

## UV complete theory of SUSY radiative seesaw scenarios



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S. Kanemura, T.S, and T. Yamada, PRD86,055023
S. Kanemura, E. Senaha, T.S,T. Yamada,JHEP1305,066
S. Kanemura, N. Machida, T.S, T. Yamada, arXiv:1309.3207 14/11/2013 LCWS2013 @ Univ. of Tokyo Japan

## Physics beyond the SM

Discovery of a Higgs boson\&measurements of properties

Essence of the electroweak symmetry breaking
New Physics at TeV scale
It's quite interesting,
if the NP provides solutions on the problems in the SM:

- Baryon asymmetry of the Universe
- Origin of the neutrino mass
- DM candidate


## Electroweak Baryogenesis

Electroweak Baryogenesis $\longleftrightarrow$ essence of EWSB


1st order electroweak transition

$$
V_{\mathrm{eff}}(v ; T)-V_{\mathrm{eff}}(0, T)
$$

 transition is necessary

## To get strong 1st order EWT

Strong 1st order EWPT requires extension of the SM In the SM, the condition is satisfied only when $m_{h}<50 \mathrm{GeV}$ ( $\varphi_{c} / T_{c}$ is suppressed by $m_{h}$ ) conflict with LHC data

Extra boson loop can enhance $\varphi_{c} / T_{c}$
Extended Higgs sector!
e.g. 2HDM

$$
\mathcal{L}=\frac{\lambda_{i}}{2} h^{2}\left|\Phi_{i}\right|^{2}
$$

$$
m_{\Phi}^{2}(\varphi)=M^{2}+\lambda_{i} \varphi^{2}
$$

Extra Higgs bosons as $\mathrm{H}, \mathrm{A} \mathrm{H}^{ \pm}$

Contour plot of $\Delta \lambda_{\text {hhh }} / \lambda_{\text {hhh }}$ and $\varphi_{C} T_{c}$ in the $m_{\Phi}-\mathrm{M}$ plane


## In SUSY case

## In the MSSM, there is no such a large coupling with SM-like Higgs

(The light stop scenario is the only possibility but it's almost dead)
The simplest example of strong but light Higgs scenario is SUSY 4HD+charged singlets $\phi_{\mathrm{c}} / T_{\mathrm{c}}>1$ with $\mathrm{m}_{\mathrm{h}}=126 \mathrm{GeV}$
S.Kanemura, E. Senaha, T.S, PLB706,40


## Fundamental theory?

- Electroweak baryogenesis
- Enhancement of EWPT by bosonic loop requires strong Higgs coupling( $>1$ ) but light( 125 GeV ) Higgs
- What is the fundamental theory of such models?
- Large coupling constant $\rightarrow$ Landau pole (cutoff)
- What is the origin of Higgs force?


## Fundamental theory?

- Electroweak baryogenesis
- Enhancement of EWPT by bosonic loop requires strong Higgs coupling(>1) but light(125GeV) Higgs
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- Large coupling constan Our expectation:
- What is the origin of Hig



## SUSY SU(2)н model

## In SUSY QCD: <br> $\mathrm{N}_{\mathrm{f}}=\mathrm{N}_{\mathrm{c}}+1 \Rightarrow$ confinement

See e.g. Intriligator, Seiberg, hep-th/9509006 Let us consider the simplest case $\left(\mathrm{N}_{\mathrm{c}}=2 \& \mathrm{~N}_{\mathrm{f}}=3\right)$
SUSY SU(2) $\mathrm{H} \times \mathrm{SU}(2)\llcorner\times U(1)$ Y s.Kanemura, T.S, and T. Yamada, PRD86,055023 It's asymptotic free! It's the same setup as the minimal SUSY fat Higgs R Harnik, et al., PRD70, 015002

| Fields | $\mathrm{SU}(2)_{L}$ | $\mathrm{U}(1)_{Y}$ |
| :---: | :---: | :---: |
| $\left(\begin{array}{l}T_{1} \\ T_{2}\end{array}\right.$ | 2 | 0 |
| $T_{3}$ | 1 | $+1 / 2$ |
| $T_{4}$ | 1 | $-1 / 2$ |
| $T_{5}$ | 1 | $+1 / 2$ |
| $T_{6}$ | 1 | $-1 / 2$ |

cf. In the minimal SUSY fat Higgs, only $H_{u}, H_{d}$, and $N$ are made light
(The effective theory is "minimal")

