

Precise cosmological predictions

from NLO calculations and
measurements at the LC

work in progress

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DER FORSCHUNG | DER LEHRE | DER BILDUNG

In collaboration with

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ECFA LC2013

Intro

and Structure

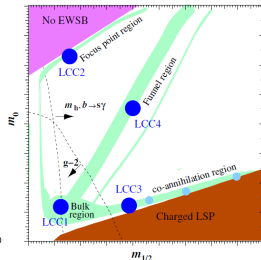
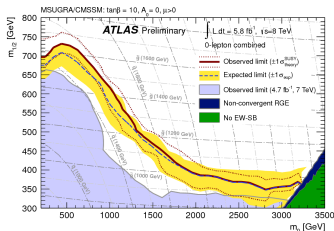
- Normally $\Omega_\chi h^2$ is used as a *constraint for BSM models*. Can we turn this around and, assuming we find SUSY, use it to see if the relic density predicted for the LSP found at colliders agrees with this measurement?
- Can predictions of $\Omega_\chi h^2$ from colliders reach accuracies comparable to the incredible 0.1196 ± 0.0031 (Planck2013) to learn about cosmology?¹

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- These work towards obtaining crucial information about DM in complementary ways: **ISR searches**, masses from thresholds and endpoints, (polarized) production cross-sections: staus, neutralinos...

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$\Delta\Omega_\chi h^2$ using parameters from $e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-$ @NLO:

- Our recent NLO analysis showed precise determination of M_1 , M_2 , $\tan\beta + m_{\tilde{t}_1}$, $\cos\theta_t$ M_A possible at LC^a
- In pMSSM, $\Delta\Omega_\chi h^2$ requires ΔX_i , $i = 1$ to 19^b
- Aim: combining our results with $\Delta X_i^{\text{LHC/LC}}$, how precisely could we predict $\Omega_\chi h^2$?

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- Background and further motivation

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- Background and further motivation
- Discuss errors on two complementary scenarios considered
- Present results of $\Delta\Omega_\chi h^2$ and conclusions

Background

The LC: an ideal machine for predicting $\Omega_\chi h^2$ -not a new idea

- Here LC determination of the fundamental parameters M_1 , M_2 , μ and $\tan \beta$ plays a primary role
- LO analysis showed precise determination of these parameters possible via $e^+e^- \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^0 \tilde{\chi}_2^0, \tilde{\chi}_2^0 \tilde{\chi}_2^0$ ³

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- pMSSM study by Peskin et. al. for 4 scenarios combining LHC and LC results ($\Delta\Omega_\chi h^2 \sim 14\%$ for a similar scenario)⁵

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The LHC has given us invaluable insight about the Higgs mass as well as the masses of the squarks and gluinos, motivating redoing these analyses

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Why incorporate loop effects?

SUSY loop effects can be **large**:

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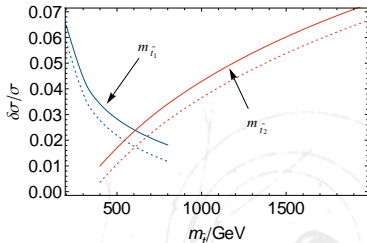
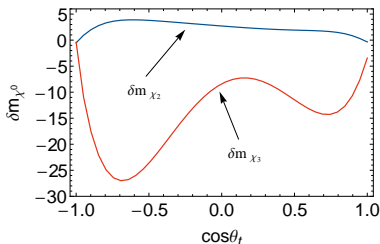
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- Use corrected **cross-sections**, A_{fb} and $m_{\tilde{\chi}_2^0}$, $m_{\tilde{\chi}_3^0}$ as observables
- Fit to M_1 , M_2 , μ , $\tan\beta$, + **stop sector** $m_{\tilde{t}_1}$, $m_{\tilde{t}_2}$ and $\cos\theta_t$

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- Sensitive to the stop sector and M_A , important in DM annihilation calculations!

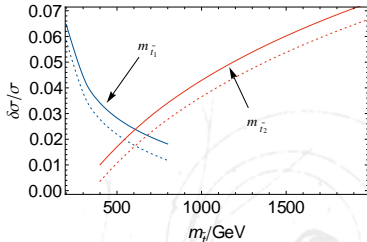
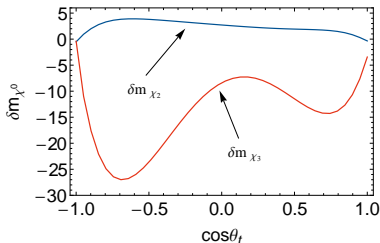


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Realistic % level determination of fundamental electroweak parameters possible: interesting to see implications for DM!

Scenarios

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- In the MSSM $\Omega_\chi h^2$ depends strongly on region, i.e. $m_{\tilde{\chi}_1^0}$, $m_{\tilde{\tau}_1}$ (stau coannihilation), M_A (funnel), μ (focus point)
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Scenario 1: Focus point region

Heavy sleptons (m_{ν_e} accessible in the fit) and charged Higgs (1 TeV) - inaccessible: creates uncertainty

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Scenario 2: Hybrid (focus point/bulk) region

Light sleptons, accessible at LHC and/or LC(500) and Lighter charged Higgs (~ 500 GeV). NLSP is $\tilde{\tau}_1$.

