



# Precise luminosity measurement at 3 TeV CLIC

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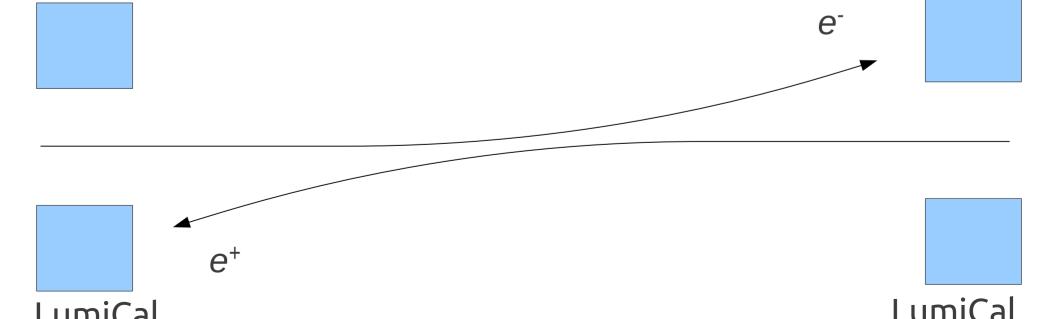




## Luminosity measurement de



- Counting Bhabha pairs in coincidence  $L = \frac{N(\Xi(E_{1,2}^{lab}, \Omega_{1,2}^{lab}))}{\sigma(Z(E_{1,2}^{CM}, \Omega_{1,2}^{CM}))}$ Precision ~0.6 permille at LEP
- A number of systematic effects limiting precision, at future colliders notably the beam-beam effects
  - Luminosity > 2 orders of magnitude higher than @LEP
  - Higher energy

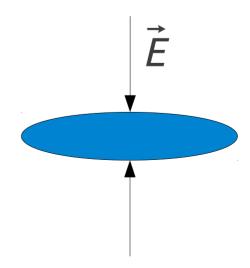




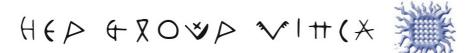
#### Beam-beam effects



- EM interaction at bunch crossing
  - $-\gamma \approx 6 \times 10^6$
  - "Pinch" effect strong focusing of the bunches
  - Beamstrahlung (before collision)
    - Energy loss
    - Shift of CM in the phase space Counting loss
  - EM deflection of the final charged particles – minor additional counting loss





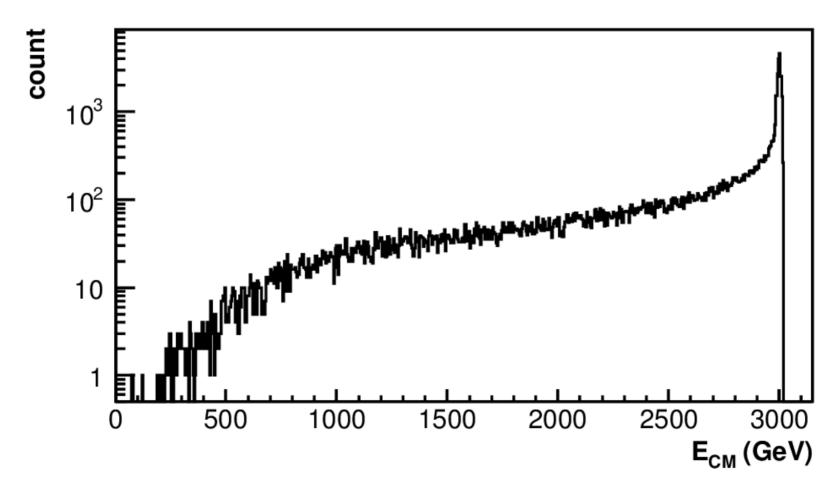




## $E_{CM}$ spectrum at CLIC



CM energies of colliding e<sup>-</sup>e<sup>+</sup> pairs in Guinea-Pig



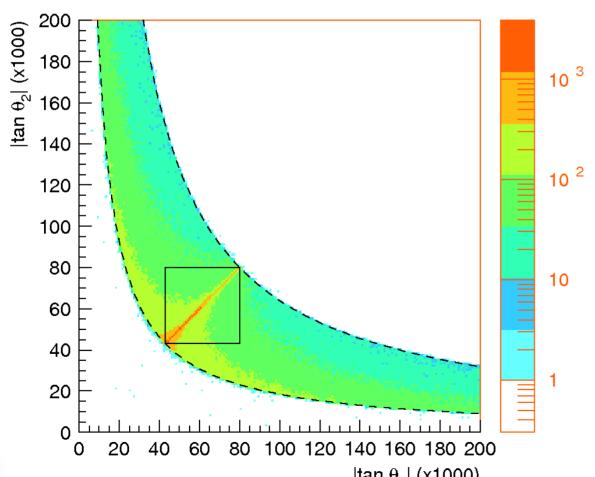




#### Angular loss at CLIC



 Distortion of polar angles of the outgoing leptons due to the beamstrahlung emission