## Effective theory of $\operatorname{SU}(2)_{\text {н }}$ model

S.Kanemura,E. Senaha, T.S,T.Yamada,JHEP1305,066

$$
\begin{aligned}
& W=-\mu \overleftarrow{H_{u} H_{d}-\mu_{\Phi} \Phi_{u} \Phi_{d}-\mu_{\Omega}\left(\Omega_{+} \Omega_{-}-\zeta \eta\right)} \text { MSSM-like Higgs doublets } \\
&+\hat{\lambda}\left\{H_{d} \Phi_{u} \zeta+H_{u} \Phi_{d} \eta-H_{u} \Phi_{u} \Omega_{-}-H_{d} \Phi_{d} \Omega_{+}\right\} \\
& \hat{\lambda}\left(\Lambda_{H}\right) \simeq 4 \pi \text { (Naive dimensional analysis) }
\end{aligned}
$$



## 1st order EWPT

## Benchmark:

$$
\begin{aligned}
& \tan \beta=15, m_{H^{+}}=350 \mathrm{GeV}, \mu=200 \mathrm{GeV}, M_{\tilde{t}}=M_{\tilde{q}}=2000 \mathrm{GeV} \\
& \bar{m}_{\Omega^{+}}^{2}=\bar{m}_{\Phi_{d}}^{2}=\bar{m}_{\zeta}^{2}=(1500 \mathrm{GeV})^{2}, \bar{m}_{\eta}^{2}=(2000 \mathrm{GeV})^{2}, \mu_{\Phi}=\mu_{\Omega}=550 \mathrm{GeV} \\
& m_{0}^{2} \equiv \bar{m}_{\Phi_{u}}^{2}=\bar{m}_{\Omega_{-}}^{2}(\text { Scanned }) \\
& \left(m_{\phi}^{2}=\bar{m}_{\phi}^{2}+c_{\phi} \lambda^{2} v^{2}\right)
\end{aligned}
$$



$\varphi_{c} / T_{c}>1$ can be satisfied!!

## Lightest $Z_{2}$ odd masses

## 1st order EWPT



## Contribution to hyp

S.Kanemura,E. Senaha, T.S, T.Yamada,JHEP1305,066

$\sim 20 \%$ deviation is possible in the region of $v_{c} / T_{c}>1$

## hhh coupling


$\sim 20 \%$ deviation is possible in the region of $v_{c} / T_{c}>1$

## How about neutrino mass?

Origin of the neutrino mass at TeV scale Alternative to the well-known seesaw model: Idea of loop induced neutrino mass
Especially, radiative seesaw scenarios are interesting
Loop diagram with RH neutrinos give tiny neutrino mass

$$
\left(Z_{2}-o d d\right) \longleftarrow \text { To avoid tree level contribution }
$$

Some new scalars are introduced!



Lightest $\mathrm{Z}_{2}$-odd neutral particle can be a DM

## AKS model

:Āōk̄i-K̄ānēmürä-S̄ēto mōdē
Aoki, Kanemura, Seto, PRL102, 051805
( $2 \mathrm{HD}+\mathrm{Z}_{2}$-odd charged and neutral singlet+Z2-odd RHN)


As a phenomenological model, this is quite interesting But ...
Large couplings $\longrightarrow$ Landau pole at low energy scale Many extra scalars $\longrightarrow$ It seems artificial What is the fundamental theory of this model?