Projected errors from the LHC and LC

- Errors from NLO EWino production analysis LC
- Larger errors from continuum measurements at LC
- 10% errors on sfermion masses < 1 TeV from LHC
- Errors on accessible slepton masses from LC

Parameter	Scenario A	Scenario B
M_1	123 ± 0.3 (0.6)	105 ± 0.3
M_2	250 ± 0.6 (1.6)	211 ± 0.5
μ	182 ± 0.4 (0.7)	181 ± 0.4
$\tan \beta$	10 ± 0.5 (1.3)	11 ± 0.3
m_{A^0}	1000 ± 500	500 ± 150
M_3	1000 ± 100	1500 ± 150
$m_{\tilde{t}_1}$	400 ± 40	430 ± 43
$m_{\tilde{t}_2}$	800 ± 80	1520^{+200}_{-300}
$\cos \theta_t$	0.46 ± 0.15	$0.15^{+0.08}_{-0.06}$
$m_{\tilde{b}_1}$	400 ± 40	450 ± 45
$\cos \theta_b$	0 ± 0.06	0 ± 0.01
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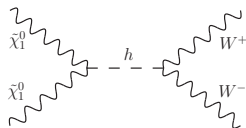
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Scenario A

$$\tilde{\chi}_1^0 \sim 0.83\tilde{B} - 0.18\tilde{W} + 0.44\tilde{h}_1 - 0.29\tilde{h}_2$$

Annihilation channels: W^+W^- (68%),
 ZZ (12%), hh (7%), Zh (6%) etc



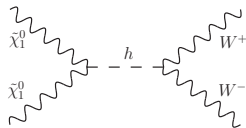
- The most important parameter here is $m_{\tilde{\chi}_1^0}$, determined via endpoints or monojet searches⁶
- Sfermions \sim decoupled (except stops via loop corrections to m_h)
- % level precision on fundamental EW parameters critical

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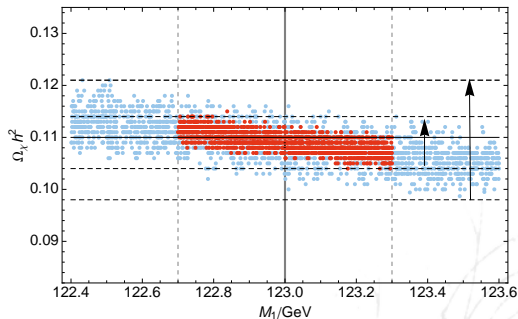
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Threshold vs. continuum masses, Scenario A



$m_{\tilde{\chi}_1^0}$ from endpoints improved via threshold scans

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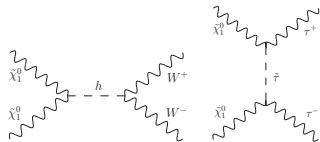
Scenario B

Results I

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W^+W^- (24%), $\tau^+\tau^-$ (23%), $\mu^+\mu^-$

(10%), e^+e^- (8%), $b\bar{b}$ (7%)



- Most important parameters are $m_{\tilde{\chi}_1^0}$ and $m_{\tilde{\tau}_1}$.
- The stop, sbottom, selectron and muon sectors will also be influential

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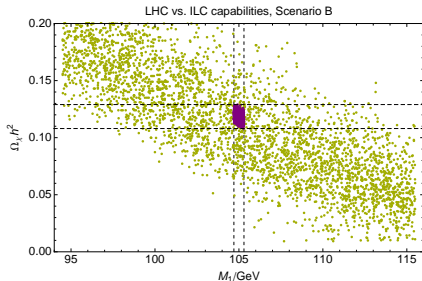
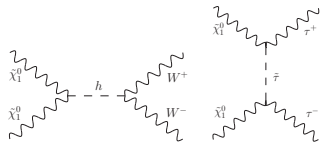
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We show how this improves at the LC wrt LHC determinations,⁷
and the effect of $\delta m_{\tilde{\tau}}^{\text{LC}}$: stau-coannihilation region ruled out⁸

⁷ Disclaimer!! LHC errors might be pessimistic as no valid dedicated study exists

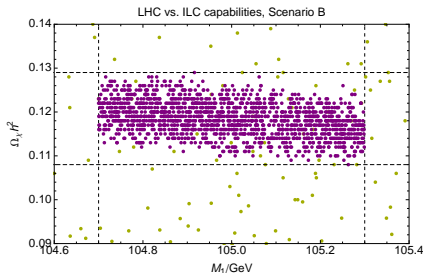
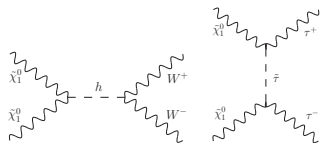
⁸ see e.g. P. Bechtle, M. Berggren, J. List, P. Schade and O. Stempel, Phys. Rev. D **82** (2010) 055016 [arXiv:0908.0876 [hep-ex]], E. Boos, H. U. Martyn, G. A. Moortgat-Pick, M. Sachwitz, A. Sherstnev and P. M. Zerwas, Eur. Phys. J. C **30** (2003) 395 [hep-ph/0303110].

