

Light higgsino models and parameter determination

Krzysztof Rolbiecki
IFT-CSIC, Madrid



in collaboration with:
Mikael Berggren, Felix Brümmer, Jenny List,
Gudrid Moortgat-Pick, Hale Sert and Tania Robens

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- What does LHC tell us about 1st/2nd gen. squarks? → quite heavy
- Gaugino and stop searches model dependent – limits weaker

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: LHCp 2013

ATLAS Preliminary

$$\int L dt = (4.4 - 20.7) \text{ fb}^{-1} \quad (\sqrt{s} = 7, 8 \text{ TeV})$$

Model	$\epsilon, \mu, \tau, \gamma$	Jets	E_T^{miss}	$\int L dt \text{ (fb}^{-1}\text{)}$	Mass limit	Reference	
Inclusive searches	MSUGRA/CMSM	0	2-6 jets	Yes	20.3	$m_0 = m_{1/2}$ 1.8 TeV	ATLAS-CONF-2013-047
	MSUGRA/CMSM	1, μ	4 jets	Yes	20.3	1.9 TeV	ATLAS-CONF-2013-104
	MSUGRA/CMSM	0	7-10 jets	Yes	20.3	1.3 TeV	ATLAS-CONF-2013-054
	GL $\tilde{g} \rightarrow q\bar{q}$	0	2-6 jets	Yes	20.3	740 GeV	$m_{21} = 0 \text{ GeV}$ ATLAS-CONF-2013-047
	GL $\tilde{g} \rightarrow q\bar{q}$	0	2-6 jets	Yes	20.3	900 GeV	$m_{21} = 0 \text{ GeV}$ ATLAS-CONF-2013-047
	Clusio med. $\tilde{g} \rightarrow q\bar{q} + \tilde{g} \rightarrow q\bar{q}$	1, μ	2-4 jets	Yes	4.7	900 GeV	$m_{21} = 300 \text{ GeV}, m_{12}^2 = 0.5(m_{21}^2 + m_{12}^2)$ ATLAS-CONF-2013-077
	GG $\rightarrow q\bar{q}g\tilde{g}$	2, μ (SS)	3 jets	Yes	20.7	1.3 TeV	$m_{21} = 650 \text{ GeV}$ ATLAS-CONF-2013-007
	GMSB (L NLSB)	2, μ	2-4 jets	Yes	4.7	1.24 TeV	$\text{tag} \neq \text{tag}$ ATLAS-CONF-2013-025
	GMSB (L NLSB)	1, 2, τ	0-2 jets	Yes	4.8	1.4 TeV	$m_{21} = 50 \text{ GeV}$ ATLAS-CONF-2013-025
	GGM (bino NLSB)	2, τ	0	Yes	4.8	1.07 TeV	$m_{21} = 50 \text{ GeV}$ 1209.0753
	GGM (bino NLSB)	1, $\mu, \tau + \gamma$	0	Yes	4.8	619 GeV	$m_{21} = 50 \text{ GeV}$ ATLAS-CONF-2012-144
	GGM (higgsino-bino NLSB)	1, τ	1 b	Yes	4.8	900 GeV	$m_{21} = 320 \text{ GeV}$ 1211.1167
GGM (higgsino NLSB)	2, $\mu, \nu(\tilde{e})$	0-3 jets	Yes	5.8	900 GeV	$m_{21} = 200 \text{ GeV}$ ATLAS-CONF-2012-152	
Gravitino LSP	0	mono-jet	Yes	10.5	640 GeV	$m_{21} = 10^4 \text{ eV}$ ATLAS-CONF-2012-147	
3 rd gen. \tilde{g} med.	$\tilde{g} \rightarrow q\bar{q}$	0	3 b	Yes	12.8	1.34 TeV	$m_{21} = 300 \text{ GeV}$ ATLAS-CONF-2012-145
	$\tilde{g} \rightarrow q\bar{q}$	2, μ (SS)	0-3 b	No	20.7	900 GeV	$m_{21} = 600 \text{ GeV}$ ATLAS-CONF-2013-007
	$\tilde{g} \rightarrow q\bar{q}$	0	7-10 jets	Yes	20.3	1.14 TeV	$m_{21} = 600 \text{ GeV}$ ATLAS-CONF-2013-054
	$\tilde{g} \rightarrow q\bar{q}$	0	3 b	Yes	12.8	1.05 TeV	ATLAS-CONF-2013-145
3 rd gen. squarks direct production	$\tilde{t}_1 \rightarrow t\tilde{g}$	0	2 b	Yes	20.1	150-630 GeV	$m_{21} = 100 \text{ GeV}$ ATLAS-CONF-2013-053
	$\tilde{t}_1 \rightarrow t\tilde{g}$	2, μ (SS)	0-3 b	Yes	20.7	430 GeV	$m_{21} = 2 \text{ GeV}$ ATLAS-CONF-2013-007
	$\tilde{t}_1 \rightarrow t\tilde{g}$	1, 2, μ	1-2 b	Yes	4.7	187 GeV	$m_{21} = 55 \text{ GeV}$ 1208.4305, 1209.2102
	$\tilde{t}_1 \rightarrow t\tilde{g}$	2, μ	0-2 jets	Yes	20.3	230 GeV	$m_{21} = m_{12} = m_{13} = 50 \text{ GeV}, m_{23}^2 \ll m_{21}^2$ ATLAS-CONF-2013-048
