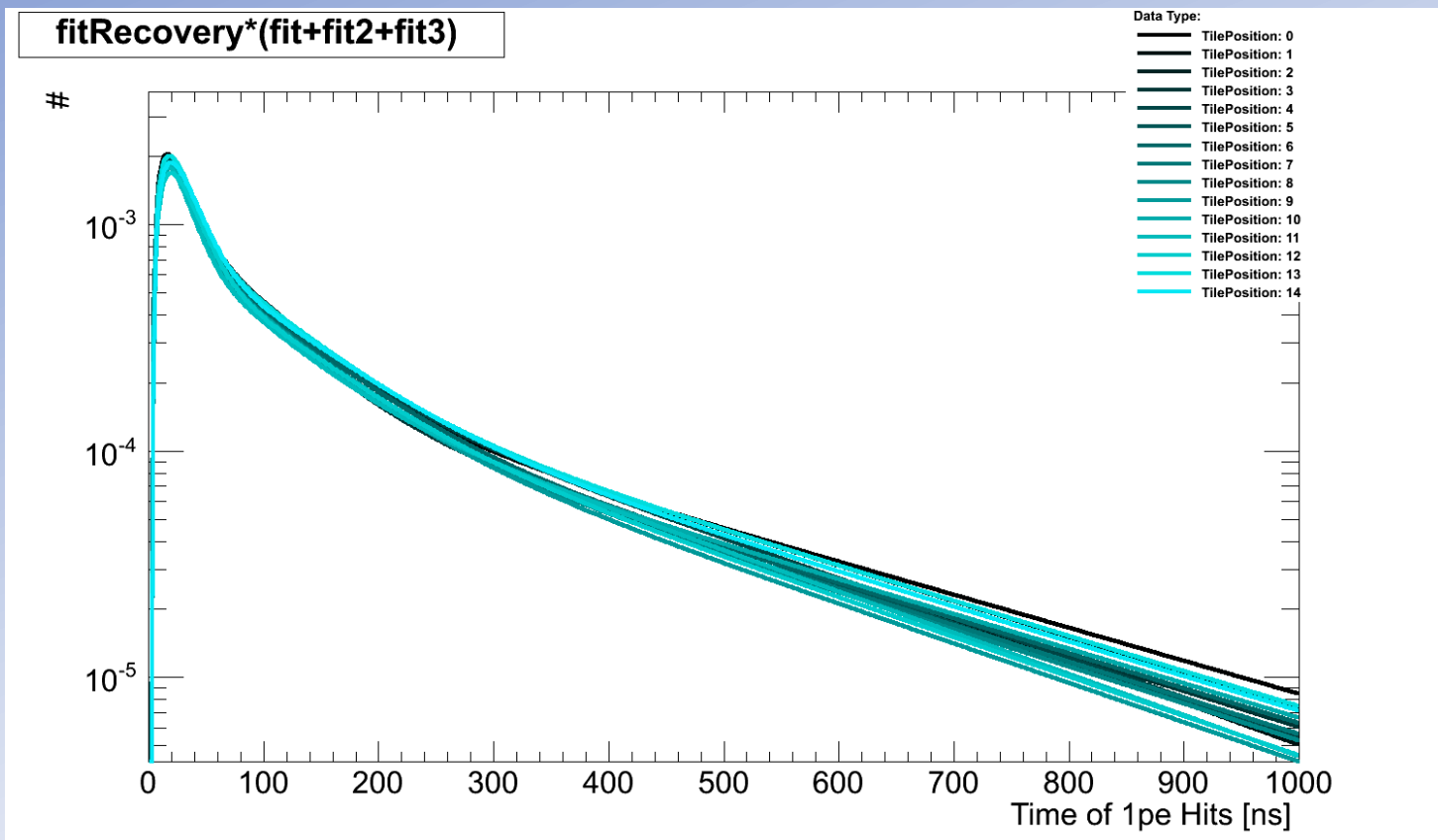


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→ Obtain a stable fit to the AP distribution for all T3B tiles → use as template to correct for AP





Afterpulsing Analysis

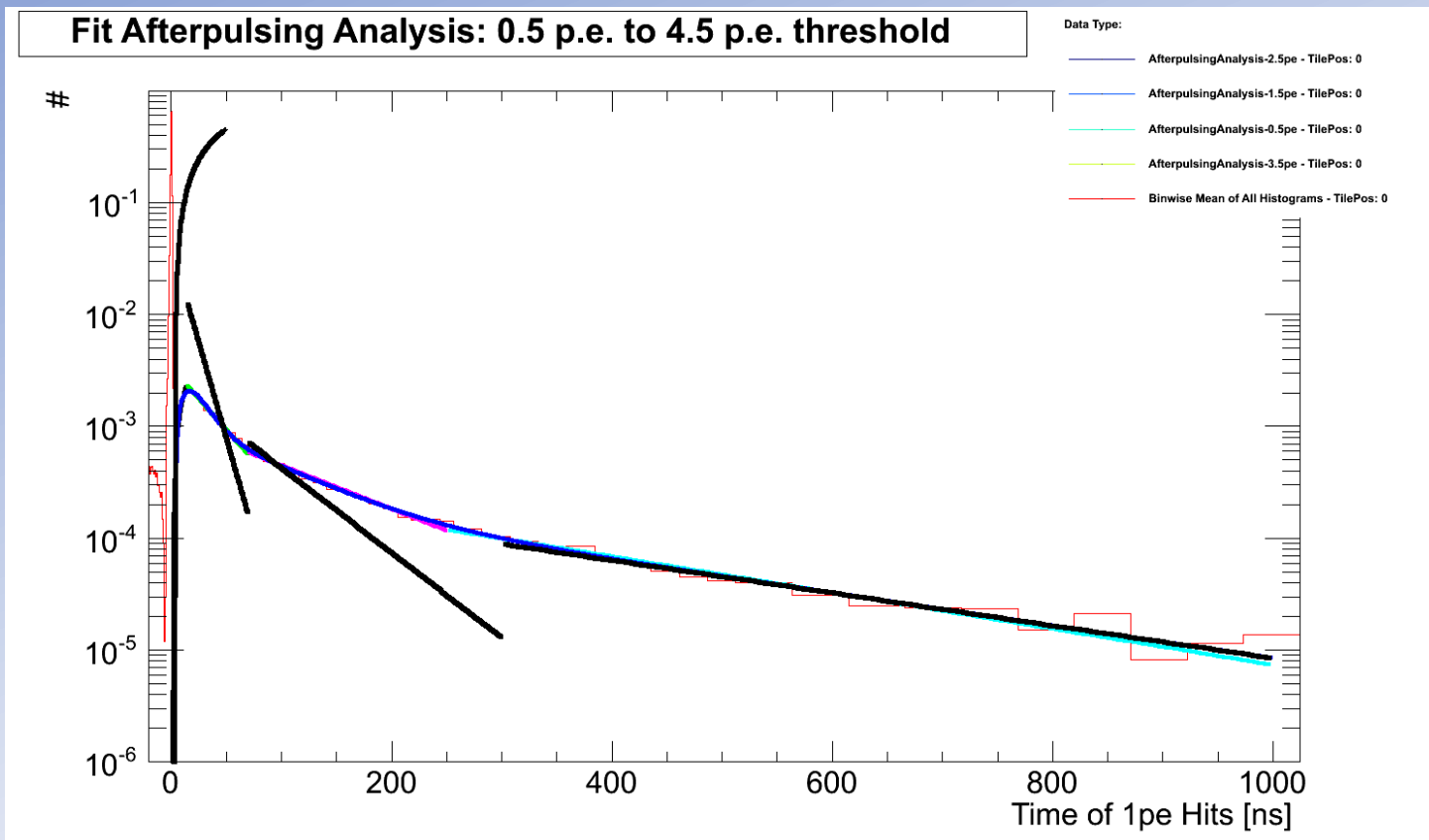


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Afterpulsing Analysis

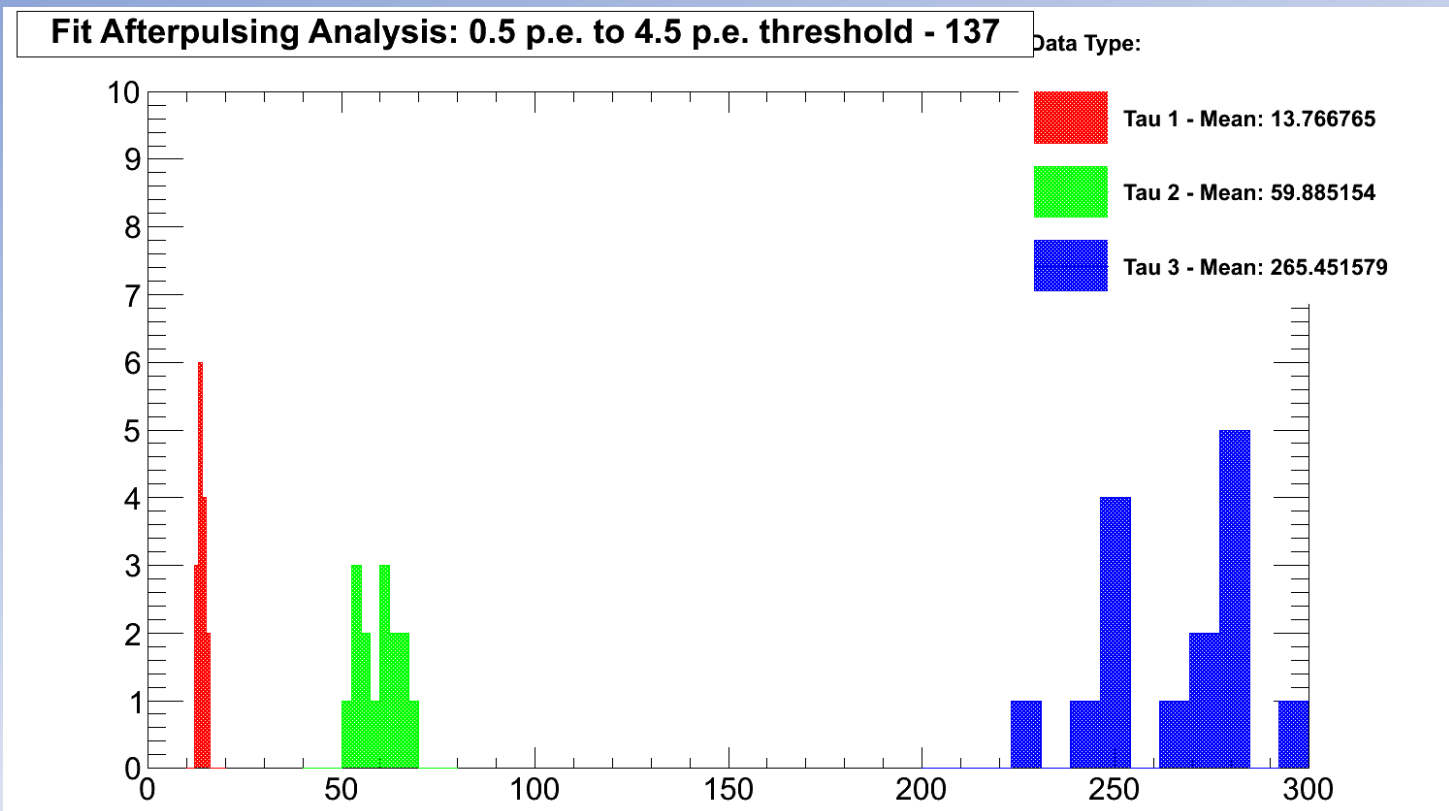


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- Obtain a stable fit to the AP distribution for all T3B tiles → use as template to correct for AP
- Attempt to extract the time constants of the AP distribution





TIME OF HIT ANALYSIS



Time of Hit Analysis



Analysis Criteria:

- Events Rejected for analysis:
 - Overshoots of vertical range of T3B Oscilloscopes (400mV)
 - Events without or with multiple particles (scintillator coincidence signal)
 - Events with < 15 p.e. in the whole acquisition window of 2.4 microseconds

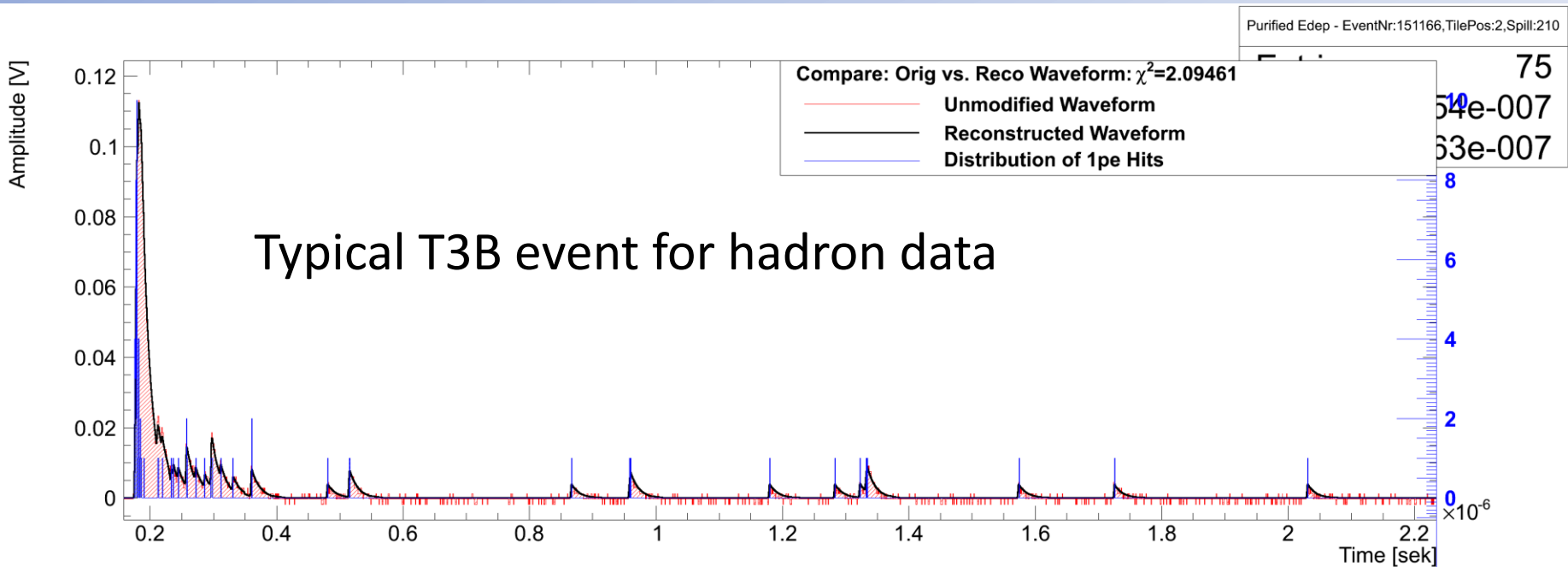


Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

- Fill the time of every single fired pixel into one timing histogram (per E and tile)



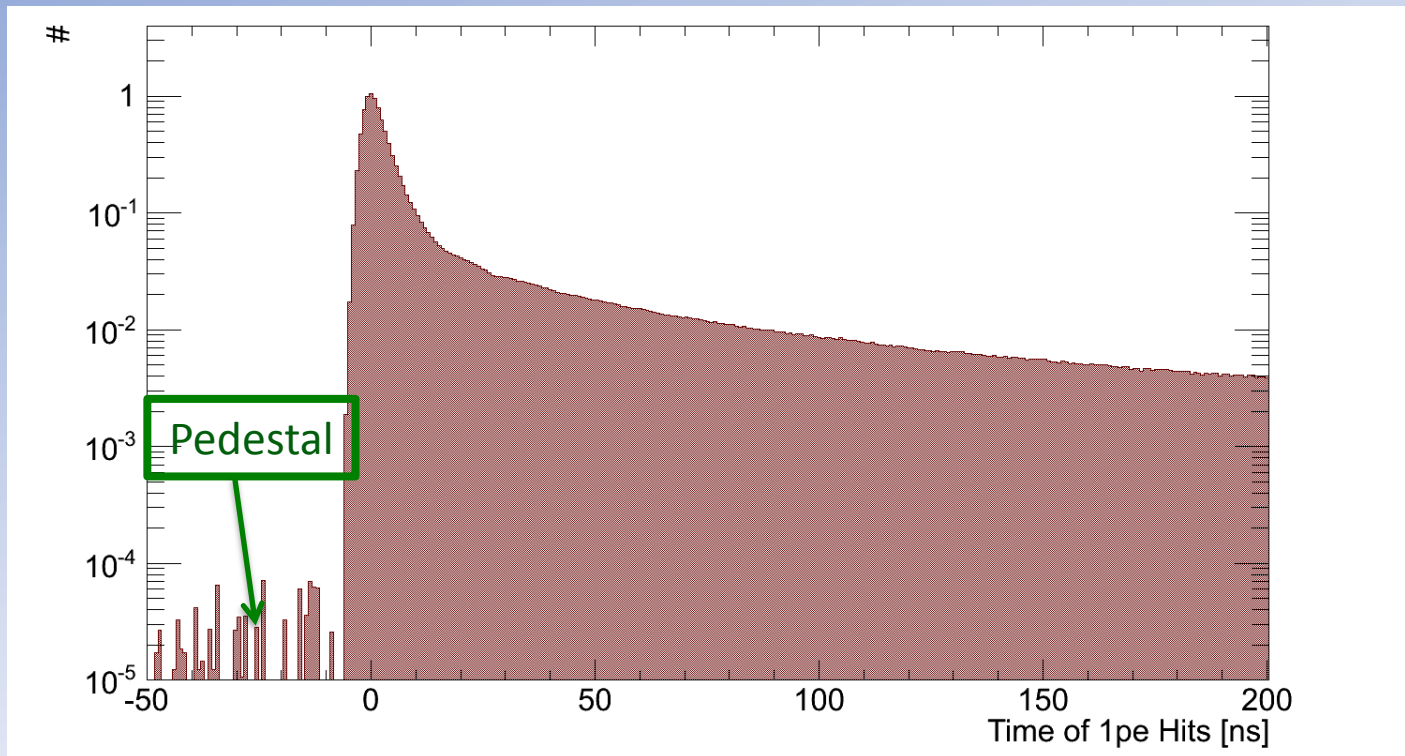


Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

- Fill the time of every single fired pixel into one timing histogram (per E and tile)