Scenario B

Results I

$$\tilde{\chi}_1^0 \sim 0.87\tilde{B} - 0.18\tilde{W} + 0.41\tilde{h}_1 - 0.23\tilde{h}_2$$

W^+W^- (24%), $\tau^+\tau^-$ (23%), $\mu^+\mu^-$ (10%), e^+e^- (8%), $b\bar{b}$ (7%)



We show how this improves at the LC wrt LHC determinations,⁷
and the effect of $\delta m_{\tilde{\tau}}^{\text{LC}}$: stau-coannihilation region ruled out⁸

⁷ Disclaimer!! LHC errors might be pessimistic as no valid dedicated study exists

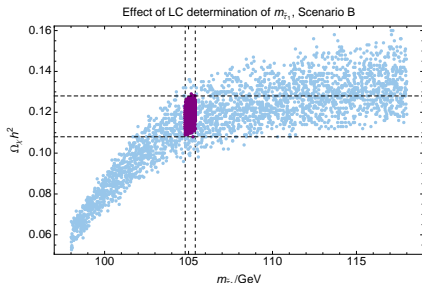
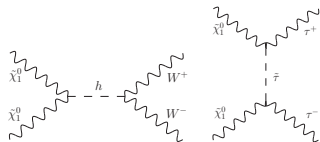
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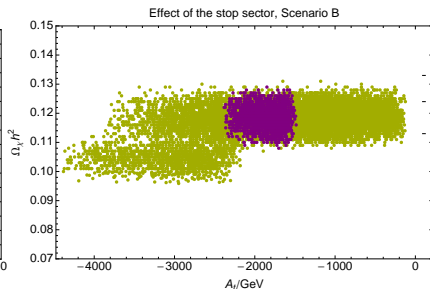
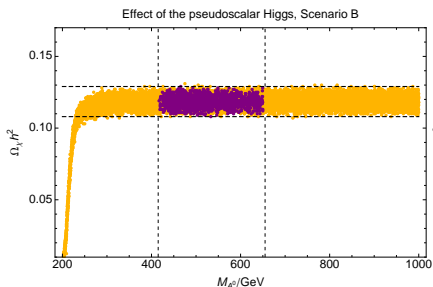
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Scenario B

Results II

- NLO chargino production analysis gives information about the stop sector and M_A
- Unique info? Bounds on $\cos\theta_t$, $m_{\tilde{t}_2}$ and upper bound on M_A



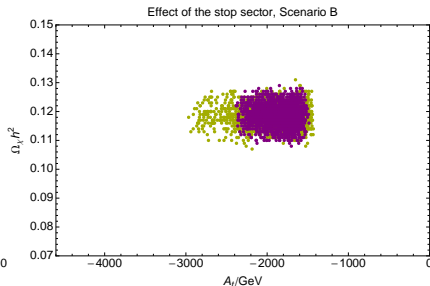
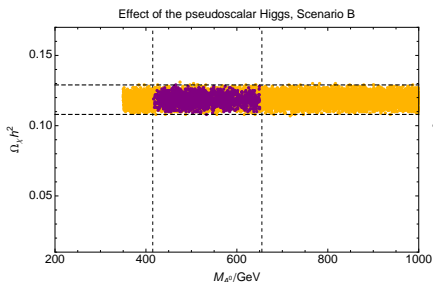
The constraints on M_A could be interesting in the funnel region + stop constraints for large A_t

The situation would probably somewhat change if loop corrections were included in the relic density calculation

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Higgs measurement (annihilation to Higgs ruled out) + $b \rightarrow s\gamma$
result in strong constraints on M_A +stops

The situation would probably somewhat change if loop corrections were included in the relic density calculation

What does this mean?

Summary

- Precise predictions $\sim 10\%$ for $\Omega_\chi h^2$ possible via LC measurements
- NLO parameter determination possible up to $\mathcal{O}(\%)$ level at a LC via $\tilde{\chi}^0/\tilde{\chi}^\pm$ production (with only light spectrum)
- Extract parameters \mathbf{M}_1 , \mathbf{M}_2 , μ , $\tan \beta$, $\mathbf{m}_{\tilde{\tau}_i}$ and $\cos \theta_t$ from fit to **NLO** predictions for masses, polarised σ s and A_{fb}
- Increased sensitivity to larger number of parameters compared to LO analyses \Rightarrow Precise determination of $\Omega_\chi h^2$ possible
- **Crucial** role played by LC: tune-able cms energy, threshold scans, polarization crucial
- If stau light, determination of stau masses at LC v. imp.

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Outlook

- Investigate case when non-zero **CP phases** e.g. ϕ_t
- Re-evaluation necessary when **full LHC study** of error projections available