## SLOOPS

# An automatic program for full one-loop calculations in the SM/MSSM

Guillaume CHALONS LAPTH, Annecy-le-Vieux

The SloopS team : Fawzi BOUDJEMA, Guillaume DRIEU LA ROCHELLE, Sun Hao (LAPTH, Annecy-le-Vieux) Nans BARO (ITTK, RWTH Aachen University) <u>Andreë SEMENOV</u> (JINR Dubna)

## **IWLC 2010**

## **NEED FOR PRECISE THEORETICAL PREDICTIONS**

#### RELIC DENSITY OF DARK MATTER

- WMAP :  $0.0975 < \Omega_{DM} h^2 < 0.1223$  (10% precision)
- PLANCK : 2% precision



#### PRECISION MEASUREMENTS



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Automatic calculations in the SM/MSSM : The SloopS project

## **COSMOLOGY AND PARTICLE PHYSICS**

#### RELIC DENSITY IN THE STANDARD SCENARIO

$$\Omega_{DM}h^2\simeq rac{3 imes 10^{-27}cm^3s^{-1}}{\langle\sigma(\chi\chi o SM)v
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## COSMOLOGY AND PARTICLE PHYSICS

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#### PRECISION

- Need for precise theoretical predictions w.r.t experimental measurements.
- Precision needed at the level of  $\sigma \Rightarrow$  One-loop calculations (at least).
- Reconstruction of fundamental underlying parameters at LHC and LC.
- Radiative corrections must be under control to be able to constrain the cosmological underlying scenario.
- What precision required at colliders and theory to constrain cosmology?









## A PARADIGM OF BSM MODELS : SUSY AND THE MSSM





16 18

## SUPERSYMMETRY AND THE MSSM

#### COMPLICATIONS

- Not observed yet, neither the Higgs boson...
- $\mathcal{L}_{soft}$  unkown.
- Lots of free parameters ( $\simeq 105$ ).
- Calculations become extremely tedious and involved.



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#### BEYOND LEADING ORDER IN SUSY

- At LO : m<sub>h</sub> < m<sub>Z</sub> but no Higgs found !
- LEP Bound on Higgs mass  $| m_h > 114 \text{GeV} |$
- At higher orders : Higgs mass can get large corrections.
- Generically SUSY processes get large radiative corrections.
- Calculations become even more complicated...



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## RADIATIVE CORRECTIONS ARE IMPORTANT

#### AUTOMATION NEEDED



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## AUTOMATIC TOOL FOR ONE-LOOP CALCULATIONS : SLOOPS



- Evaluation of one-loop diagrams including a complete and coherent renormalisation of each sector of the MSSM with an OS scheme.
- Modularity between different renormalisation schemes.
- Non-linear gauge fixing.
- Checks : results UV, IR finite and gauge independent.

http://code.sloops.free.fr/



#### FERMION + GAUGE SECTOR

Input parameters as in the Standard Model  $M_f, \alpha(0), M_W, M_Z$ 



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Input parameters : M

$$I_{A0}, t_{\beta} = v_2/v_2$$

 $v_1$  . Several definitions for  $\delta t_\beta$  :

•  $\overline{DR}$  :  $\delta t_{\beta}$  is a pure divergence

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Several definitions for  $\delta t_{\beta}$  :

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Input parameters : 3 sfermions masses  $m_{\tilde{d}_1}, m_{\tilde{d}_2}, m_{\tilde{u}_1}$  and 2 conditions for  $A_{u,d}$ 

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#### NEUTRALINOS/CHARGINOS SECTOR

Input parameters : 2 charginos  $m_{\tilde{\chi}_1^{\pm}}, m_{\tilde{\chi}_2^{\pm}}$  and 1 neutralino  $\tilde{\chi}_1^0$ 

#### Linear gauge fixing

$$\mathcal{L}_{GF} = -\frac{1}{\xi_W} |\partial_\mu W^{\mu +} + i\xi_W \frac{g}{2} vG^+|^2$$
$$-\frac{1}{2\xi_Z} (\partial_\mu Z^\mu + \xi_Z \frac{g}{2c_w} vG^0)^2$$
$$-\frac{1}{2\xi_A} (\partial_\mu A^\mu)^2$$

$$\Gamma^{VV} = rac{-i}{q^2 - M_V^2 + i\epsilon} \left[ g_{\mu
u} + (\xi_V - 1) rac{q_\mu q_
u}{q^2 - \xi_V M_V^2} 
ight]$$



