Stau NLSP Pair Production @ ILC 500GeV

Takuaki Mori, Ryo Katayama, Tomohiko Tanabe,
Satoru Yamashita (The University of Tokyo),
Sho Iwamoto, Norimi Yokozaki, Shigeki Matsumoto (Kavli IPMU),
Taikan Suehara (Kyushu University),
Keisuke Fujii (KEK)

TALK CONTENTS

Introduction **Constraints of mass & lifetime** Physics motivation of Stau pair production Analysis **Event selection** Cut variables in detail Results Significance Lifetime & Mass sensitivity

Future plan

stau lifetime&mass

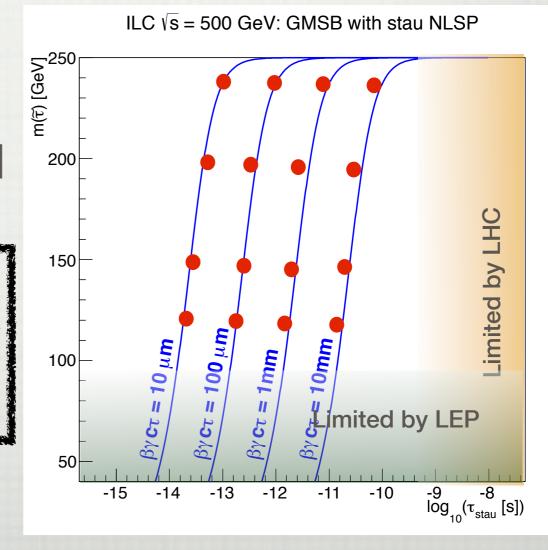
Lifetime upper limit @LHC
Squark, Gluino > 1TeV
Stable stau < 360GeV is excluded

-> Stau decays promptly

Mass lower limit @ LEP stau < 90 GeV is excluded

Benchmark point

- Mass 120GeV~240GeV
- Lifetime $10 \mu \text{m} \sim 10 \text{mm}$



Supersymmetry Breaking



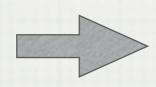
No Breaking \longleftrightarrow $M_{\rm SUSY} \simeq M_{\rm SM}$



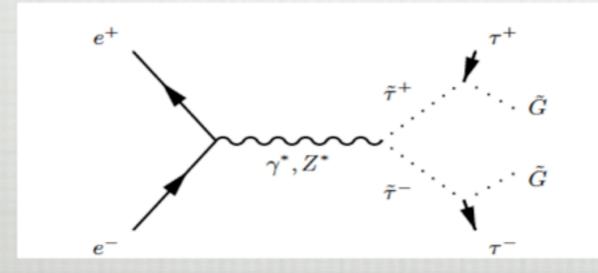
→ Soft Breaking term



Gauge Mediated Symmetry Breaking (GMSB)

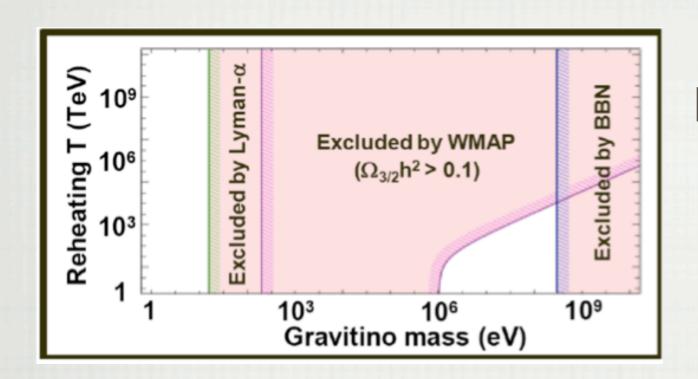


- Solve SUSY Flavor Problem
- Gravitino LSP
- Decay mode: Stau->Tau+Gravitino



We observe particles decayed from tau lepton

Low Scale Gravitino mass $m_{3/2} \sim \mathcal{O}(10) \mathrm{eV}$



Free from constraints of *reheating Temperature *LHC experiment

In GMSB...

Stau lifetime and mass

Gravitino Mass

$$\tau_{\tilde{\tau}} = 48\pi M_{pl}^2 m_{3/2}^2/m_{\tilde{\tau}}^5 \simeq 5.9 \times 10^{-12} \times (\frac{m_{3/2}}{10 \mathrm{eV}})^2 (\frac{m_{\tilde{\tau}}}{100 \mathrm{GeV}}^{-5})$$
 [arXiv 1104.3624v1]

Sensitivity of Stau mass and lifetime is important for determination of susy breaking scale.

Other decay modes can be considered..

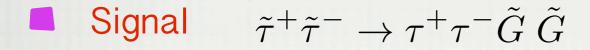
[Bhattacharyya Bhattacherjee, Yanagida, Yokozaki arciv 1304.2508v2]

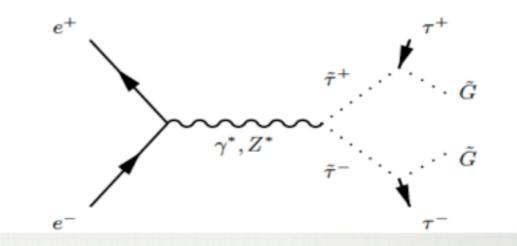
- Stau \longrightarrow Tau + Axino $ilde{ au} o au ilde{a}$
- Stau \longrightarrow Tau + Neutrino (RPV) $ilde{ au} o au
 u_e, au
 u_{\mu}$

Stau Search strategy is applicable in these modes.