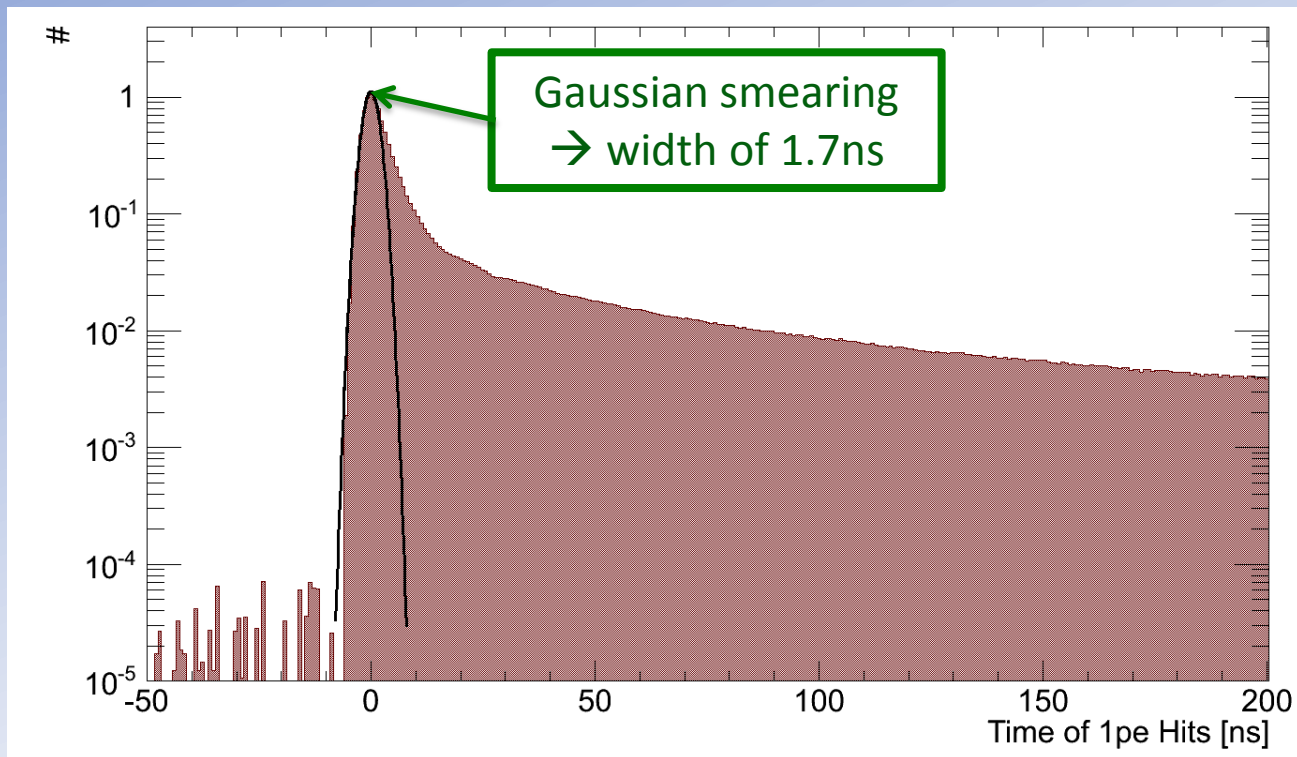


Time of Hit Analysis



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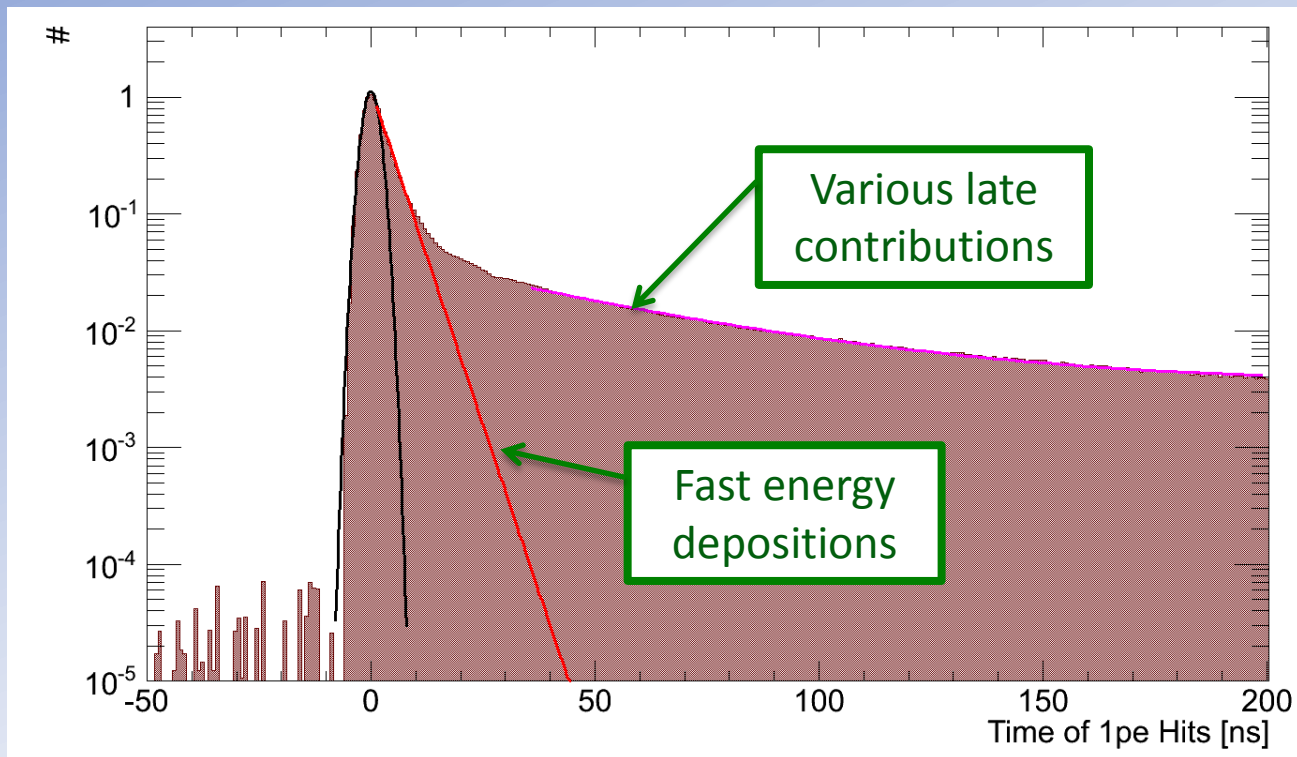


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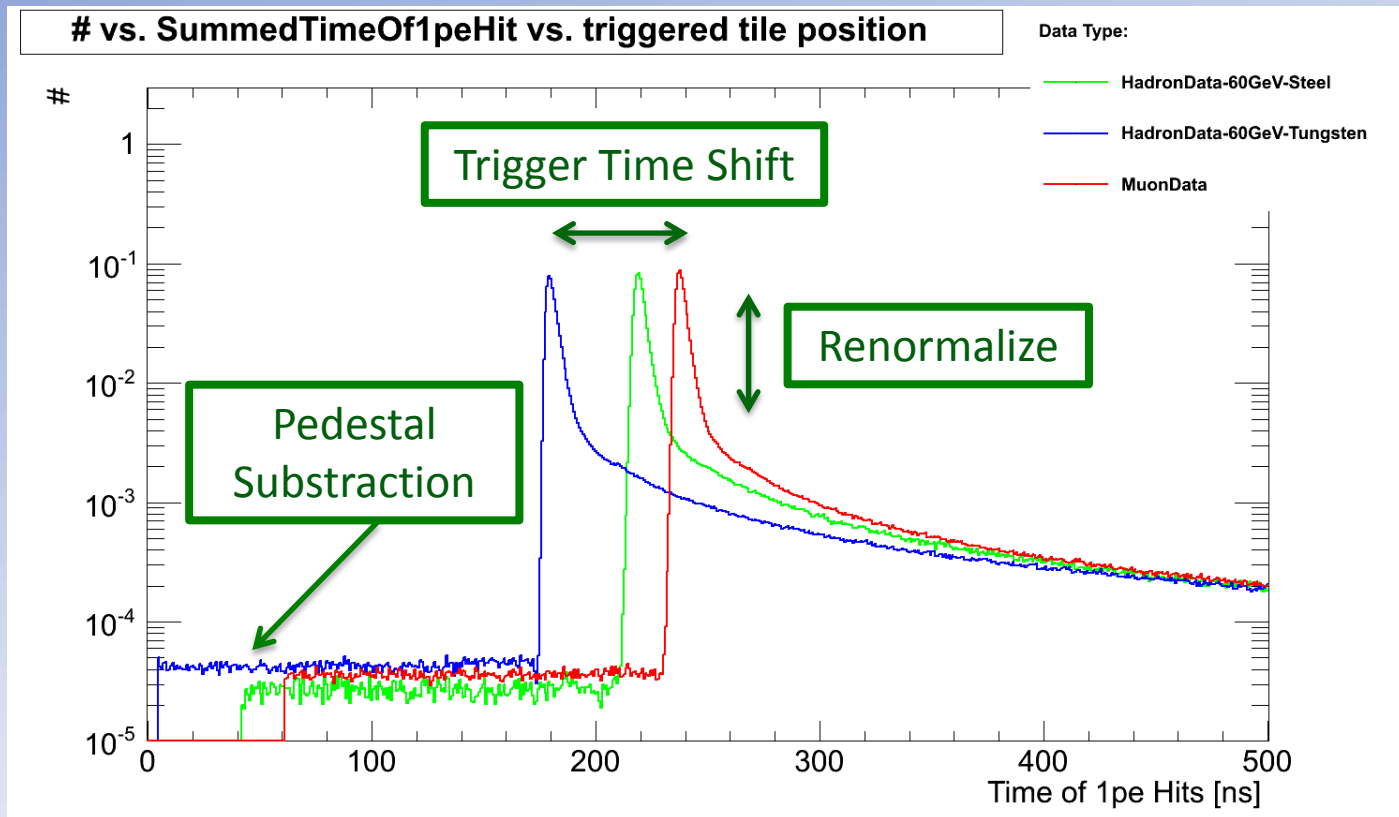


Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

- Fill the time of every single fired pixel into one timing histogram (per E and tile)
- Match different run periods → Pedestal sub., Renormalization, time shift...



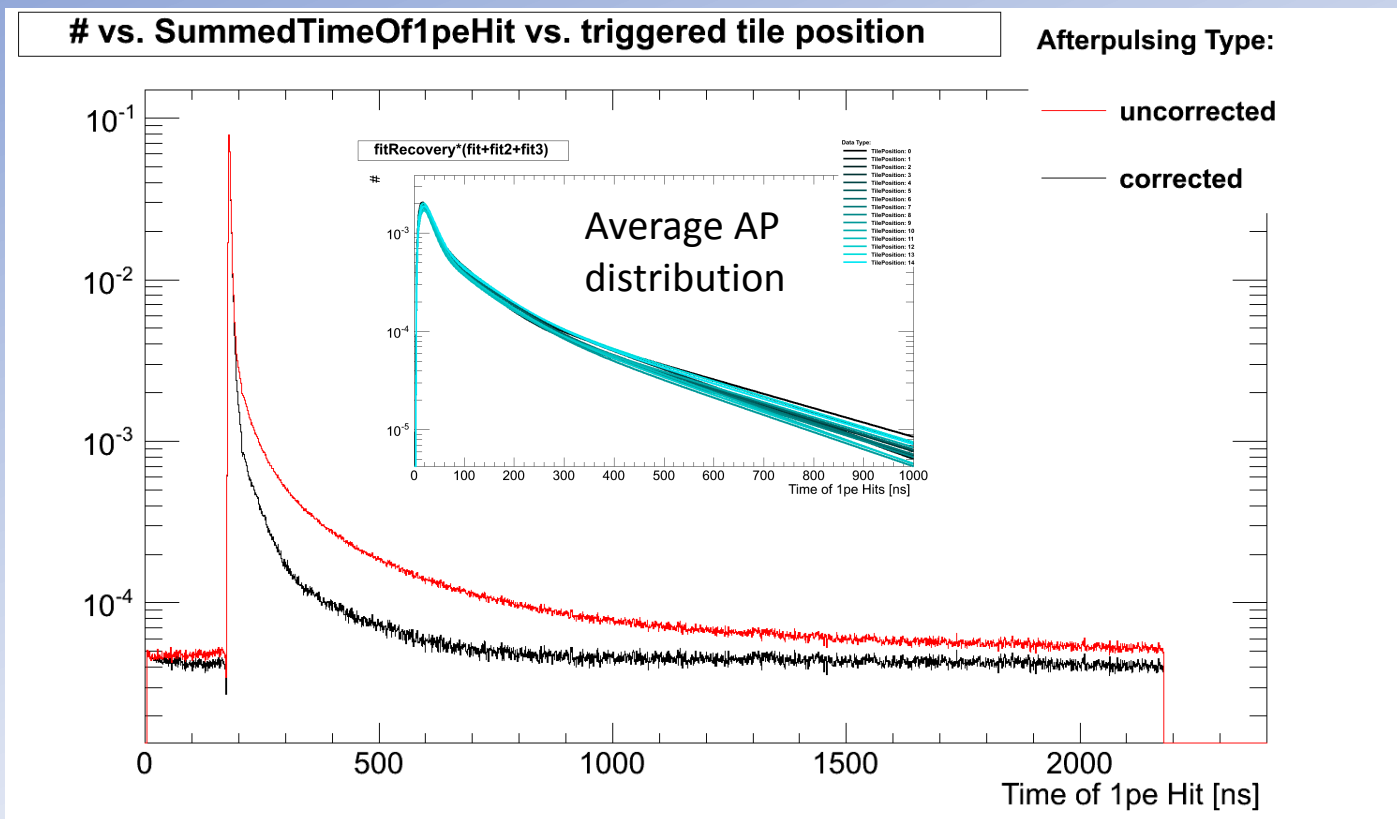


Time of Hit Analysis



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- Fill the time of every single fired pixel into one timing histogram (per E and tile)
- Match different run periods → Pedestal sub., Renormalization, time shift...
- Correct Afterpulsing → subtract tile corresp. AP distribution for every fired pixel





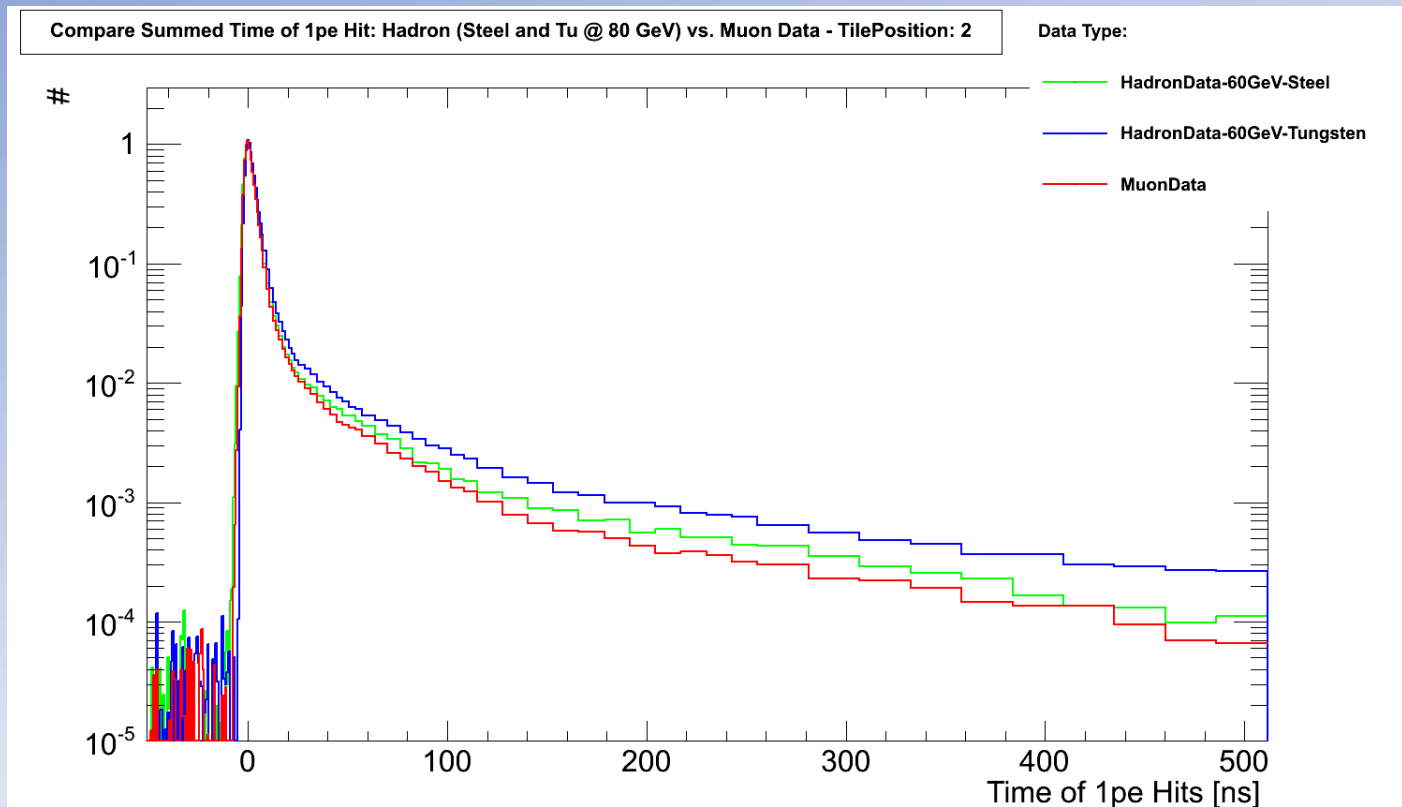
Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

Muon Data does not drop to zero after AP correction:

- Additional long time components: Mirror Foil, long time constant of scintillator?
- AP behaviour must be adjusted to current SiPM gain (T, V(bias))
- Contributions should be in Hadron AND Muon data identially
- Comparison possible



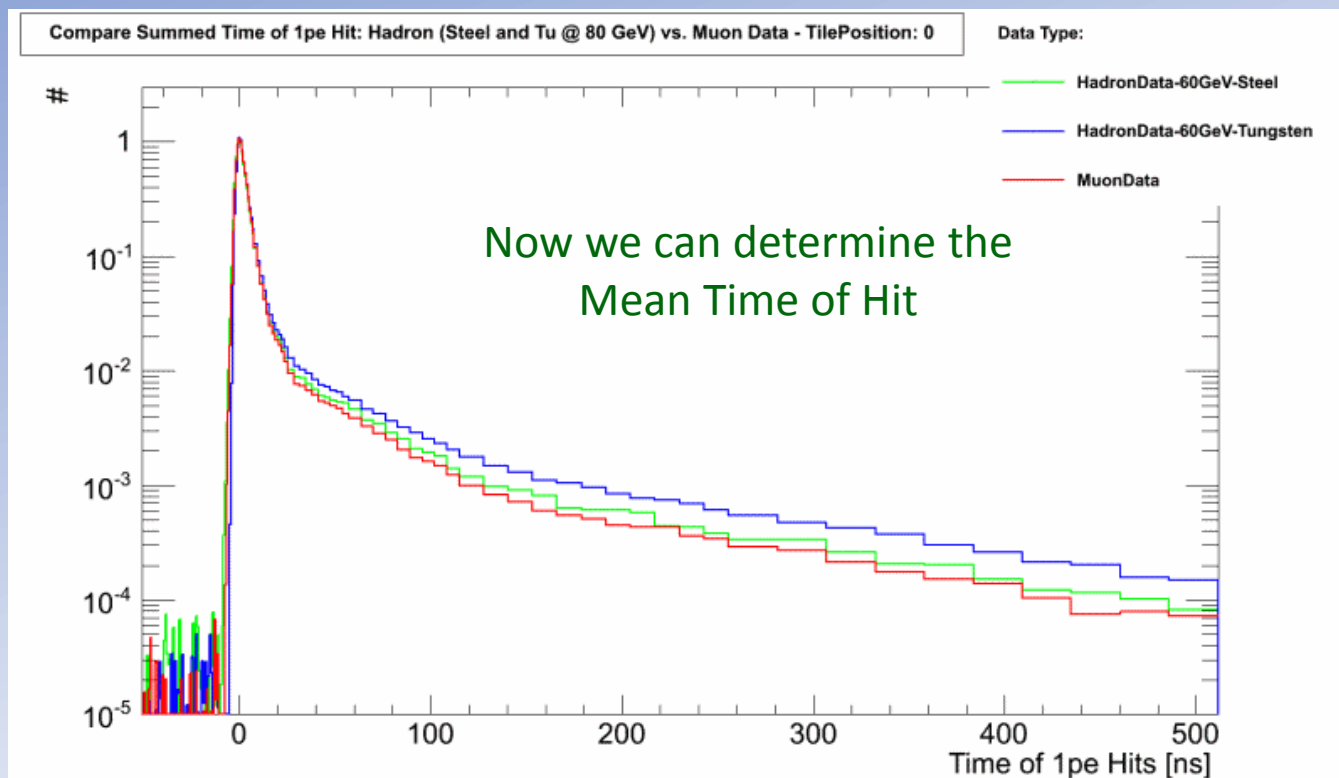


Time of Hit Analysis



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- Use only muon data of central tile due to high statistics (when necessary)