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$$\Gamma^{VV} = \frac{-i}{q^2 - M_V^2 + i\epsilon} \left[ g_{\mu\nu} + (\xi_V - 1) \frac{q_{\mu} q_{\nu}}{q^2 - \xi_V M_V^2} \right]$$

 $\xi_{W,Z,A} = 1$  (Feynman gauge)



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#### Non-Linear gauge fixing

$$\mathcal{L}_{GF} = -\frac{1}{\xi_{W}} |(\partial_{\mu} - ie\tilde{\alpha}A_{\mu} - igc_{w}\tilde{\beta}Z_{\mu})W^{\mu +} \\ + i\xi_{W}\frac{g}{2}(v + \tilde{\delta}h^{0} + \tilde{\omega}H^{0} + i\tilde{\kappa}G^{0} + i\tilde{\rho}A^{0})G^{+}|^{2} \\ - \frac{1}{2\xi_{Z}}(\partial_{\mu}Z^{\mu} + \xi_{Z}\frac{g}{2c_{w}}(v + \tilde{\epsilon}h^{0} + \tilde{\gamma}_{H}^{0})G^{0})^{2} \\ - \frac{1}{2\xi_{A}}(\partial_{\mu}A^{\mu})^{2}$$

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 $\xi_{W,Z,A} = 1$  (Feynman gauge)

 $\rightarrow$  Gauge parameter dependence in gauge/Goldstone/ghost vertices.

→ No "unphysical " threshold, no higher rank tensor.



#### TREE LEVEL CALCULATIONS

## Comparison with public codes : ${\tt Grace}$ and ${\tt CompHEP}$ Nans Baro PhD Thesis

Cross-section [pb]	SloopS	CompHEP	Grace	
$h^0 h^0 \rightarrow h^0 h^0$ 3	$3.932 \times 10^{-2}$	$3.932 \times 10^{-2}$	$3.929 \times 10^{-2}$	
$W^+W^- \rightarrow \tilde{t}_1\bar{\tilde{t}}_1$ 7	$7.082 \times 10^{-1}$	$7.082 \times 10^{-1}$	$7.083 \times 10^{-1}$	
$e^+e^- \rightarrow \tilde{\tau}_1 \bar{\tilde{\tau}}_2$ 2	$2.854 \times 10^{-3}$	$2.854 \times 10^{-3}$	$2.854 \times 10^{-3}$	
$H^+H^- \rightarrow W^+W^-$ 6	$5.643 \times 10^{-1}$	$6.643 \times 10^{-1}$	$6.644 \times 10^{-1}$	
Decay [GeV]				# 200 processes checked
$A^0 \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-$ 1	.137×10 <sup>0</sup>	1.137×10 <sup>0</sup>	1.137×10 0	
$\tilde{\chi}_1^+ \rightarrow t \tilde{\tilde{b}}_1$ 5	5.428×10 <sup>0</sup>	5.428×10 0	5.428×10 0	
$H^{0} \rightarrow \tilde{\tau}_1 \bar{\tilde{\tau}}_1 = 7$	7.579×10 <sup>-3</sup>	$7.579 \times 10^{-3}$	$7.579 \times 10^{-3}$	
$H^+ \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^0$ 1	$1.113 \times 10^{-1}$	$1.113 \times 10^{-1}$	$1.113 \times 10^{-1}$	

#### ONE-LOOP PROCESSES THAT DO NOT NEED RENORMALISATION

Comparison with public codes : PLATON and DarkSUSY Implementation of a special routine for loop integrals at v = 0Boudjema,Semenov,Temes, *Phys. Rev.* **D72**, 055024 (2005)

• 
$$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \to \gamma \gamma$$

• 
$$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow gg$$

• 
$$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \gamma Z^0$$

## **APPLICATIONS IN THE HIGGS SECTOR**

N. Baro., F. Boudjema, A. Semenov, Phys. Lett. B660 (2008) 550, 0710.1821 [hep-ph]

 One-loop corrections to Higgs masses H<sup>+</sup>, h<sup>0</sup> Freitas, Stockinger, Phys. Rev. D66 (2002) 095014, hep-ph/0205281

$t_{eta}=3$	mhmax	large $\mu$	nomix				
Tree Level	72.51	72.51	72.51				
DCPR	134.28	97.57	112.26				
MH	140.25	86.68	117.37				
A au au	134.25	97.59	112.27				
$\overline{\mathrm{DR}}\ \overline{\mu} = m_{A^0}$	134.87	98.10	112.86				
1 * 1 * 1 *							