common signature: Tau + missing particle

EVENTSELECTION





• 1-prong decay (85%) $\tau \to e \nu \bar{\nu} (17.82\%) \qquad \tau \to \mu \nu \bar{\nu} (17.39\%)$

$$au
ightarrow e
u \overline{
u} (17.82\%)$$

$$au o \pi \nu \nu (10.91\%)$$

$$au o \mu \nu \bar{\nu} (17.39\%)$$

$$\tau \to \pi \nu \nu (10.91\%)$$
 $\tau \to \pi \pi^o \nu (25.51\%)$

- 3-prong decay (15%)
- Background
- main background

$$e^{+}e^{-} \to W^{+}W^{-} \to l^{+}l^{-}\nu\bar{\nu}$$
 $e^{+}e^{-} \to ZZ \to l^{+}l^{-}\nu\bar{\nu}$ $e^{+}e^{-} \to Z/\gamma \to l^{+}l^{-}$

large cross section

$$e\gamma \to el^+l^ e^+e^- \to e^+e^-$$
 (bhabha) $\gamma\gamma \to l^-l^+$

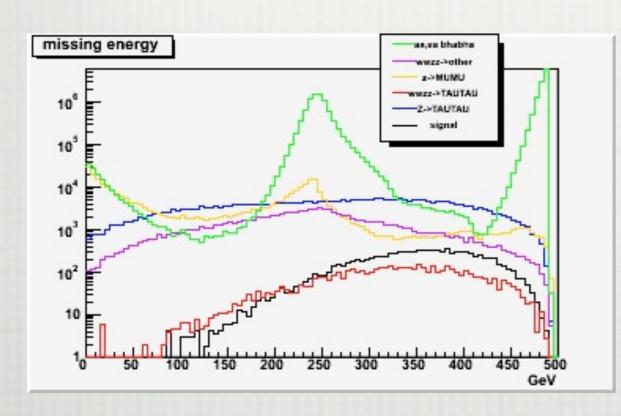
Beam polarization (e,p)=(+0.8, -0.3)

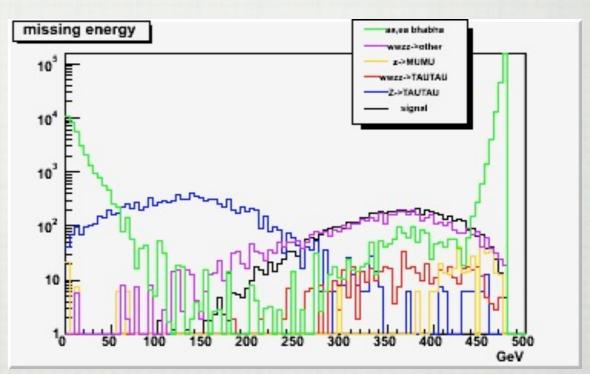
CUT FLOW (ex. 200GeV 100um)

- \square 0. number of tracks==2 \longrightarrow cut multi-prong decay
- 2. Track Energy < 125GeV for each track → cut Z->μμ
- □ 3. $-0.9 < \cos \theta_{-} < 0.8 \& -0.8 < \cos \theta_{+} < 0.9 \longrightarrow \text{cut t-channel BG}$
- \square 4. acoplanarity $-0.9 < \cos(\phi_- \phi_+) \longrightarrow \text{cut Z->II}$
- \Box 5. $|\cos \theta_{\rm miss}| < 0.9$ cut ey, yy BG
- 6. missing energy >250GeV → cut Z->ττ BG
- □ 7. Hcal Deposit > 3 % → cut WWZZ->ee,μμ,eμ + vv BG
- □ 8. Track Energy / Calorimeter Deposit > 0.03 → cut Z-> μμ
- ☐ 9. Cosine angle between two tracks > -0.9 —select signal
- 10. Impact parameter significance > 2.5 → cut WWZZ-> Ilvv

MISSING ENERGY>250GEV

Reduce Z->tauTau process (blue line).
Gravitino has large missing energy (signal: black line)



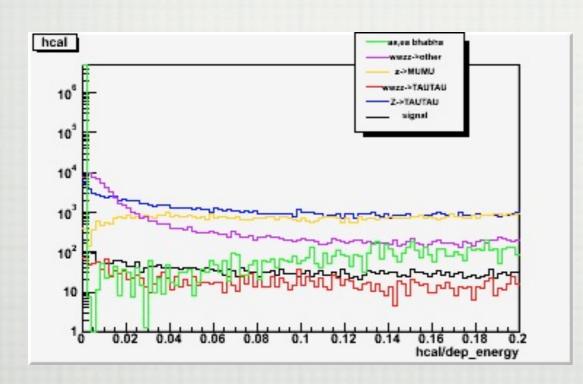


number of tracks==2

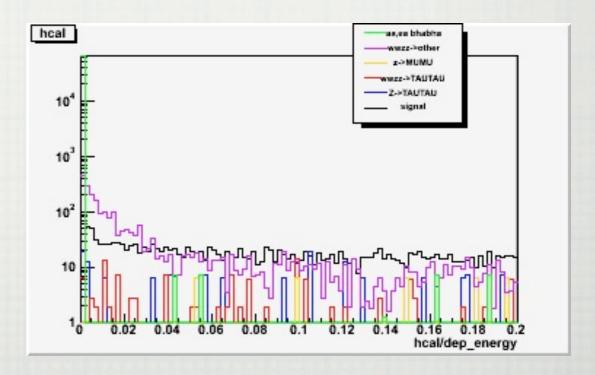
cut variable 0~5

HCAL Deposit / Energy Depotit > 0.03

 $WW, ZZ \rightarrow \mu\mu\nu\nu, ee\nu\nu, e\mu\nu\nu$ is reduced for its low energy deposit in hadron calorimeter.(purple line)



number of tracks==2



Cut Variable 0~6

impact parameter cut (ex 200GeV, 100um)