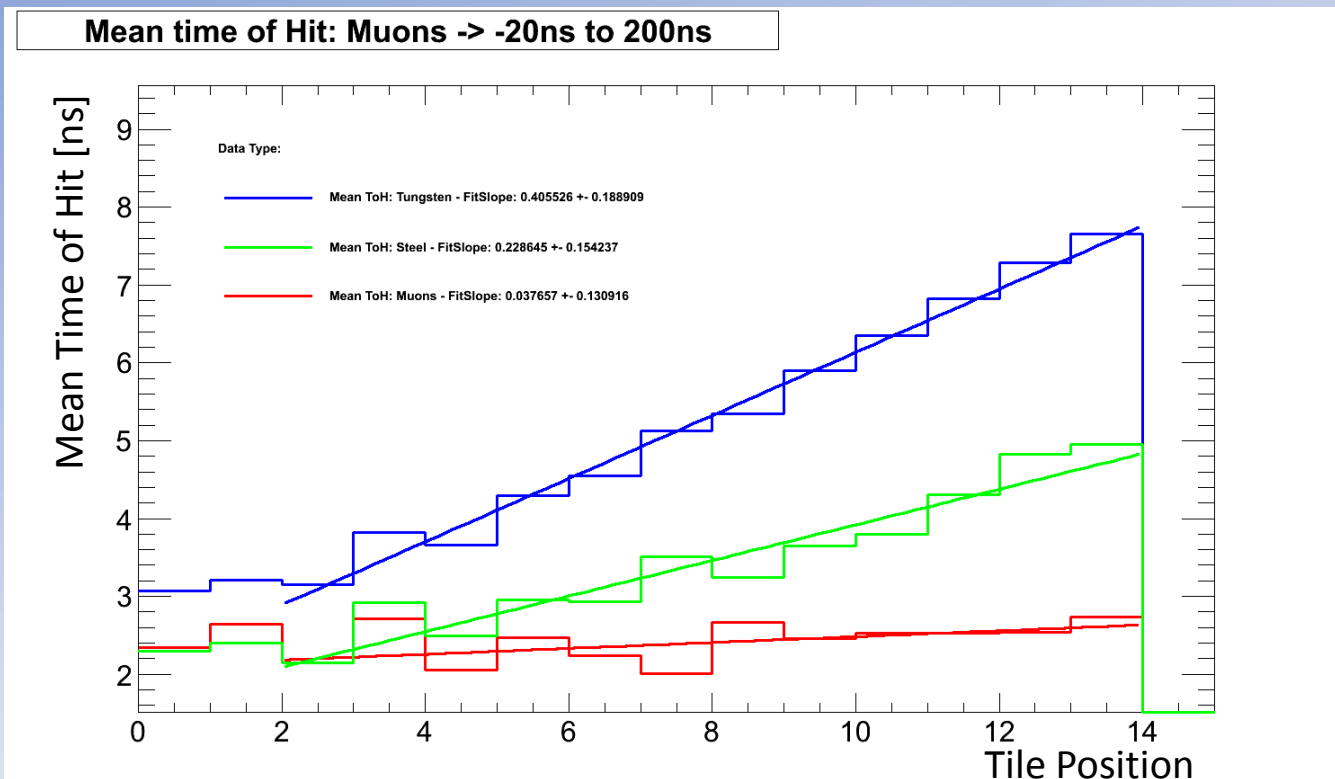


Time of Hit Analysis



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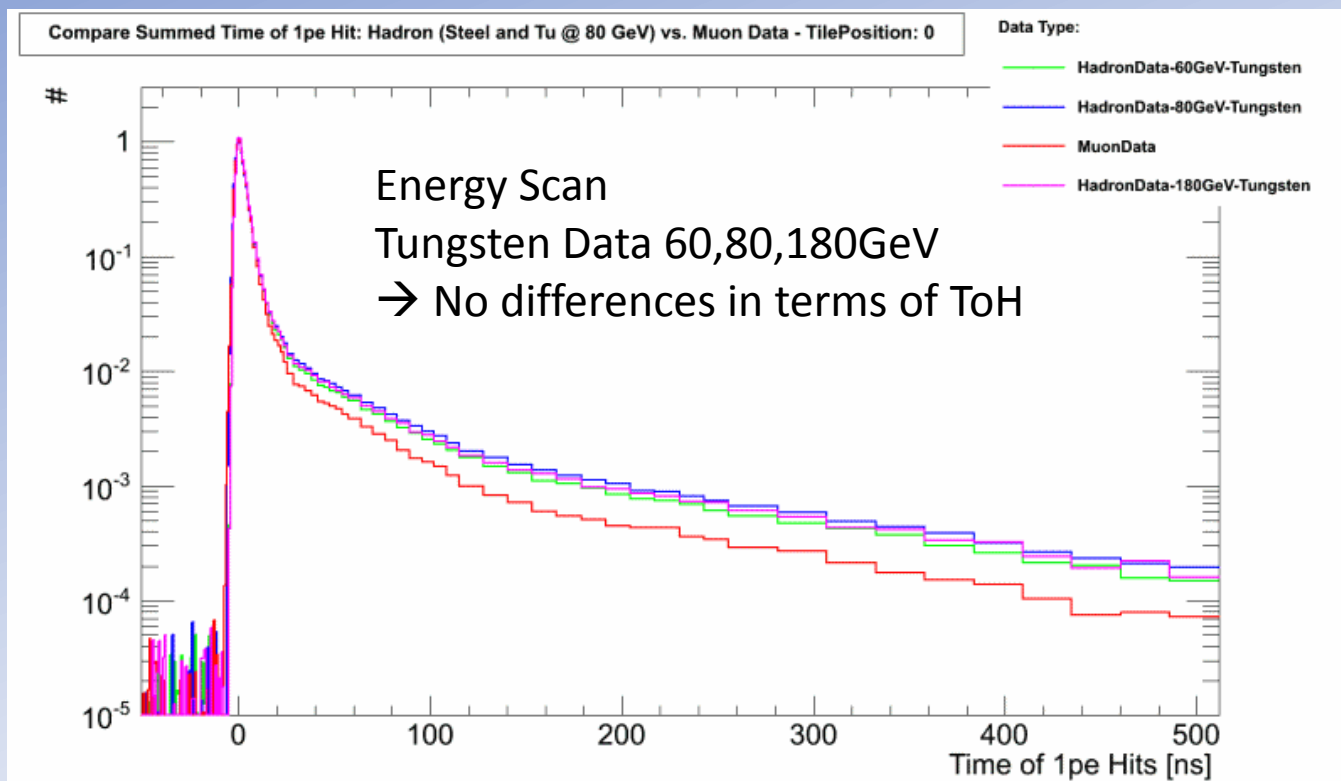


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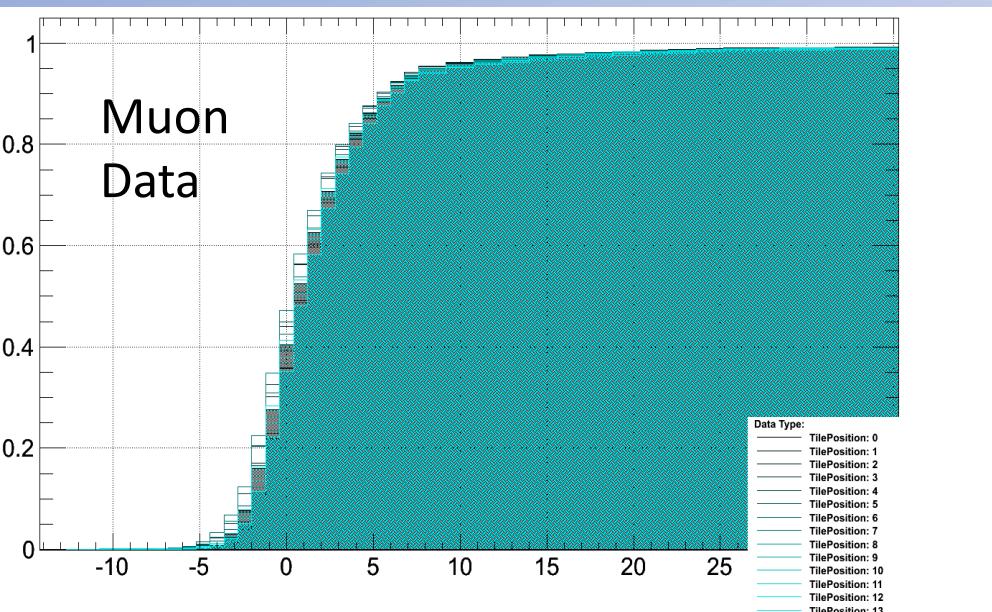


Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

- Which fraction of the total Energy is deposited at which time?
 - Tilewise
 - T3B Overall



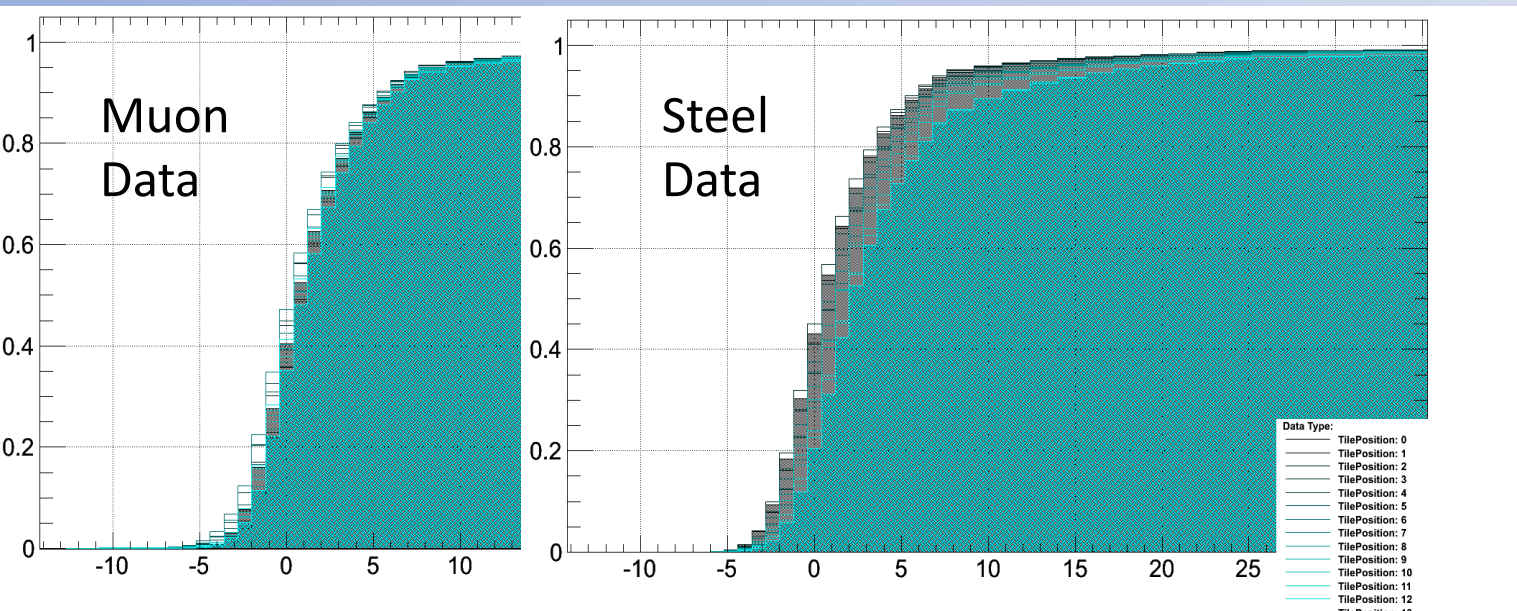


Time of Hit Analysis



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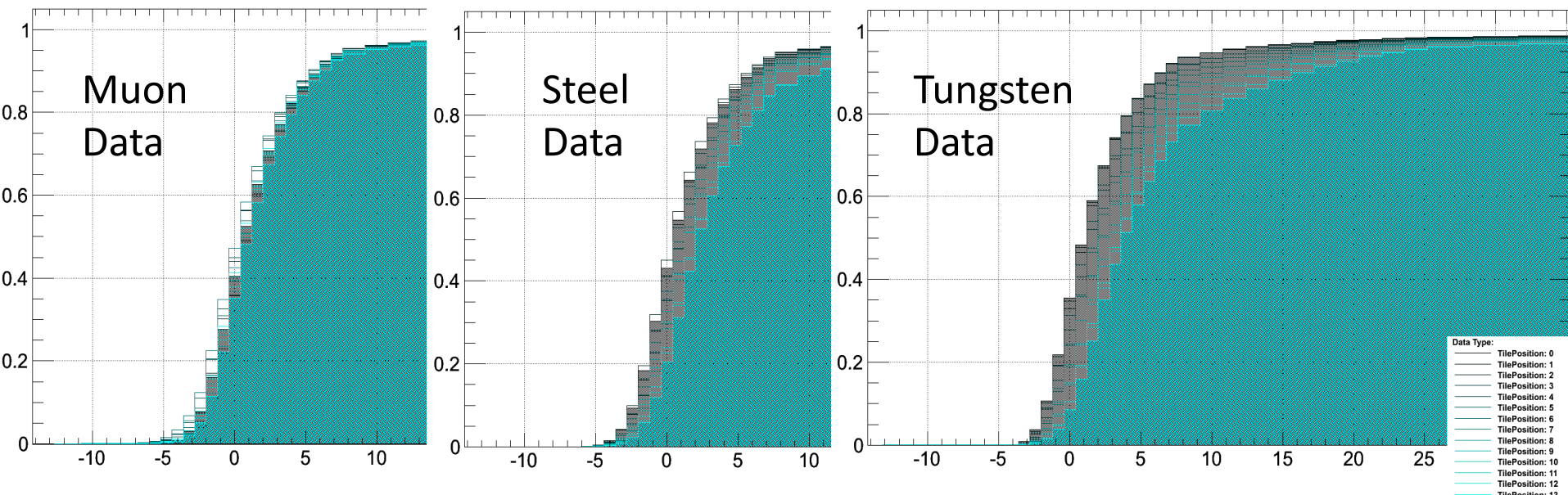


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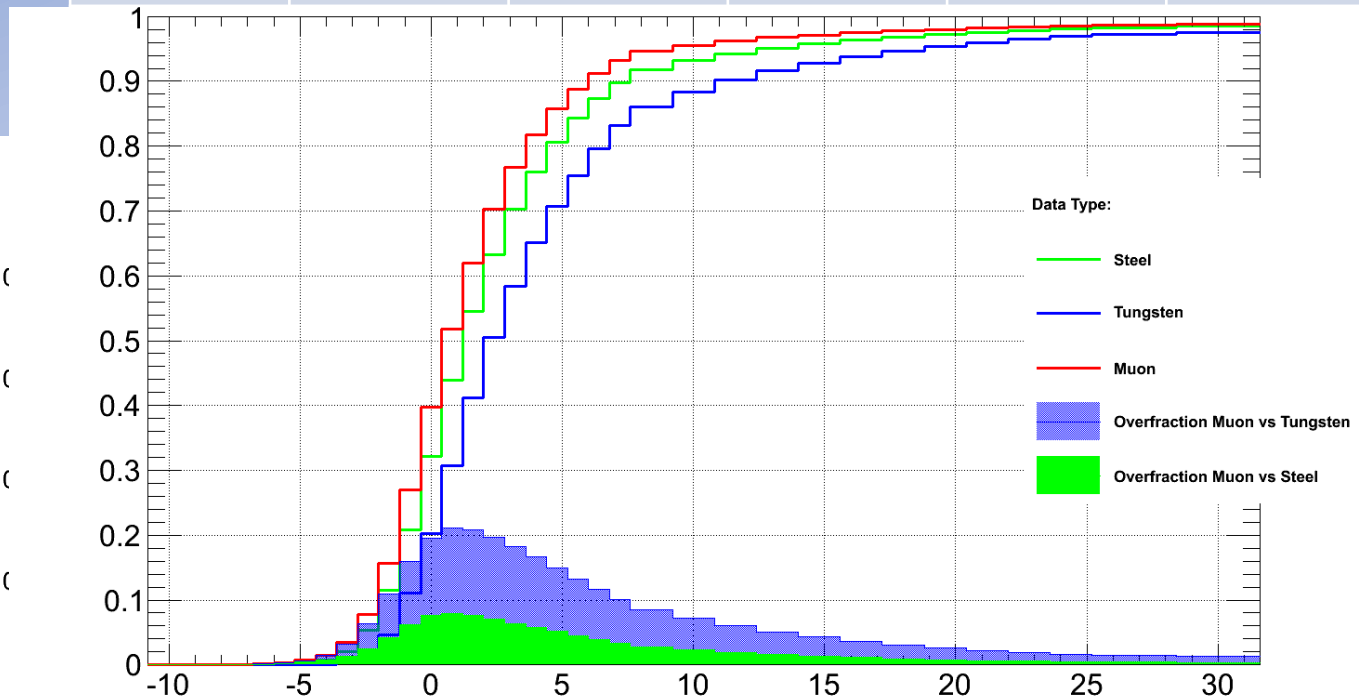
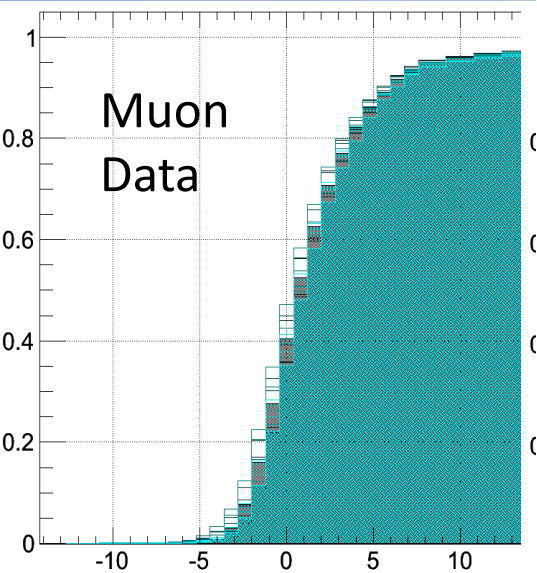
Time of Hit Analysis



Analysis Steps (All plots for Hadron Data, 60GeV, Tungsten, 4mio Events):

- Which fraction of the total Energy is deposited at which time?
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 - T3B Overall

Time [ns]	Fraction Mu [T3B]	Fraction Steel [T3B]	Fraction Steel [CDR]	Fraction W [T3B]	Fraction W [CDR]
6	89%	84%	90%	76%	n.A.
10	96%	93%	n.A.	88%	n.A.
25	99%	98%	n.A.	97%	82%





Time of Hit Analysis



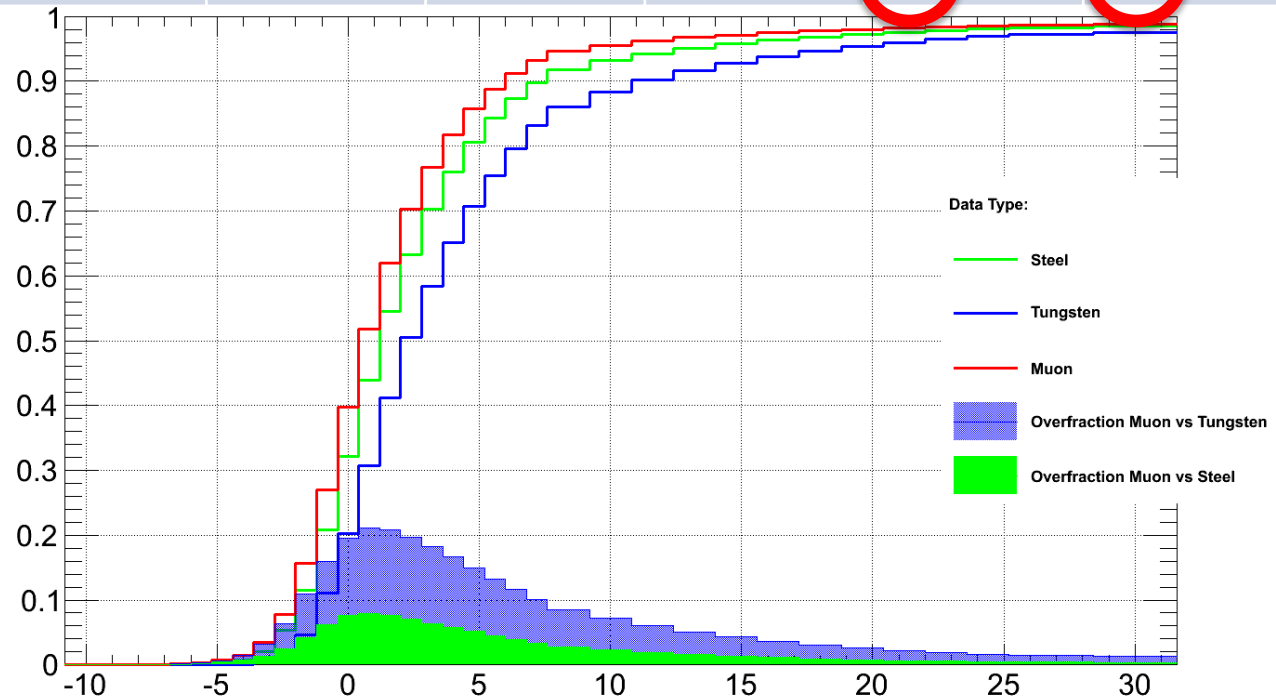
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10	96%	93%	n.A.	88%	n.A.
25	99%	98%	n.A.	97%	82%

Discrepancies due to:

- Integrating over only a small strip instead of whole shower?
- Position behind the HCAL? → Answer through Calice-T3B sync. Information
- Missing correction factors (e.g. Overshoot, Saturation correction...)





SUMMARY AND OUTLOOK



Summary and Outlook



Summary:

We presented results on the time development of hadronic showers

Mean time of first hit:

- We see significant differences in the lateral shower timing between Tu and Steel, but no differences for different energies
- Calibration to the MIP scale allows us to quantify late energy depositions
- Late (first) energy depositions are small for Steel but can be high for Tungsten

Mean time of Hit:

- Device specific templates could be used to correct for SiPM afterpulsing
- We see significant differences in the lateral shower timing between Tu and Steel, but no differences for different energies
- The timing of the fraction of total energy deposited within the T3B strip shows differences wrt. the values in the CLIC CDR



Summary and Outlook



Outlook:

Geant4: Need extensive simulation study to compare to data results and identify discrepancies

Calibration:

- Implement Afterpulsing correction depending on SiPM gain (Overvoltage, Temperature)
- Use synchronization information to clean data from punch throughs and muon contamination
- ...