## For radiative seesaw

S.Kanemura, N. Machida, T.S, T.Yamada, arXiv:1309.3207

| Fields | $\mathrm{SU}(2)_{L}$ | $\mathrm{U}(1)_{Y}$ | $Z_{2}$ |
| :---: | :---: | :---: | :---: |
| $\binom{T_{1}}{T_{2}}$ | 2 | 0 | + |
| $T_{3}$ | 1 | $+1 / 2$ | + |
| $T_{4}$ | 1 | $-1 / 2$ | + |
| $T_{5}$ | 1 | $+1 / 2$ | - |
| $T_{6}$ | 1 | $-1 / 2$ | - |

We can use the $\mathrm{SU}(2)_{\mathrm{H}}$ model

| Field | $\mathrm{SU}(2)_{L}$ | $\mathrm{U}(1)_{Y}$ | $Z_{2}$ |
| :---: | :---: | :---: | :---: |
| $H_{u}=\binom{H_{13}}{H_{23}}$ | 2 | $+1 / 2$ | + |
| $H_{d}=\binom{H_{14}}{H_{24}}$ | 2 | $-1 / 2$ | + |
| $N=H_{56}, N_{\Phi}=H_{34}, N_{\Omega}=H_{12}$ | 1 | 0 | + |
| $\Phi_{u}=\binom{H_{15}}{H_{25}}$ | 2 | $+1 / 2$ | - |
| $\Phi_{d}=\binom{H_{16}}{H_{26}}$ | 2 | $-1 / 2$ | - |
| $\Omega_{+}=H_{35}$ | 1 | +1 | - |
| $\Omega_{-}=H_{46}$ | 1 | -1 | - |
| $\zeta=H_{36}, \xi=H_{45}$ | 1 | 0 | - |

In the low energy effective theory,

$$
W_{N}=\left(y_{N}\right)_{i} N_{i}^{c} L_{j} \Phi_{u}+\left(h_{N}\right)_{i j} N_{i}^{c} E_{j}^{c} \Omega^{-}+\frac{M_{i}}{2} N_{i}^{c} N_{i}^{c}
$$

## Neutrino mass generation

S.Kanemura, N. Machida, T.S, T.Yamada,arXiv:1309.3207

Two different types of contributions are possible
1-loop
driven by $\mathrm{y}_{\mathrm{N}}$


It corresponds to SUSY Ma model


They correspond to SUSY AKS model

## Comment on SUSY AKS

S.Kanemura, N. Machida, T.S, T.Yamada,arXiv:1309.3207
e.g. Āoki-Kanemura-Seto model Aoki, Kanemura, Seto, PRL102, 051805 ( $2 \mathrm{HD}+\mathrm{Z}_{2}$-odd charged and neutral singlet+Z2-odd RHN)


In SUSY version, $\mathrm{H}_{\mathrm{u}}, \mathrm{H}_{\mathrm{d}}$ (MSSM-like Higgs)

$$
\begin{array}{cl}
\Omega^{+}, \Omega^{-} & \Phi u_{u}, \phi_{\mathrm{d}} \\
\zeta & \mathrm{~N}^{c}(\mathrm{RHN})
\end{array}
$$ fields are required

provides all the fields in the Higgs sector!!

## Benchmark points

(A):1-loop dominant point (B):3-loop dominant point

| Case | $\lambda$ | $\tan \beta$ | $m_{H^{ \pm}}$ | $m_{\tilde{W}}$ | $\mu$ | $\mu_{\Phi}$ | $\mu_{\Omega}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (A) | 1.8 | 15 | 350 GeV | 500 GeV | 100 GeV | 550 GeV | -550 GeV |
| (B) | 1.8 | 30 | 350 GeV | 500 GeV | 100 GeV | 550 GeV | -550 GeV |


| Case | $\bar{m}_{\Phi_{u}}^{2}$ | $\bar{m}_{\Phi_{d}}^{2}$ | $\bar{m}_{\Omega^{+}}^{2}$ | $\bar{m}_{\Omega^{-}}^{2}$ | $\bar{m}_{\zeta}^{2}$ | $\bar{m}_{\eta}^{2}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (A) | $(100 \mathrm{GeV})^{2}$ | $(1500 \mathrm{GeV})^{2}$ | $(1500 \mathrm{GeV})^{2}$ | $(100 \mathrm{GeV})^{2}$ | $(1500 \mathrm{GeV})^{2}$ | $(2000 \mathrm{GeV})^{2}$ |
| (B) | $(1500 \mathrm{GeV})^{2}$ | $(1500 \mathrm{GeV})^{2}$ | $(1500 \mathrm{GeV})^{2}$ | $(30 \mathrm{GeV})^{2}$ | $(1410 \mathrm{GeV})^{2}$ | $(30 \mathrm{GeV})^{2}$ |