	$\tilde{t}_1 \rightarrow t\tilde{g}$	2, μ	2-6 jets	Yes	20.3	150-340 GeV	$m_{21} = 5 \text{ GeV}, m_{12}^2 = m_{13}^2 = 10 \text{ GeV}$ ATLAS-CONF-2013-048
	$\tilde{t}_1 \rightarrow t\tilde{g}$	0	2 b	Yes	20.1	150-330 GeV	$m_{21} = 300 \text{ GeV}, m_{12}^2 = m_{13}^2 = 5 \text{ GeV}$ ATLAS-CONF-2013-053
	$\tilde{t}_1 \rightarrow t\tilde{g}$	1, μ	1 b	Yes	20.7	200-610 GeV	$m_{21} = 0 \text{ GeV}$ ATLAS-CONF-2013-037
	$\tilde{t}_1 \rightarrow t\tilde{g}$	1, μ	0 b	Yes	20.5	320-660 GeV	$m_{21} = 0 \text{ GeV}$ ATLAS-CONF-2013-024
	$\tilde{t}_2 \rightarrow t\tilde{g}$	2, μ (Z)	1 b	Yes	20.7	300 GeV	$m_{21} = 100 \text{ GeV}$ ATLAS-CONF-2013-025
	$\tilde{t}_2 \rightarrow t\tilde{g}$	3, μ (Z)	1 b	Yes	20.7	530 GeV	$m_{21} = m_{12}^2 = 100 \text{ GeV}$ ATLAS-CONF-2013-025
EW direct	$\tilde{W}_2 \rightarrow W\tilde{g}$	0	Yes	20.3	1	85-315 GeV	$m_{21} = 0 \text{ GeV}$ ATLAS-CONF-2013-049
	$\tilde{W}_2 \rightarrow W\tilde{g}$	2, μ	0	Yes	20.3	125-450 GeV	$m_{21} = 0 \text{ GeV}, m_{12}^2 = 0.5(m_{21}^2 + m_{12}^2) + m_{13}^2$ ATLAS-CONF-2013-049
	$\tilde{W}_2 \rightarrow W\tilde{g}$	2, τ	0	Yes	20.7	180-330 GeV	$m_{21} = 0 \text{ GeV}, m_{12}^2 = 0.5(m_{21}^2 + m_{12}^2) + m_{13}^2$ ATLAS-CONF-2013-028
	$\tilde{W}_2 \rightarrow W\tilde{g}$	3, μ	0	Yes	20.7	300 GeV	$m_{21} = 0 \text{ GeV}, m_{12}^2 = 0.5(m_{21}^2 + m_{12}^2) + m_{13}^2$ ATLAS-CONF-2013-025
	$\tilde{W}_2 \rightarrow W\tilde{g}$	3, μ	0	Yes	20.7	315 GeV	$m_{21} = m_{12}^2 = m_{13}^2 = 0$, δ leptons decoupled ATLAS-CONF-2013-035
	$\tilde{W}_2 \rightarrow W\tilde{g}$	3, μ	0	Yes	20.7	600 GeV	$m_{21} = m_{12}^2 = m_{13}^2 = 0$, δ leptons decoupled ATLAS-CONF-2013-035
Long-lived particles	Direct $\tilde{g} \rightarrow \text{prod. long-lived } \tilde{g}$	0	1 jet	Yes	4.7	230 GeV	$1 + m_{21}^2 = 10 \text{ m}$ 1210.2852
	Stable \tilde{g} , R-hadrons	0-2, μ	0	Yes	4.7	300 GeV	$\delta = \text{tag} + 20$ 1211.1557
	GMSB, stable \tilde{g} , low β	2, μ	0	Yes	4.7	300 GeV	$0 < \alpha - 10 \text{ GeV} < 10 \text{ m}$ 1208.6310
	GMSB, $\tilde{g} \rightarrow q\bar{q}$ long-lived \tilde{g}	2, τ	0	Yes	4.4	230 GeV	$1 \text{ m} < \alpha < 1 \text{ m}$, decoupled 1210.7451
RPV	LFV $pp \rightarrow X, \tau, \nu, e\mu$	1, μ	0	-	4.4	700 GeV	$1 \text{ m} < \alpha < 1 \text{ m}$, decoupled 1212.1272
	LFV $pp \rightarrow X, \tau, \nu, e\mu$	2, $\mu, \tau + 0$	-	-	4.6	1.8 TeV	$\lambda_{1,2,3} < 0, \lambda_{4,5,6} < 0.05$ 1212.1272
	Bilinear RPV CMSM	1, μ	7 jets	Yes	4.7	1.3 TeV	$m_{21} = m_{12}^2 = m_{13}^2 = 1 \text{ mm}$ ATLAS-CONF-2013-140
	Bilinear RPV CMSM	4, μ	7 jets	Yes	20.7	700 GeV	$m_{21} = 300 \text{ GeV}, \lambda_{1,2,3} = 0$ ATLAS-CONF-2013-026
Other	$\tilde{g} \rightarrow q\bar{q}$	0	6 jets	-	4.6	350 GeV	$m_{21} = 80 \text{ GeV}, \lambda_{1,2,3} = 0$ 1210.4813
	$\tilde{g} \rightarrow q\bar{q}$	2, μ (SS)	0-3 b	Yes	20.7	600 GeV	ATLAS-CONF-2013-007
	Scalar gluon	0	4 jets	-	4.6	100-207 GeV	incl. from 1115.2693 1210.4826
WMP interaction (DS, Dirac \tilde{g})	0	mono-jet	Yes	10.5	700 GeV	$m_{21} = 80 \text{ GeV}$, $\text{tag} < 4 - 40 \text{ GeV}$ for DS ATLAS-CONF-2012-147	