Now pass to Lars with his successful T3B-Calice synchronization studies





BACKUP

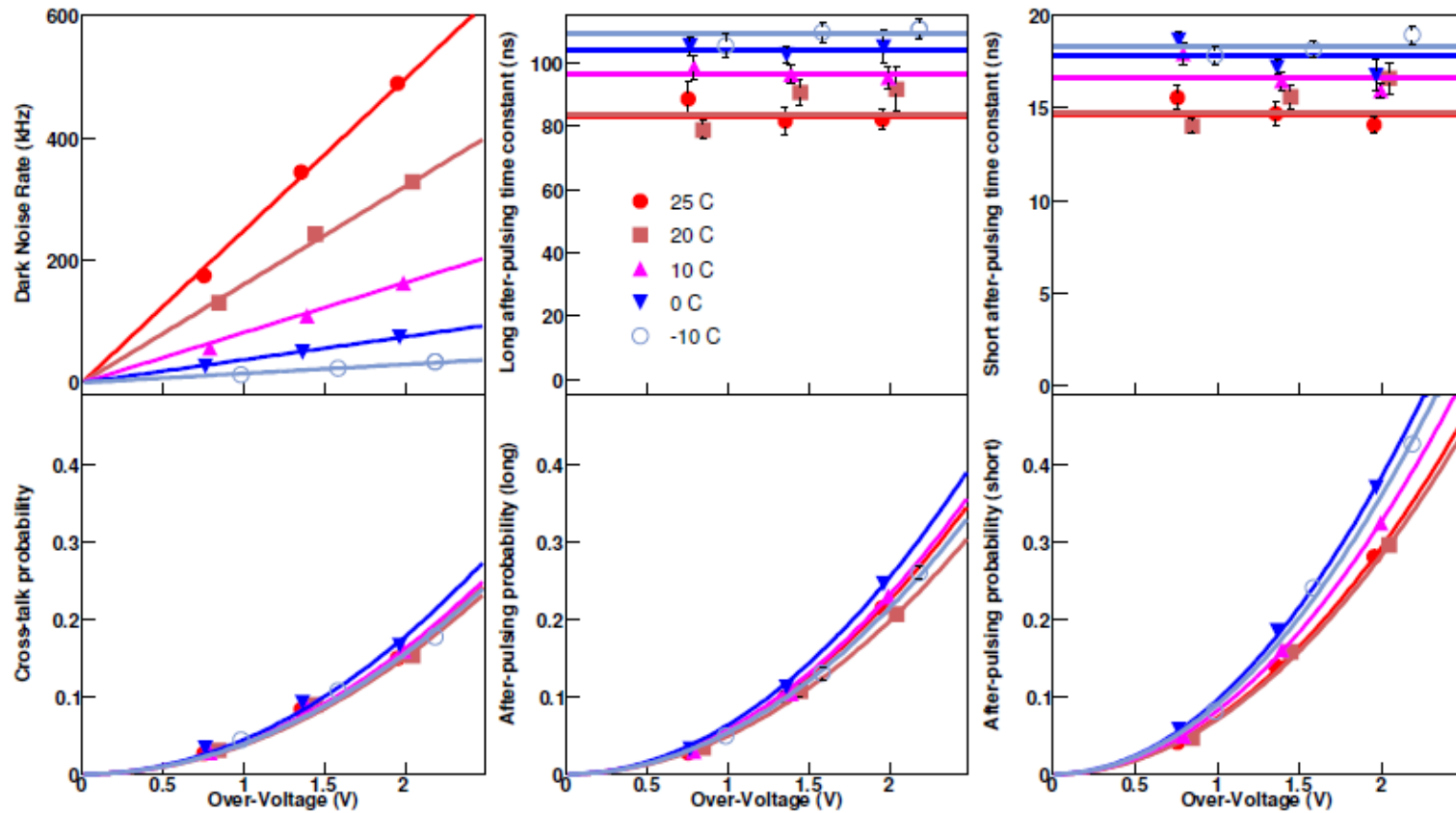


Fig. 1. Dependence of the dark noise rate, cross-talk probabilities and after-pulsing parameters on over-voltage and temperature, measured by triggering on thermally generated pulses



T3B Calibration to the MIP Scale: Sr90 Data



During the Test Beam T3B monitors the SiPM Gain continuously

→ This data can be used to calibrate energy depositions to the MIP Scale

Assumption: The MIP MPV depends in first order only(!) on the Gain



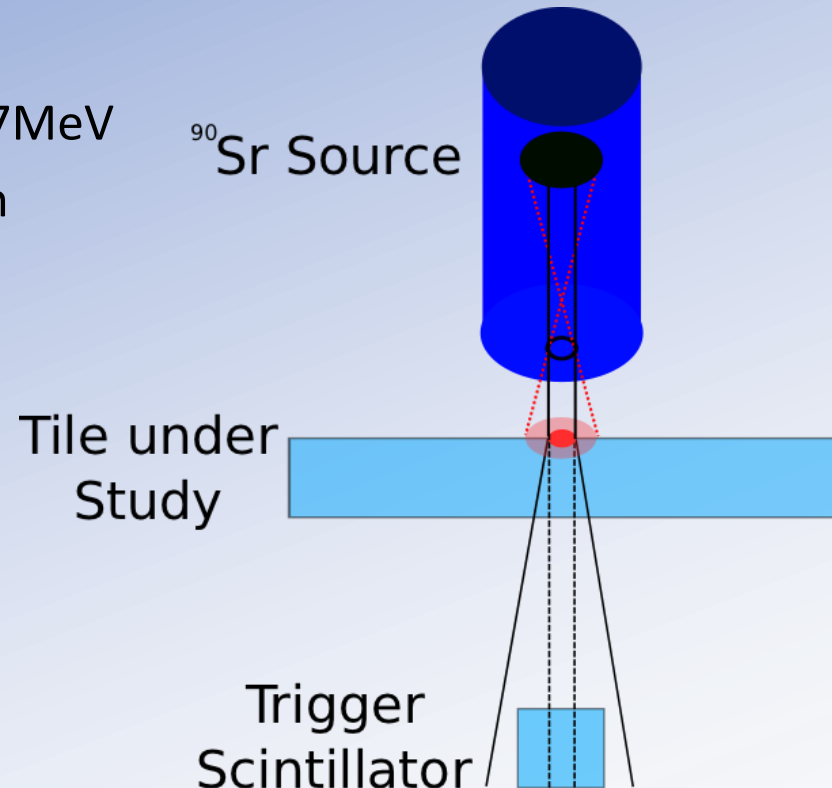
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Offline Calibration Setup:

- Sr90 Source with end point energy of 2.27MeV
- Coincidence trigger to ensure penetration of tile under study
- Consecutive calibration of all T3B cells individually
- Use T3B DAQ: Acquire Sr90 and SiPM gain data at the same time
- Use climate chamber to ensure temperature stability





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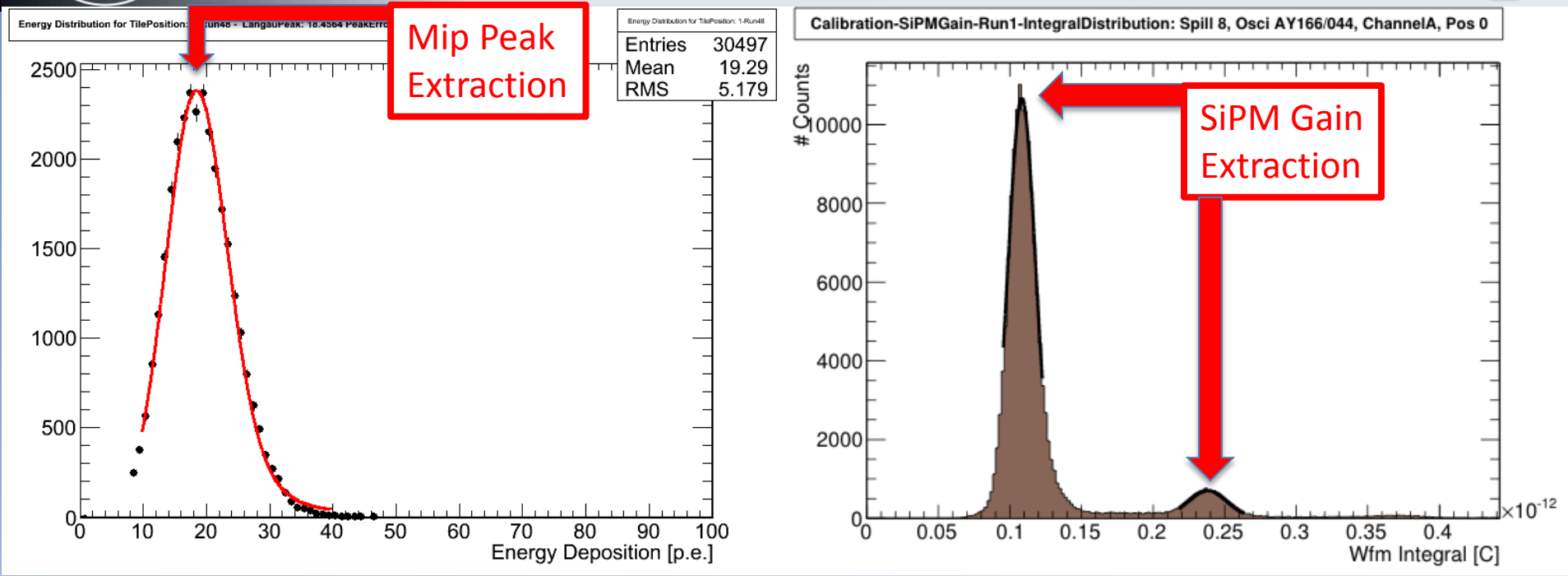
GEANT4 Simulation:
MPV (mu) = MPV(e)*0.825

$$MPV_{\mu} = MPV_e \cdot 0.825$$

Note: Electrons are no perfect MIPs → need scale factor



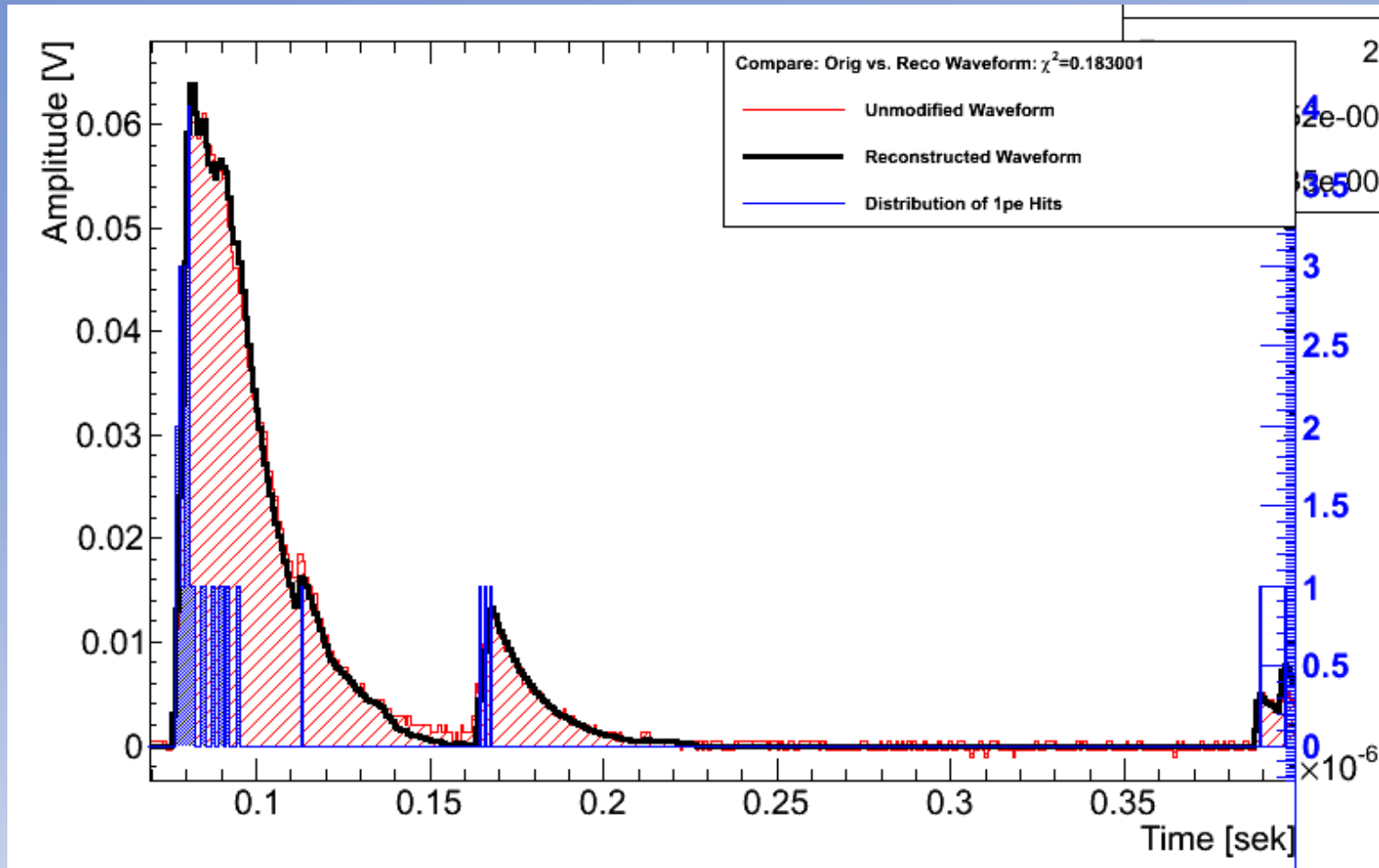
T3B Calibration to the MIP Scale: Sr90 Data



Simultaneous extraction of
SiPM Gain and **most probable value**
of energy deposition of Sr90 electrons



The Time Integration Window



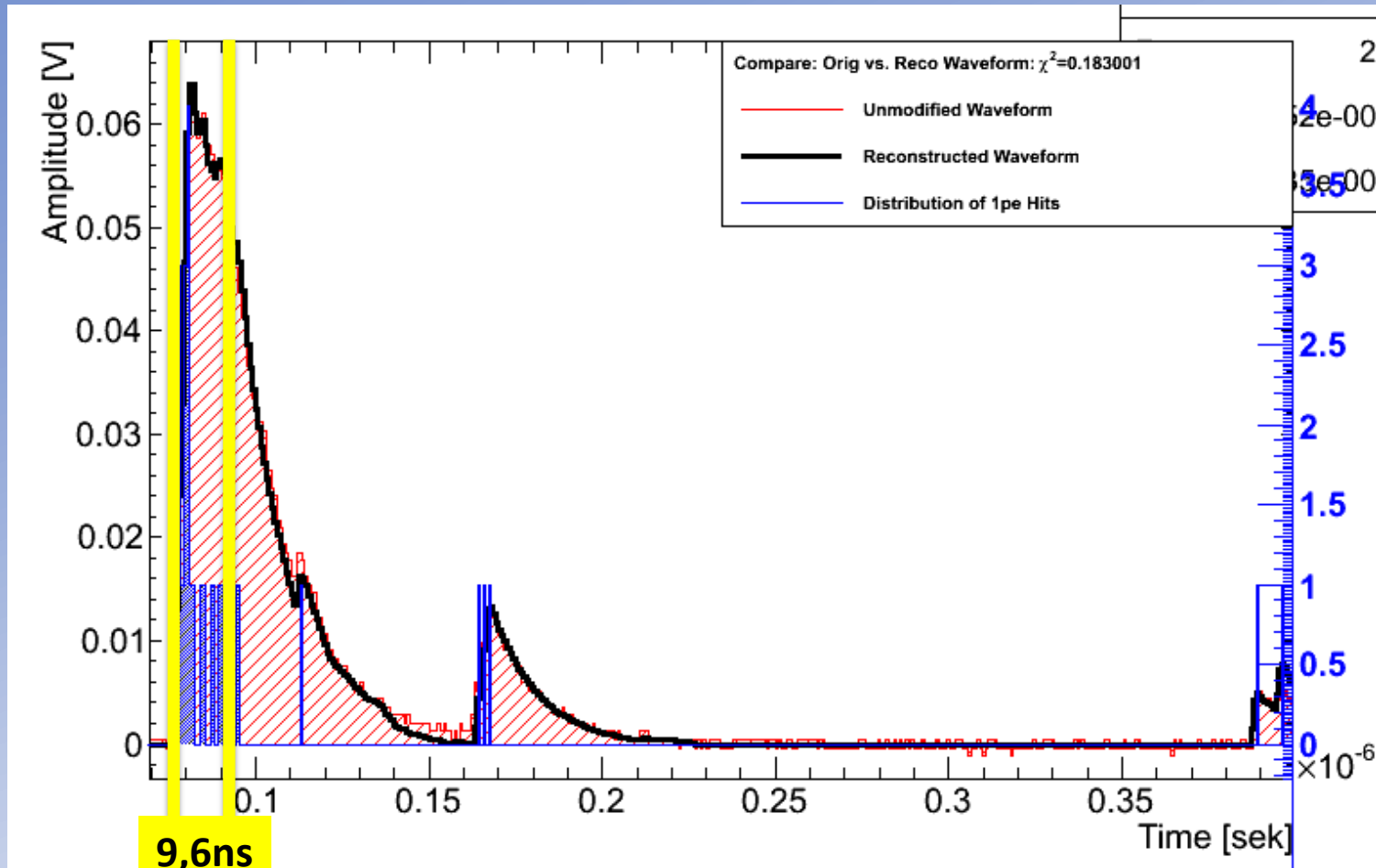
The MPV is very sensitive on the Time Integration Window

→ Dominant effect: SiPM Afterpulsing

- Separate afterpulsing from energy depositions
- Study the effect of afterpulsing