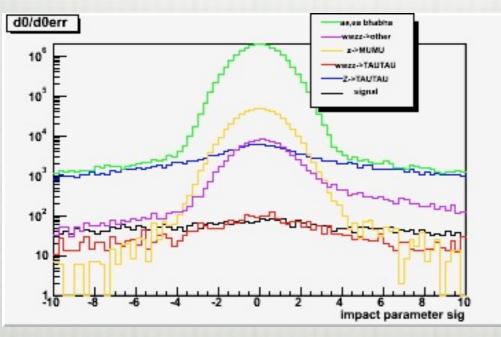
+ STAU IMPACT PARAMETER

• |D0/D0_err| > 2.5

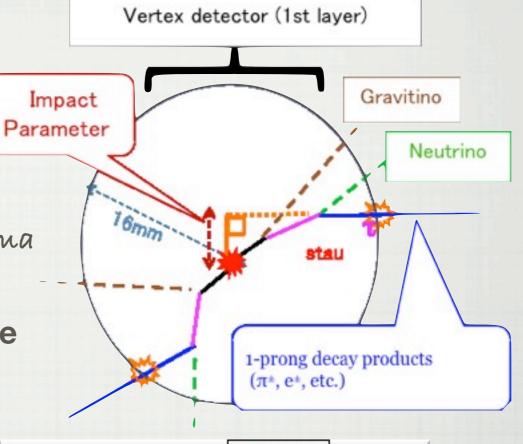
Do=impact parameter, Do_err = Do sigma

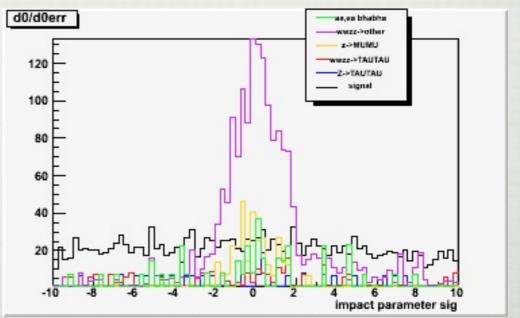
impact parameter related with stau lifetime

 $\longrightarrow WW, ZZ \rightarrow ll\nu\bar{\nu}$



number of tracks==2





Cut Variable 0~9

Statistical Significance Lifetime vs Mass RESULT

	I20GeV	I50GeV	200GeV	240GeV
I0μm	113.2	98.6	53.4	6.05
I00µm	117.1	101.5	55.3	6.00
Imm	121.5	105.3	59.1	8.14
I0mm	123.1	109.5	62.4	12.3

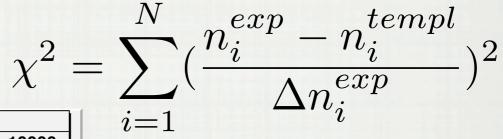
 $>5 \sigma$ for all benchmark points

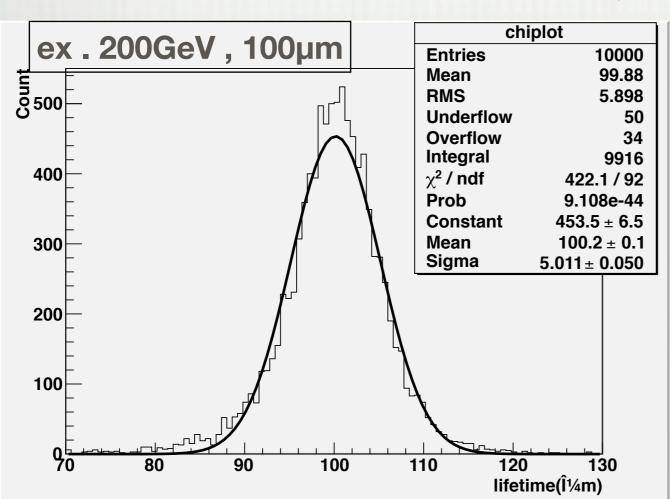
Lifetime Analysis

Perform the toy MC experiment with each experiment distribution with poisson statistics folded in , and compare with the high statistics samples

with various lifetime by calculation of chi square quantity

of chi square quantity.





Stau Mass 200GeV Stau Lifetime 100µm

$$c\tau = 99.8 \pm 5.0 \; (\mu m)$$

5% precision

Mass Analysis (kinematic cut)

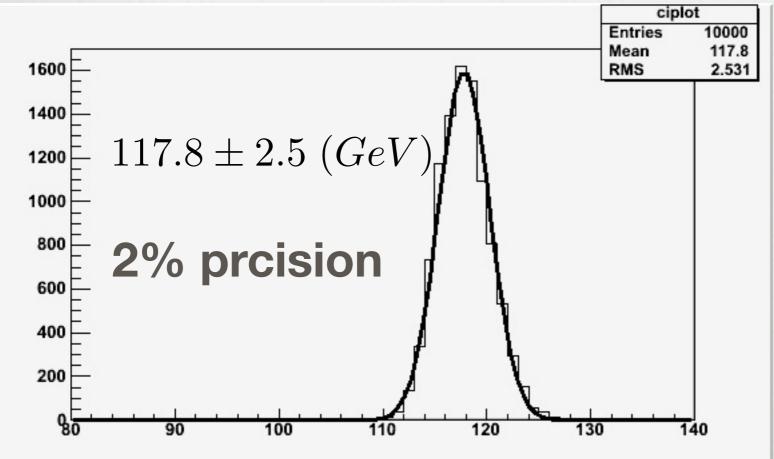
Kinematic cut is applied for measurement of stau mass via the maximum track energy.