- Beam-beam effects simulated by Guinea-PIG
- Polar angles undergo the Lorentz boost along the beam axis (to a very good approximation)





# Bhabha scattering and the beam-beam effects



- ISR and Beamstrahlung before collision escape detection
- FSR and Beamstrahlung of the final particles summed with electrons in the calorimeter
- Reconstructed kinematics of the collision frame  $(E_{CM}, \beta, \theta)$



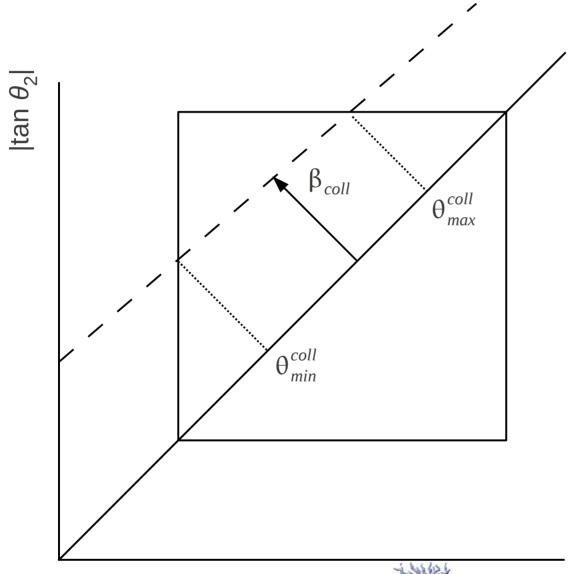






## Boost of the polar angles 🔠





- Among events with a given  $\beta_{coll}$  (dashed line), the angular counting loss can be analytically calculated
- Correct by the weighting factor

$$w(\beta_{coll}) = \frac{\int_{\theta_{min}}^{\theta_{max}} \frac{d\sigma}{d\theta} d\theta}{\int_{\theta_{min}}^{coll} \frac{d\sigma}{d\theta} d\theta}$$





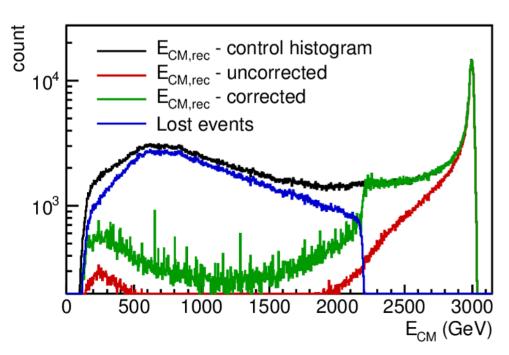


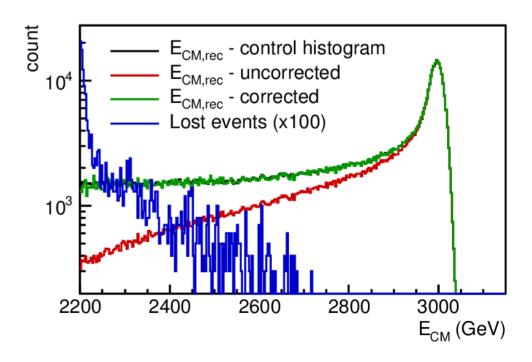


### Test of the angularloss correction



- Guinea-Pig + BHLUMI
- CM energies of Bhabha events (after emission of ISR, LumiCal energy response included)









### Test of the angularloss correction



- To quantify the agreement, the integral count in the top 5% of CM energy and in the tail from 80% to 90% was compared to the control histogram:
  - top 5%:  $\Delta L/L = (0.1 \pm 0.4) \times 10^{-3}$
  - 80%-90%:  $\Delta L/L = (-3.6 \pm 1.8) \times 10^{-3}$  (corrected "lost events"):  $\Delta L/L = (-0.9 \pm 1.8) \times 10^{-3}$
- No significant deviation in the corrected peak
- Deviation in the tail mostly due to the "lost events" (off-axis ISR)