| Case | $B_{\zeta}^{2}$ | $B_{\eta}^{2}$ | $m_{\zeta \eta}^{2}$ |
| :---: | :---: | :---: | :---: |
| (A) | $(100 \mathrm{GeV})^{2}$ | $(100 \mathrm{GeV})^{2}$ | $(100 \mathrm{GeV})^{2}$ |
| (B) | $(1400 \mathrm{GeV})^{2}$ | 0 | 0 |


| Case | $M_{1}$ | $M_{2}$ | $M_{3}$ | $m_{\tilde{\nu}_{R 1}}$ | $m_{\tilde{\nu}_{R 2}}$ | $m_{\tilde{\nu}_{R 3}}$ | $m_{\tilde{e}_{R i}}(i=1,2,3)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $(\mathrm{A})$ | 60 GeV | 120 GeV | 180 GeV | 60 GeV | 120 GeV | 180 GeV | 6000 GeV |
| $(\mathrm{B})$ | 100 GeV | 2000 GeV | 4000 GeV | 100 GeV | 4000 GeV | 8000 GeV | 6000 GeV |


| Case | $\left(y_{N}\right)_{i j}$ | $\left(h_{N}\right)_{i j}$ |
| :---: | :---: | :---: |
| (A) | $\left(\begin{array}{ccc}-0.45 & -0.44 & 0.51 \\ 0.23 & 0.23 & -0.26 \\ 0.19 & 1.37 & 1.37\end{array}\right) \times 10^{-4}$ |  |
| (B) | $\sim\left(\begin{array}{lll}0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0\end{array}\right)$ | $\sim\left(\begin{array}{lll}0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0\end{array}\right)$ |
| $\left(\begin{array}{ccc}0.001 & 0 & 0 \\ -0.0624+0.16 i & -0.0314-0.0016 i & -0.0022+0.000297 i \\ 0.902+2.46 i & 0.000681-0.00126 i & -0.000755-0.00161 i\end{array}\right)$ |  |  |


| Case | $m_{1}$ | $m_{2}$ | $m_{3}$ | $\sin ^{2} \theta_{12}$ | $\sin ^{2} 2 \theta_{23}$ | $\left\|\sin \theta_{13}\right\|$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $(\mathrm{A})$ | 0.0 eV | 0.0090 eV | 0.050 eV | 0.31 | 1.0 | 0.1 |
| $(\mathrm{~B})$ | 0.0 eV | 0.0089 eV | 0.050 eV | 0.31 | 1.0 | 0.1 |

The neutrino mass and angles are reproduced

| Case | $\mathrm{B}(\mu \rightarrow e \gamma)$ | $\mathrm{B}(\mu \rightarrow$ eee $)$ |
| :---: | :--- | :--- |
| (A) | $4.6 \times 10^{-19}$ | $7.2 \times 10^{-21}$ |
| (B) | $5.2 \times 10^{-14}$ | $4.7 \times 10^{-13}$ |

## $\phi_{c} / T_{c}>1$ is realized!

Serious LFV constraints are also satisfied

## Comments on direct detection

Our model is characterized by the $Z_{2}$ odd sector

## $\underline{Z}_{2}$-odd particle search is important colorless

Case (A): light inert doublet

$$
\begin{aligned}
& e^{+} e^{-} \rightarrow H^{\prime} A^{\prime} \rightarrow Z H^{\prime} H^{\prime} \\
& e^{+} e^{-} \rightarrow H^{+\prime} H^{-1} \rightarrow W^{+} W^{-} H^{\prime} H^{\prime} @ \text { ILC }
\end{aligned}
$$

Mass determination can be done with a few GeV accuracy
M. Aoki, S. Kanemura and H. Yokoya, PLB725,302.