The Time Integration Window



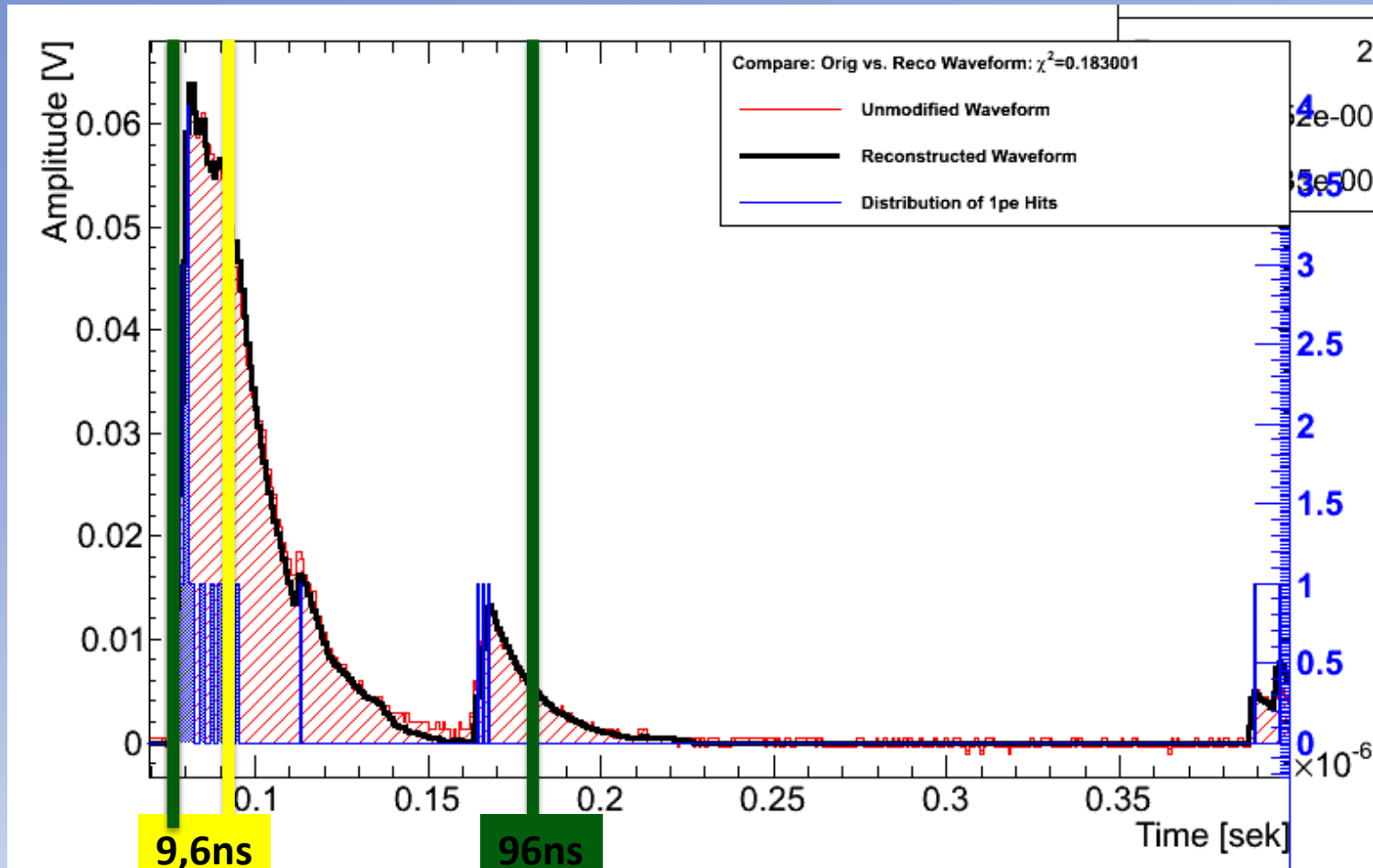
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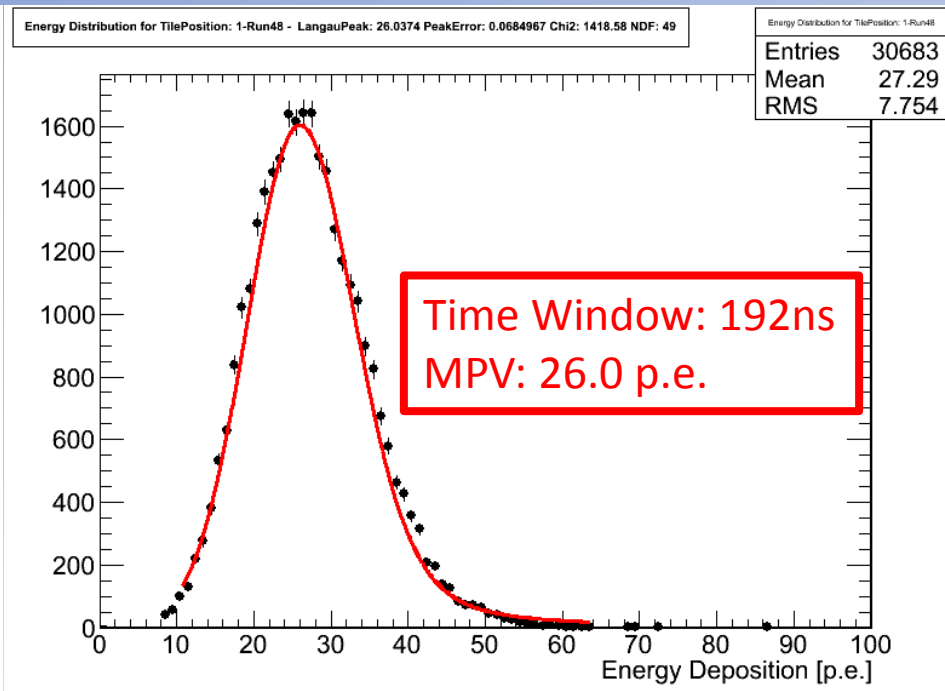
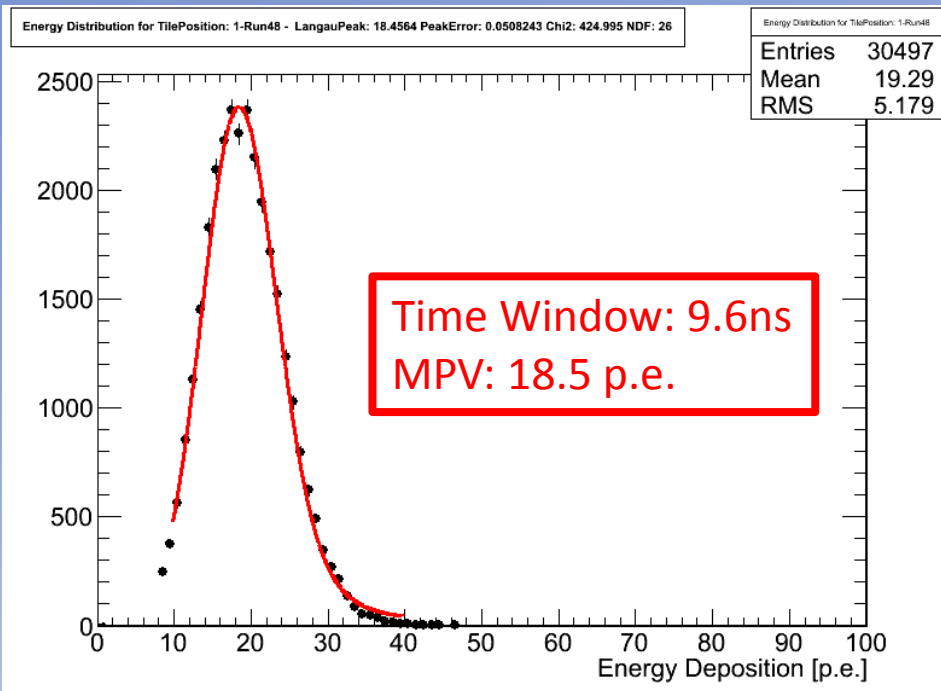
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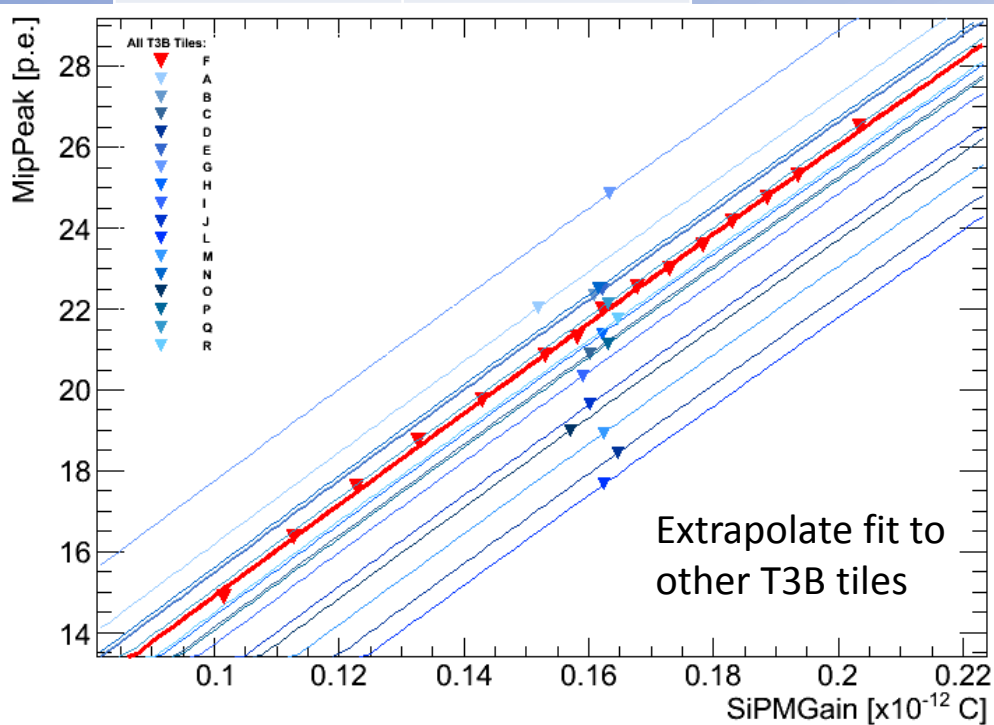


The Effect of Afterpulsing



Bias Voltage Scan for one T3B
“Master Tile”
One Measurement for
all other T3B tiles

Time Window	MPV-Gain dependence
9.6 ns	linear
307.2 ns	quadratic



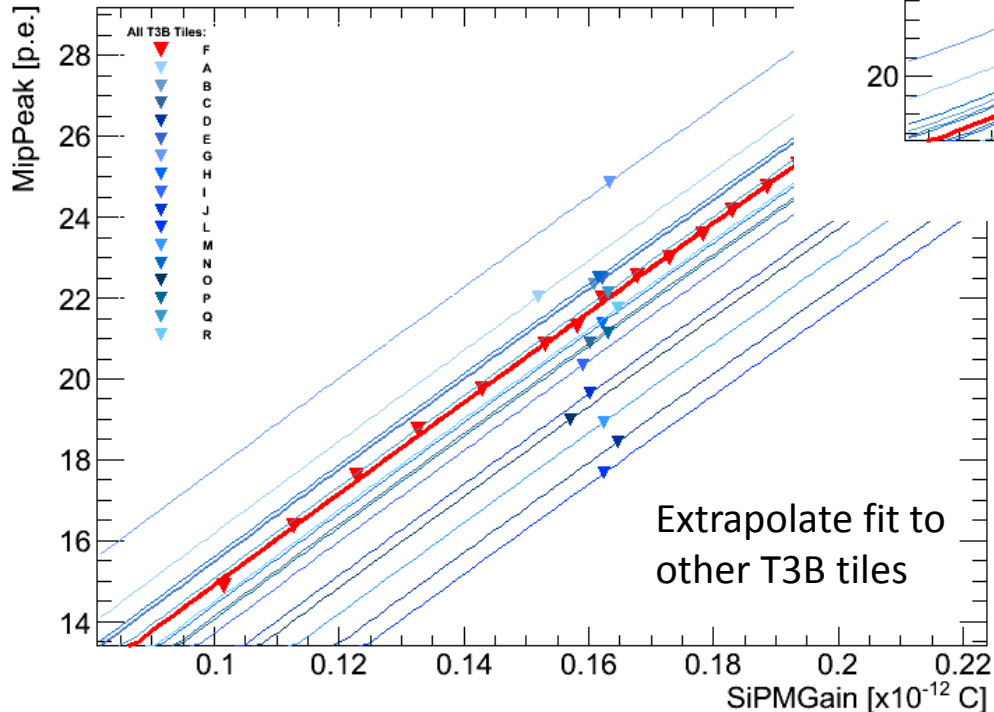
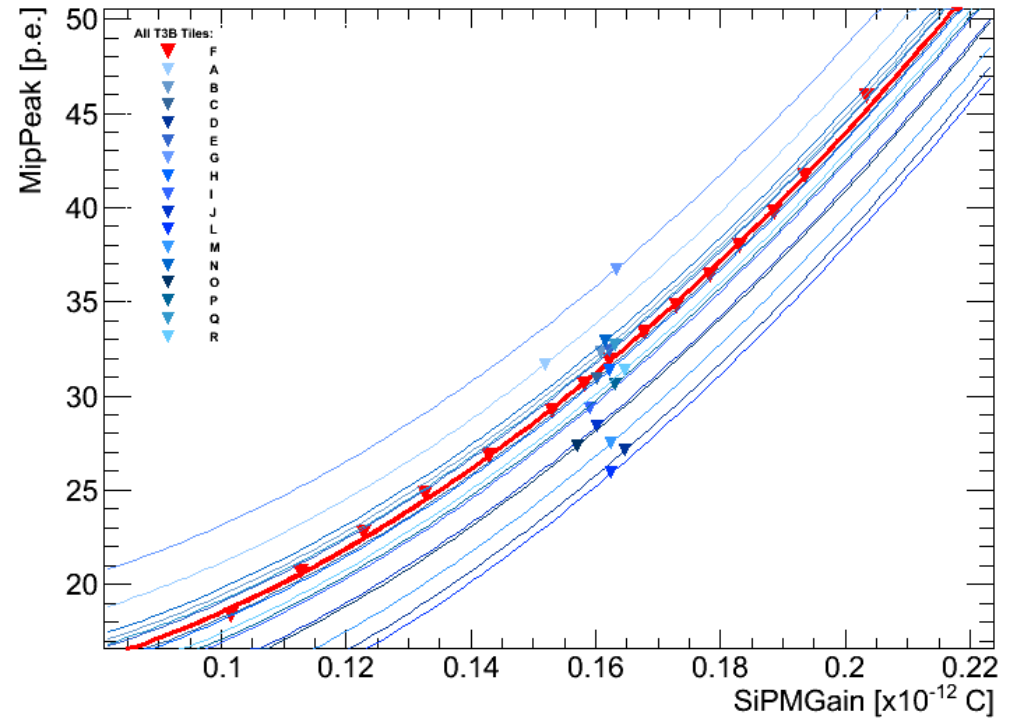


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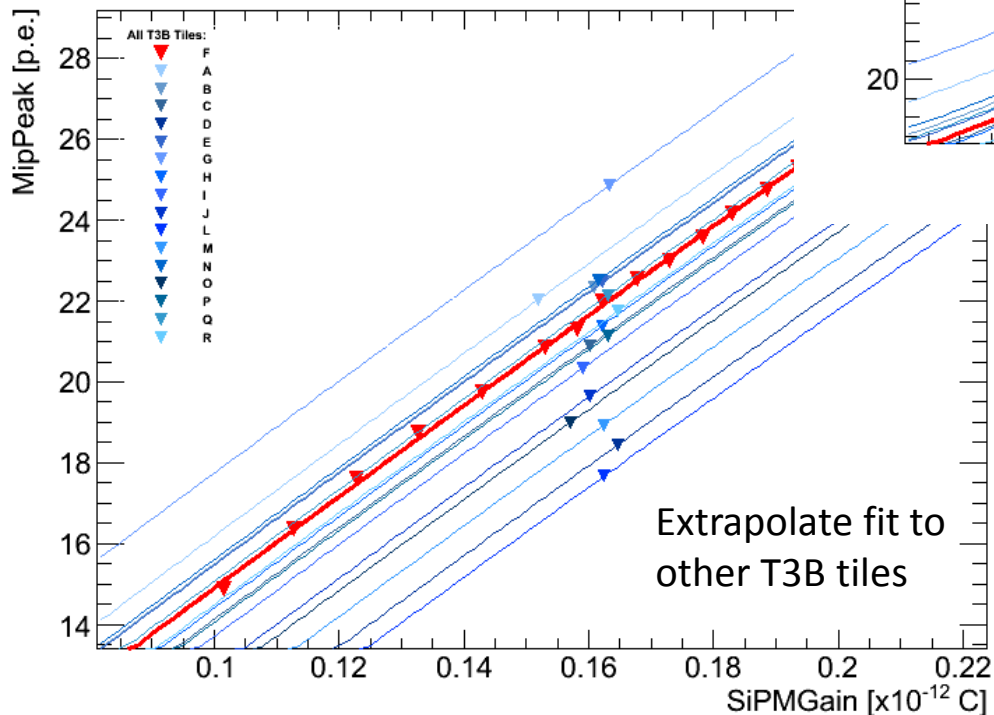
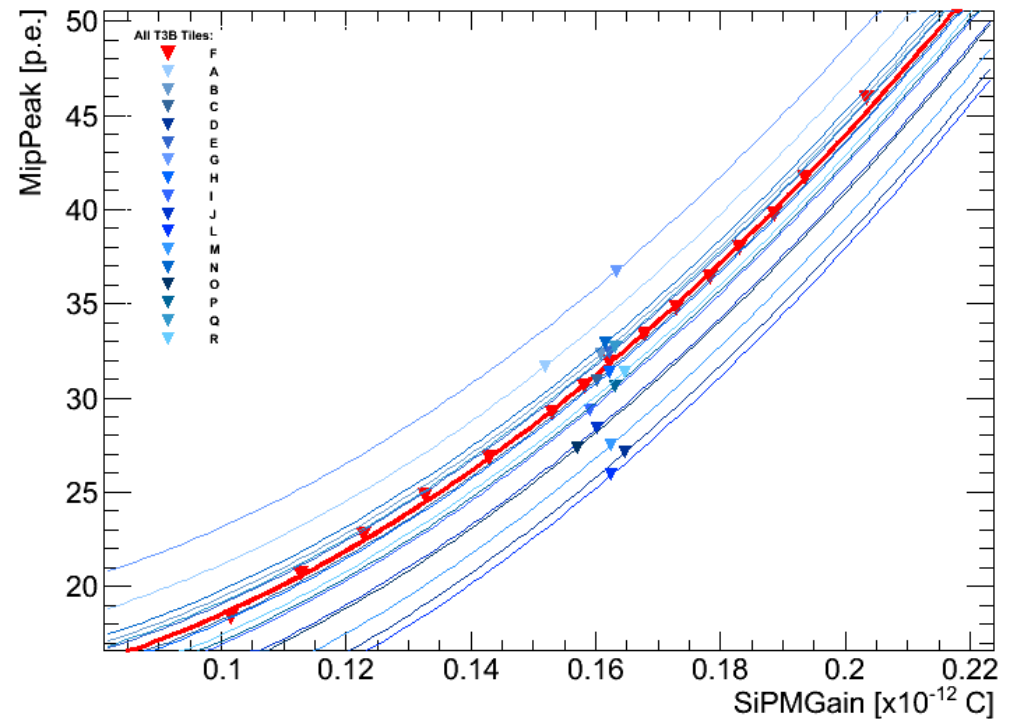


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One Measurement for
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9.6 ns	linear
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Interpretation:

1. More afterpulsing is integrated
→ would just result in a constant offset
2. Higher Gain
→ Afterpulsing and Crosstalk probability increased
→ Increased MPV dependence

Needs to be taken into account in Calib



The Effect of Afterpulsing



Bias Voltage Scan for one T3B
"Master Tile"

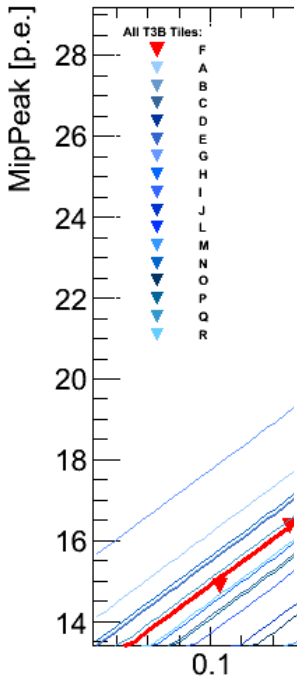
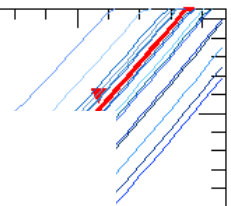
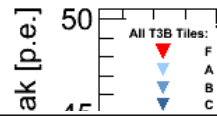
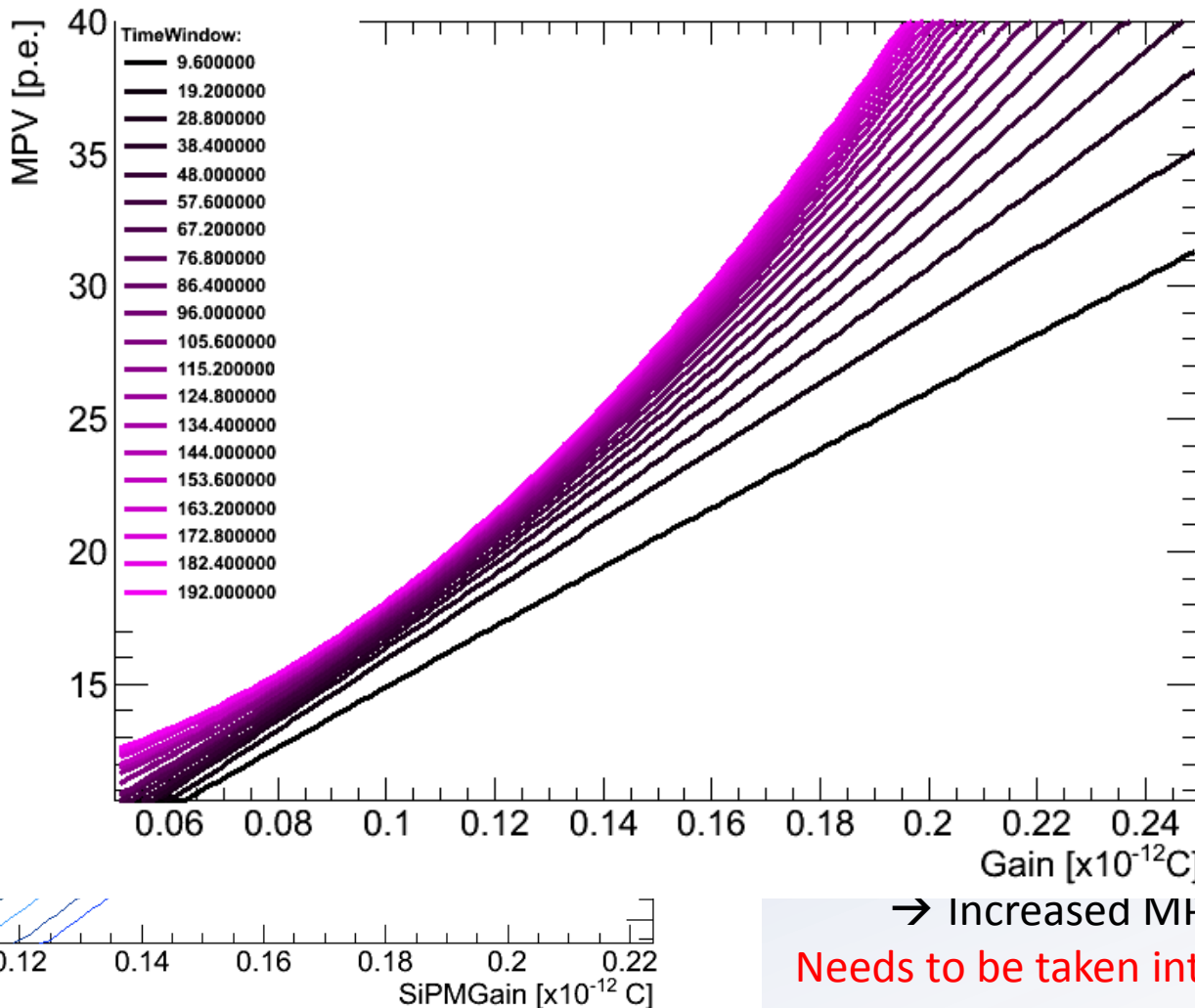
One Master Tile
all of