### Systematic effects



- Off-axis ISR  $(\vec{\beta}_{coll} \neq \beta \hat{z})$  MC correction reliable
- Most energetic shower occasionally not containing the final electron
- Simplified expression for  $d\sigma/d\theta$  ( $\sim\theta^{-3}$ )
- Assumption of clean separation of ISR from FSR

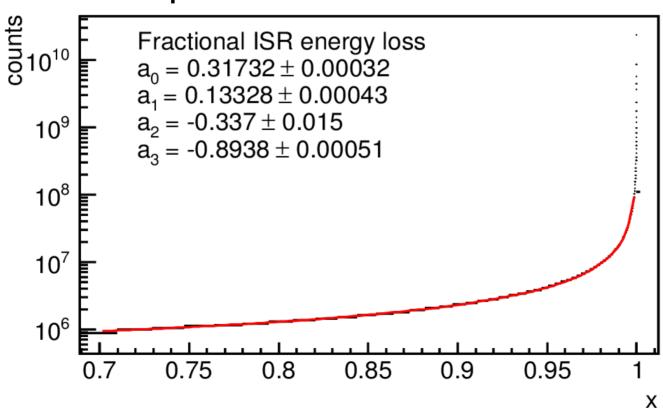


# ISR energy loss deconvolution



$$h(E_{CM,rec}) = \int_{0}^{E_{max}} B(E_{CM}) \frac{1}{E_{CM}} \Im(\frac{E_{CM,rec}}{E_{CM}}) dE_{CM}$$

ISR distribution parametrized from BHLUMI events



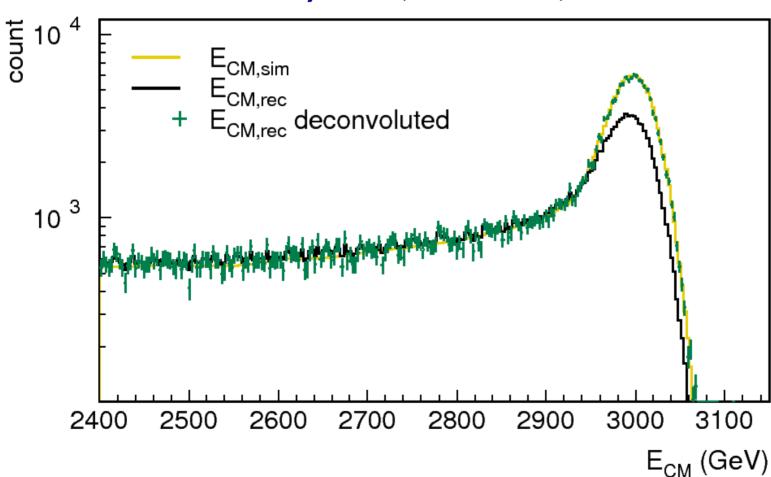




# ISR energy loss deconvolution



- Top 5%:  $\Delta L/L = (+1.3 \pm 2.0) \times 10^{-3}$
- 80% 90%:  $\Delta L/L = (-2.3 \pm 3.9) \times 10^{-3}$









### **CLIC - Summary**



Step	Top 5% Δ <i>L/L</i> (10 <sup>-3</sup> )	80% - 90% Δ <i>L/L</i> (10 <sup>-3</sup> )
BS+ISR correction	$-0.1 \pm 0.4$	-3.6 ± 1.8
Deconvolution	1.3 ± 2.0	-2.3 ± 3.9
Energy resolution	$0.05 \pm 0.03$	$0.09 \pm 0.09$
EMD (uncorrected)	$0.50 \pm 0.05$	1.08 ± 0.08
high $oldsymbol{eta}_{coll}$	< 0.1	2.7 ± 0.1
total	1.4 ± 2.0	2.7 ± 4.3









#### Conclusions



- The luminosity spectrum at CLIC extends down to almost zero CM energy
- Bhabha events at lower energies mostly invisible to the LumiCal
- Above 2200 GeV, the luminosity can be measured with precision of few permille
- The luminosity spectrum reconstructed to within the energy resolution of the luminometer

S. Lukić et al., JINST 8 (2013) P05008

S. Lukić, LCD-Note-2012-008

ILC case:

S. Lukić and I. Smiljanić, arXiv:1211.6869

I. Božović-Jelisavčić et al., arXiv:1304.4082







#### Thank you for your attention