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

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$$\mathcal{L}_{\text{MSSM}} = \mu \tilde{H}_u \tilde{H}_d + \text{h.c.} + (m_{H_u}^2 + |\mu|^2) |H_u|^2 + (m_{H_d}^2 + |\mu|^2) |H_d|^2 + \dots$$

- Higgsino mass parameter μ is special: supersymmetric

A priori μ is unrelated to the scale of SUSY breaking

- μ cannot be too small (LEP chargino bound: $m_{\chi_1^\pm} \gtrsim 100$ GeV)
- μ should not be too large:

$$m_Z^2 = -2m_{H_u}^2 - 2|\mu|^2 + \mathcal{O}(\cot^2 \beta)$$

If $|m_{H_u}^2|, |\mu|^2 \gg m_Z^2 \Rightarrow$ large cancellation needed \Rightarrow Fine-tuning!

Two approaches:

- μ generated supersymmetrically, around EW scale by coincidence
- effective μ generated by SUSY breaking
in calculable models: μ/B_μ **problem** $\Rightarrow \mu$ still special

Naturalness wants μ around 100 GeV:

$$m_Z^2 = -2 m_{H_u}^2 - 2|\mu|^2 + \mathcal{O}(\cot^2 \beta)$$

LHC bounds want squarks and gluinos above 1 TeV.