Light Higgs mass m<sub>h0</sub>

• $A^0 \rightarrow \tau^+ \tau$	$^-$ , $A^0  ightarrow Z^0 h^0$ ,	$H^0  ightarrow Z^0 Z^0$ , $I$	$H^0  ightarrow  au^+  au^-$				
	$t_{eta}=3$	mhmax	large $\mu$	nomix			
	Tree Level	$9.35 \times 10^{-3}$	$9.35 \times 10^{-3}$	9.35×10 <sup>-3</sup>			
	DCPR	$-1.09 \times 10^{-4}$	$-7.96 \times 10^{-5}$	$-1.09 \times 10^{-4}$			
	MH	$+6.28 \times 10^{-3}$	$-7.91 \times 10^{-3}$	$+4.47 \times 10^{-3}$			
	A  au  au	$-1.45 \times 10^{-4}$	$-7.09 \times 10^{-5}$	$-1.01 \times 10^{-4}$			
	$\overline{\mathrm{DR}}\ \overline{\mu} = m_{A^0}$	$+5.08 \times 10^{-4}$	$+3.24 \times 10^{-4}$	$+4.17 \times 10^{-4}$			
	$H^0 \rightarrow \tau^+ \tau^-$ at one-loop with no QED						

• Theoretical issue due to non-linear gauge fixing and modified Ward-Slavnov-Taylor Identity in the Higgs sector :

$$m_{A^0}^2 \times A^0 \dashrightarrow \bigcirc \neg \rightarrow \Diamond \neg \rightarrow \land A^0 \dashrightarrow \bigcirc \neg \rightarrow G^0 = (m_{A^0}^2 - m_{Z^0}^2) \frac{i\epsilon}{s_{2W}} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow \bigcirc (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{h^0}^{G^0} \dashrightarrow \land A^0 + \tilde{\gamma} \times \circlearrowright_{H^0}^{G^0} \dashrightarrow (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{H^0}^{G^0} : (\tilde{\epsilon} \times \circlearrowright_{H^0}^{G^0} : (\tilde{\epsilon}) \xrightarrow{i\epsilon} [\tilde{\epsilon} \times \circlearrowright_{H^0}^{G^0} : (\tilde{\epsilon} \times \circlearrowright_{H^0}^{G^0} : (\tilde$$

## **APPLICATIONS IN THE CHARGINO/NEUTRALINO SECTOR**

N. Baro, F. Boudjema, Phys. Rev. D80 (2009) 076010, arXiv :0906.1665[hep-ph].



 Chargino decays at one-loop (comparison with J. Fujimoto et al., Phys. Rev. D75 (2007) 113002, hep-ph/0701200.)



## **APPLICATIONS TO COLLIDER PHYSICS**

e<sup>+</sup>e<sup>-</sup> → χ̃<sub>1</sub><sup>+</sup> χ̃<sub>1</sub><sup>-</sup> J. Fujimoto *et al.*, *Phys. Rev.* D75 (2007) 113002, hep-ph/0701200.
 e<sup>+</sup>e<sup>-</sup> → τ̃<sub>i</sub>τ̃<sub>i</sub><sup>\*</sup> K. Kovarik *et al.*, *Phys. Rev.* D72 (2005) 053010, hep-ph/0506021.

![](_page_23_Figure_2.jpeg)

![](_page_23_Picture_3.jpeg)

## APPLICATIONS TO PURE STANDARD MODEL PROCESSES

F. Boudjema, Le Duc Ninh, Sun Hao, M. M. Weber, Phys. Rev. D81 073007 (2010)

- $e^+e^- \rightarrow W^+W^-Z^0$
- $e^+e^- \rightarrow Z^0 Z^0 Z^0$

• Important processes to test the quartic gauge couplings and Higgs mechanism

![](_page_24_Figure_5.jpeg)

See also Su Ji-Juan et al. Phys. Rev D78 016007, Sun Wei et al. Phys. Lett B680, 321

![](_page_24_Picture_7.jpeg)

## **APPLICATIONS TO DARK MATTER**

#### THERMAL RELIC

- $\Omega_{\chi} h^2 \propto 1/(\sigma(\chi\chi \to {
  m SM}))$
- Relic density calculated through the interface of SloopS with micrOMEGAs (Bélanger et al.)