Time Window

9.6 ns

307.2 ns

Sr90 Calibration Data - Langau MPV vs. Gain



$\times 10^{-12}$ C]

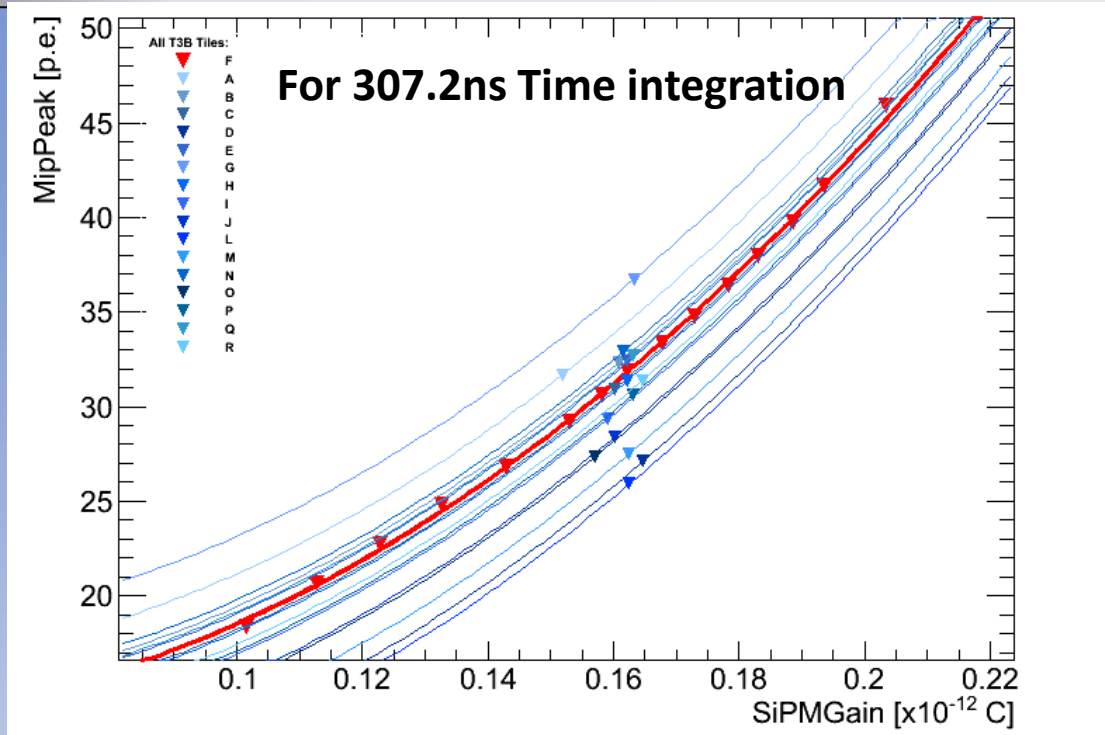
grated
stant offset

stalk
d

→ Increased MPV dependence
Needs to be taken into account in Calib



T3B Calibration to the MIP Scale



Obtain a dictionary:

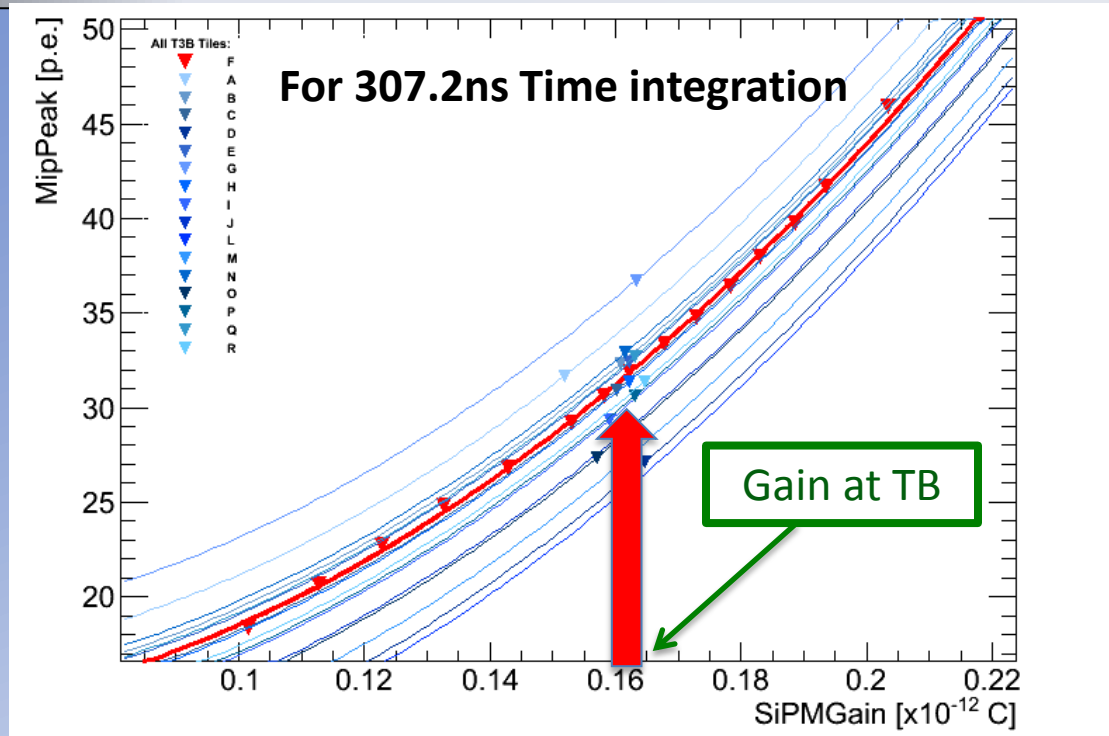
Determine live SiPM Gain from testbeam data

Select MPV-Gain dependence for distinct time integration window

Obtain corresponding MPV of MIP distrib.



T3B Calibration to the MIP Scale



Obtain a dictionary:

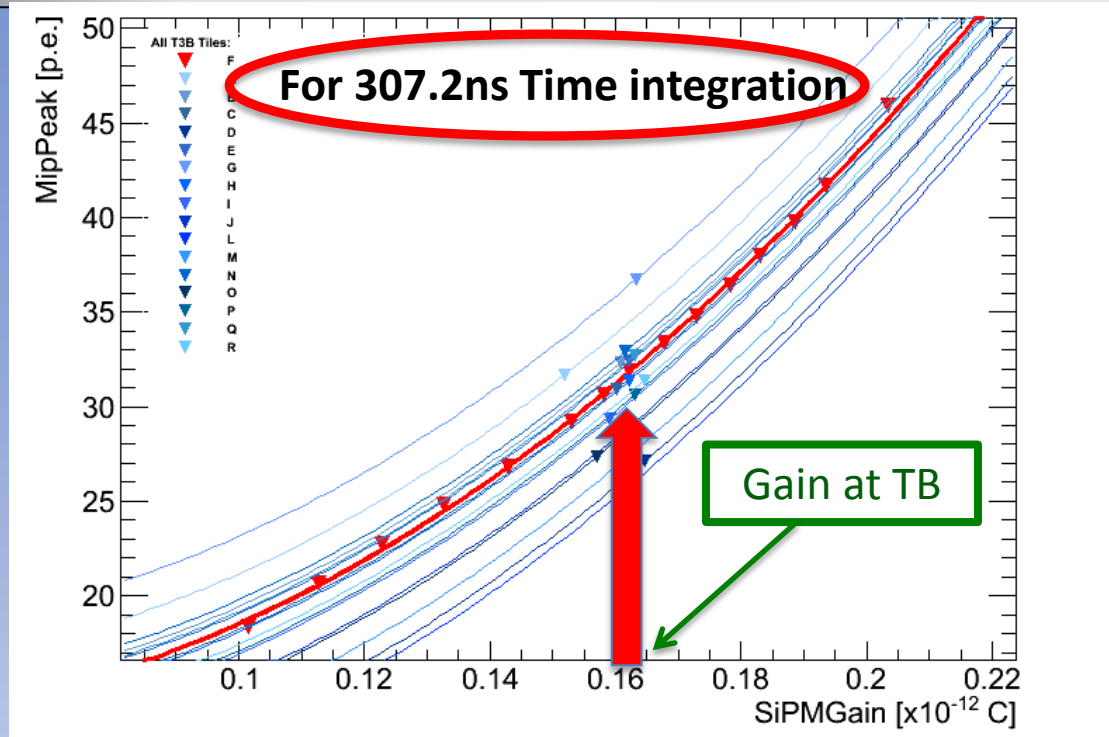
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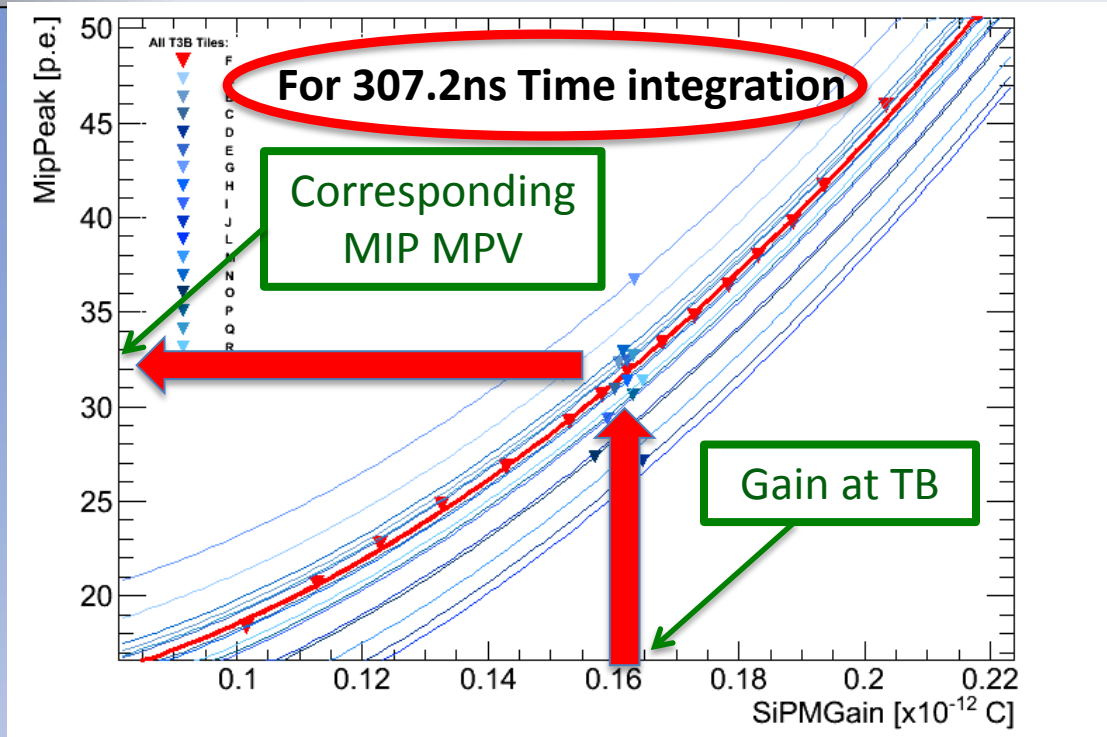


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T3B Calibration to the MIP Scale



Obtain a dictionary:

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VERIFY CALIBRATION PRINCIPLE: TESTBEAM MUON DATA



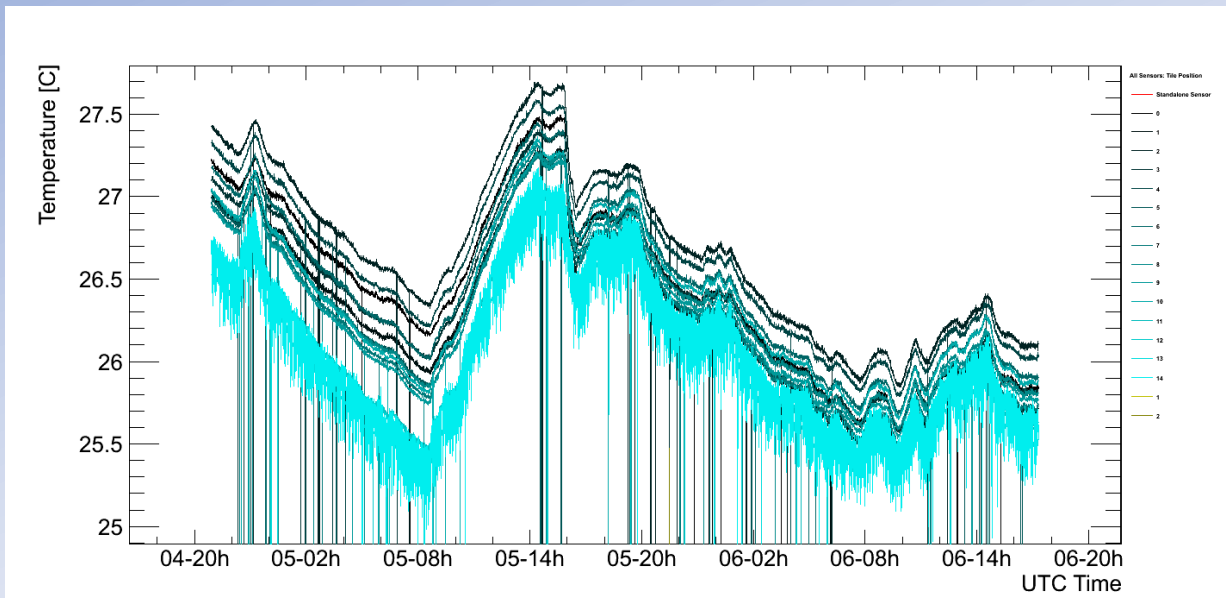
Verification of the Calibration

Principle: Muon Data



During the commissioning of the SDHCAL we could take an excessive amount of muon data:

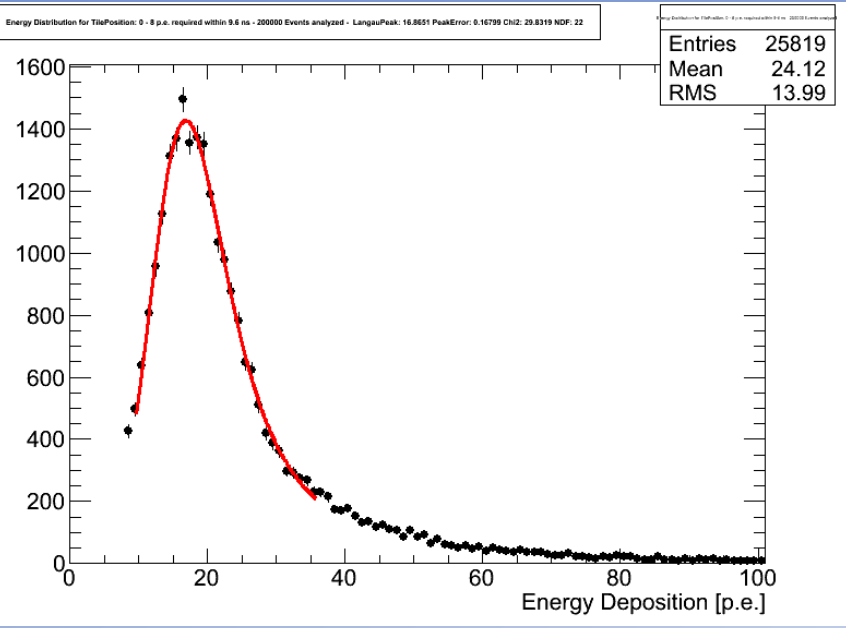
- 14 mio Muon Events
- 40 hours without interruption
- Day-night-cycle Temperature Range: $\sim 25.5\text{C}$ to 27.5C
- Enough to extract the Mip MPV-Temperature dependence
- Then: Apply correction factor from Sr90 Data to eliminate the dependence (remember: We assume the MPV depends in first order only on the SiPM gain)





Verification of the Calibration

Principle: Muon Data

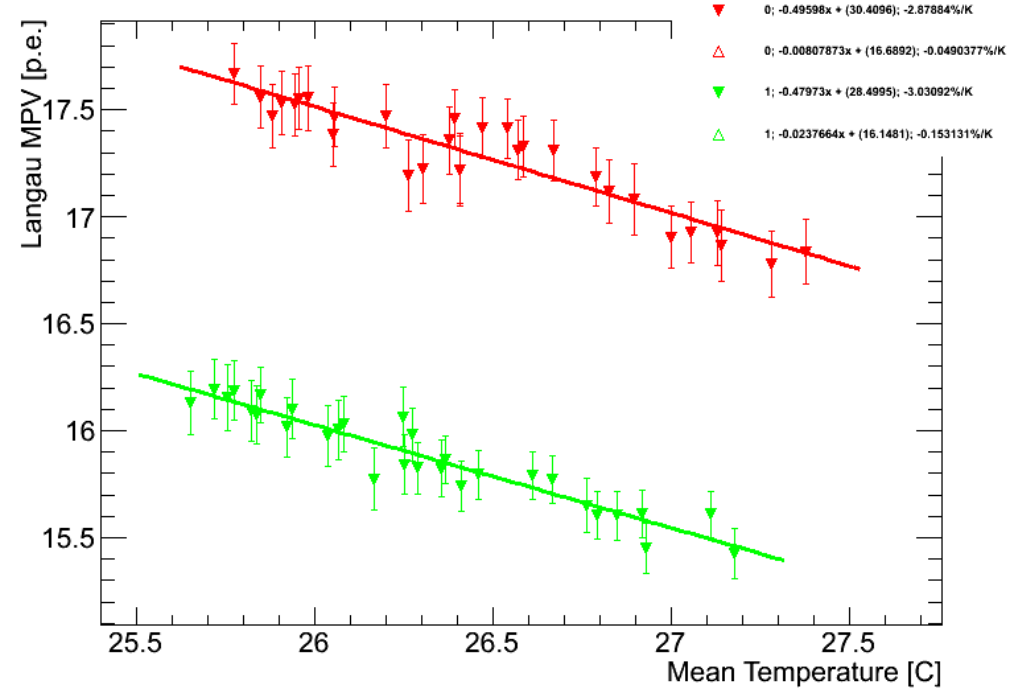
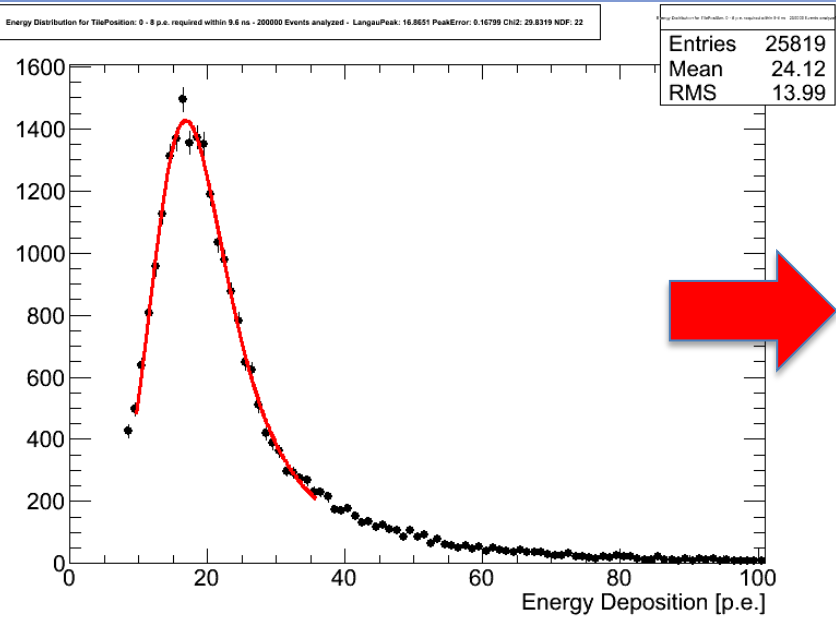


- T3B tiles hit in a small fraction of triggers
→ Determine MIP MPV every 200k events
- Time window of 9.6ns selected