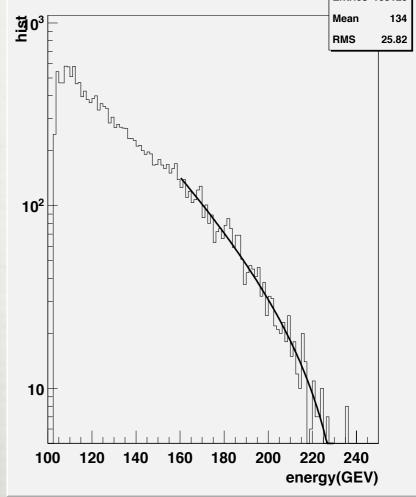
Track energy distribution is modeled with the following

function $f(x) = \alpha(\beta - x)exp(-\gamma x)\theta(\beta - x)$.

120GeV mass corresponds to the maximum kinematic

energy of 234GeV.

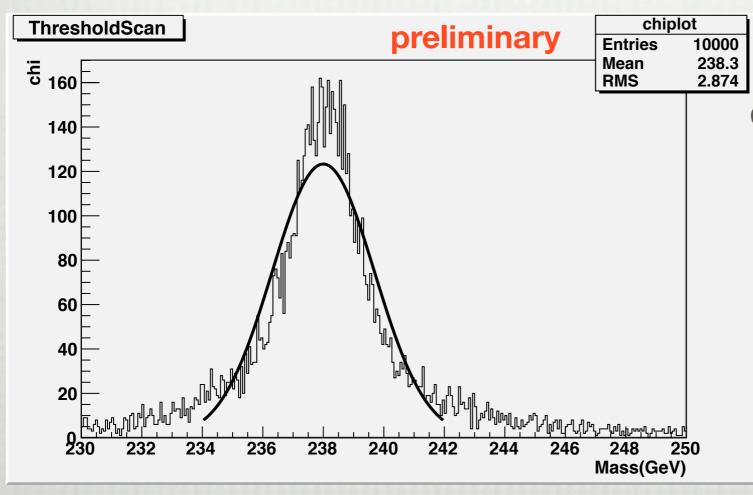




track hist

Mass Analysis (Threshold Cut)

The technique of Threshold scan is used for heavy stau, for its low sensitivity of kinematic mass measurement.



ex. M=240GeV ct=100µm

 $238.3 \pm 2.8 \; (GeV)$

1.2 % precision

Conclusions & Future Plans

Conclusions

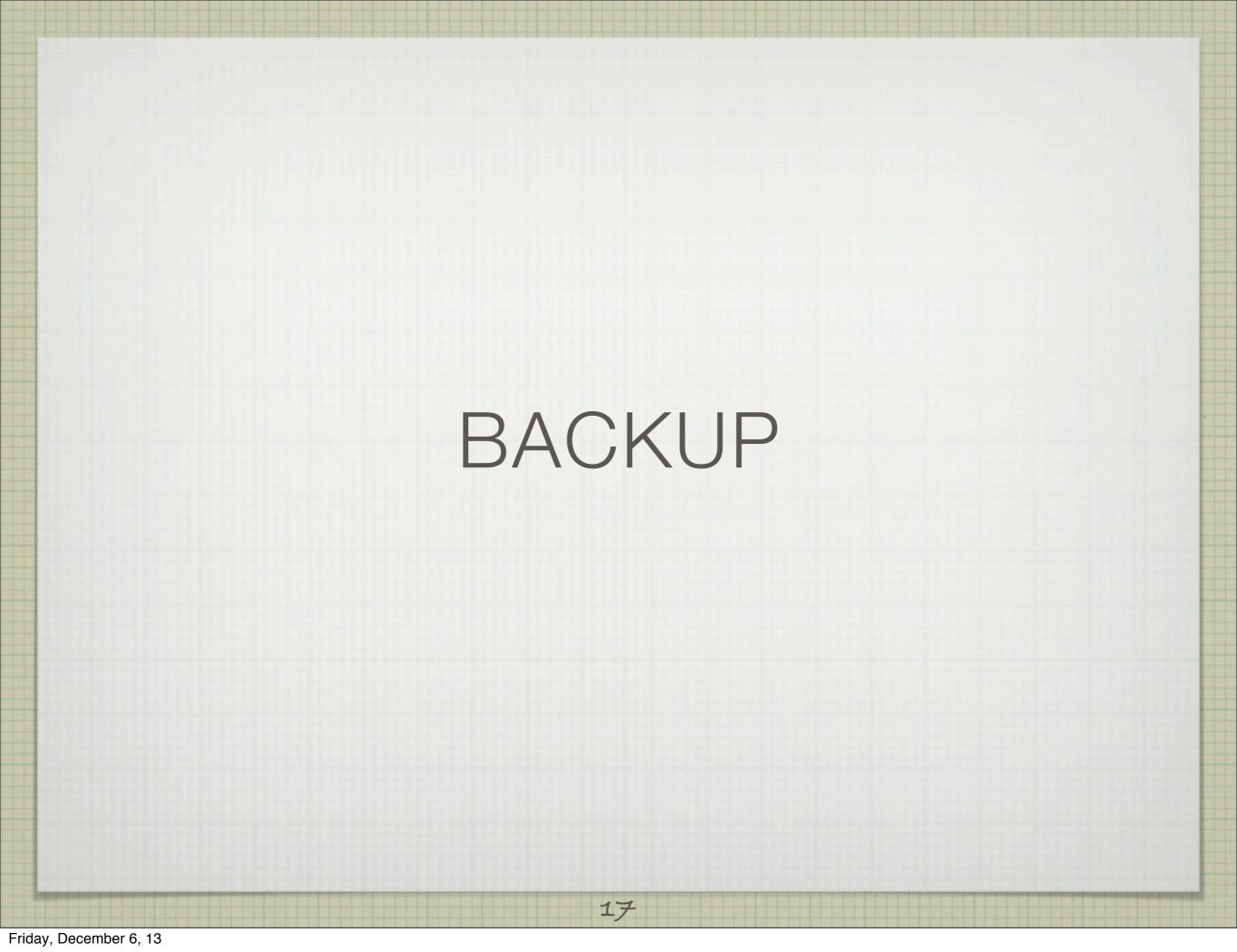
We studied the ILC sensitivity for measuring the wide range of benchmark points of stau mass and lifetime and obtain over 5 σ significance.