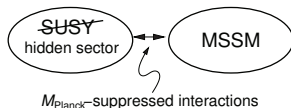
Motivates studying scenarios where **higgsinos are light** (EW scale) while **everything else is heavy** (multi-TeV) except maybe 3rd generation

light higgsinos = near-degenerate $\chi_1^0, \chi_1^\pm, \chi_2^0$ around 100–200 GeV

Light higgsinos from hybrid gauge-gravity mediation

Gravity med.: $\mu \sim m_{3/2}$, \rightarrow Giudice/Masiero '88

$$m_{\text{soft}} \sim m_{3/2}$$



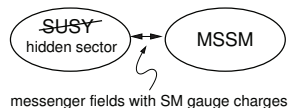
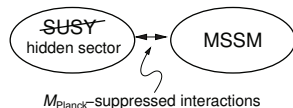
Light higgsinos from hybrid gauge-gravity mediation

Gravity med.: $\mu \sim m_{3/2}$, \rightarrow Giudice/Masiero '88

$$m_{\text{soft}} \sim m_{3/2}$$

Gauge med.: $\mu = 0$,

$$m_{\text{soft}} \sim m_{3/2} \cdot N_{\text{mess.}} \cdot \frac{M_{\text{Planck}}}{M_{\text{mess.}}} \cdot \frac{g^2}{16\pi^2}$$



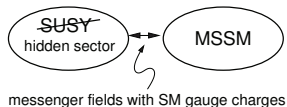
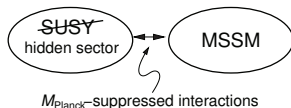
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Models with GUT-scale extra dimensions:

- typically include superheavy "exotic matter": candidate messengers

- masses: $M_{\text{mess.}} \approx M_{\text{GUT}} \approx M_{\text{Planck}} \cdot \frac{g^2}{16\pi^2}$

- multiplicities: $N_{\text{mess.}} \sim \mathcal{O}(\text{few tens})$

Hybrid gauge-gravity mediation in higher-dim. GUTs: \rightarrow Brümmer, Buchmüller '11,'12

$$\mu \sim m_{3/2} \sim \mathcal{O}(100 \text{ GeV}), \quad m_{\text{soft}} \sim N_{\text{mess.}} \cdot m_{3/2} \sim \mathcal{O}(\text{TeV})$$

A mass spectrum from hybrid gauge-gravity mediation

particle	mass [GeV]
h^0	124
χ_1^0	164
χ_1^\pm	166
χ_2^0	167
χ_3^0	2700
χ_4^0	4100
χ_2^\pm	4100
H_0	2200
A_0	2200
H^\pm	2200
\tilde{g}	4200
$\tilde{\tau}_1$	1900
other sleptons	2500 – 3600
squarks	2700 – 5000

$$\tan \beta = 44$$

Quick recap: chargino and neutralino Sector

$$\begin{aligned}\mathcal{L}_{\tilde{\chi}} = & \overline{\tilde{\chi}_i^-} (\not{p} \delta_{ij} - \omega_L (U^* X V^\dagger)_{ij} - \omega_R (V X^\dagger U^T)_{ij}) \tilde{\chi}_j^- \\ & + \frac{1}{2} \overline{\tilde{\chi}_i^0} (\not{p} \delta_{ij} - \omega_L (N^* Y N^\dagger)_{ij} - \omega_R (N Y^\dagger N^T)_{ij}) \tilde{\chi}_j^0\end{aligned}$$

$$X = \begin{pmatrix} M_2 & \sqrt{2} M_W \sin \beta \\ \sqrt{2} M_W \cos \beta & \mu \end{pmatrix}$$

diagonalised via
 $\mathbf{M}_{\tilde{\chi}^+} = U^* X V^\dagger$

⁰where we define $\omega_{L/R} = \frac{1}{2}(1 \mp \gamma_5)$

Quick recap: chargino and neutralino Sector

$$\mathcal{L}_{\tilde{\chi}} = \overline{\tilde{\chi}_i^-} (\not{p} \delta_{ij} - \omega_L (U^* X V^\dagger)_{ij} - \omega_R (V X^\dagger U^T)_{ij}) \tilde{\chi}_j^- \\ + \frac{1}{2} \overline{\tilde{\chi}_i^0} (\not{p} \delta_{ij} - \omega_L (N^* Y N^\dagger)_{ij} - \omega_R (N Y^\dagger N^T)_{ij}) \tilde{\chi}_j^0$$

$$X = \begin{pmatrix} M_2 & \sqrt{2} M_W \sin \beta \\ \sqrt{2} M_W \cos \beta & \mu \end{pmatrix}$$

diagonalised via
 $\mathbf{M}_{\tilde{\chi}^+} = U^* X V^\dagger$

$$Y = \begin{pmatrix} M_1 & 0 & -M_Z c_\beta s_W & M_Z s_\beta s_W \\ 0 & M_2 & M_Z c_\beta c_W & -M_Z s_\beta c_W \\ -M_Z c_\beta s_W & M_Z c_\beta c_W & 0 & -\mu \\ M_Z s_\beta s_W & -M_Z s_\beta c_W & -\mu & 0 \end{pmatrix}$$