Verification of the Calibration

Principle: Muon Data

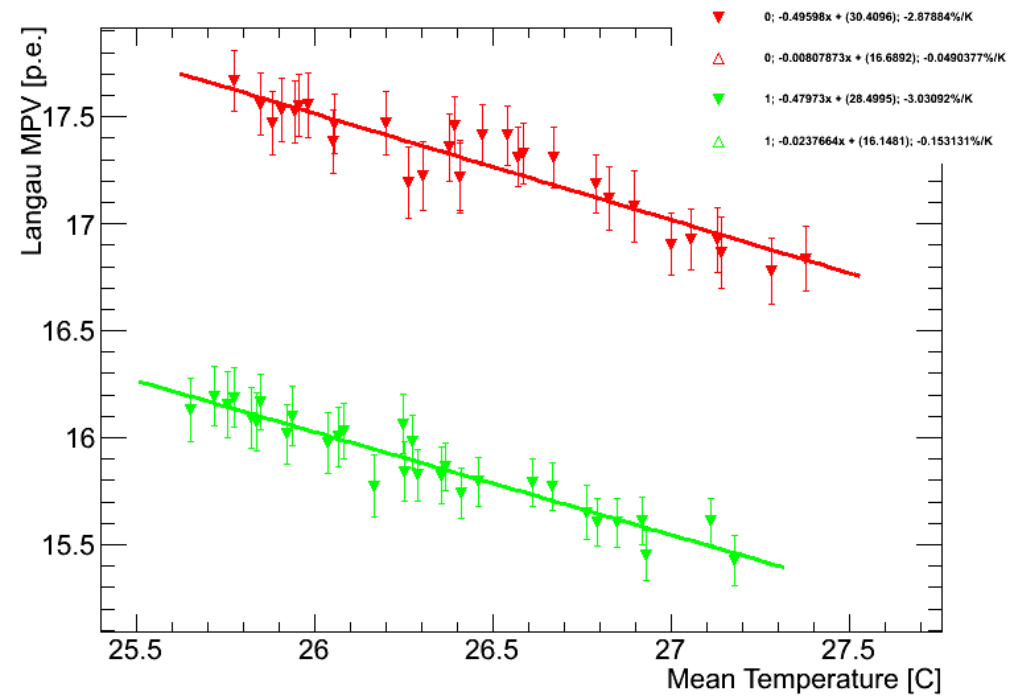
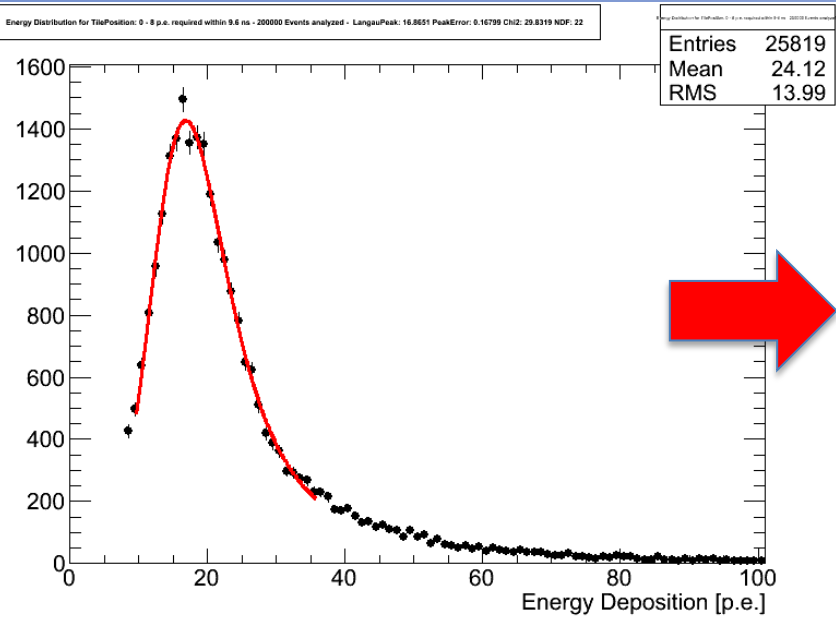


- Time window of 9.6ns



Verification of the Calibration

Principle: Muon Data



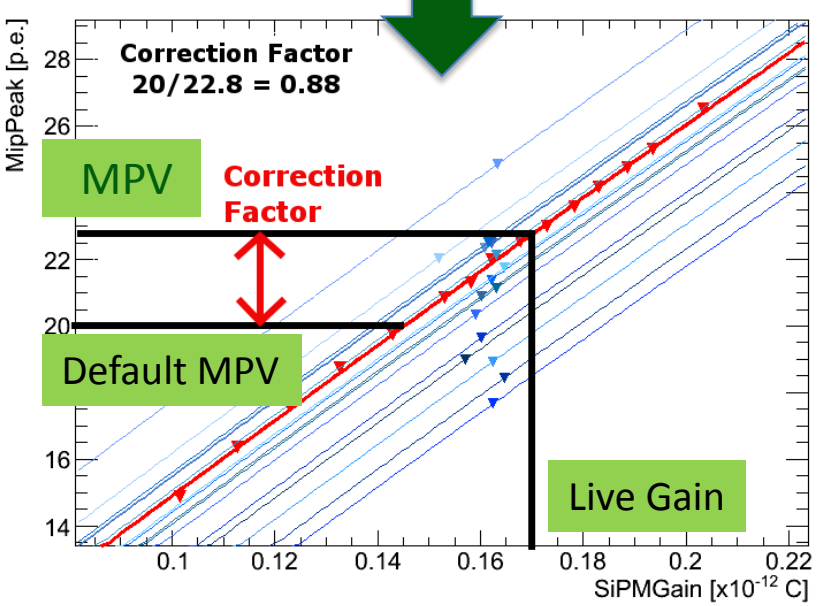
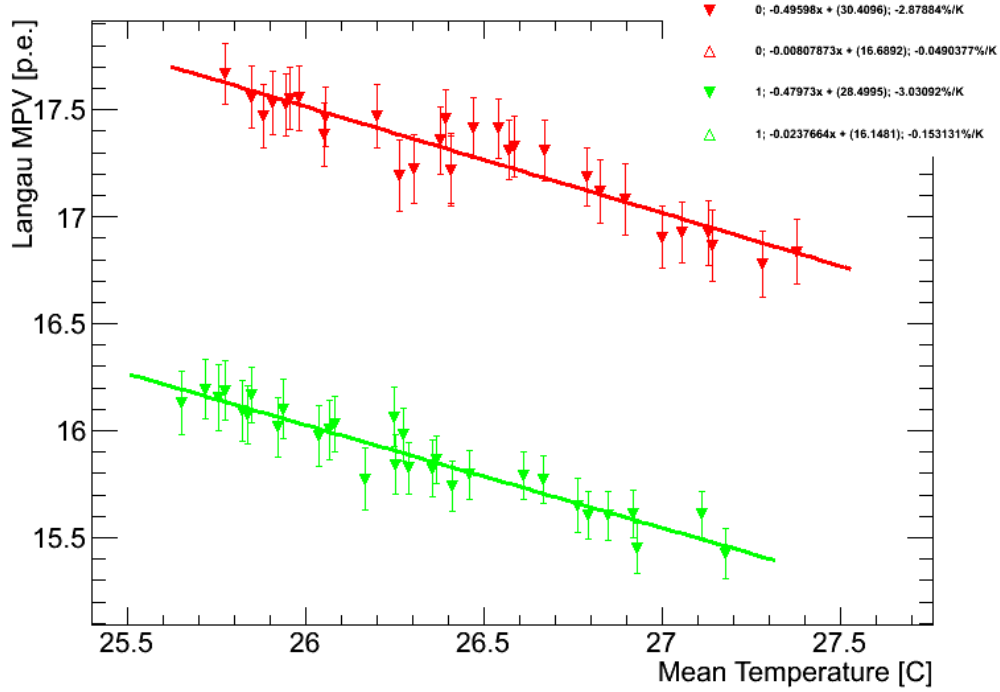
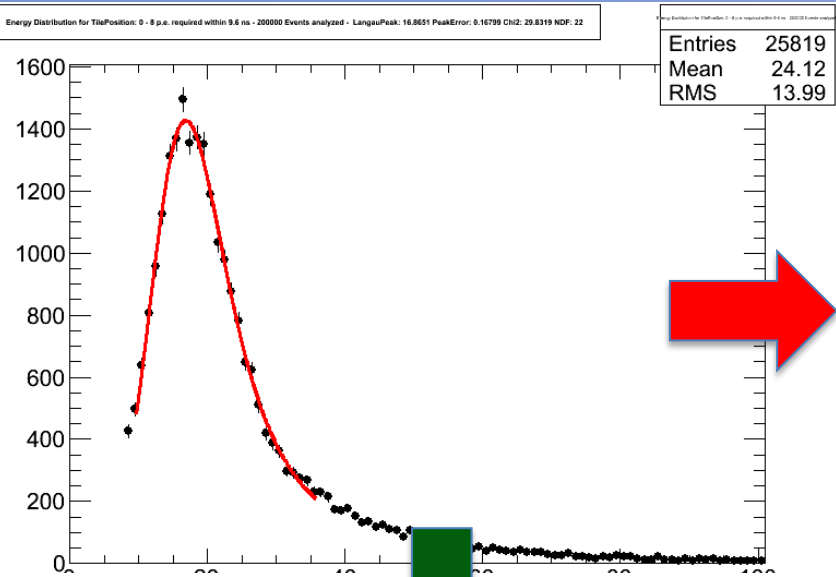
- Time window of 9.6ns

T3B Tile	MPV Drop	Slope
Center	-2.9 %/K	-0.5 p.e./K
Center + 1	-3.0 %/K	-0.48 p.e./K



Verification of the Calibration

Principle: Muon Data



- Time window of 9.6ns

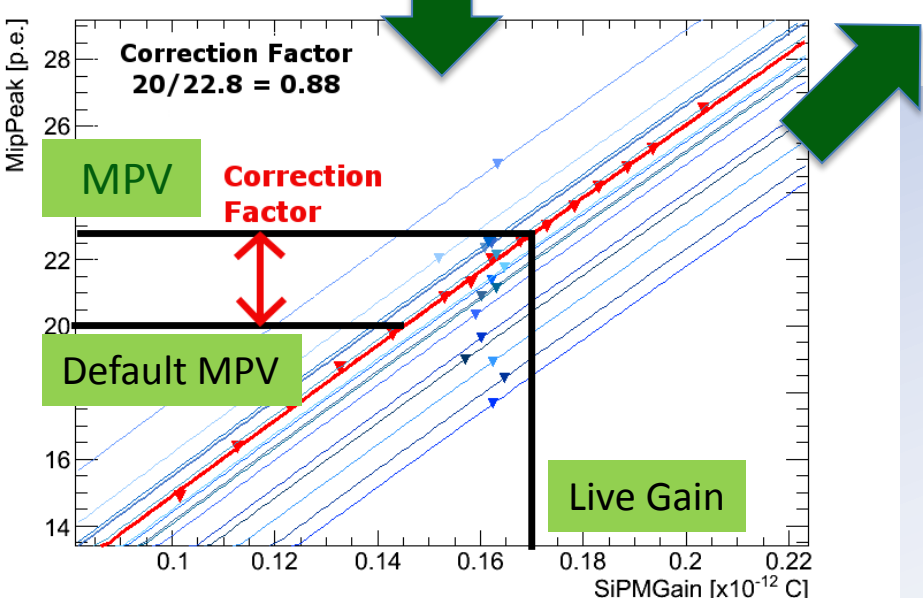
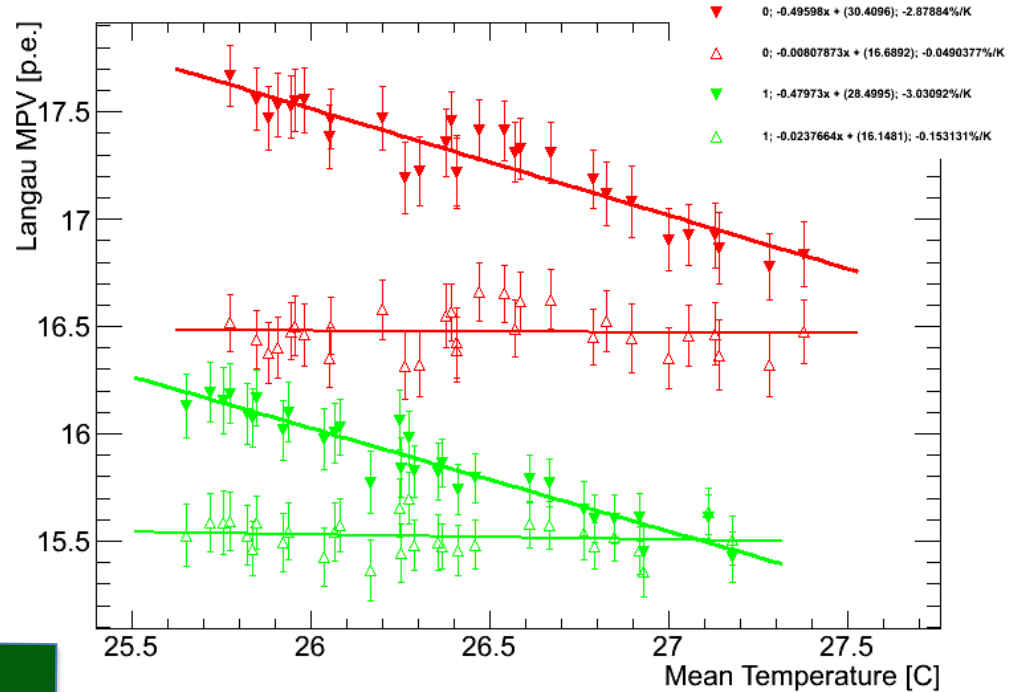
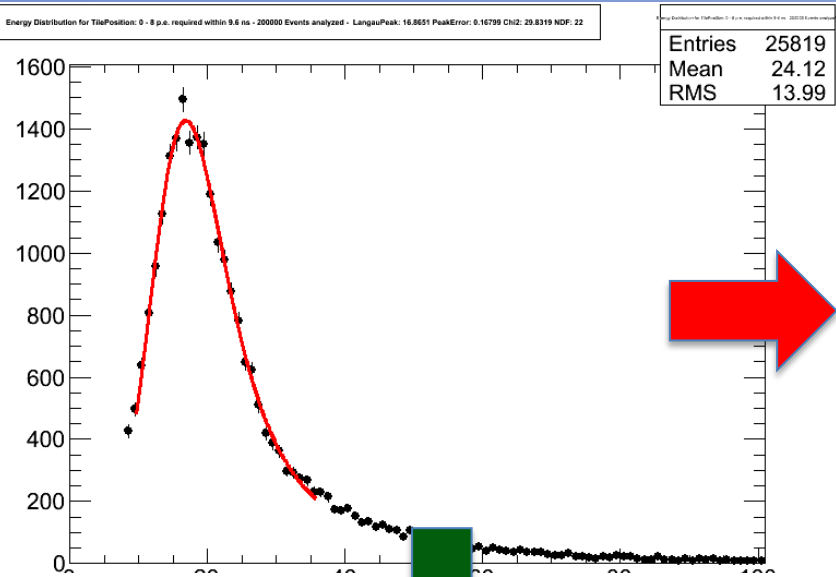
T3B Tile	MPV Drop	Slope
Center	-2.9 %/K	-0.5 p.e./K
Center + 1	-3.0 %/K	-0.48 p.e./K

- Get live gain from Intermediate RM
- Determine corresponding Sr90 MPV
- Choose default MPV of 20 p.e. → later 1MIP
- Obtain Correction factor



Verification of the Calibration

Principle: Muon Data



- Time window of 9.6ns

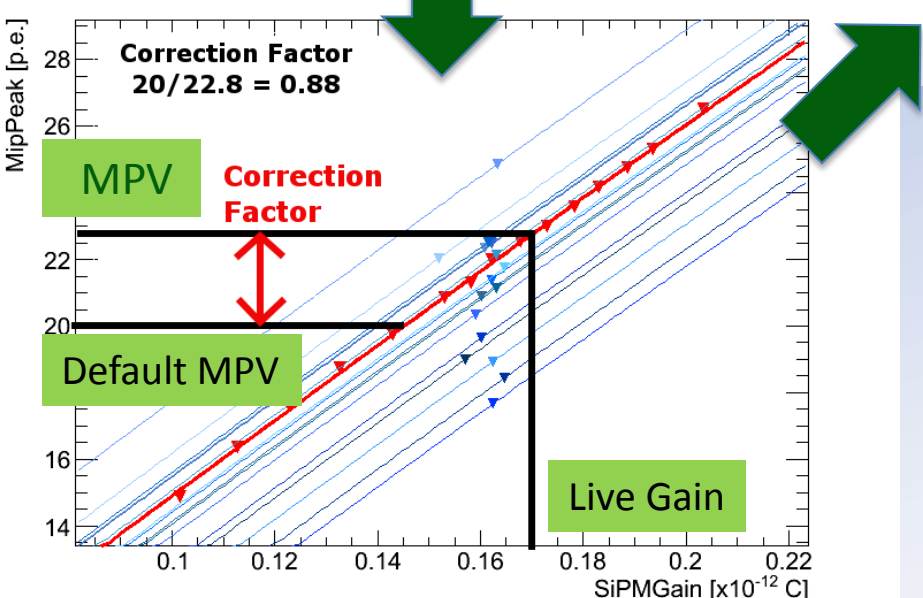
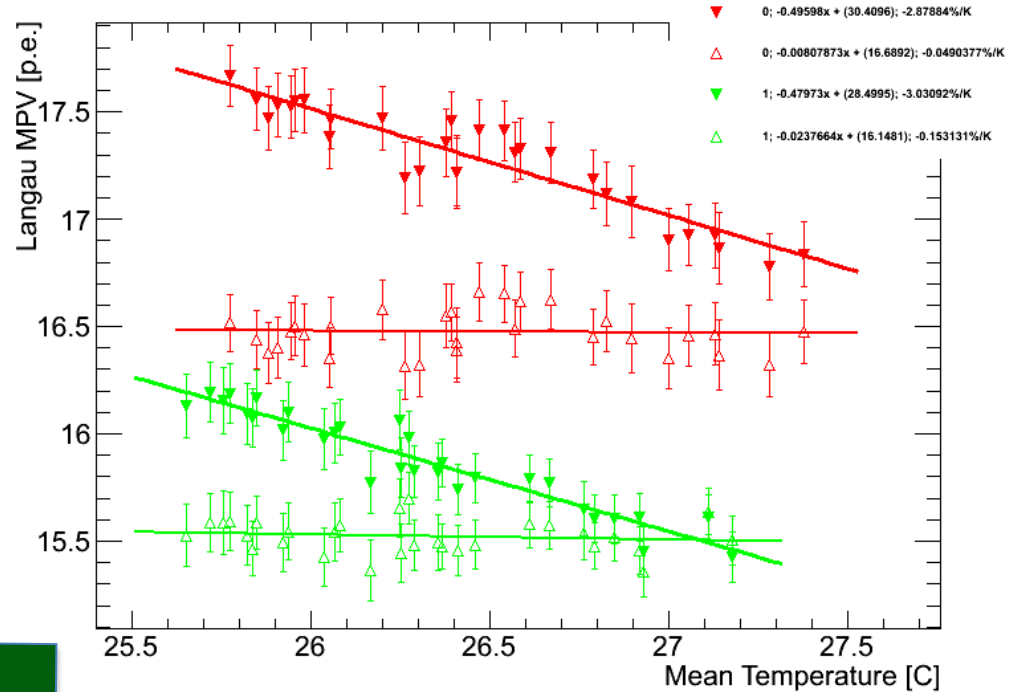
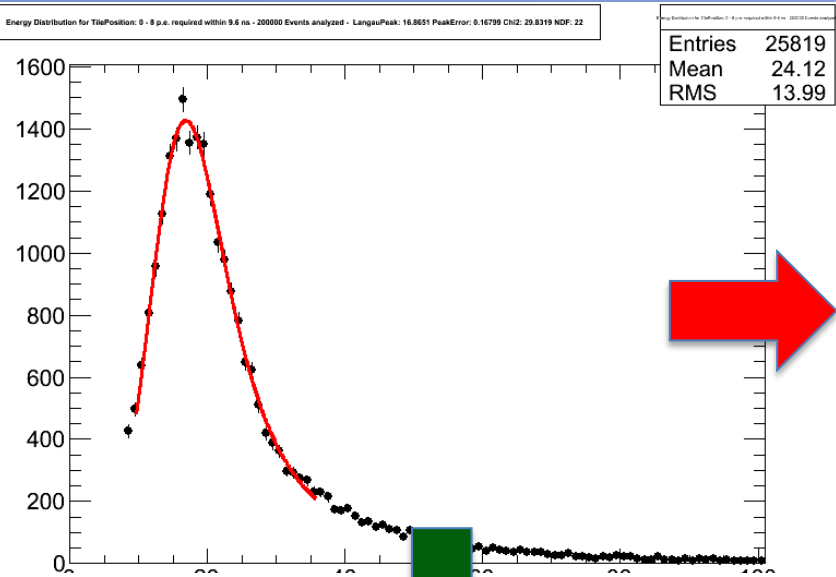
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Verification of the Calibration

Principle: Muon Data



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T3B Tile	MPV Drop	Slope
Center	-2.9 %/K	-0.5 p.e./K
Center + 1	-3.0 %/K	-0.48 p.e./K
Center corrected	-0.05 %/K	-0.008 p.e./K
Center + 1 corrected	-0.15 %/K	-0.024 p.e./K

- Obtain Correction factor



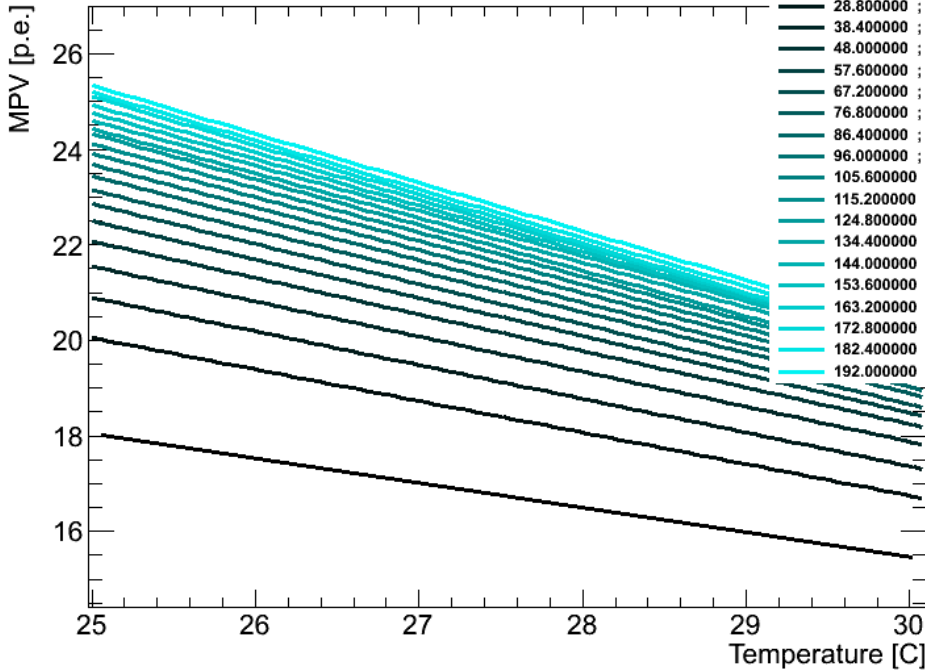
Verification of the Calibration

Principle: Muon Data



TBMuonData - Langau MPV vs. Temperature

TimeWindow: [ns]	Slope: [p.e./C]
9.600000	-0.515722
19.200000	-0.661304
28.800000	-0.706164
38.400000	-0.734528
48.000000	-0.765289
57.600000	-0.805177
67.200000	-0.839533
76.800000	-0.852818
86.400000	-0.884584
96.000000	-0.896156
105.600000	-0.915146
115.200000	-0.926719
124.800000	-0.956111
134.400000	-0.935413
144.000000	-0.946444
153.600000	-0.968623
163.200000	-0.991851
172.800000	-1.028012
182.400000	-1.021254
192.000000	-1.019814



Extracted MPV-Temperature dependence

Time integration window: 9.6 ns – 192 ns

→ Lower Temperature equivalent to higher gain

→ As before: Results in higher Afterpulsing and Crosstalk Probability

Linearity due to low T-Range (2C)!?