Case (B): Singlet-like charged particle $\Omega^{+}$
$e^{+} e^{-} \rightarrow \Omega_{1}^{+} \Omega_{1}^{-}$
$e^{-} e^{-} \rightarrow \Omega_{1}^{-} \Omega_{1}^{-} \longleftarrow$ Strong evidence of the model Aoki\&Kanemura\&Seto, PRD80,033007; Aoki\&Kanemura, PLB689,28.

## Light inert doublet @ ILC

M. Aoki, S. Kanemura and H. Yokoya, PLB725,302.


## Light inert doublet @ ILC

M. Aoki, S. Kanemura and H. Yokoya, PLB725,302.

$$
e^{+} e^{-} \rightarrow H^{+\prime} H^{-\prime} \rightarrow W^{+} W^{-} H^{\prime} H^{\prime}
$$




The masses can be precisely determined

## Singlet-like scalar @ ILC

Aoki\&Kanemura, PLB689,28.

$$
e^{+} e^{-} \rightarrow S^{+} S^{-} \rightarrow \tau^{+} \tau^{-}+\text {missing }
$$



A signal can be seen at the ILC@1TeV

## Singlet-like scalar @ILC-e-e-



The signal is quite clear evidence of the Majorana nature and the scenario

## Summary

- It is quite interesting, NP in the Higgs sector provides solutions for baryogenesis, neutrino mass, DM.
- Electroweak baryogenesis, radiative generation of neutrino mass,...
- It can be tested at collider experiments
- Many models have been considered but they have been developed purely phenomenologically
- We have succeeded to provide a candidate of fundamental theory of such models
- SUSY SU(2) ${ }_{\text {H }}$ with $N_{f}=3+Z_{\downarrow}$-odd RHN is attractive simple candidate

It provides new DM candidate

- It's very different from GUT beyond the grand desert Rich field will be there!


## Back up

## Top Yukawa coupling

Introducing several new fields ( $\mathrm{SU}(2)_{\mathrm{H}}$ singlets) as
conformal

$$
\begin{aligned}
W_{f}= & M_{f}\left(\varphi_{u} \bar{\varphi}_{u}+\bar{\varphi}_{d} \varphi_{d}\right)+\bar{\varphi}_{d} T T_{4}+\bar{\varphi}_{u} T T_{3} \\
& +h_{u}^{i j} Q_{i} u_{j} \varphi_{u}+h_{d}^{i j} Q_{i} d_{j} \varphi_{d}+h_{e}^{i j} L_{i} e_{j} \varphi_{d}
\end{aligned} T=\binom{T_{1}}{T_{2}}
$$

Q,L,u,d,e: Matter fields in the SM
enhancement
$\varphi_{u, d}$ and $\bar{\varphi}_{u, d}$ are integrated out

$$
W=\frac{4 \pi}{M_{f}}\left\{h_{u}^{i j} Q_{i} u_{j}\left(T T_{3}\right)+h_{d} Q_{i} d_{j}\left(T T_{4}\right)+h_{e} L_{i} e_{j}\left(T T_{4}\right)\right\}
$$

Below $\wedge_{H}$

$$
\begin{aligned}
& \left(T T_{3}\right) \rightarrow \frac{\Lambda_{H}}{4 \pi} H_{u} \quad\left(T T_{4}\right) \rightarrow \frac{\Lambda_{H}}{4 \pi} H_{d} \\
& W=h_{u}^{i j} Q_{i} u_{j} H_{u}+h_{d}^{i j} Q_{i} d_{j} H_{d}+h_{e}^{i j} L_{i} e_{j} H_{d}
\end{aligned}
$$