diagonalised via
 $\mathbf{M}_{\tilde{\chi}^0} = N^* Y N^\dagger$

$$M_1, M_2 \sim \mathcal{O}(\text{TeV}), \mu \sim \mathcal{O}(100 \text{ GeV})$$

⁰where we define $\omega_{L/R} = \frac{1}{2}(1 \mp \gamma_5)$

Approximate mass eigenstates

$$\begin{aligned}\tilde{\chi}_1^0 &= \frac{1}{\sqrt{2}} (\tilde{h}_d^0 - \tilde{h}_u^0) + \frac{\sin \beta + \cos \beta}{\sqrt{2}} \frac{m_Z}{M_1} \sin \theta_w \tilde{B} - \frac{\sin \beta + \cos \beta}{\sqrt{2}} \frac{m_Z}{M_2} \cos \theta_w \tilde{W}^0 \\ \tilde{\chi}_2^0 &= \frac{1}{\sqrt{2}} (\tilde{h}_d^0 + \tilde{h}_u^0) - \frac{\sin \beta - \cos \beta}{\sqrt{2}} \frac{m_Z}{M_1} \sin \theta_w \tilde{B} + \frac{\sin \beta - \cos \beta}{\sqrt{2}} \frac{m_Z}{M_2} \cos \theta_w \tilde{W}^0 \\ \tilde{\chi}_1^\pm &= \tilde{h}_u^\pm - \sqrt{2} \sin \beta \frac{m_W}{M_2} \tilde{W}^\pm\end{aligned}$$

with masses given by

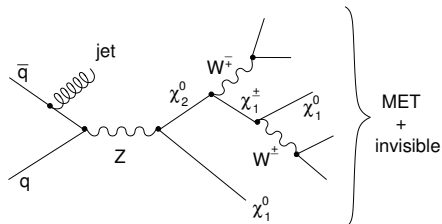
$$\begin{aligned}M_{\tilde{\chi}_{1,2}^0} &= |\mu| \mp \frac{m_Z^2}{2} (\cos \beta \pm \text{sign}(\mu) \sin \beta)^2 \left(\frac{\sin^2 \theta_w}{M_1} + \frac{\cos^2 \theta_w}{M_2} \right) \\ M_{\tilde{\chi}_1^\pm} &= |\mu| - \sin 2\beta \text{sign}(\mu) \cos^2 \theta_w \frac{m_Z^2}{M_2}\end{aligned}$$

and the mass difference between chargino $\tilde{\chi}_1^\pm$ and the LSP

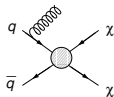
$$M_{\tilde{\chi}_1^\pm} - M_{\tilde{\chi}_1^0} = m_Z^2 \left[\frac{\sin 2\beta}{2} \left(\frac{\sin^2 \theta_w}{M_1} - \frac{\cos^2 \theta_w}{M_2} \right) + \frac{1}{2} \left(\frac{\sin^2 \theta_w}{M_1} + \frac{\cos^2 \theta_w}{M_2} \right) + \mathcal{O} \left(\frac{\mu}{M_i^2} \right) \right]$$

- 1 Light higgsinos from gauge-gravity mediation
- 2 Light higgsinos at colliders

Higgsinos + monojets



- Monojet (and -photon) signal at ATLAS and CMS
- Usual interpretation of $j+\text{MET}$: generic WIMP with contact interaction



- Should also provide mass limits for higgsinos!

Detailed parameter measurements require linear collider

→ Baer/Barger/Huang '11; Berggren/Brümmer/List/Moortgat-Pick/Rolbiecki/Sert, in preparation