Mass and lifetime sensitivity are also studied, and for example we can obtain stau Mass with 5% precision at M=200GeV.

Future Plans

- Optimization of cut flow for kinematic mass measurement to improve its sensitivity.
- 3-prong decays are being studied.

THANK YOU FOR YOUR ATTENTION.



physics motivation

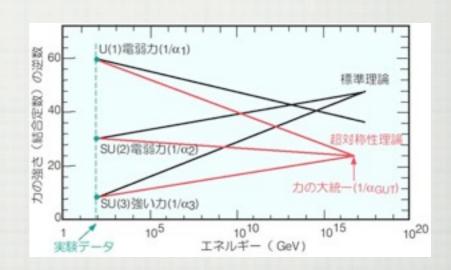
WHY SUSY?

FINE TUNING

quadric divergence become logarithmic divergence

GUT SCALE UNIFICATION

Coupling constant unification at high scale region.

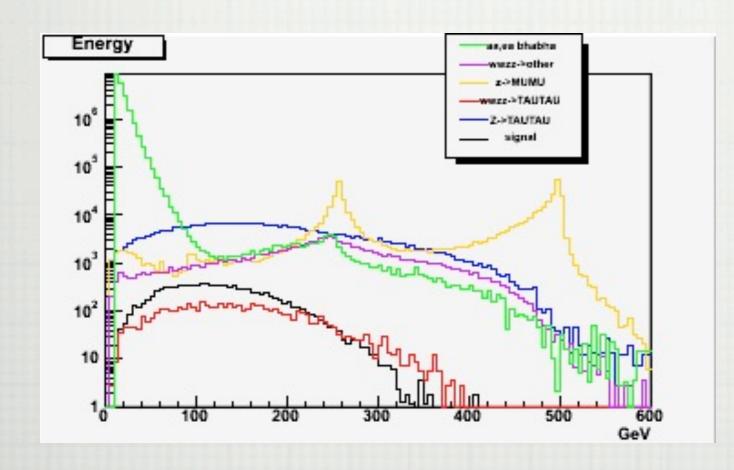


DARK MATTER

LSP(lightest supersymmetric particle) is a prominent candidate

Energy >20GeV

Suppress low energy beam related $\gamma\gamma \rightarrow l^+l^-$ process



Signal

$$--- e^+e^- \rightarrow \tilde{\tau}^-\tilde{\tau}^+$$

BackGround

$$\gamma \gamma \rightarrow l^- l^+$$
, bhabha 散乱

WW,
$$ZZ \rightarrow l^+ l^- \nu \bar{\nu} (\tau \tau \nu \bar{\tau}$$
 除く)

$$--- Z \rightarrow \mu\mu$$

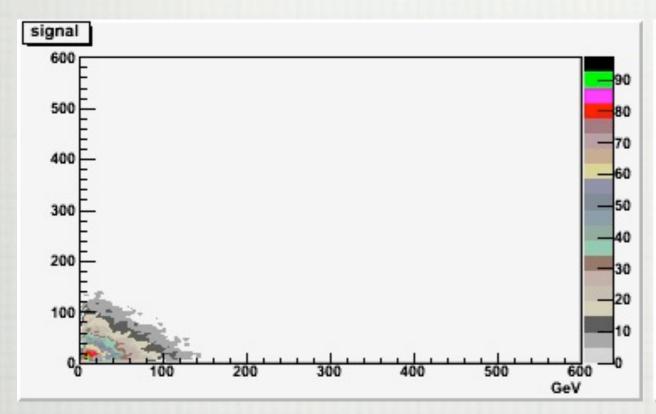
$$WW, ZZ \to \tau^+ \tau^- \nu \bar{\nu}$$

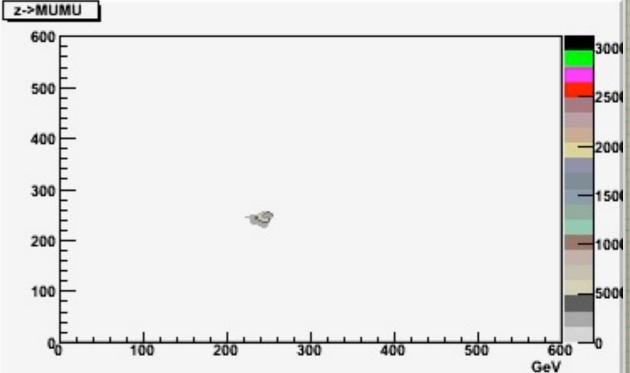
$$- Z \rightarrow \tau \tau$$

number of tracks==2

Track Energy < 125 GeV (for each track)

Supress high energy muon from $Z o \mu \mu$.





number of tracks==2

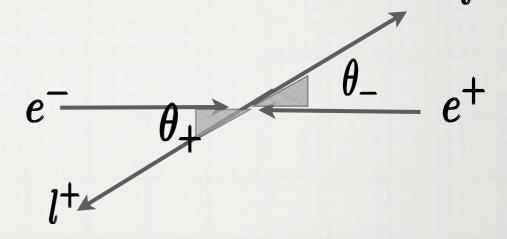
$$e^+e^- \to \tilde{\tau}^-\tilde{\tau}^+$$