## EWBG in the SM

In the high temperature approximation,

$$
V(\varphi, T) \simeq D\left(T^{2}-T_{0}^{2}\right) \varphi^{2}-E T \varphi^{3}+\frac{\lambda_{T}}{4} \varphi^{4}+\cdots
$$

$$
\varphi_{c} / T_{c}=2 E / \lambda_{T_{c}} \quad 1 \text { st order PT is possible }
$$

$$
E=\frac{1}{12 \pi v^{3}}\left(6 m_{W}^{3}+3 m_{Z}^{3}\right)
$$ due to the cubic term

$\lambda_{T}=\frac{m_{h}^{2}}{2 v^{2}}+\log$ corrections

$$
\varphi_{c} / T_{c} \propto 1 / m_{h}^{2}
$$

## Light Higgs is required !!

In SM, Higgs should be lighter than 50 GeV
excluded by NEW CP phases are also necessary for successful baryogenesis LEP data

## Extension of the SM at TeV scale is necessary

It can be tested by experiments

- New bosonic loop contribution

Q Higher dim. term in the potential

## EWBG in the MSSM

Lighter stop loop can contribute

$$
\varphi_{c} / T_{c}=2 E / \lambda_{T_{c}}>1
$$

$$
E \simeq \frac{1}{12 \pi v^{3}}\left(6 m_{W}^{3}+3 m_{Z}^{3}\right)+\frac{m_{t}^{3}}{2 \pi v^{3}}\left(1-\frac{\left|A_{t}+\mu \cot \beta\right|^{2}}{M_{\tilde{q}}^{2}}\right)
$$

where the maximal contribution case is considered;

$$
m_{\tilde{t}_{1}}^{2}(\varphi, \beta)=\left(M_{T_{R}}^{2}+\frac{y_{t}^{2} s_{\beta}^{2}}{2}\left(1-\frac{\left|A_{t}+\mu \cot \beta\right|^{2}}{M_{\tilde{q}^{2}}}\right) \varphi^{2}\right.
$$

For larger $M_{T R}$, the effect is smaller
Light stop is necessary $\leftrightarrow$ No new coloured particles at LHC $\cdots$ Even with such a maximal case, it's not easy to get $\varphi_{c} / T_{c}>1$

MSSM should be also modified at TeV scale for EWBG

## What kind of modification?

 $\begin{aligned} & \text { Small } m_{h} \text { is } \\ & \text { preferable } \\ & \text { support }\end{aligned} m_{h}=126 \mathrm{GeV} @ 1$A Good point of MSSM : $\mathrm{h}^{4}$ coupling is from gauge coupling $\rightarrow$ Light Higgs

## Large bosonic loop contribution

A strong Higgs coupling with additional bosons ( $\mathrm{h}-\Phi^{\prime \prime}-\Phi^{\prime \prime}$ )
'Q Mass of $\phi^{\prime}$ ' is dominated by vev $m_{\Phi^{\prime}}^{2}=M^{2}+\lambda^{2} v^{2}$
A natural realization of "strong but light" in SUSY model:

MSSM Higgs $\quad Z_{2}$ odd new fields
$W=\lambda \Phi_{u, d} \Phi_{1}^{\prime} \Phi_{2}^{\prime} \longrightarrow \Delta V=|\lambda|^{2} h^{2} \varphi_{1,2}^{\prime \dagger} \varphi_{1,2}^{\prime}$
strong but light!

It provides strong coupling but $m_{h}$ is kept small!

## Tests of the scenario

Enhancement
of $\varphi_{c} / T_{c}$ $\uparrow$
destructive

## Ino loop

negative contribution

## contribution to hhh coupling $\leftarrow$ Linear Collider

Inert scalar mass: $m_{\Phi^{\prime}}^{2}=M^{\prime 2}+\lambda^{2} v^{2}$
Inert ino mass: $m_{\tilde{\Phi}^{\prime}}=\mu^{\prime}+\lambda v$
The loop contributions are significant when $\lambda v$ dominates the masses.
$\mathrm{Z}_{2}$ odd scalars as light as $\sim \lambda v$

Large $\mu^{\prime}$ and small $M^{\prime 2}$ provides large deviation in hhh and large $\varphi_{c} / T_{c}$