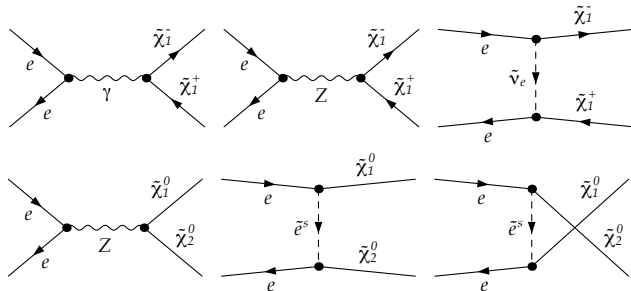
Two benchmark scenarios with hybrid gauge-gravity mediation:

m_{h^0}	124
$m_{\chi_1^0}$	164.2
$m_{\chi_1^\pm}$	165.8
$m_{\chi_2^0}$	166.9
M_1	1700
M_2	4400
μ	166

m_{h^0}	127
$m_{\chi_1^0}$	166.6
$m_{\chi_1^\pm}$	167.4
$m_{\chi_2^0}$	167.6
M_1	5300
M_2	9500
μ	166

- production of $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ and $\tilde{\chi}_2^0 \tilde{\chi}_1^0$ already at $\sqrt{s} = 350$ GeV
- small mass splittings \Rightarrow pions, soft γ s

Production at the ILC



- only $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ produced: Z couplings forbid $\tilde{\chi}_1^0 \tilde{\chi}_1^0$ and $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ for higgsinos
- the t and u slepton exchange proportional to Yukawa coupling \Rightarrow negligible for higgsinos \Rightarrow heavy sleptons in our scenarios but even for light ones they would not contribute

- neutralino decays via off-shell Z boson or two-body $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \gamma$
- for low mass difference the latter loop induced dominates
- chargino decays with off-shell W boson $\tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 W^* \rightarrow \tilde{\chi}_1^0 f f'$
- lifetime up to $\sim 10^{-12}$ s \Rightarrow decay length $c\tau \sim 10^{-3}$ m
- in scenarios with small mass difference one can look for displaced vertices
- two approaches to calculate decay modes:
 - calculate parton level matrix element (e.g. with `Whizard`) and hadronize with `Pythia`
 - use hadronic currents like for τ decays, implemented in `Herwig++`
 - give very different results...

Chargino decays

Pythia hadronization vs Herwig++ hadronic currents

$m_h = 124 \text{ GeV}; \Delta m = 1.6 \text{ GeV}$

$m_h = 127 \text{ GeV}; \Delta m = 0.7 \text{ GeV}$

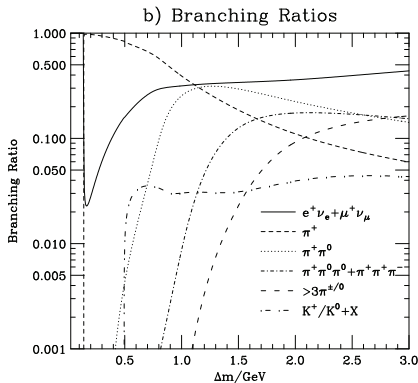
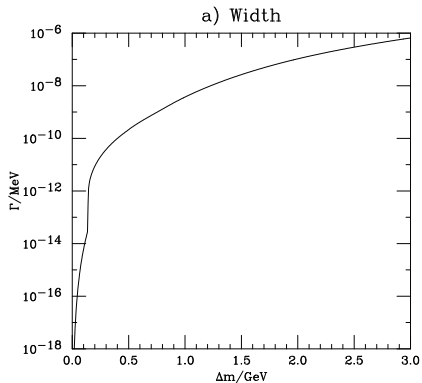
decay	Pythia	Herwig++
$\tilde{\chi}_1^0 \ell \nu$	39%	29%
$\tilde{\chi}_1^0 \pi^+$	11%	60%
$\tilde{\chi}_1^0 \pi^+ \pi^0$	23%	7%
$\tilde{\chi}_1^0 K^+$	0%	3.5%

decay	Pythia	Herwig++
$\tilde{\chi}_1^0 \ell \nu$	40%	33%
$\tilde{\chi}_1^0 \pi^+$	10%	17%
$\tilde{\chi}_1^0 \pi^+ \pi^0$	20%	29%
$\tilde{\chi}_1^0 \pi^+ \pi^0 \pi^0$	6%	8%
$\tilde{\chi}_1^0 \pi^+ \pi^+ \pi^-$	6%	7%
$\tilde{\chi}_1^0 \pi^+ \pi^+ \pi^- \pi^0$	11%	2.5%

- Pythia also predicts decay modes not present in Herwig++ that should otherwise be highly suppressed
- to be fair, Herwig++ hadronization is not doing much better here...