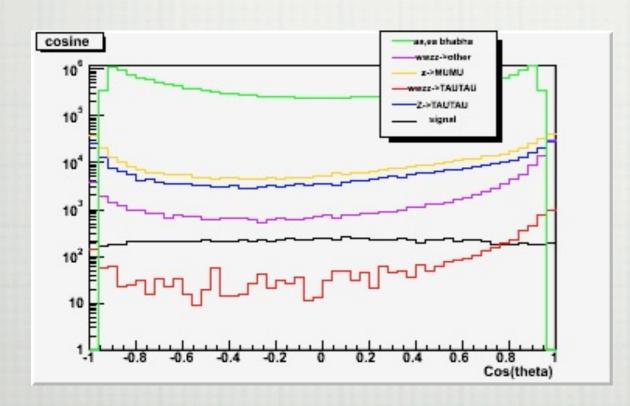
number of tracks==2

$$Z \to \mu^+ \mu^-$$

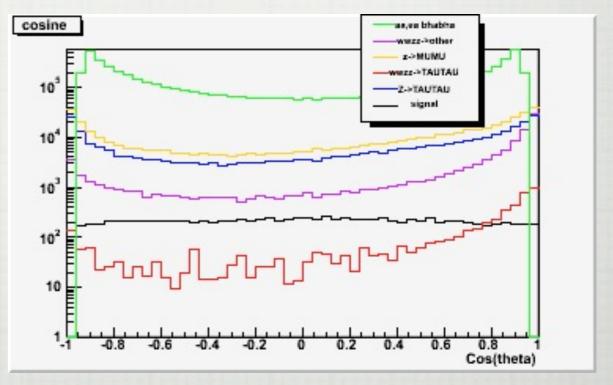
$$-0.9 < \cos \theta_{-} < 0.8 \& -0.8 < \cos \theta_{+} < 0.9$$

Cut forward-backword dominant W boson t-channel decays (purple line)







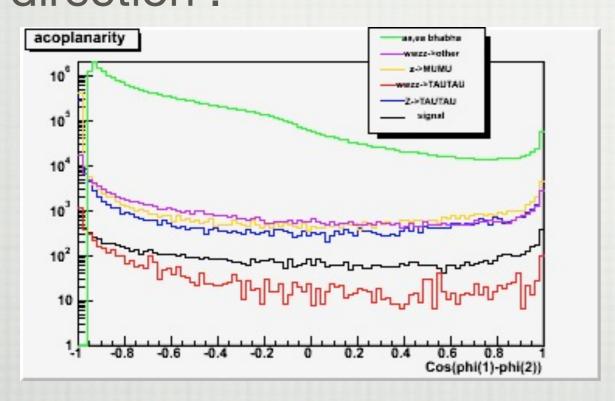


Cut Variable 0~2

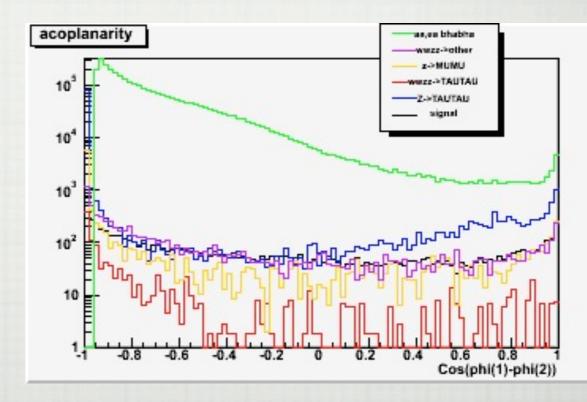
Acoplanarity $-0.9 < \cos(\phi_- - \phi_+)$

Cut tracks in opposite direction.

z->tautau background can be reduced for its low missing energy and highly boosted in opposite direction.



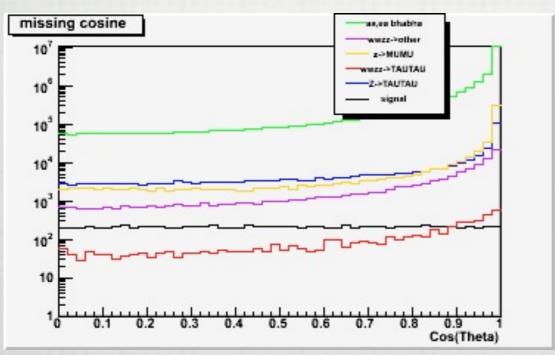
number of tracks==2



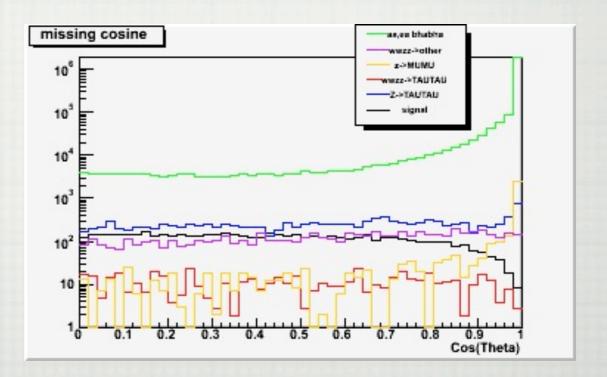
Cut Variable 0~3

$$|\cos \theta_{\rm miss}| < 0.9$$

Reduces beam-related background for bhabha scattering, each of whose tracks is towards the undetected beam pipe direction, and it appears as large missing energy towards forward-backward direction.