![](_page_25_Figure_4.jpeg)

#### BUNCH OF FULL ONE-LOOP PROCESSES CALCULATED

Baro,Boudjema,Semenov, *Phys. Lett* **B660** Baro, Boudjema, G.C, Sun Hao, *Phys. Rev.* **D81** 015005 (2010) (2008) 550

• 
$$\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow f \overline{f}$$
 (bino)

• 
$$ilde{\chi}_1^0 ilde{ au}_1^+ 
ightarrow au^+ \gamma(Z^0)$$
 (bino)

- $\tilde{\tau}_1^+ \tilde{\tau}_1^+ \rightarrow \tau^+ \tau^+$  (bino)
- $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \to W^+ W^-, Z^0 Z^0$  (bino-wino, bino-higgsino, higgsino, higgsino-bino, wino)
- $\tilde{\chi}_1^0 \tilde{\chi}_1^+ \rightarrow u \bar{d}, t \bar{b}$  (bino-wino, higgsino, higgsino-bino, wino)
- $\tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow W^+ W^-, Z^0 Z^0$  (wino)

![](_page_25_Picture_13.jpeg)

## INTERPLAY BETWEEN ILC AND COSMOLOGY

- What is required from collider data to get a precise prediction of  $\Omega_{\chi}h^2$ ? (see Allanach *et al.* JHEP 0412 :020,2004.)
- What are the relevant observables to control the uncertainty on the predicted  $\Omega_{\chi} h^2$  ?
- What is the required accuracy in order to achieve PLANCK precision ?

#### Basic example as an illustration

- Typical mSUGRA scenario : LSP is  $\tilde{\chi}_1^0$  in the bulk region
- ${ ilde \chi}^0_1 { ilde \chi}^0_1 o \ell {ar \ell}$  through R-sleptons for  $\Omega_\chi h^2$

• At tree-level  $\boxed{m_{\tilde{\chi}_1^0}, m_{\tilde{\ell}}(m_{\tilde{\tau}_1})}_{\text{needed}}$  + mixing matrix of neutralinos and in the  $\tilde{\tau}_1$  sector needed to reconstruct  $\Omega_{\chi} h^2$ .

#### At one-loop level?

- Sensitivity to parameters entering in loops? (non decoupling corrections, thresholds...)
- Sensitivity to renormalisation schemes?

## SENSITIVITY TO SQUARKS MASSES

	$M_1$	$M_2$	$\mu$	$t_{\beta}$	M <sub>ẽ</sub> <sub>R</sub>	M <sub>ẽ</sub>	$M_3$	$M_{A^0}$
Masses (GeV)	90	200	-600	5	110	250	800	500
		250	${\rm GeV} \leq$	$M_{\widetilde{Q}} \leq$	< 800 Ge	V		

• At tree-level  $\Omega_{\chi} h^2$  sensitive to  $M_{\widetilde{Q}}$  when  $M_{\widetilde{Q}} \simeq 300$  GeV (new channels open).

• Effects of squarks relevant in loops for  $M_{\widetilde{Q}} \ge 300 \text{ GeV}$ ?

![](_page_27_Picture_4.jpeg)

## SENSITIVITY TO SQUARKS MASSES

	$M_1$	$M_2$	$\mu$	$t_{\beta}$	M <sub>ẽ</sub> <sub>R</sub>	M <sub>ẽ</sub>	$M_3$	$M_{A^0}$
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![](_page_28_Figure_4.jpeg)

![](_page_28_Picture_5.jpeg)

## SENSITIVITY TO SQUARKS MASSES

	$M_1$	$M_2$	$\mu$	$t_{\beta}$	M <sub>ẽ</sub> <sub>R</sub>	M <sub>ẽ</sub> ,	$M_3$	$M_{A^0}$
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- At tree-level  $\Omega_{\chi} h^2$  sensitive to  $M_{\widetilde{Q}}$  when  $M_{\widetilde{Q}} \simeq 300$  GeV (new channels open).
- Effects of squarks relevant in loops for  $M_{\widetilde{O}} \ge 300$  GeV?

![](_page_29_Figure_4.jpeg)

More or less the same conclusion at one-loop level. This may not be the case for other set of parameters.

![](_page_29_Picture_6.jpeg)

![](_page_30_Figure_1.jpeg)

![](_page_30_Picture_2.jpeg)

- Complete EW renormalisation of the MSSM and modularity with different schemes.
- One-loop corrections to masses, decays, cross sections at colliders.
- One-loop corrections to neutralino annihilation relevant for relic density and indirect detection.
- First steps done for the connection with micrOMEGAs.
- In any case for  $\Omega_{\chi} h^2 \otimes 1-2\% \Rightarrow$  one-loop corrections mandatory.
- Then at one-loop level more input is needed for an efficient reconstruction of parameters compared to the tree level case.
- Gather all available data to construct efficient renormalisation schemes.
- Implementation of QCD renormalisation in SloopS ongoing.

![](_page_31_Picture_9.jpeg)