Verification of the Calibration

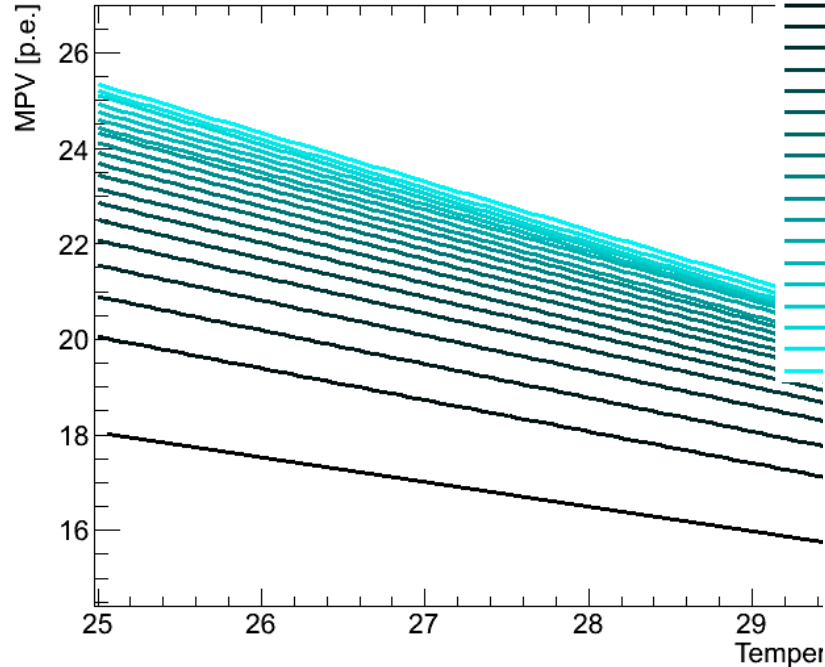
Principle: Muon Data



TBMuonData - Langau MPV vs. Temperature

TimeWindow: ; Slope:

9.600000	-0.515722
19.200000	-0.661304
28.800000	-0.706164
38.400000	-0.734528
48.000000	-0.765289
57.600000	-0.805177
67.200000	-0.839533
76.800000	-0.852818
86.400000	-0.884584
96.000000	-0.896156
105.600000	-0.915146
115.200000	-0.926719
124.800000	-0.956111
134.400000	-0.935413
144.000000	-0.946444
153.600000	-0.968623
163.200000	-0.991851
172.800000	-1.028012
182.400000	-1.021254
192.000000	-1.040814

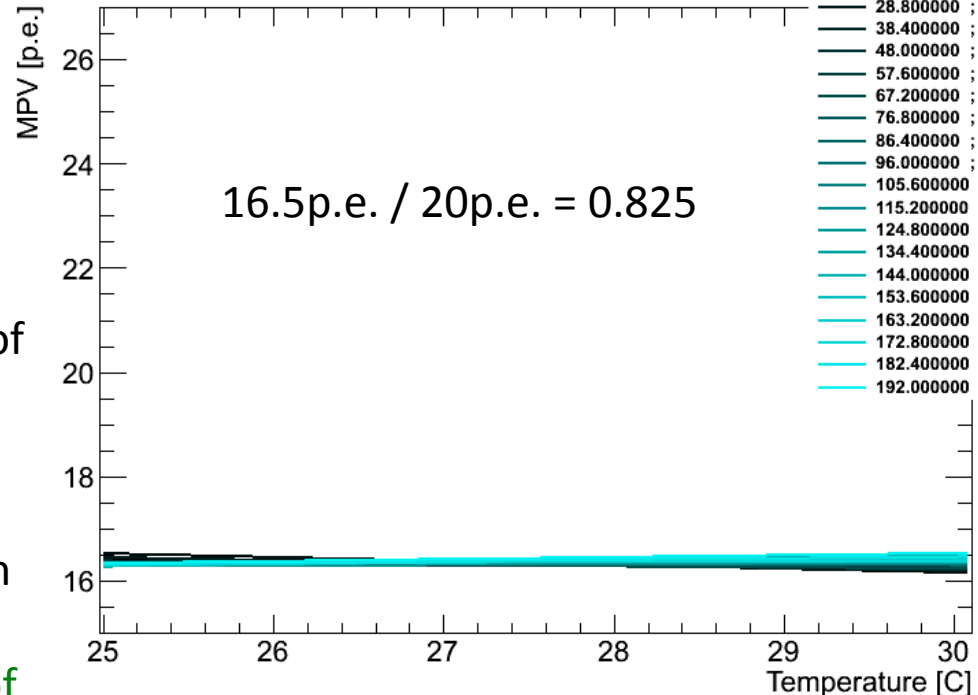


Extracted MPV-Temperature dependence
 Time integration window: 9.6 ns – 192 ns
 → Lower Temperature equivalent to higher gain
 → As before: Results in higher Afterpulsing and Crosstalk Probability
Linearity due to low T-Range (2C)!

TBMuonData - Sr90Corr - MPV vs. Temperature

TimeWindow: ; Slope:

9.600000	-0.025680
19.200000	-0.073377
28.800000	-0.051413
38.400000	-0.034544
48.000000	-0.022324
57.600000	-0.021153
67.200000	-0.020850
76.800000	0.003104
86.400000	-0.003405
96.000000	0.005223
105.600000	0.006223
115.200000	0.013547
124.800000	0.006016
134.400000	0.031616
144.000000	0.044222
153.600000	0.029342
163.200000	0.024105
172.800000	0.005804
182.400000	0.029401
192.000000	0.038111



Corrected MPV-Temperature dependence

Calibration results in efficient elimination of the dependence

Note: Corrected MPV values at ~16.5 p.e., not at the 20 p.e. we corrected to.

Interpretation: 0.825 is the Sr90 ↔ Muon MPV conversion factor

Matches simulations → Experimental proof



Verification of the Calibration

Principle: Muon Data



TBMuonData - Langau MPV vs. Temperature

TimeWindow	Slope
9.600000	-0.515722
19.200000	-0.661304
28.800000	-0.706164
38.400000	-0.734528
48.000000	-0.765289
57.600000	-0.805177

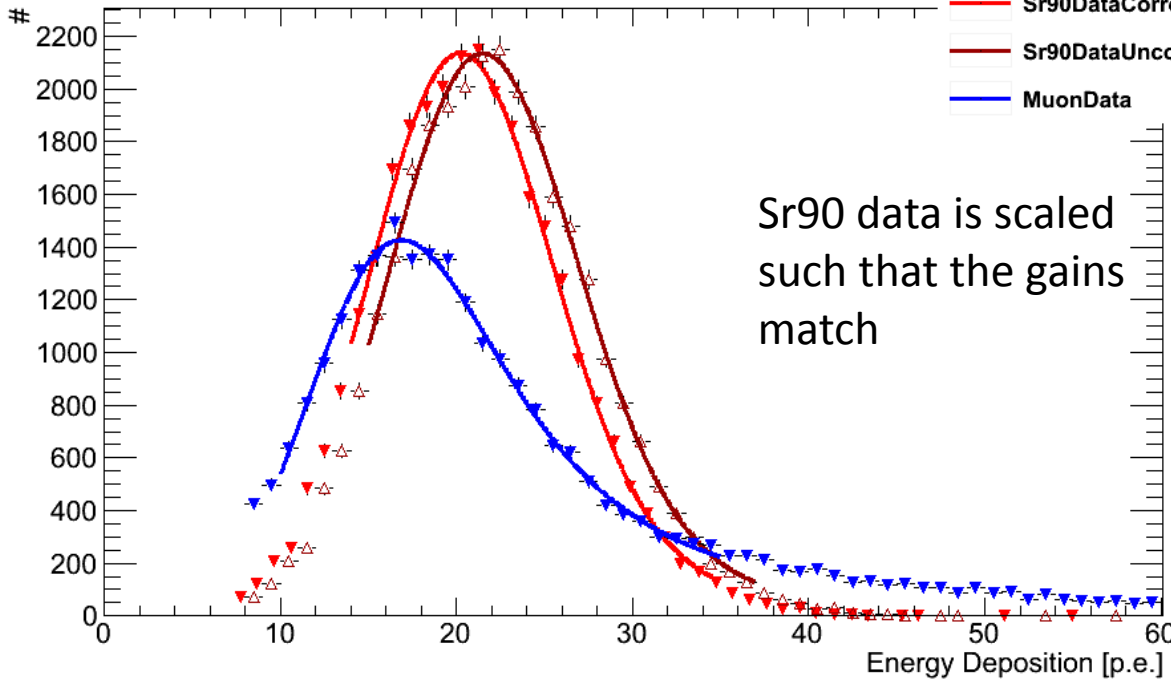
Extracted MPV-Temperature dependence
Time integration window: 9.6 ns – 192 ns

ivalent to
her
lk Probability
e (2C)!

Compare Energy Distribution: Sr90 vs. Muon Data - 8peIn12TimeBins

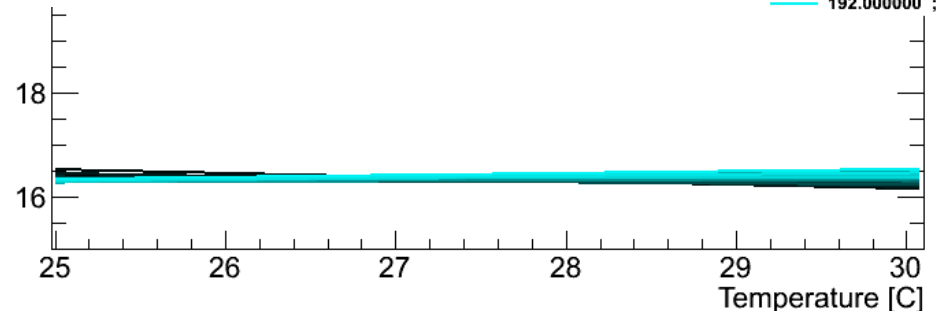
Data Type:

- Sr90DataCorrected
- Sr90DataUncorrected
- MuonData



Sr90 data is scaled such that the gains match

TimeWindow	Slope
9.600000	-0.025680
19.200000	-0.073377
28.800000	-0.051413
38.400000	-0.034544
48.000000	-0.022324
57.600000	-0.021153
67.200000	-0.020850
76.800000	0.003104
86.400000	-0.003405
96.000000	0.005223
105.600000	0.006223
115.200000	0.013547
124.800000	0.006016
134.400000	0.031616
144.000000	0.044222
153.600000	0.029342
163.200000	0.024105
172.800000	0.005804
182.400000	0.029401
192.000000	0.038111



Corrected MPV

Calibration re:
the dependence

Note: Corrected MPV values at ~16.5 p.e., not at the 20 p.e. we corrected to.

Interpretation: 0.825 is the Sr90 ↔ Muon MPV conversion factor

Matches simulations → Experimental proof



ROADMAP



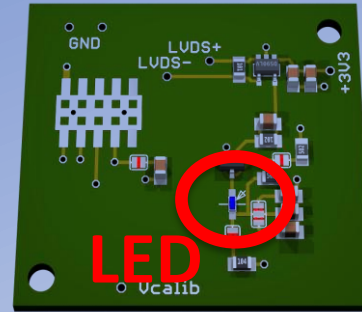
Roadmap:

Missing Calibration Steps



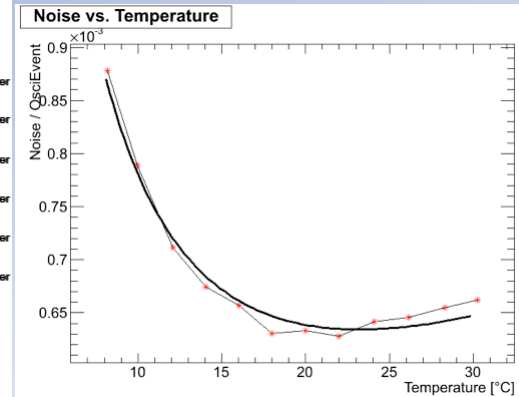
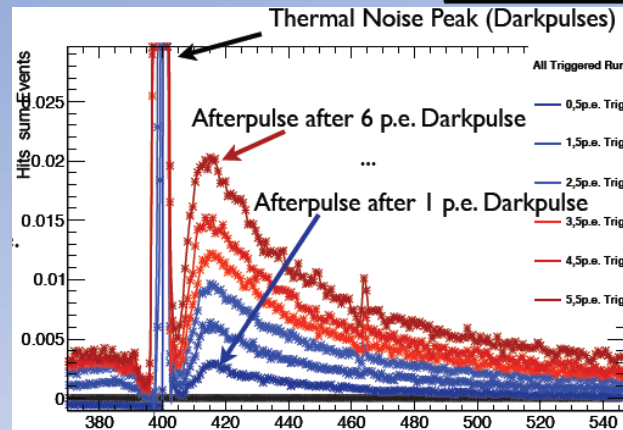
- **SiPM Saturation correction:**

Very promising results from Marco with Wuppertal LED board and T3B tiles



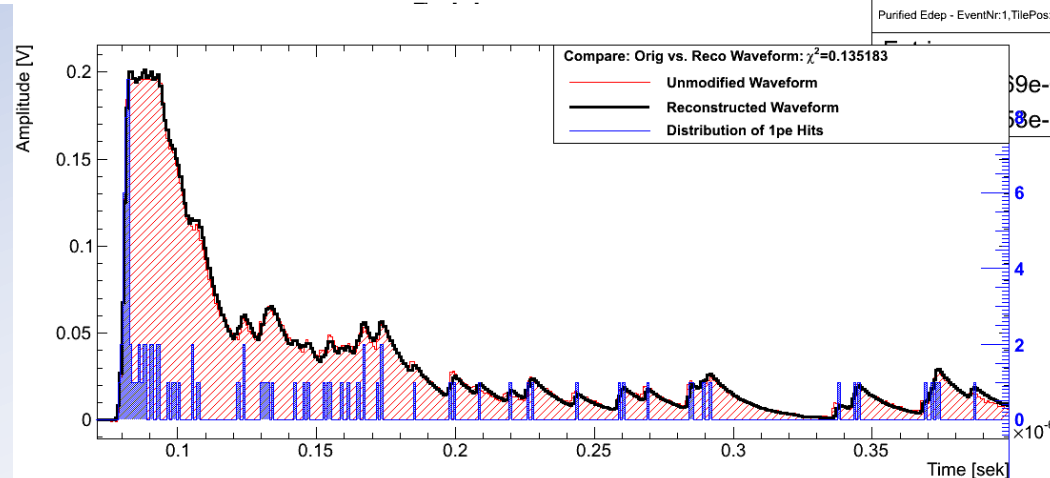
- **Correction for Afterpulsing:**

→ Need a dictionary: Which pulse height causes on average which afterpulsing contribution at a certain time after the initial pulse?
→ Promising results by Simon (also correction for darkrate)



- **Clipping Correction:**

Waveform decomposition can only work up to +/-200mV range with an 8bit ADC
→ Higher energy depositions clipped
→ Original waveform probably recoverable from the signal shape





Roadmap: Run Quality Checks

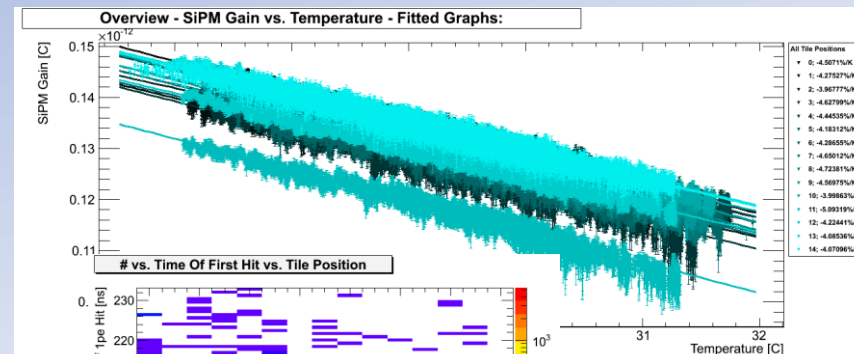
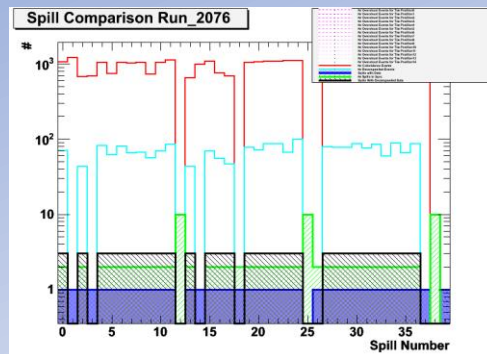
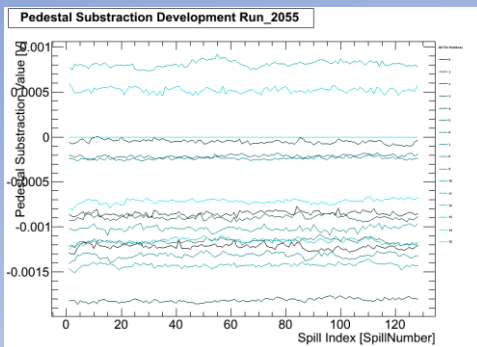


T3B is a very high statistics experiment → need to concatenate all Runs at one energy

Processing power is no issue: Analyze ~ 15min/million events on a standard CPU

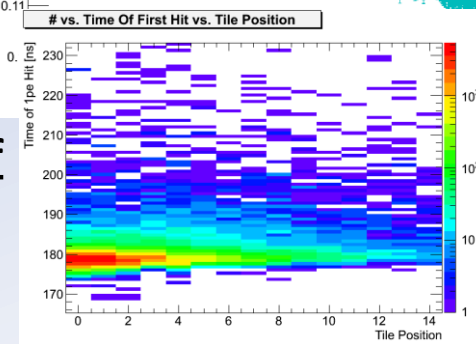
Developing procedure to identify suboptimal run conditions:

- CALICE Runlog → by eye ☹️
- Use Particle ID (from Cerenkovs), Beam profile
→ needs T3B-Calice synchronization for most of the data → Lars ongoing...
- T3B Hardware (e.g. pedestal jumps...) → automated “Calibration Quality Check” exists



Final step → obtain timing results that are bullet proof

- Energy deposition vs. time
- Shower timing vs. particle energy
- Longitudinal timing of hadron showers
- ...



There is still big potential in the T3B data → we look forward to a successful year 2012