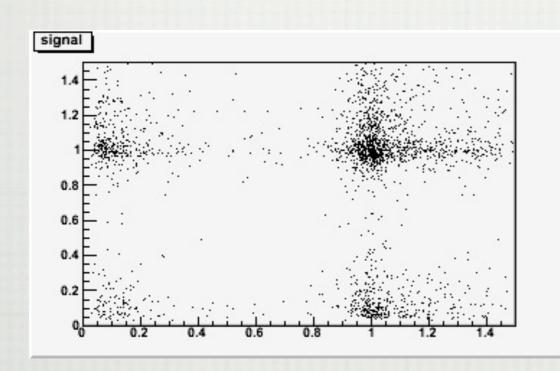


Cut Variable 0~4

number of tracks==2

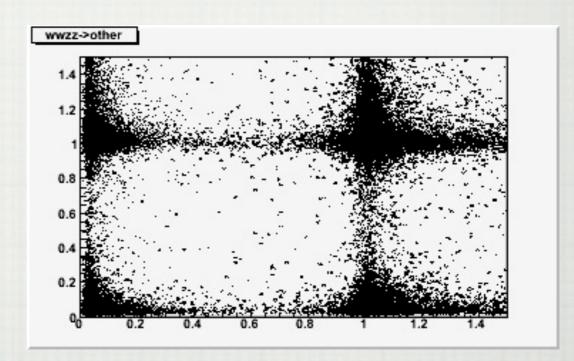
Energy deposit/Track energy > 0.3 for each track

Reduce $Z \to \mu\mu$ process for its low deposit in calorimeters and highly energetic track.



number of tracks==2

$$e^+e^- \to \tilde{\tau}^-\tilde{\tau}^+$$



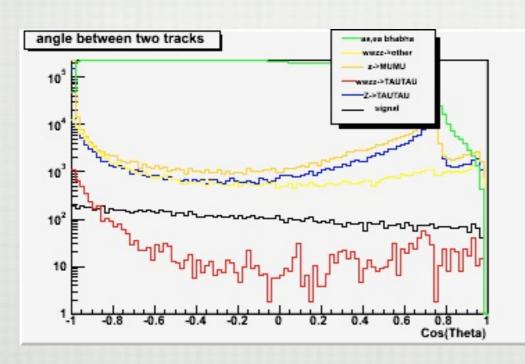
number of tracks==2

$$WW, ZZ \rightarrow ll\nu\nu \ (l \neq \tau)$$

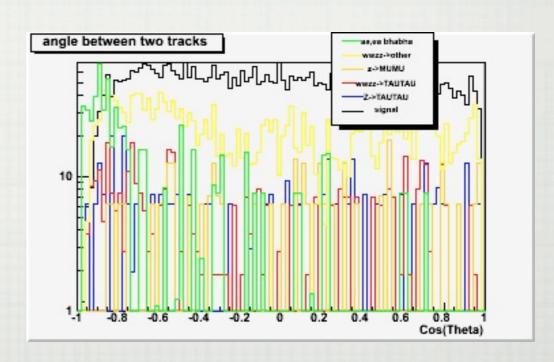
Cosine angle between two tracks > -0.9

 \widetilde{T} の質量が重い場合ほぼ静止しているので、 $\widetilde{\tau} \to \tau \widetilde{G}$ におけるTは等方的に飛びやすい。

一方 $WW \to l ar l
u ar
u$ はブーストされ、正反対方向に飛びやすいためカットされる。(黄色、赤線)



number of tracks==2

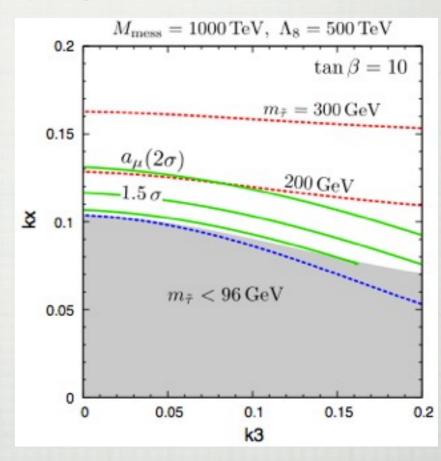


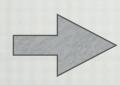
Cut Variable 0~8

- 重いSquark、軽いSlepton を示唆し、g-2の実験結果を説明する新たなモデル [lbe, Matsumoto, Yanagida, Yokozaki arxiv 1210.3122v2]
- GUT, GMSB -> no SUSY flavor/CP problem
- ■g-2 を説明可能

Stau mass = 200GeV で 実験値とのズレを2σで説明可能

(Stau mass = muon 150GeV の場合 1.5σ)





Stau -> Tau + Gravitino の崩壊モード

Polarization (e-,e+)=(+0.8,-0.3)

SIGNAL

BACKGROUND

Mass (GeV)	Cross Section (fb)			
I20GeV	139	V		
200GeV	33.3			
I50GeV	97.2			
240GeV	3.22			

Category	Cross Section (fb)	
$WW, ZZ \rightarrow l^+ l^- \nu \bar{\nu}$	445.9	
$Z/\gamma \to \tau^+ \tau^-$	1283.4	
$\gamma\gamma \to l^+ l^-$	1.41×10^6	

PHYSICS MOTIVATION2

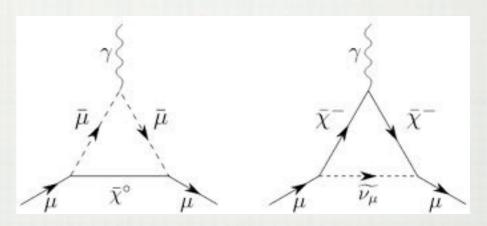
WHY STAU PRODUCTION?