Chargino decay modes from Herwig++

Grellscheid, Richardson arXiv:0710.1951



⇒ note that radiative corrections to the mass difference (typically ~ 200 MeV) will have a profound effect on branching ratios if the tree level difference $\lesssim 1$ GeV

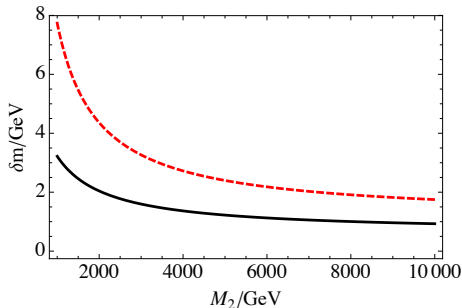
Neutralino decays

Decay mode	$m_h = 124 \text{ GeV}$	$m_h = 127 \text{ GeV}$
$\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 + \dots$		
γ	23.6%	74.0%
$\nu\bar{\nu}$	21.9%	9.7%
e^+e^-	3.7%	1.6%
$\mu^+\mu^-$	3.7%	1.5%
hadrons	44.9%	12.7%
$\chi_1^\pm + X$	1.9%	0.4%

- radiative decay to $\tilde{\chi}_1^0\gamma$ dominates for small mass differences
- cleanest experimentally
- for $m_h = 124 \text{ GeV}$ scenario hadronic decay modes become important giving low mass jets

Measurements at the ILC

- cross sections: $\tilde{\chi}_1^+ \tilde{\chi}_1^-$, $\tilde{\chi}_1^0 \tilde{\chi}_2^0$, $\sqrt{s} = 350, 500$ GeV
⇒ accuracy 1.5%–6%
- mass differences: $\Delta m(\tilde{\chi}_1^\pm - \tilde{\chi}_1^0)$
⇒ accuracy 20–70 MeV



- chargino/neutralino masses from recoil against ISR
⇒ accuracy 1.5–3.5 GeV
- ⇒ more details in Hale's talk

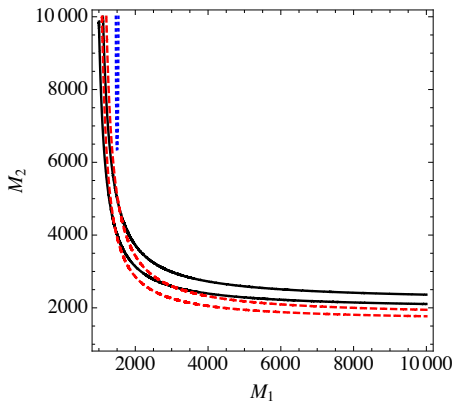
- four fitted parameters:
⇒ M_1 , M_2 , μ and $\tan\beta$
- perform a χ^2 fit

$$\chi^2 = \sum_i \left| \frac{\mathcal{O}_i^{\text{exp}} - \bar{\mathcal{O}}_i^{\text{th}}}{\delta \mathcal{O}_i^{\text{exp}}} \right|^2$$

- $\tan\beta$ remains unconstrained, will be varied in $[1, 60]$ range
- m_μ constrained by cross sections and absolute masses:
⇒ in $m_h = 124$ GeV scenario $\mu = 167_{-1.5}^{+2.1}$ GeV
⇒ in $m_h = 127$ GeV scenario $\mu = 166 \pm 1$ GeV
- mass difference between chargino and LSP strongly constraints M_1 and M_2