MUON g-2 3.6 sigma deviation between SM theory

SUSY particles 1-loop correction

$$\delta a_{\mu} \sim \frac{3}{5} \frac{g_1}{8\pi^2} \frac{m_{\mu}^2 \mu tan\beta}{m_{\rm soft}^3} F_b$$

 $m_{
m soft}$ = loop sparticle mass g is SU(2), Y(1) coupling const. $a_{\mu}\equiv \frac{g-2}{2}$

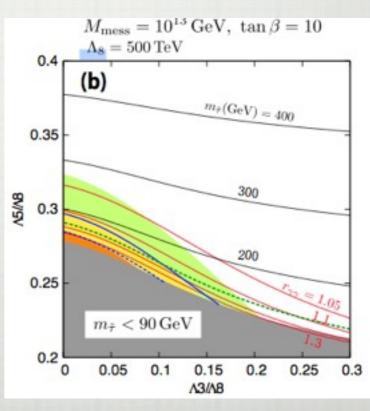


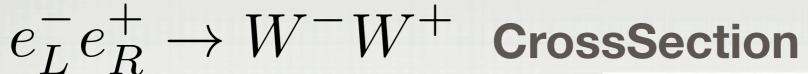
3.6 sigma deviation is explain with assumption of Sfermion exist in the region of 100GeV

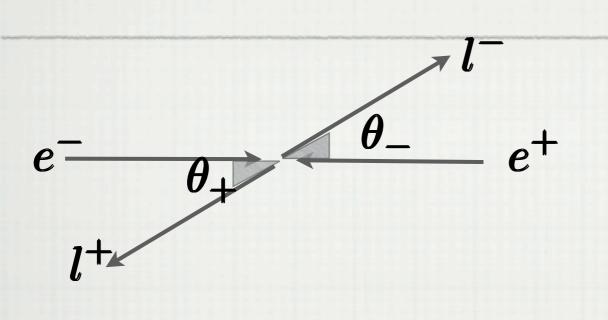
$$\mu \tan \beta \sim 4 \text{TeV}$$

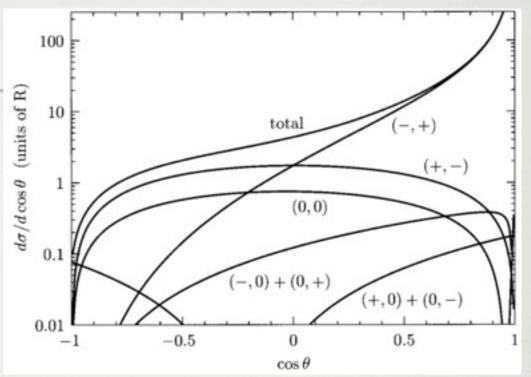
$$\mu$$
tan β -4TeV \longrightarrow large stau left-right mixing

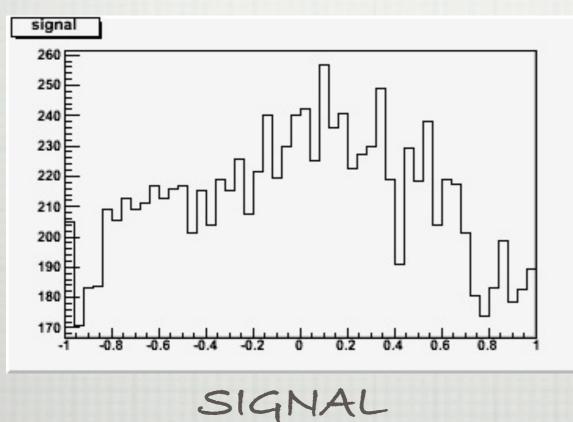
$$\longrightarrow$$
 FOR $a_{\mu} \sim 1.5\sigma \rightarrow m_{\tilde{ au}} \sim 200 GeV$

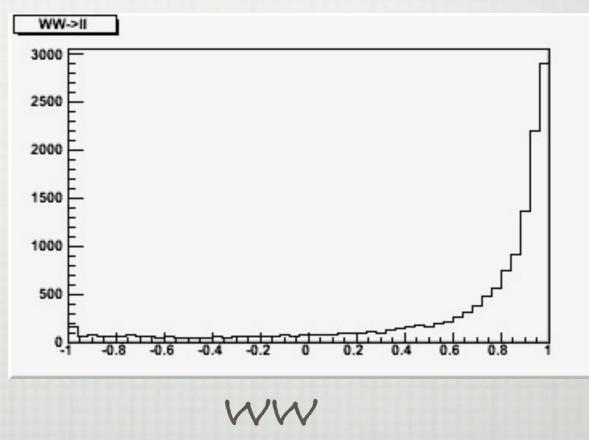












$$\frac{\Delta m_{\frac{3}{2}}}{m_{\frac{3}{2}}} = \sqrt{\left(\frac{5}{2}\frac{\Delta m_{\tilde{\tau}}}{m_{\tilde{\tau}}}\right)^2 + \left(\frac{1}{2}\frac{\Delta \tau_{\tilde{\tau}}}{\tau_{\tilde{\tau}}}\right)^2}$$

$$\frac{\Delta m_{\tilde{\tau}}}{m_{\tilde{\tau}}} \simeq 1.4\% \text{ (track energy) } \frac{\Delta \tau_{\tilde{\tau}}}{\tau_{\tilde{\tau}}} \simeq 1.4\% \text{ (lifetime result)}$$

$$\frac{\Delta m_{\tilde{\tau}}}{m_{\tilde{\tau}}} \simeq 0.5\% \text{ (threshold scan)}$$