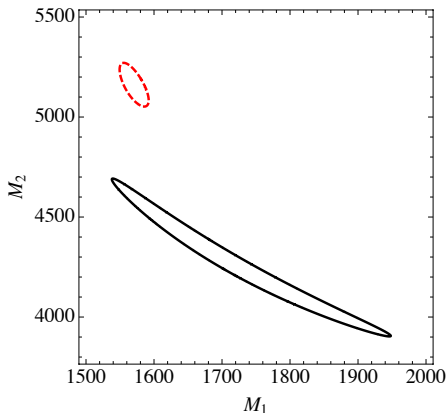
χ^2 contours $m_h = 124$ GeV scenario

$\mathcal{L} = 500 \text{ fb}^{-1}$



$\tan \beta = 50$, $\tan \beta = 10$, $\tan \beta = 1$

$\mathcal{L} = 2 \text{ ab}^{-1}$



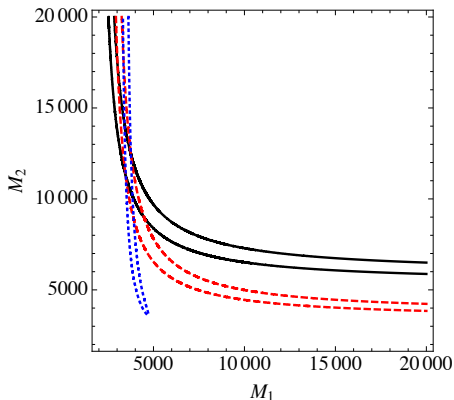
$\tan \beta = 50$, $\tan \beta = 1$

errors halved

includes $\Delta m = m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$

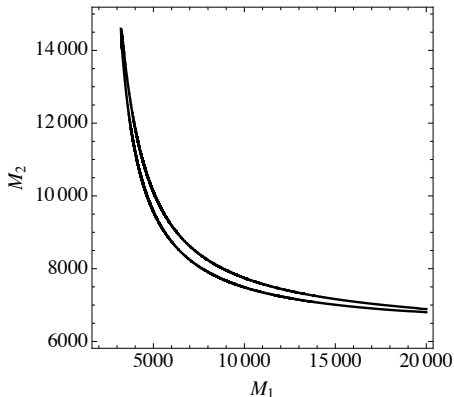
χ^2 contours $m_h = 127$ GeV scenario

$\mathcal{L} = 500 \text{ fb}^{-1}$



$\tan \beta = 50$, $\tan \beta = 5$, $\tan \beta = 2$

$\mathcal{L} = 2 \text{ ab}^{-1}$



$\tan \beta = 50$

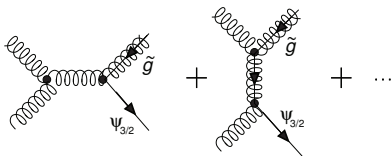
errors halved

includes $\Delta m = m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0}$

- Light higgsinos = near-degenerate $\tilde{\chi}_1^0, \tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ around 100 – 200 GeV
- Motivated from naturalness
- Motivated from model-building
E.g. hybrid gauge-gravity mediation: μ gravity-mediated, soft masses gauge-mediated
- Higgsinos hard to see: mass degeneracy \Rightarrow soft decay products
- Difficult to see at LHC if everything else heavy
- Good case for linear collider
- With precise measurements multi-TeV parameters could be resolved
- Experimental analysis ongoing \Rightarrow **Hale's talk**

Gravitino LSP is **natural dark matter candidate**

Gravitinos produced thermally during reheating at large T_R :



$$\Omega_{\psi_{3/2}} h^2 \approx 0.21 \left(\frac{T_R}{10^{10} \text{ GeV}} \right) \left(\frac{100 \text{ GeV}}{m_{3/2}} \right) \left(\frac{m_{\tilde{g}}}{1 \text{ TeV}} \right)^2$$

see e.g. \rightarrow Bolz, Brandenburg, Buchmüller '00

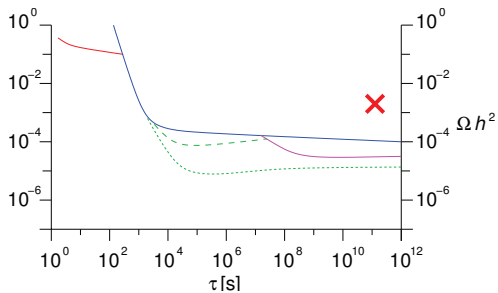
$T_R \approx 10^{10} \text{ GeV}$:

- Nicely compatible with leptogenesis
- Right order of magnitude for DM abundance

Cosmology

Problem: χ_1^0 NLSP long-lived, decays after BBN

Energetic decay products destroy nuclei, distorting light element abundances



Bounds from \rightarrow Jedamzik '06:
NLSP relic density vs. lifetime
(assuming large hadronic BR)

${}^4\text{He}$, ${}^2\text{H}$, ${}^3\text{He}$, (Li)

- Higgsino NLSP relic density **low**: coannihilation with χ_1^\pm
(recall first spectrum, $m_{\chi_1^0} = 137$ GeV, $m_{\chi_1^\pm} = 140$ GeV):

$$\Omega_{\chi_1^0} h^2 = 3 \cdot 10^{-3}$$

- ... but still in conflict with BBN bounds
- (Small) R-parity violation? Additional entropy production?